

May 27, 2002
@NEUTRINO2002

Future Experiments with Super Neutrino Beams

Tsuyoshi NAKAYA
(Kyoto University)

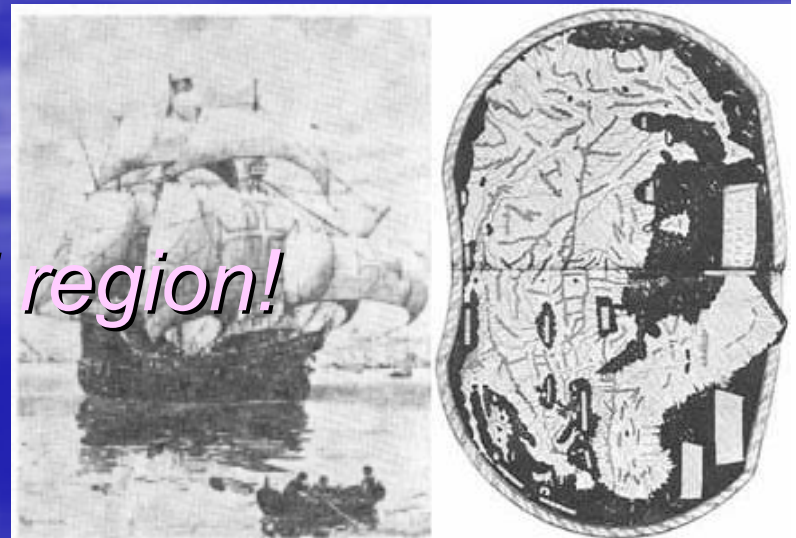
Outline

1. Introduction
2. Super Neutrino Beams
3. Experiments with a super neutrino beam
4. Physics sensitivity
5. Summary and Conclusion

1. Introduction

- A next goal of the neutrino experiments is to explore the neutrino oscillation phenomena in detail beyond the discovery phase.
 - Three generation Matrix (NMS matrix)
 - CP Violation, matter effect, the sign of Δm_{23}^2
 - May be an unexpected physics behind the oscillation phenomena.

A voyage to the unexplored region!

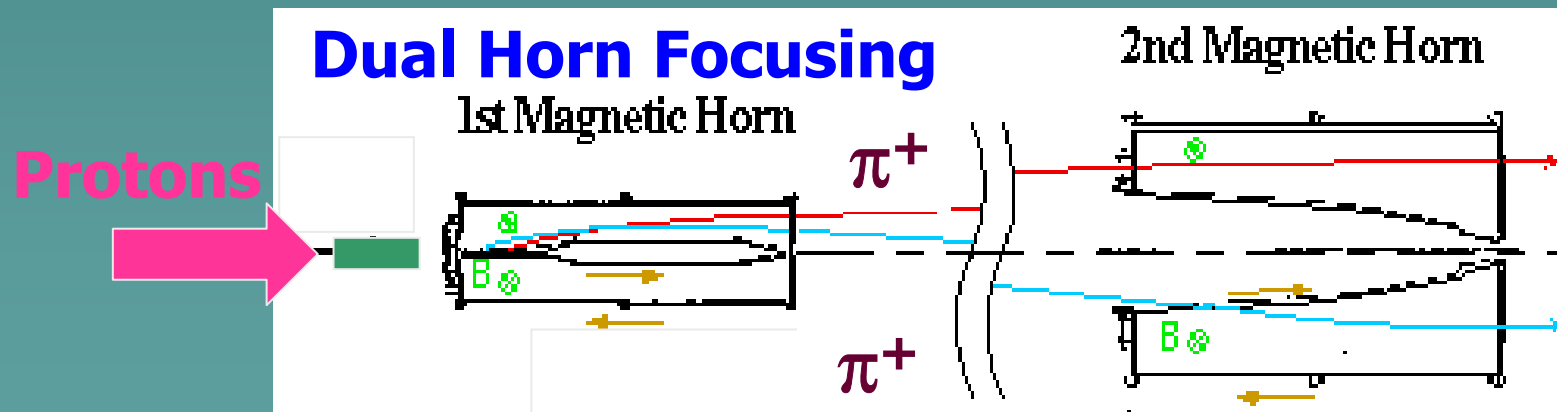


- Discovery and Confirmation.
 - Super-Kamiokande, SNO, K2K, MINOS, ICARUS,
 - These are good at discovering a surprise of the large effect (ν oscillation).
- , which should be followed by the more complete studies with high statistics,
 - more precision
 - θ_{23} , Δm^2_{23} , oscillation curve, non-oscillation scenario
 - sensitive to a rare process
 - θ_{13} ($\nu_{\mu} \rightarrow \nu_e$), CP Violation, unexpected phenomena.

to explore the underlying physics of ν oscillation.

2. Super Neutrino Beam

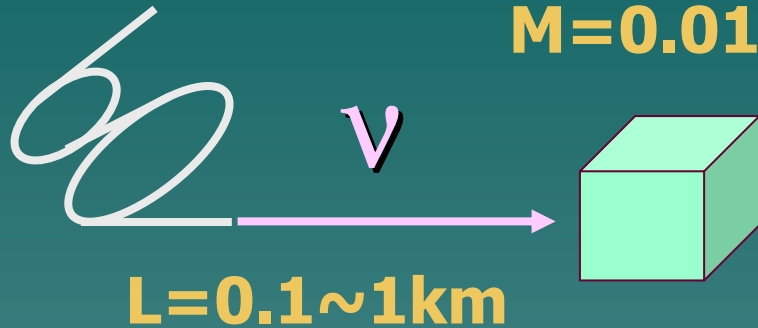
- ◆ A Super Neutrino Beam makes it possible to study ν oscillation phenomena with a great precision.
- ◆ What is the Super Neutrino Beam?
 - No Clear definition, but it is a very intense neutrino beam produced by a high power ($\sim 1\text{MW}$ or more) accelerator.
 - ◆ A conventional method.
 - ◆ Still technically challenging due to the high power and the high radiation environment, but not impossible.



- Past Accelerator Neutrino Experiment

P=0.01-0.1MW

M=0.01~1kton

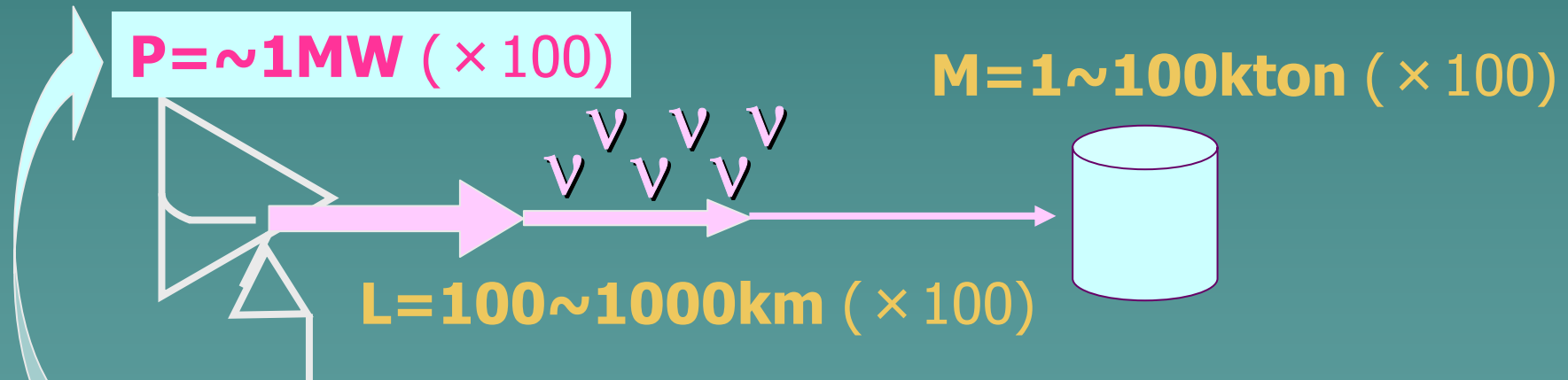


$$\# \nu \propto P \times M / L^2$$

- Accelerator Long Baseline Neutrino Experiment

P=~1MW (× 100)

M=1~100kton (× 100)



Need High Intensity accelerator ⇒ Super ν Beam

NOTE: The detector size is equally important.

(High Intensity) Proton Accelerators

| | Power (MW) | Energy (GeV) | Intensity (10^{12} ppp) | Rep. rate (Hz) |
|---------------------------------|---------------|-----------------|-------------------------------|-------------------|
| KEK-PS | 0.005 | 12 | 6 | 0.45 |
| AGS | 0.14 | 24 | 60 | 0.6 |
| FNAL-MI | 0.41 | 120 | 40 | 0.53 |
| SPS | 0.3 | 400 | 35 | 0.16 |
| JHF-I | 0.77 | 50 | 330 | 0.29 |
| Super-AGS | 1.3 | 28 | 120 | 2.5 |
| FNAL-proton driver-I | 1.2 | 16 | 30 | 15 |
| SPL | 4 | 2.2 | 230 | 50 |
| JHF-II | 4 | 50 | | |

■ Not the construction stage yet, but R&D stage.

