

An other look at Neutrino Properties

- 1) Magnetic moments
- 2) $\mu \rightarrow e + \gamma$

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May 2002

Magnetic Moments

Magnetic Moment Matrix

$$\mu_{jk}, k=1,2,3 \text{ with } \begin{cases} \mu_{jj} = 0 \text{ Majorana} \\ \mu_{jj} \neq 0 \text{ Dirac} \end{cases}$$

Astrophysics: strong constraints on Dirac moments

less strong on Majorana moments ($\leq 10^{-12} \mu_B$ stellar cooling)

$\bar{\nu}_e e^-$ scattering:

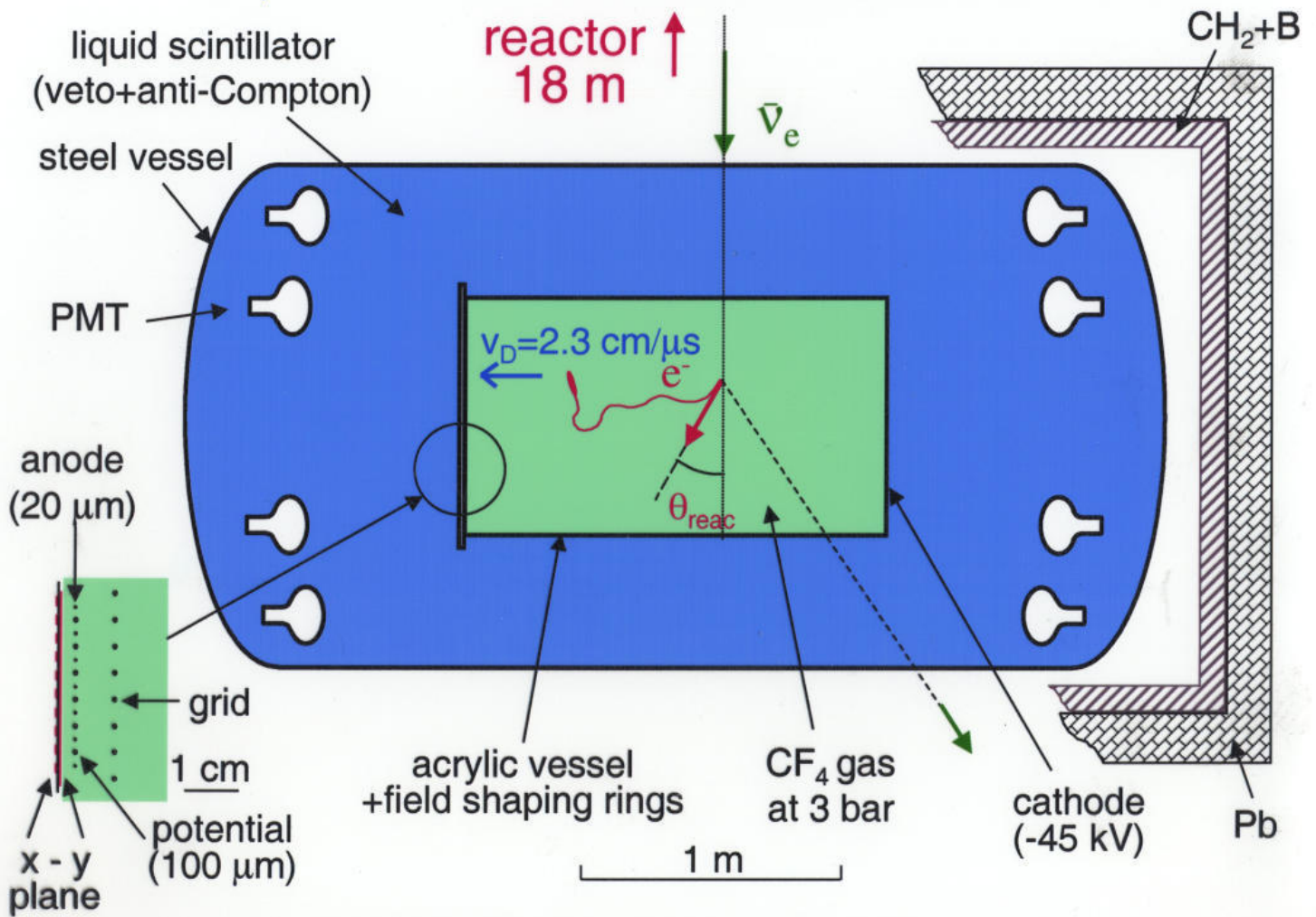
$$\text{sun: } (\mu_e^{\text{sol}})^2 = \sum_j |\sum_k U_{ek} e^{i p_k L} \mu_{jk}|^2 \quad p_i = \sqrt{E^2 - m_i^2}$$