(Super) Neutrino Beams

| | $\langle E_\nu \rangle$ (GeV) | L (km) | #CC ν /kt/yr | L/L _{osci.*} | f(ν_e) @peak |
|----------------|----------------------------------|-----------|---------------------|-----------------------|-----------------------|
| K2K | 1.3 | 250 | 2 | 0.47 | ~1% |
| NuMi (High E) | 15 | 730 | 3100 | 0.12 | 0.6% |
| NuMi (Low E) | 3.5 | 730 | 469 | 0.51 | 1.2% |
| CNGS | 17.7 | 732 | 2448 | 0.10 | 0.8% |
| JHF-I | 0.7 | 295 | 133 | 1.02 | 0.2% |
| Numi off-axis | 2.0 | 730 | ~80 | 0.89 | 0.5% |
| Super AGS | 1.5 | 2540 | 11 | 4.1 | 0.5% |
| JHF-II | 0.7 | 295 | 691 | 1.02 | 0.2% |
| SPL | 0.26 | 130 | 16.3 | 1.21 | 0.4% |
| β beam** | 0.58 | 130 | 84 | 0.54 | ----- |

(*) $L_{\text{osci.}} = \frac{\pi}{2} \cdot \frac{\langle E_\nu \rangle}{1.27 \Delta m_{23}^2} \text{ w/ } \Delta m_{23}^2 = 3 \times 10^{-3} \text{eV}^2$ (**) $\gamma=150, {}^6\text{He} (\bar{\nu}_e)$

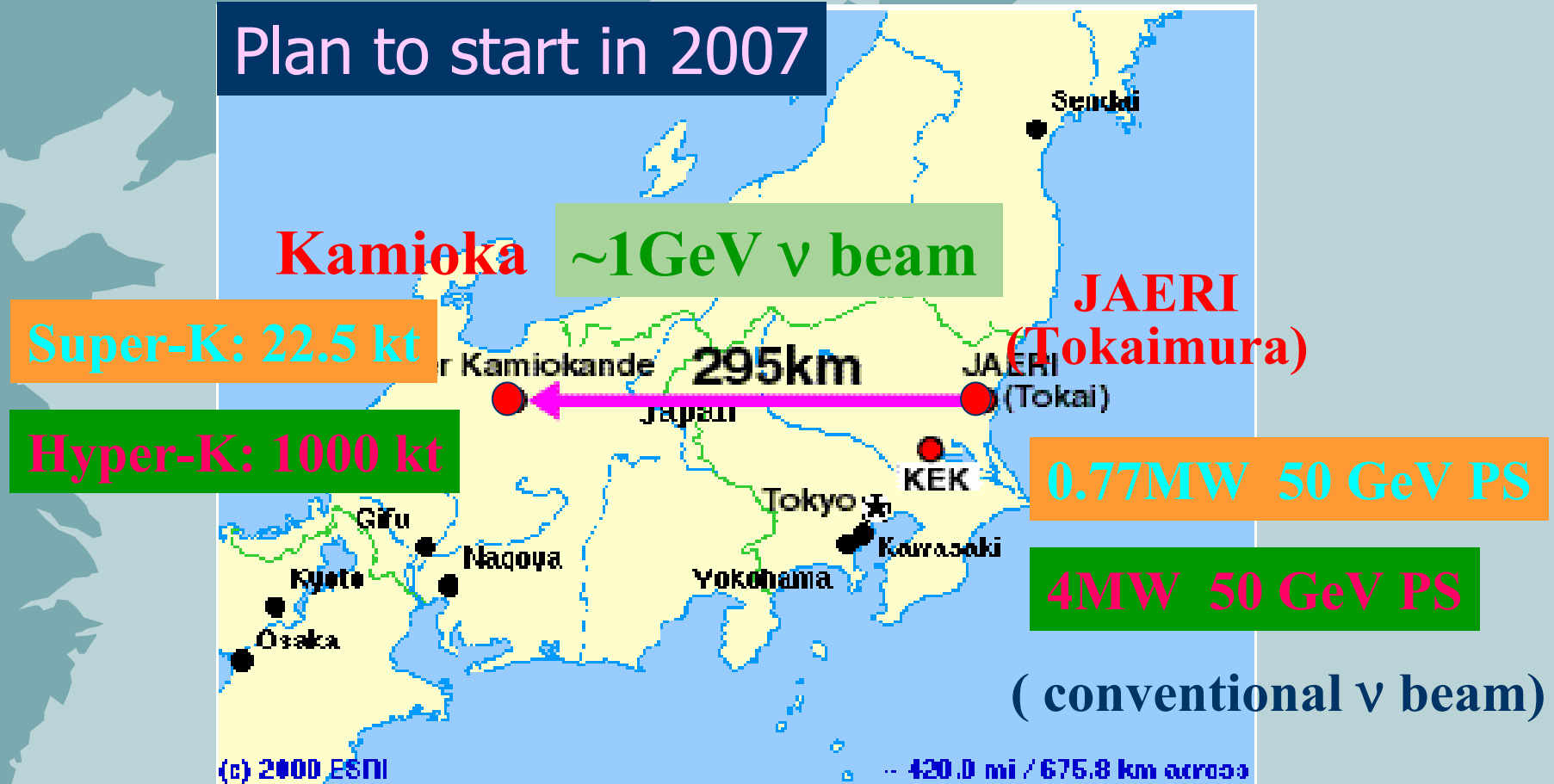
3. Experiments with the super neutrino beam

1. **Japan:** JHF-Kamioka Neutrino Project
2. **USA:** FNAL Super-NUMI w/ the proton driver upgrade and the BNL Super-AGS
3. **Europe:** SPL (or CNGS off-axis) + β beam

3.1 JHF-Kamioka Neutrino Project

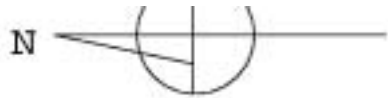
(hep-ex/0106019)

Plan to start in 2007



Phase-I (0.77MW + Super-Kamiokande)

Phase-II (4MW+Hyper-K) ~ Phase-I × 200



JHF Facility

Pacific Ocean

**JAERI@Tokai-mura
(60km N.E. of KEK)**

**Super Conducting
magnet for ν beam line
50GeV PS (0.77MW)**

**Construction
2001 ~ 2006
(approved)**

3GeV PS

50GeV PS (0.77MW)

400MeV LINAC

Neutrino facility

**Near ν detectors
@280m and
@~2km**

**Budget Request of the
 ν beam line this year**

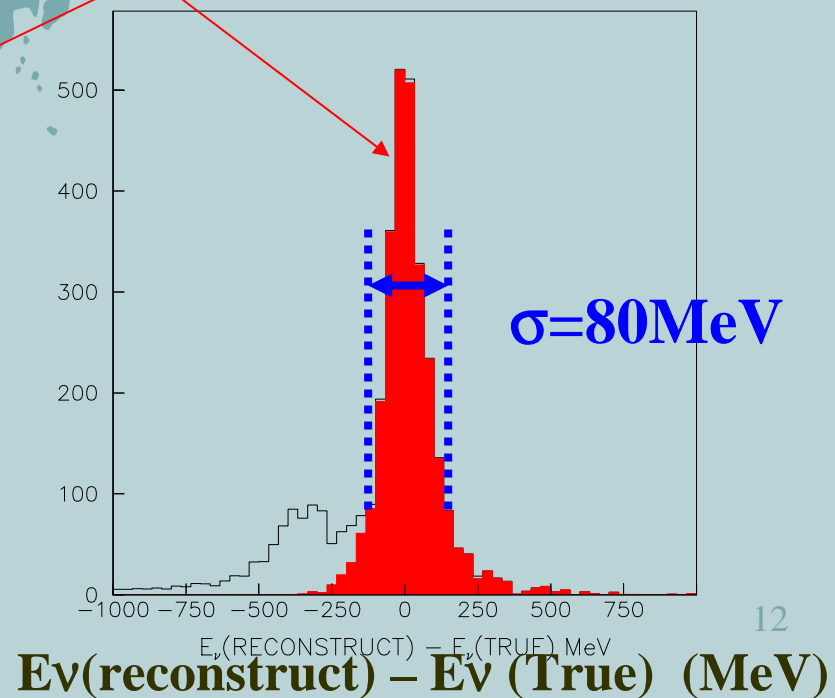
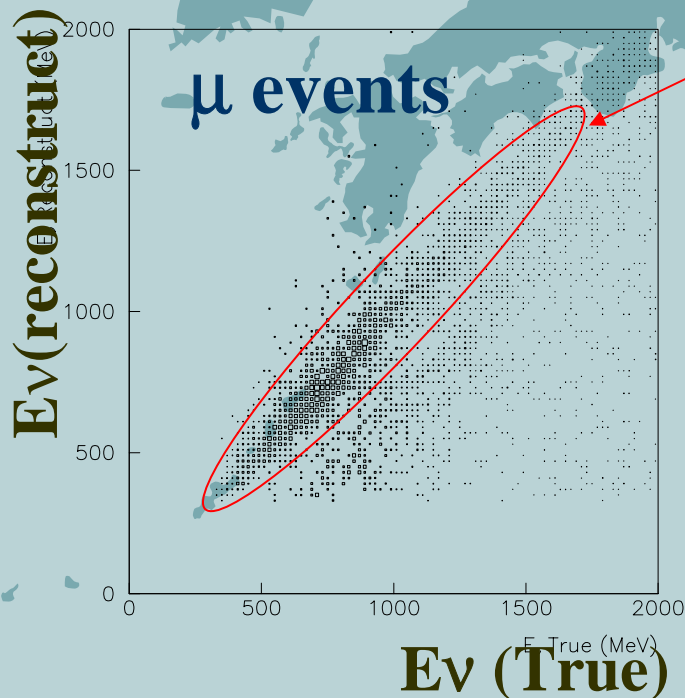
10^{21} POT(130day) \equiv "1 year"



ν beam at JHF

- Principle

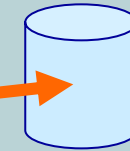
- Intense Narrow Band Beam by “off-axis”.
- Beam energy is at the oscillation maximum.
 - High sensitivity, less background
- ~ 1 GeV ν beam for Quasi-elastic interaction.



Off Axis Beam

(ref.: BNL-E889 proposal: <http://minos.phy.bnl.gov/nwg/papers/E889>)

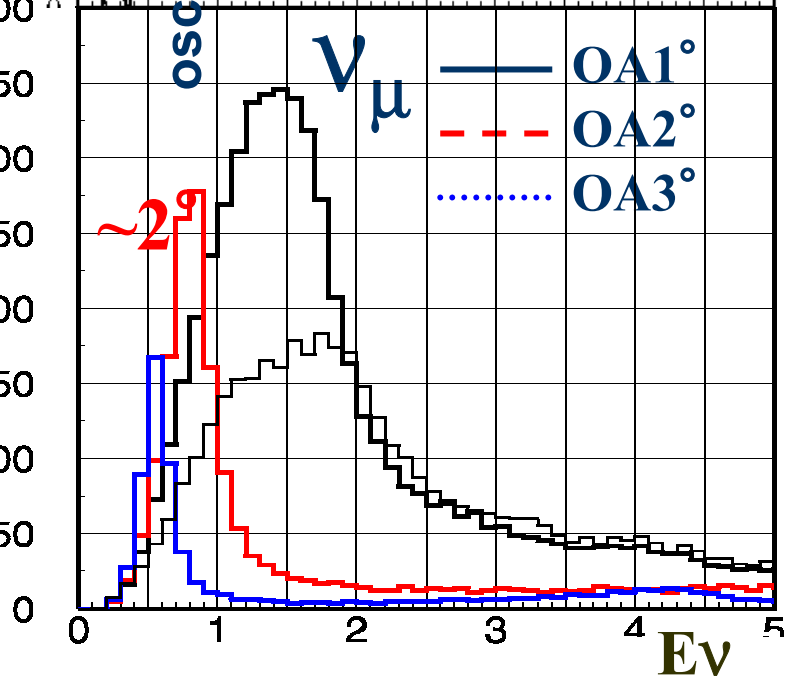
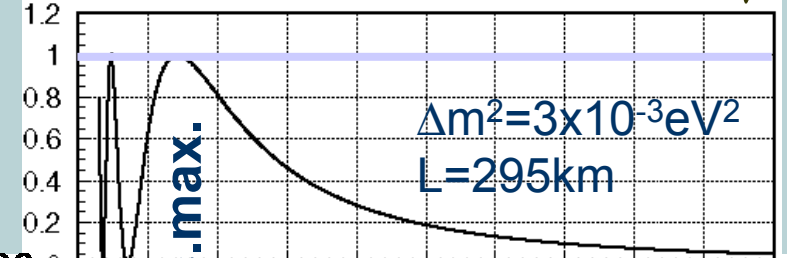
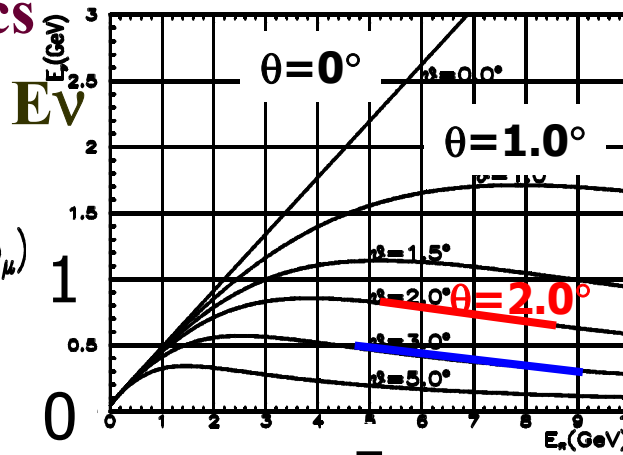
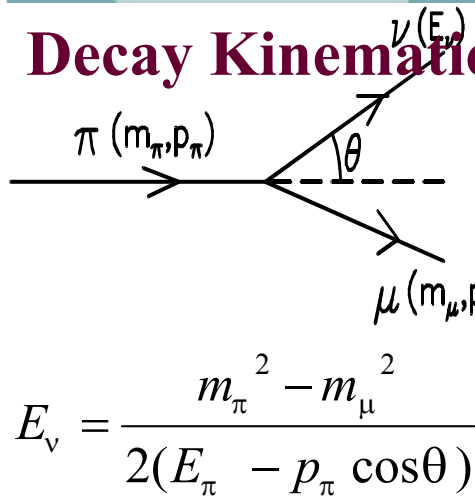
Far Det.