$$\text{SK, Beacom + Vogel: } \mu_e^{\text{sol}} < 1.5 \cdot 10^{-10} \mu_B \quad 90 \% \text{ CL}$$

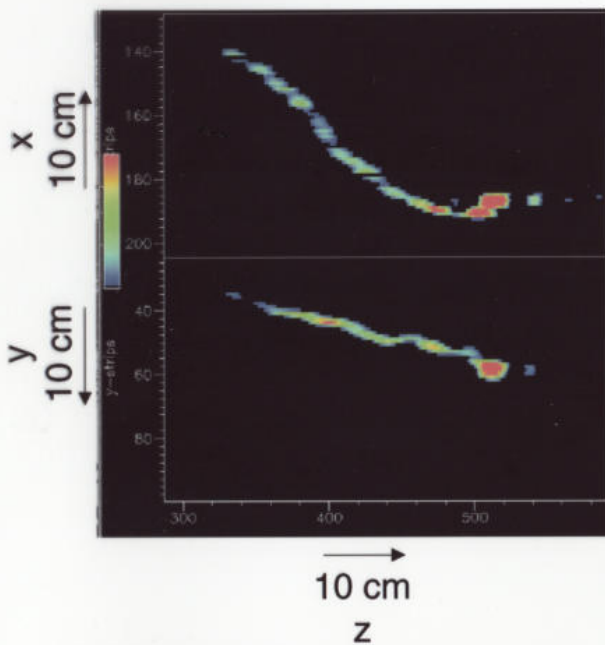
$$\text{reactor: } (\mu_e^{\text{reac}})^2 = \sum_j |\sum_k U_{ek} \mu_{jk}|^2$$

$$\text{Rovno (SB } \sim 1:100): \mu_e^{\text{reac}} < 1.9 \cdot 10^{-10} \mu_B \quad 95 \% \text{ CL}$$

μ_ν : **MUNU** (Grenoble-Neuchâtel-Padoue-Zurich)



MUNU



TPC \perp reactor-detector axis,
absolutely symmetric

Energy resolution: 13 % (1σ) at 1 MeV

Angular resolution: $\approx 20^\circ$
(\sim angle dependent)

Anti-Compton threshold: 100 keV

Live time: 60% (cosmics...)

Threshold: 300 keV (but...)

Reactor: power 2750 MWth
distance 18 m

Neutrino spectrum:

$E_\nu < 1.7$ MeV: calculations (Vogel et al.)

$E_\nu > 1.7$ MeV: from measurement of beta
decay of fission fragments
(ILL), tested in $\bar{\nu}_e + p \Rightarrow e^+ + n$
($\approx 5\%$)

MUNU: 65 days effective reactor on (40 analyzed)
25 days effective reactor off (10.7 analyzed)

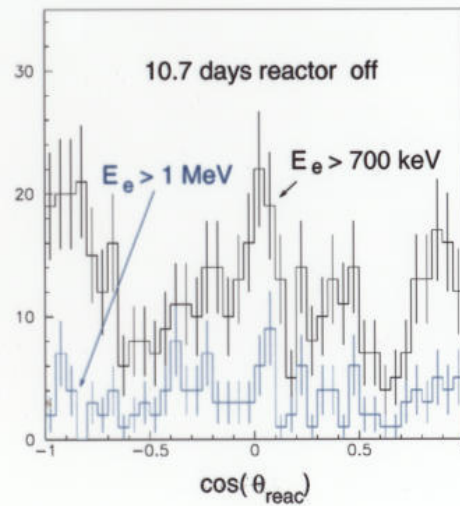
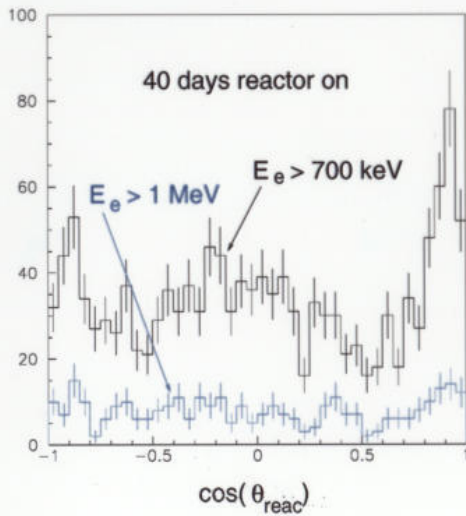
1) **Automatic filtering** eliminates:

α 's, Compton, cosmics, uncontained ($r > 42$ cm), discharges...

2) **Visual scan:**

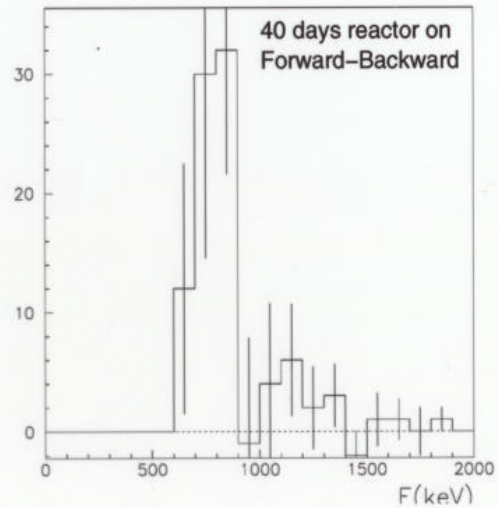
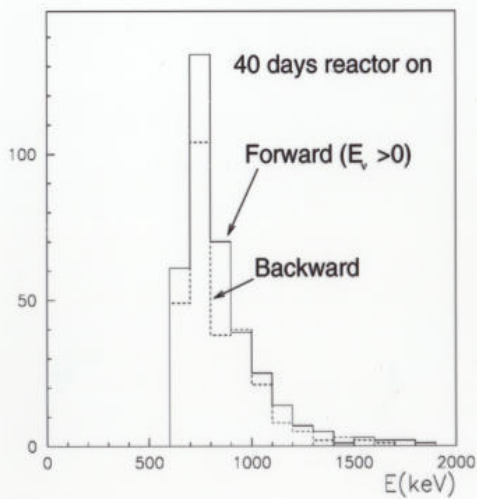
selects **contained single electrons** ($E_e > 700$ keV)

determines: $\vartheta_{det}, \varphi_{det}, \vartheta_{reac}, \varphi_{reac}$, selects: $\vartheta_{det} < 90^\circ$