WBB w/ intentionally misaligned beam line from det. axis

$$\text{Osc. Prob.} = \sin^2(1.27 \Delta m^2 L / E_\nu)$$

Decay Kinematics

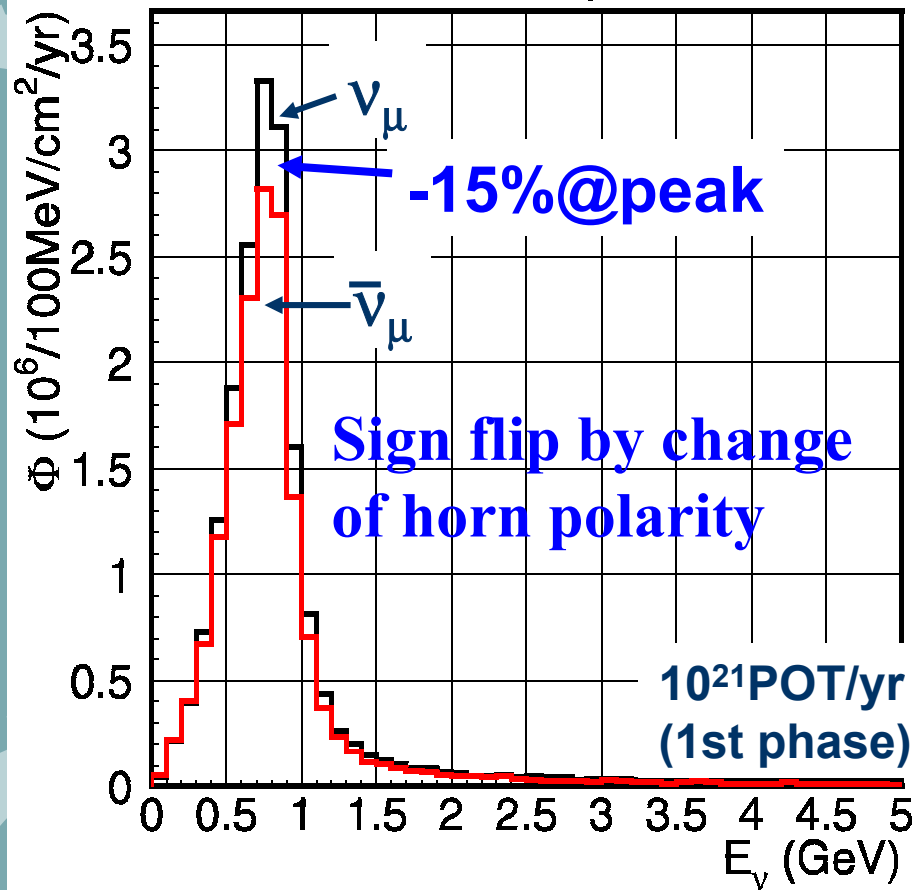


~3000 CC int./22.5kt/yr

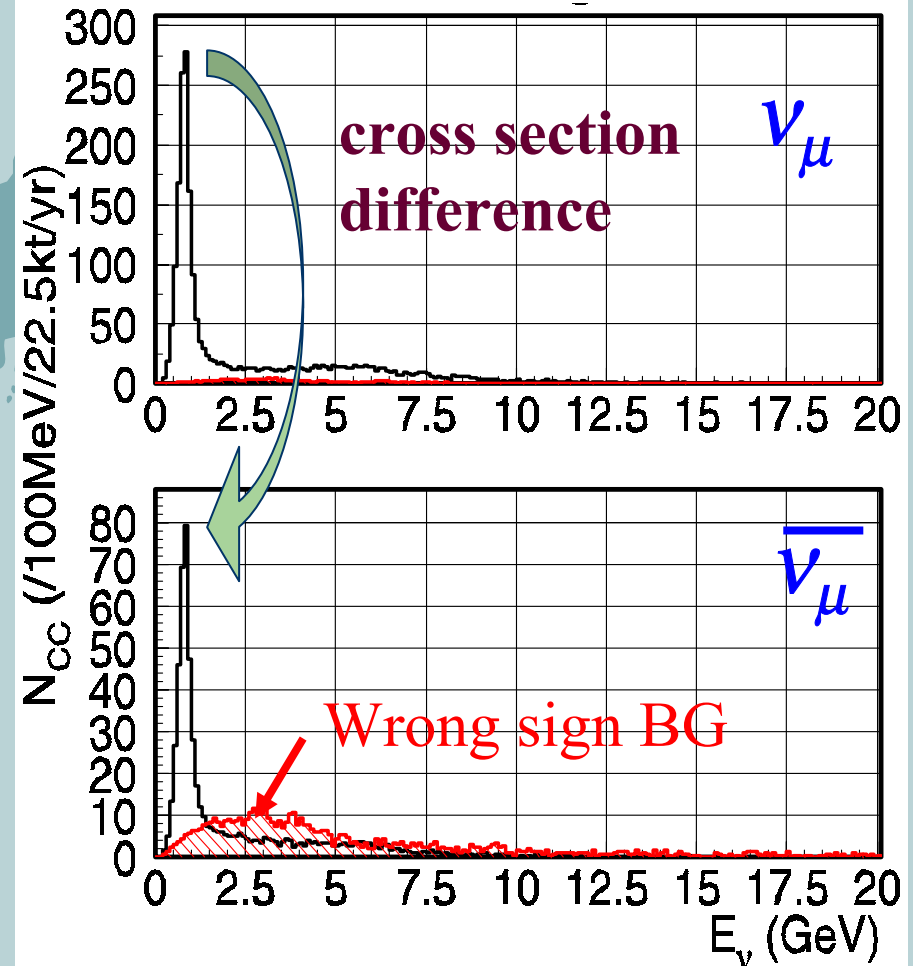
ν_e : 1.0% (0.2% @ peak);

$\bar{\nu}_\mu/\nu_\mu$ flux for CP violation search.

Flux

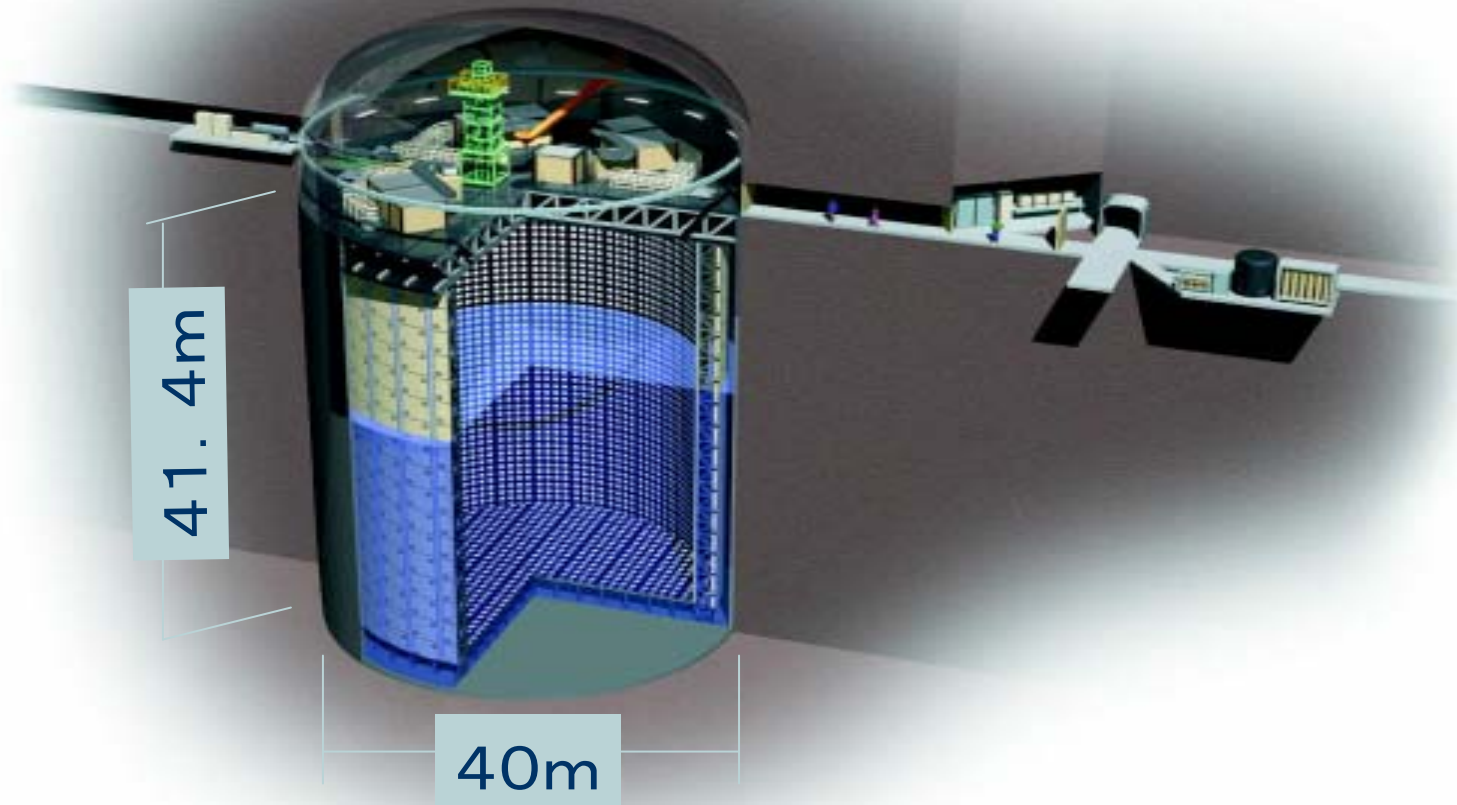


CC interaction



Super-Kamiokande

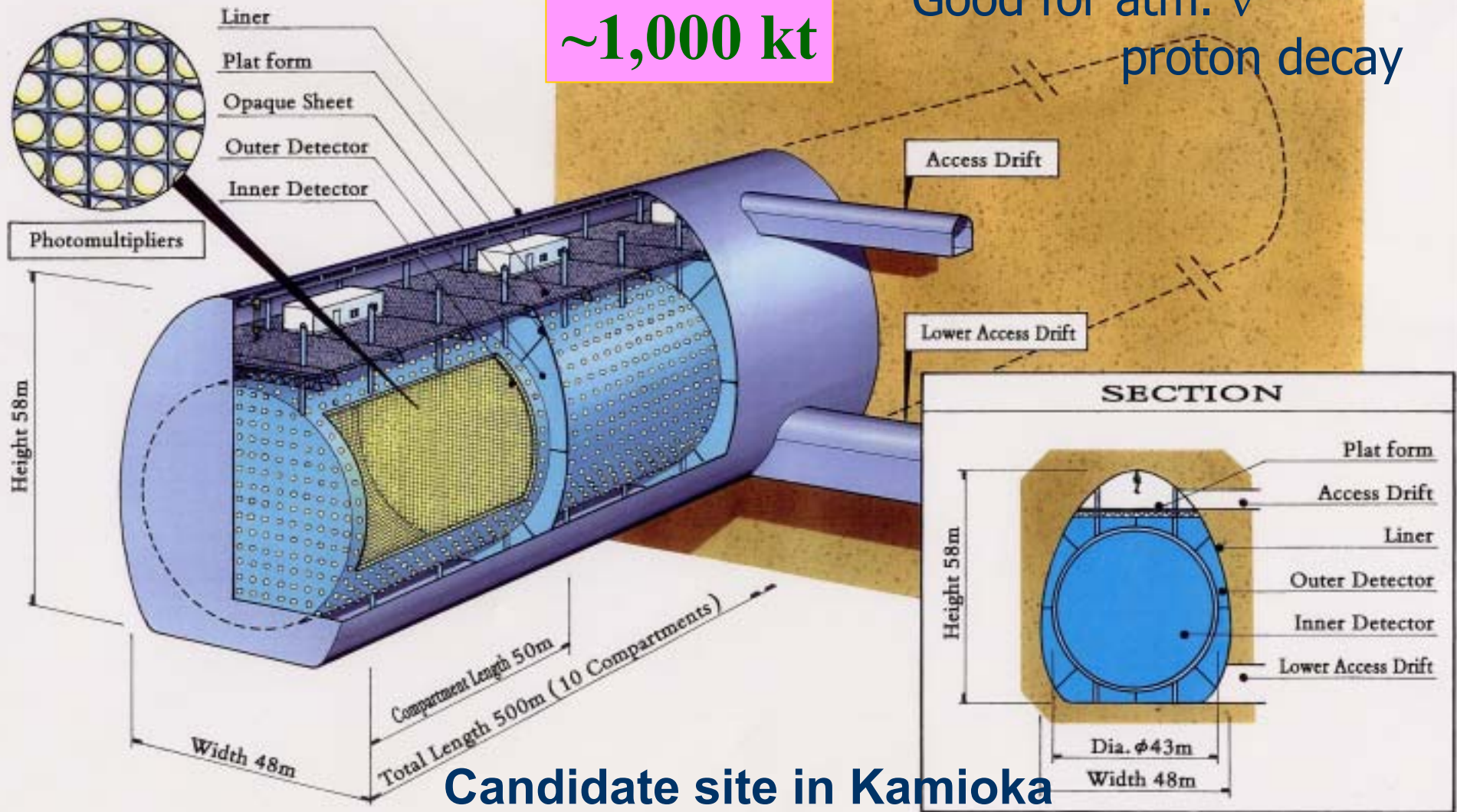
50,000 ton water Cherenkov detector
(22.5 kton fiducial volume)