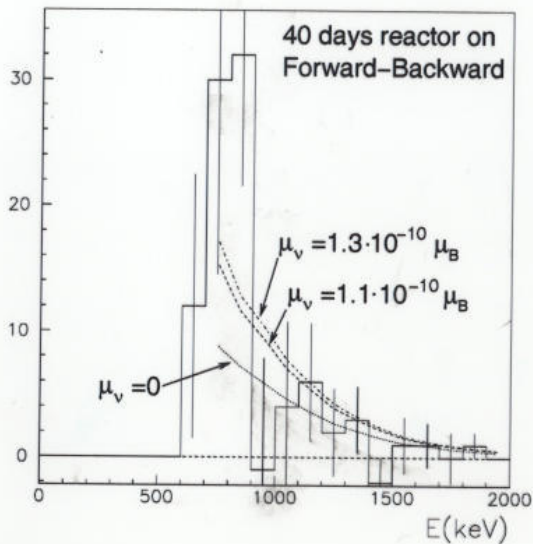


MUNU

Apply kinematical cut $E_\nu(E_e, \vartheta_{\text{reac}}) > 0 (\approx \cos(\vartheta_{\text{reac}}) > 0.7)$ **(forward)**
or $E_\nu(E_e, \pi - \vartheta_{\text{reac}}) > 0 (\approx \cos(\vartheta_{\text{reac}}) < -0.7)$ **(backward)**



MUNU



$$E_e > 700 \text{ keV: } 2.05 \pm 0.57 \text{ day}^{-1}$$

$$\text{calculated } \mu_\nu = 0: 1.03 \text{ day}^{-1}$$

$$\mu_\nu = 0: \chi^2 = 10.6/6 \text{ dof}$$

$$\mu_\nu = 1.1 \cdot 10^{-10} \mu_B: \chi^2 = 9.0/6 \text{ dof (min.)}$$

$$\mu_\nu = 1.7 \cdot 10^{-10} \mu_B: \chi^2 = 11.6/6 \text{ dof}$$

$$\rightarrow \mu_\nu < 1.7 \cdot 10^{-10} \mu_B \text{ (90\% CL)}$$

$$E_e > 1 \text{ MeV: } 0.37 \pm 0.25 \text{ day}^{-1}$$

$$\text{calculated } \mu_\nu = 0: 0.45 \text{ day}^{-1}$$

$$\mu_\nu = 0: \chi^2 = 0.54/3 \text{ dof}$$

$$\mu_\nu = 1.3 \cdot 10^{-10} \mu_B: \chi^2 = 3.01/3 \text{ dof}$$

$$\rightarrow \mu_\nu < 1.3 \cdot 10^{-10} \mu_B \text{ (90\% CL)}$$

TEXONO* Collaboration

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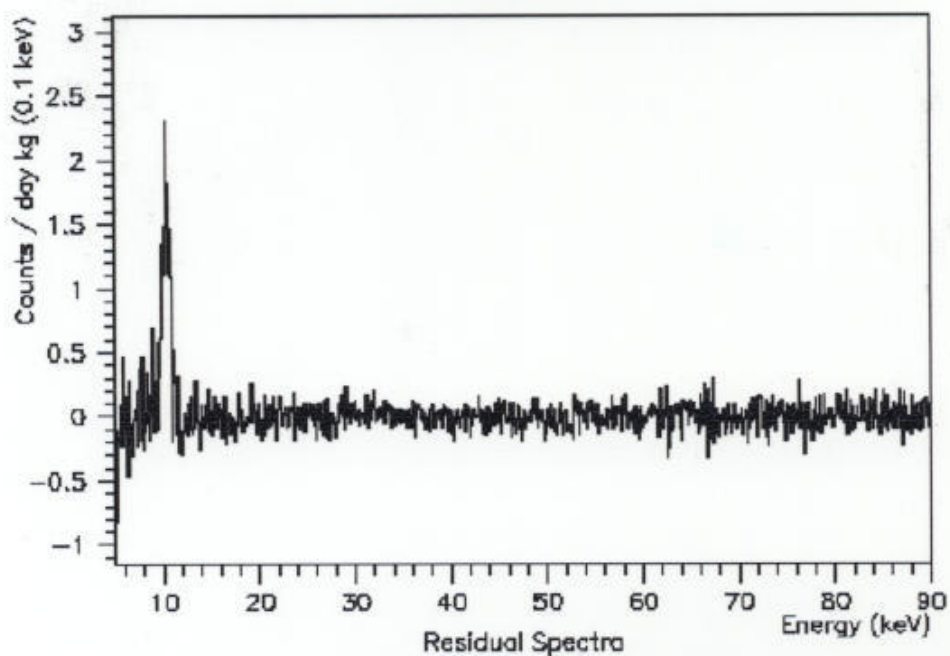
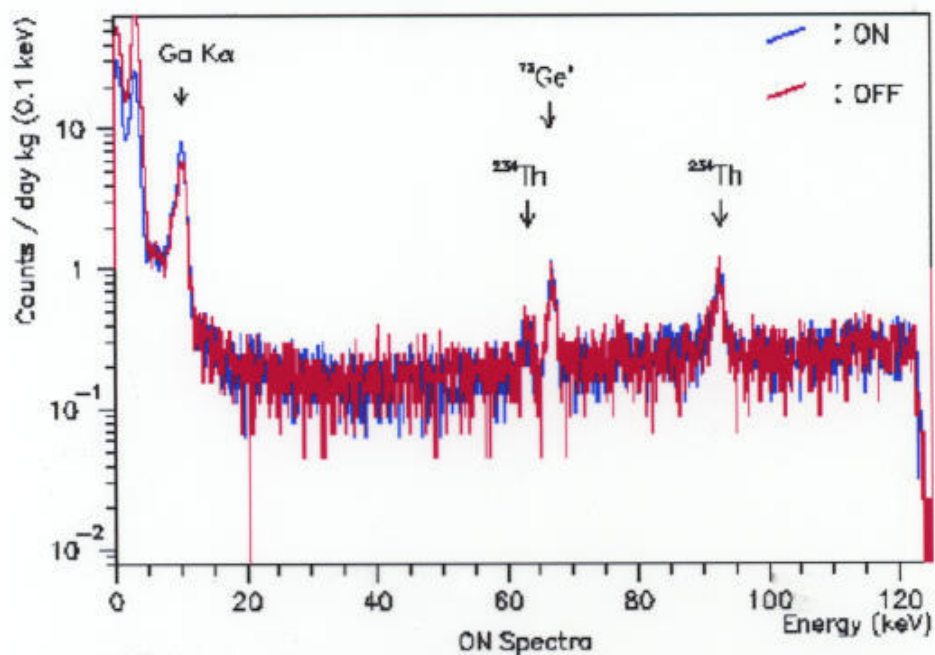
HISTORY

- ◆ Initiate : Chang Chung-Yung 1996
- ◆ First ICollaboration Meeting/Official Start : October 1997
- ◆ First Paper : October 1998
- ◆ KS Reactor Experiment Installation : June 2000
- ◆ First Ph.D. : Liu Yan , July 2000
- ◆ First Physics Data Taking : June 2001.

*Taiwan EXperiment On Neutrino

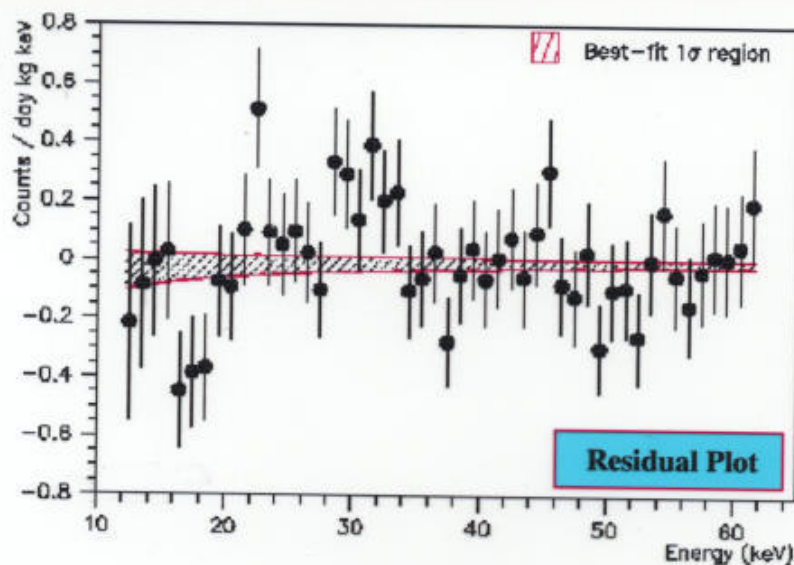
Home Page @ <http://hepmail.phys.sinica.edu.tw/~texono/>