Hyper-Kamiokande (a far detector in the 2nd phase)

~1,000 kt

Good for atm. ν
proton decay



3.2 USA: FNAL and BNL plan

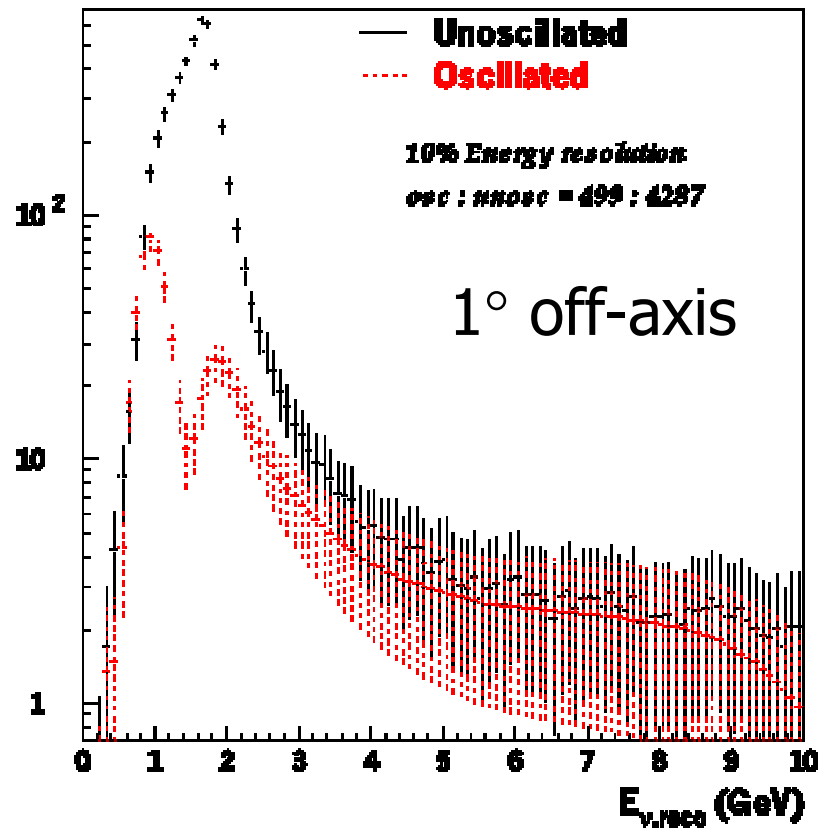
- BNL: Super AGS (1.3MW, LOI submitted)
- FNAL: Super NUMI (1.6 MW) or the new proton driver. (hep-ex/0205040,0204037,hep-ph/0204208)



FNAL, BNL to Soudan

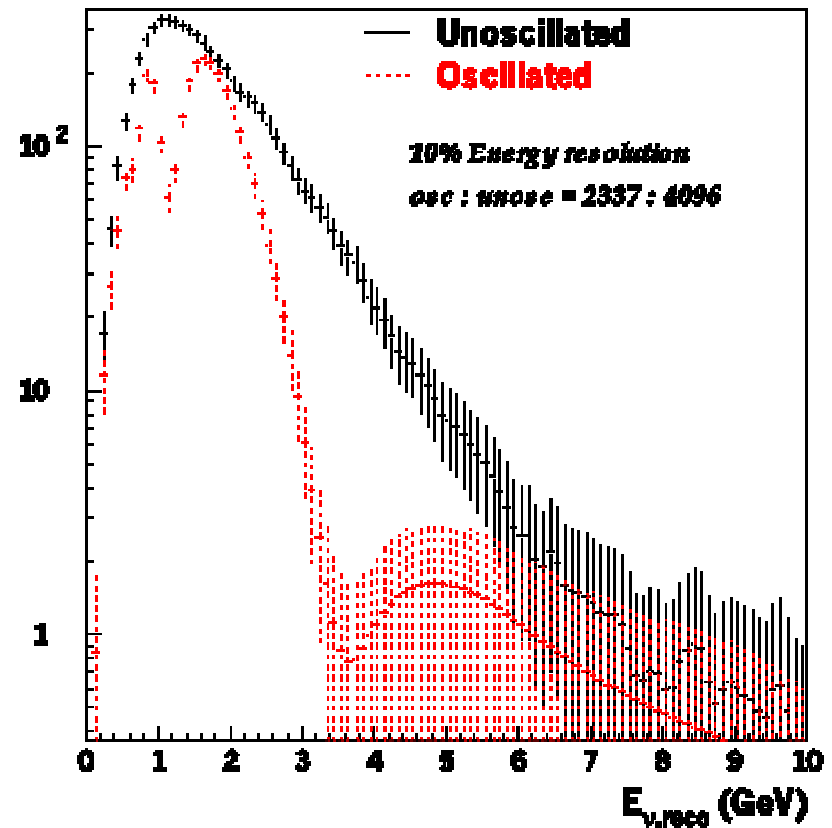
Events/0.1GeV/5years/10kTon

FNAL-Soudan: 735 km



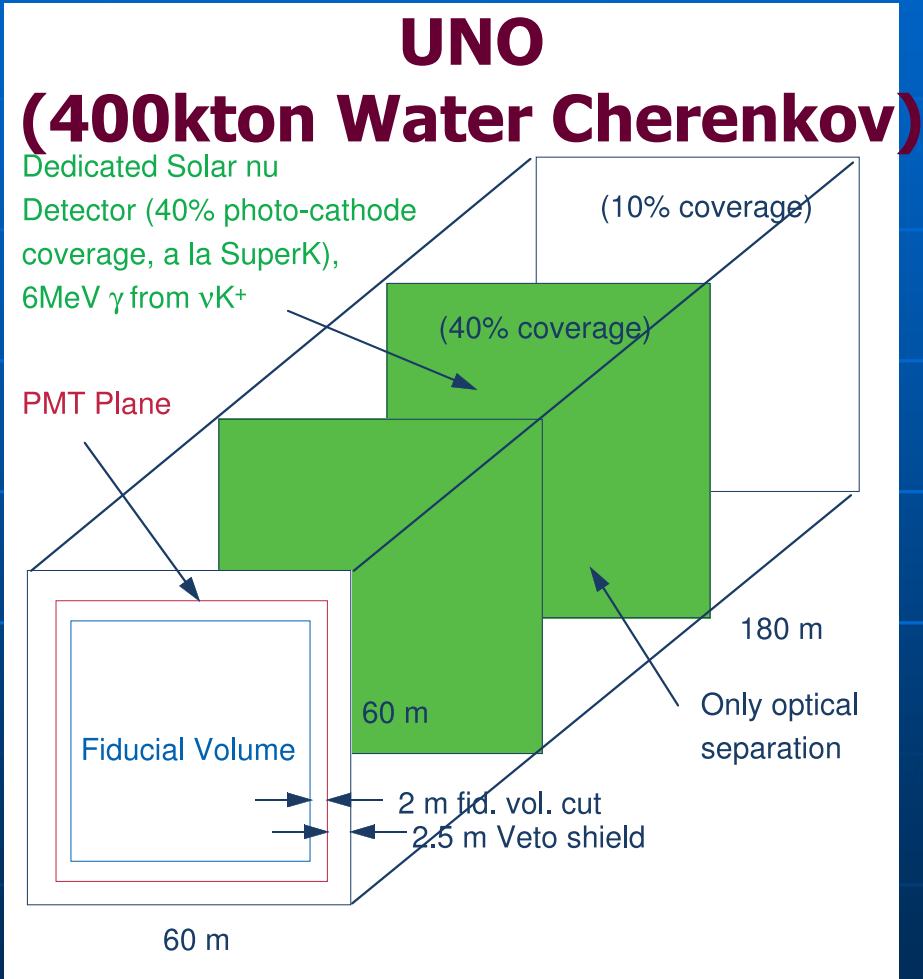
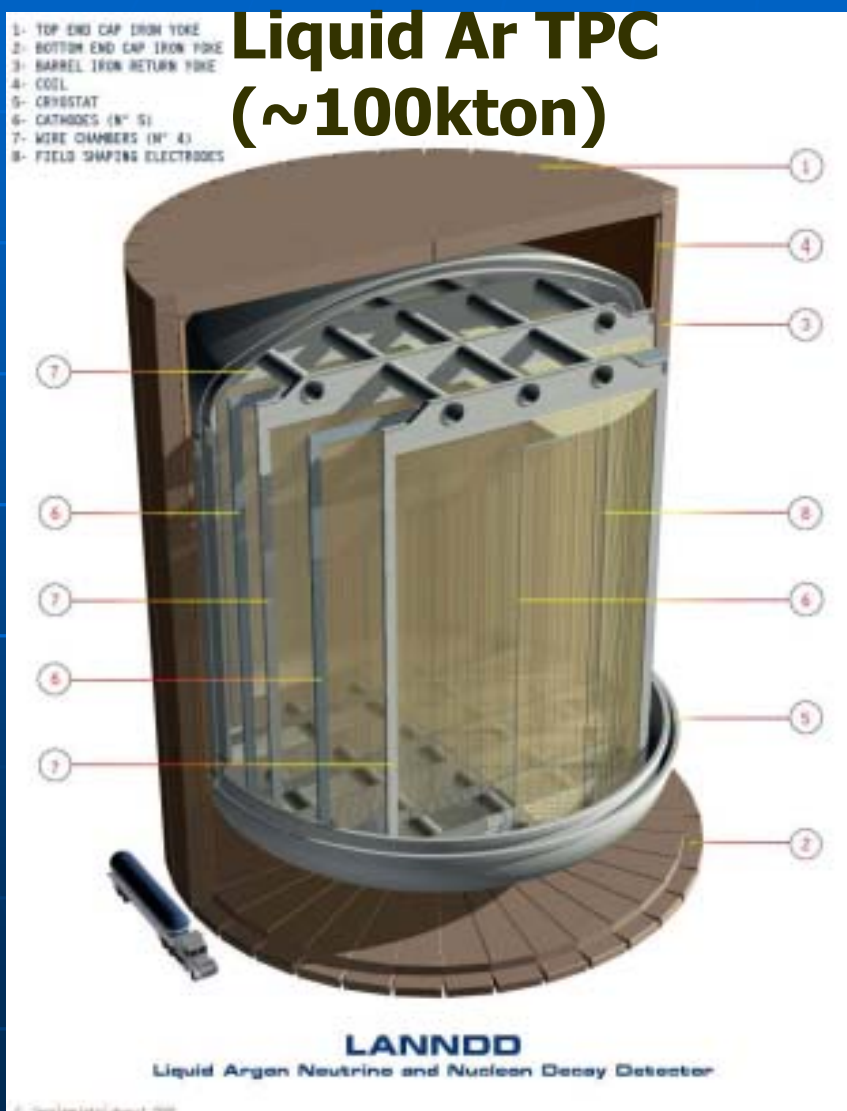
Events/0.1GeV/5years/100kTon