KS/P1/Ge : Spectra & Residual



Based on 62/46 days of Reactor ON/OFF data

KS/P1/Ge : First Results



➤ Fit OFF spectra to ϕ_{OFF}

E/keV	ϕ_{OFF}	χ^2/dof
12-61	$\alpha_1 e^{-\alpha_2 E} + \alpha_3 + \alpha_4 E$	39/46

➤ Fit ON spectra to

$$\phi_{\text{OFF}} + \phi_{\text{SM}} + \mu_\nu^2 \phi_{\text{MM}} [10^{-10} \mu_B]$$

Best Fit :

$$\mu_\nu^2 = -1.7 \pm 2.5 \quad @ \quad \chi^2/\text{dof} = 55/49$$

[Comparable sensitivities to direct-search world limits]

➤ Improvements Expected :

- ↳ more data [$\times 3^+$ ON]
- ↳ complete & improved analysis
- ↳ treatment of threshold region [e.g. timing]

[Aim : ~ 5 keV]

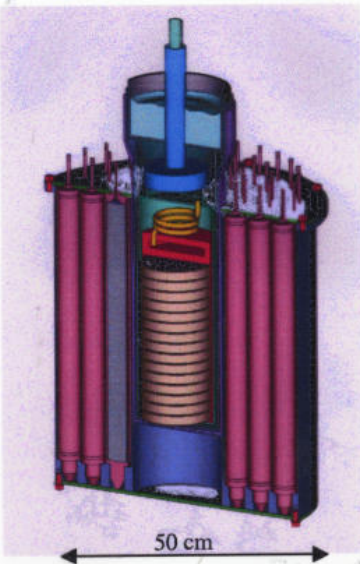
New Experiment

SEARCH FOR THE NEUTRINO MAGNETIC MOMENT

Emerging new Collaboration between:

Stanford + CWR University +... in USA and VNIIEF (Sarov) + ITEP (Moscow) + JINR (Dubna) in Russia

$\tilde{\nu}_e - e$ scattering



Conceptual layout of the ν - e scattering experiment with 40 MCi tritium source

- > New $\tilde{\nu}_e$ TRITIUM SOURCE of **40 MCi** activity (**4 kg ^3H**) with flux density of **$6 \times 10^{14} \text{ cm}^{-2} \text{ s}^{-1}$** (RFNC VNIIEF)
- > New **ULTRA-LOW-THRESHOLD DETECTORS** $E_{\text{th}} \sim 10 \text{ eV}$:
 - **SILICON CRYODETECTOR** ($T=10\text{mK}$) $15 \times 100 \text{ cm}^3$, $M=3\text{kg}$, *ionization-into-heat conversion effect (JINR-CWRU-Stanford)*
 - **HIGH-PURITY-GERMANIUM DETECTOR** $6 \times 150 \text{ cm}^3$, $M=4.8\text{kg}$, *internal proportional signal amplification by avalanche multiplication in the electric field (ITEP)*

SENSITIVITY (95% C.L.): $\mu_\nu \leq 2.5 \times 10^{-12} \mu_B$

GOALS

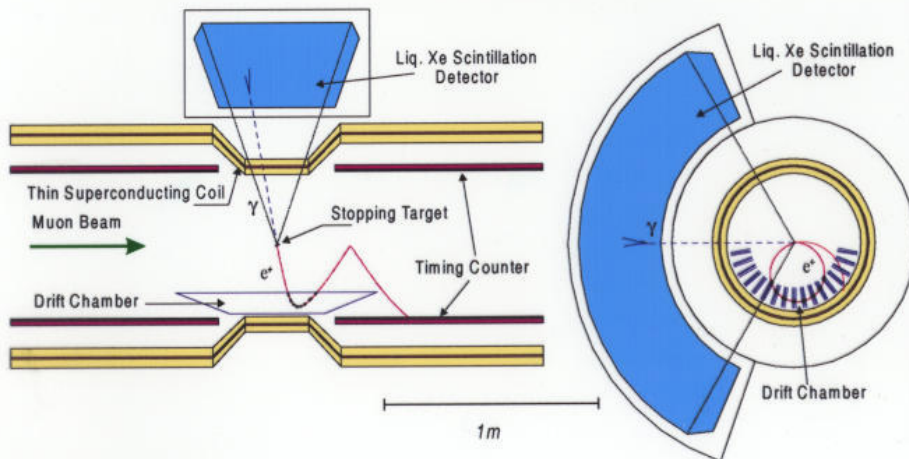
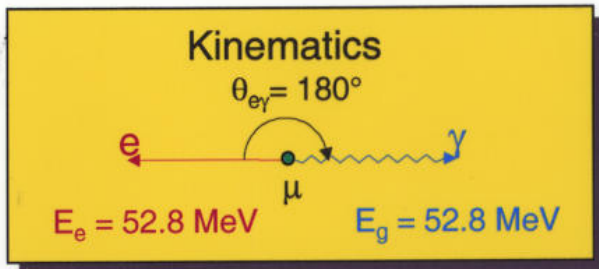
- Search for phenomena **beyond the Standard Model** $\mu_\nu \equiv 3(m_\nu/1\text{eV}) \times 10^{-19} \mu_B$
- **Increase the sensitivity** compared to previous experiments by **~ 100 times**
- **Reach the sensitivity of astrophysical/cosmological limits** $\mu_\nu \leq (1+3) \times 10^{-12} \mu_B$
- **Probe the solution to the Solar neutrino problem** (flux time-dependency) $\mu_\nu \geq 3 \times 10^{-12} \mu_B$



MEG Collaboration

Institute	Country	Main Resp.	Head	Scientists	Students
ICEPP, Univ. of Tokyo	Japan	LXe Calorimeter	T. Mori	12	2
Waseda University	Japan	Cryogenics	T. Doke	5	2
INFN, Pisa	Italy	e ⁺ counter, trigger, M.C.	C. Bemporad	4	3
IPNS, KEK, Tsukuba	Japan	Superconducting Solenoid	A. Maki	5	-
PSI	Switzerland	Drift Chamber, Beamline, DAQ	S. Ritt	4	-
BINP, Novosibirsk	Russia	LXe Tests and Purification	B. Khazin	4	-
Nagoya University	Japan	Cryogenics	K. Masuda	1	-