BNL-Soudan: 1711 km



- Water Cherenkov like Super-K

Detectors

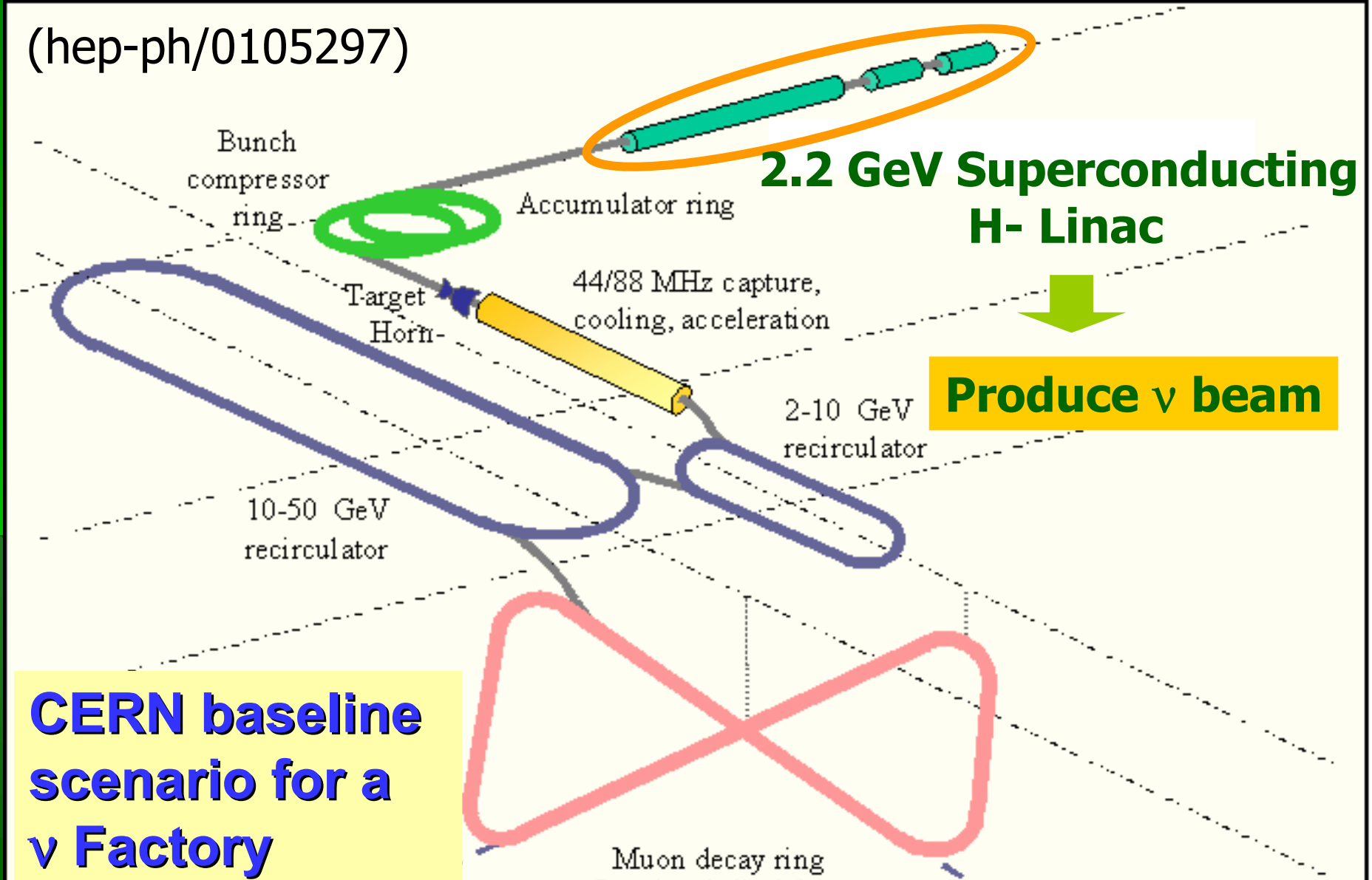


USA beams and detectors

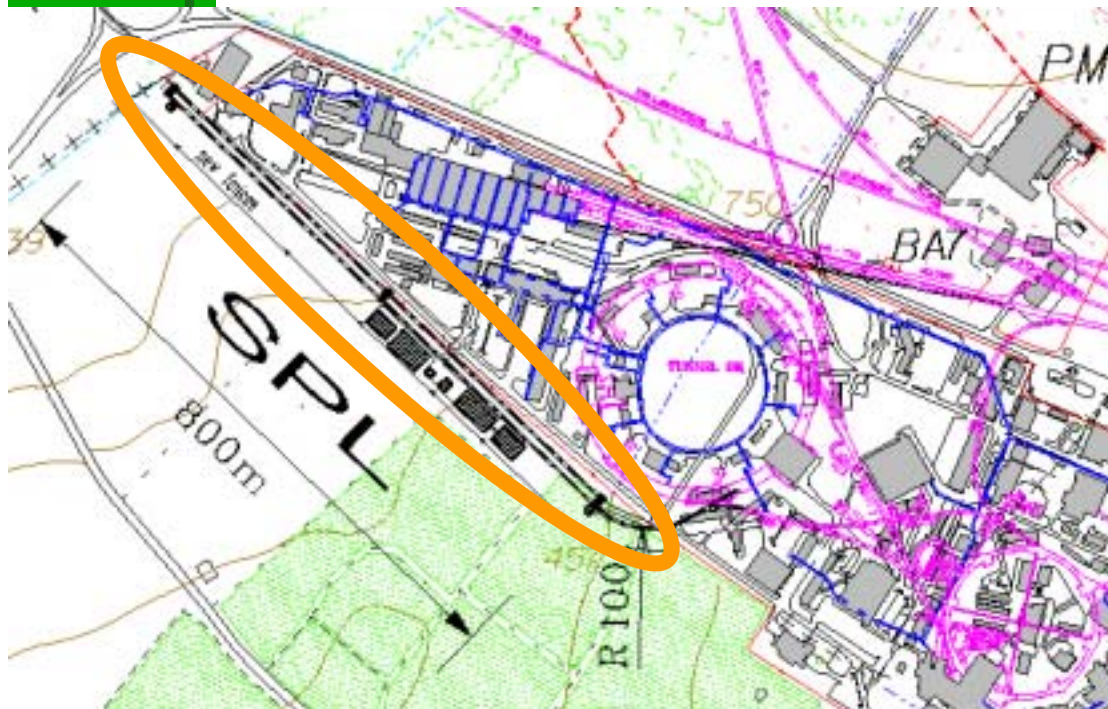
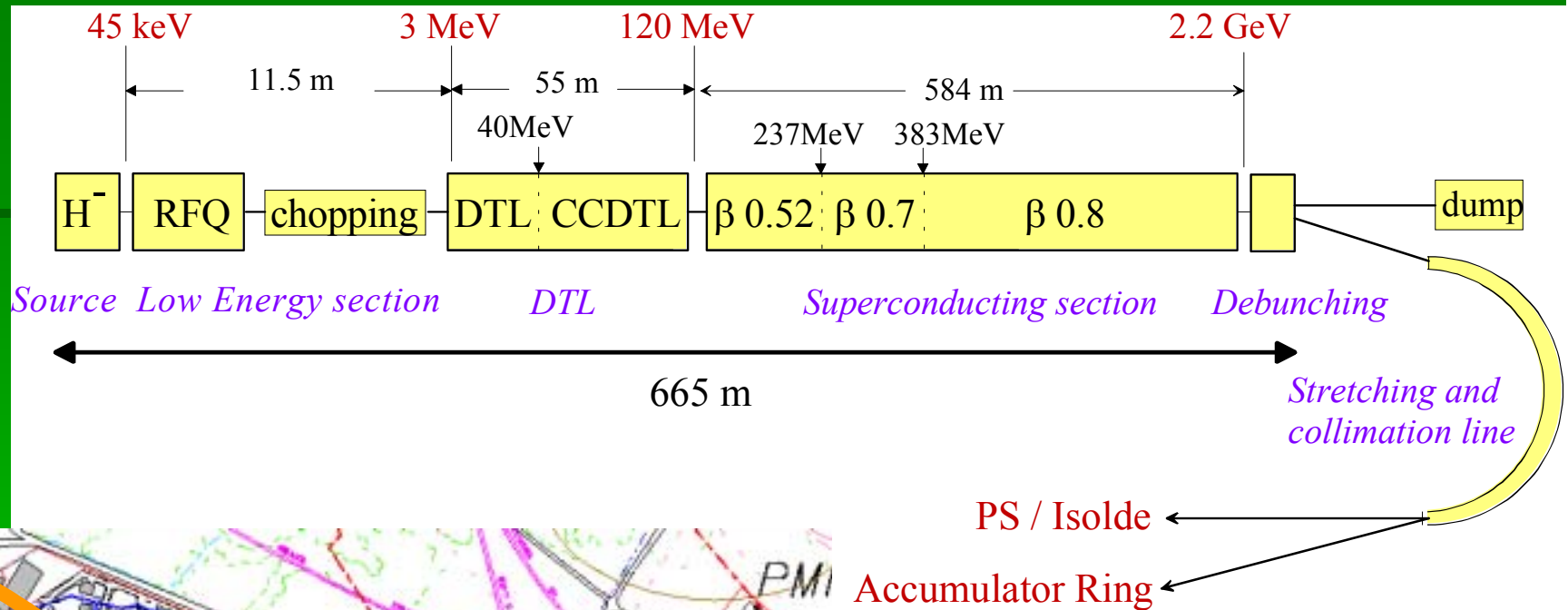
- US neutrino physicists are investigating several combinations of detectors, beams, baseline, etc. including ν -factory.
- BNL LOI is proposed to send ν beam to Homestake at the distance of $\sim 3000\text{km}$ to study matter effect and the sign of Δm_{23}^2 .
This is a unique feature of this project, which is not considered in JHF-Kamioka and CERN-SPL.
- My talk does not cover NUMI off-axis which will be presented in the later talk.

3.3 CERN: SPL + β beam

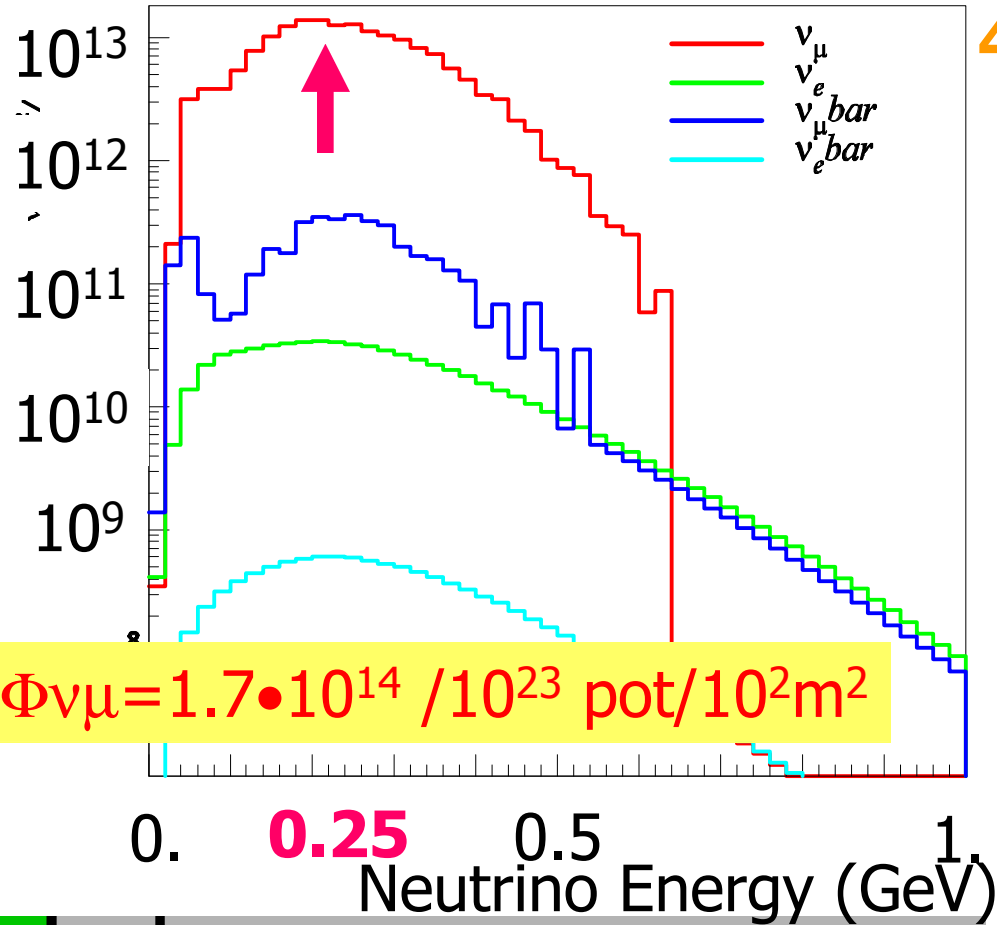
(hep-ph/0105297)



SPL (Super Proton Linac)