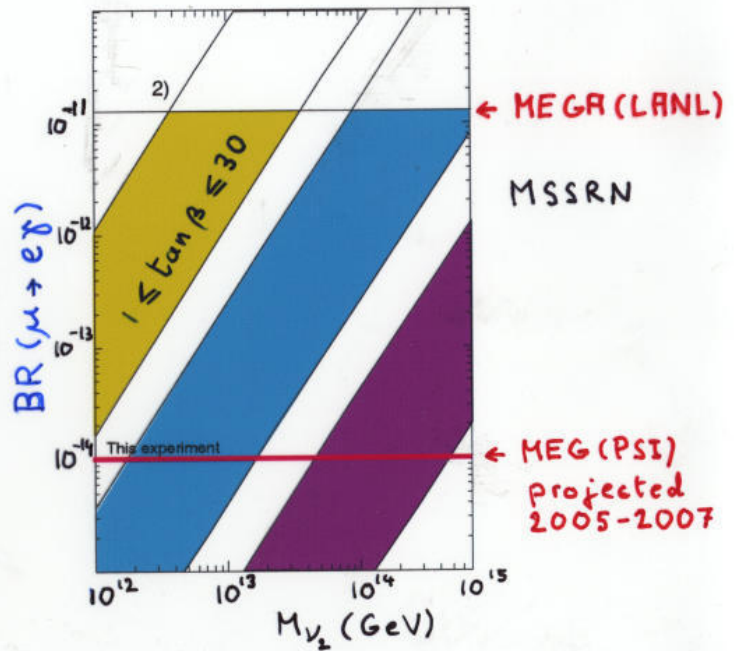
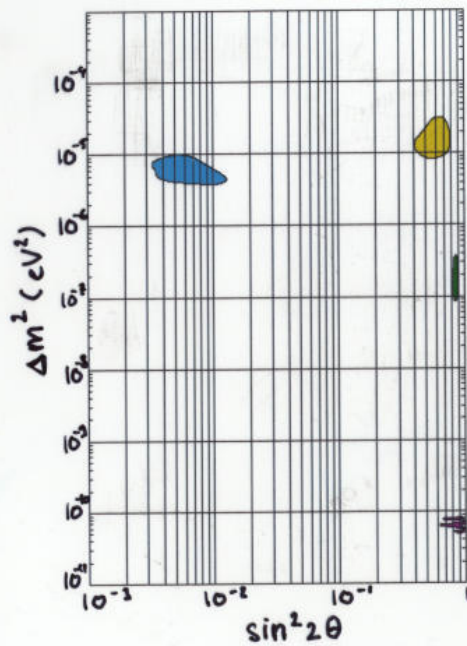
Experimental Method



- Stopped μ beam of 10^8 s^{-1} , 100% duty factor
- Liquid Xe calorimeter for γ detection
- Solenoidal magnetic spectrometer with gradient field
- Radial drift chambers for e^+ momentum determination
- Timing counter for e^+



$\mu \rightarrow e \gamma$: Connection with ν oscillations¹⁾



- 1) J. Hisano and D. Nomura, Phys. Rev. D59 (1999) 116005
- 2) MEGA collaboration, hep-ex/9905013



Sensitivity and Background Rate

N_μ	1×10^8		FWHM
T	$2.2 \times 10^7 \text{ s}$ (~50 weeks)	ΔE_e	0.7%
$\Omega/4\pi$	0.09	ΔE_γ	1.4%
ϵ_e	0.95	$\Delta\theta_{e\gamma}$	12 mrad
ϵ_γ	0.7	$\Delta t_{e\gamma}$	150 ps
ϵ_{sel}	0.8		

$$BR(\mu \rightarrow e\gamma) = (N_\mu \cdot T \cdot \Omega/4\pi \cdot \epsilon_e \cdot \epsilon_\gamma \cdot \epsilon_{sel})^{-1} = 0.94 \times 10^{-14}$$

Prompt Background $B_{pr} \cong 10^{-17}$

Accidental Background $B_{acc} \propto \Delta E_e \cdot \Delta t_{e\gamma} \cdot (\Delta E_\gamma)^2 \cdot (\Delta\theta_{e\gamma})^2 \rightarrow 5 \times 10^{-15}$

Conclusion

Constraints on magnetic moments from laboratory experiments are getting stronger.

MUNU (60 % analyzed):

$$\mu_e^{\text{reac}} < 1.3 \cdot 10^{-10} \mu_B \quad (90\% \text{ CL})$$

Texono (Kuo-Sheng) , ^3He coming,...

$\mu \rightarrow \gamma$ may provide relevant information on neutrino properties (SUSY)