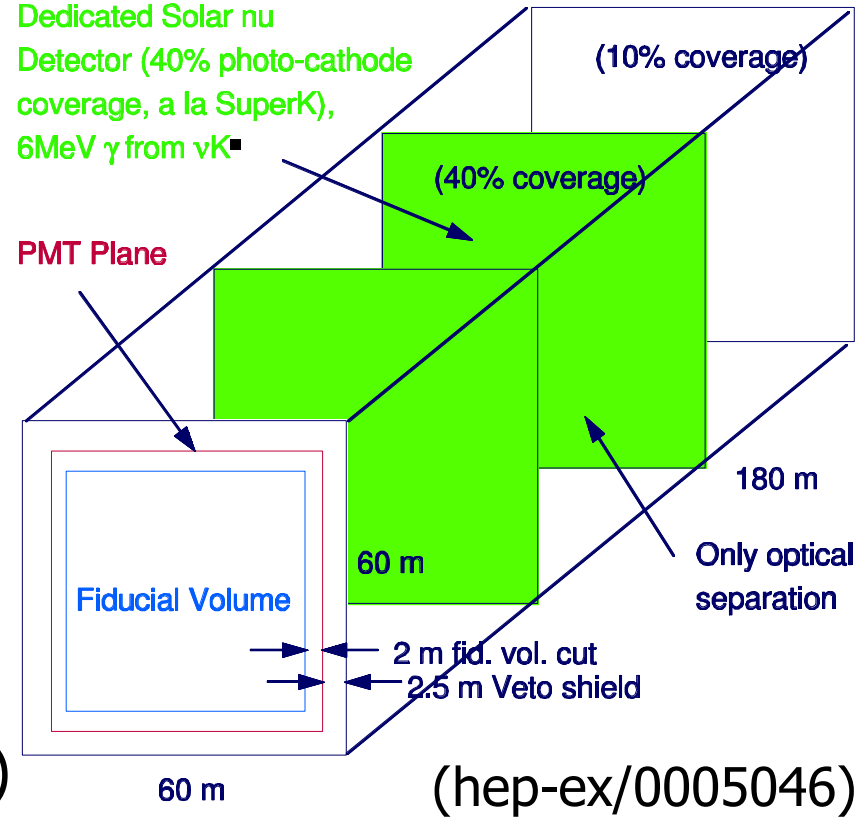


CERN SPL to UNO at Frejus



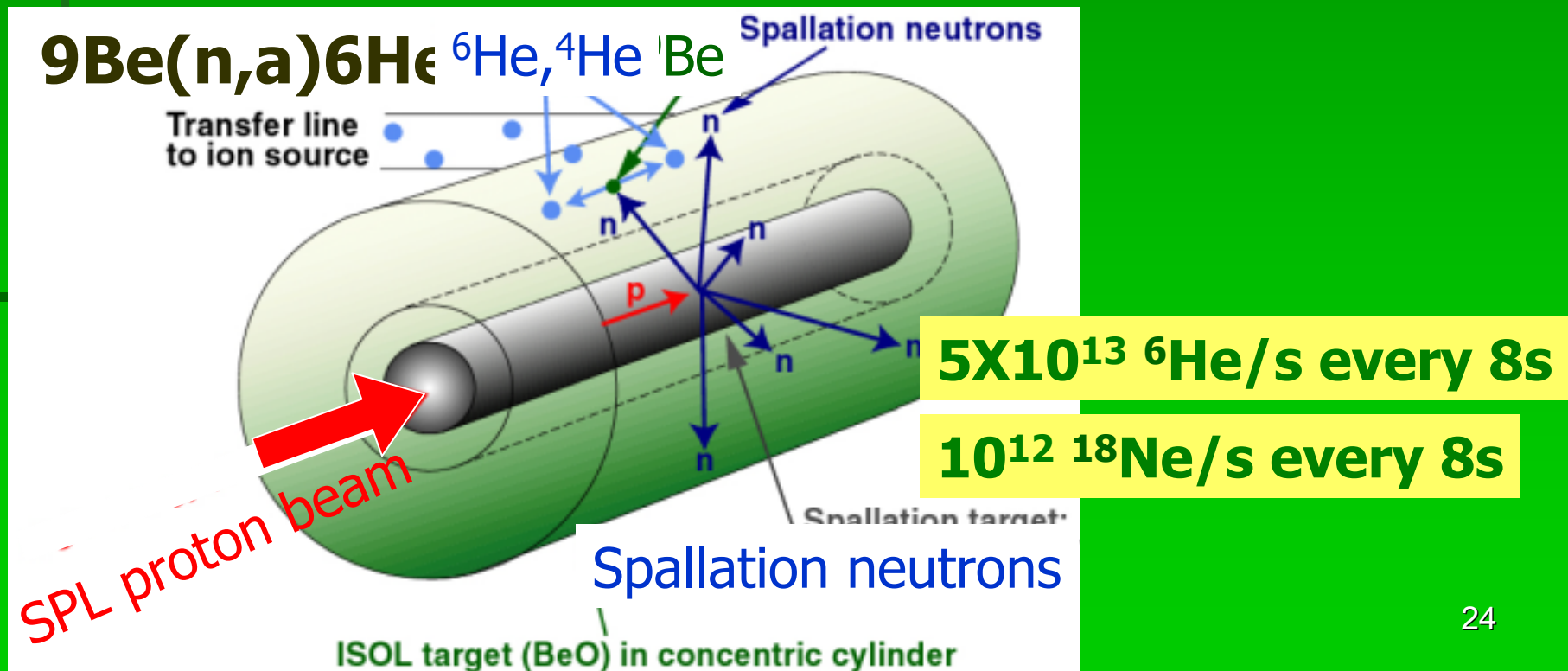
400kton Water Cherenkov

Dedicated Solar nu
 Detector (40% photo-cathode
 coverage, a la SuperK),
 6MeV γ from νK^{\pm}

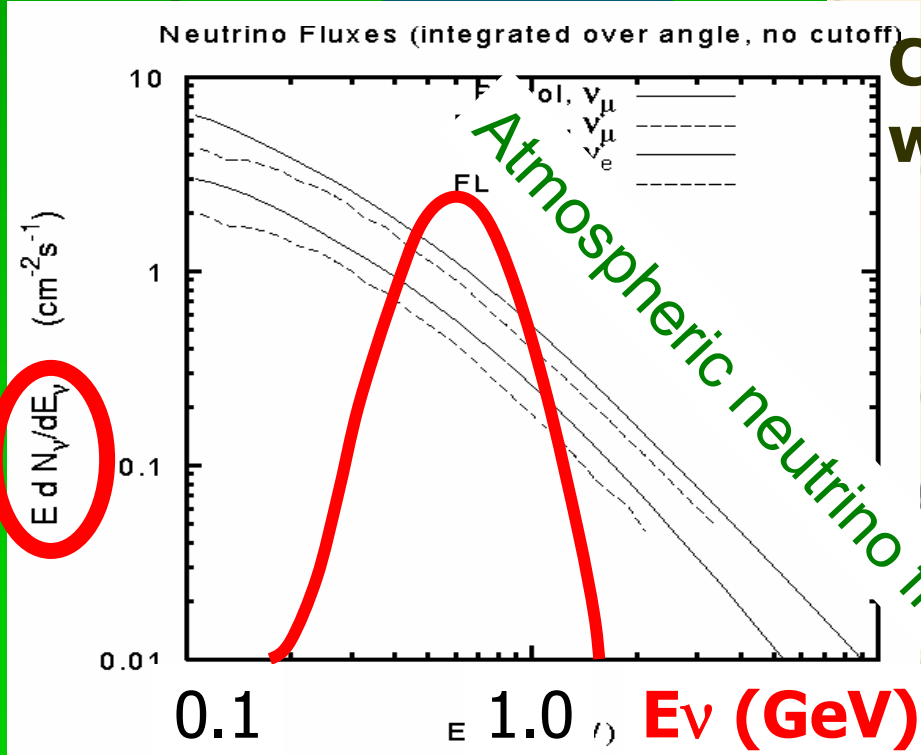
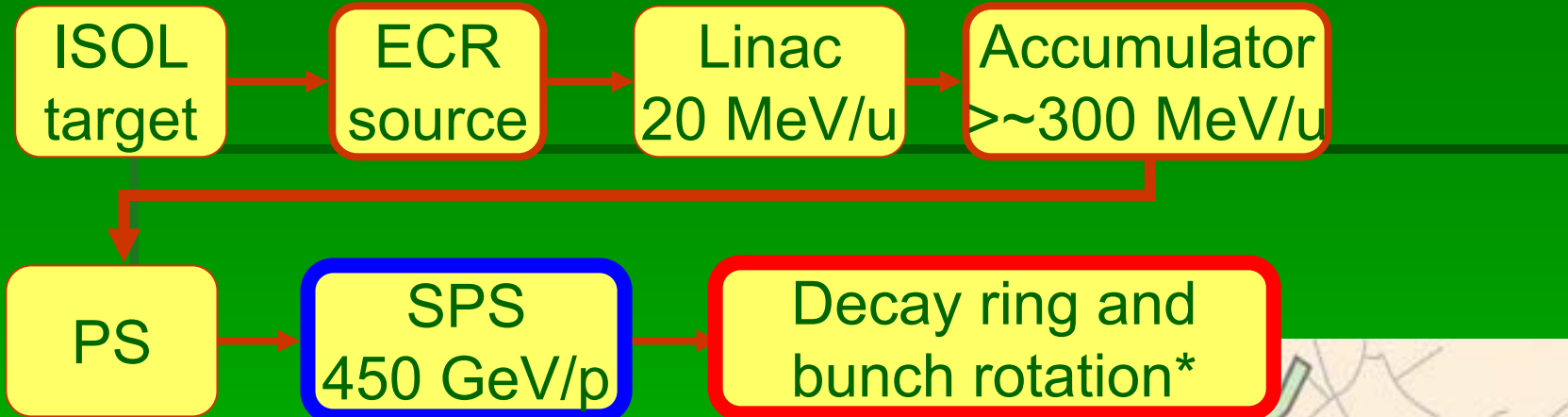


SPL β beam (ν_e and $\bar{\nu}_e$)

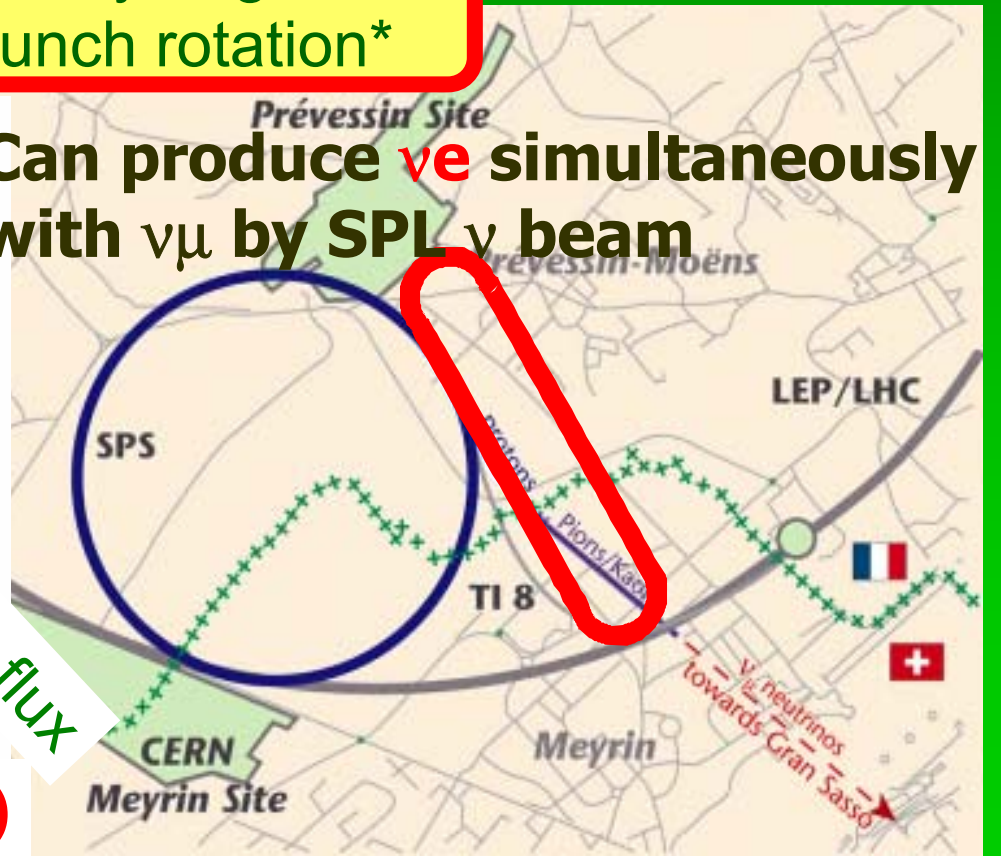
- A novel idea to generate a neutrino beam from accelerated radioactive ions.
 - Low energy ($\sim 2\text{MeV}$) β ray in the CM frame, which can produce **the forward focused ν beam**.



SPL β beam to Frejus



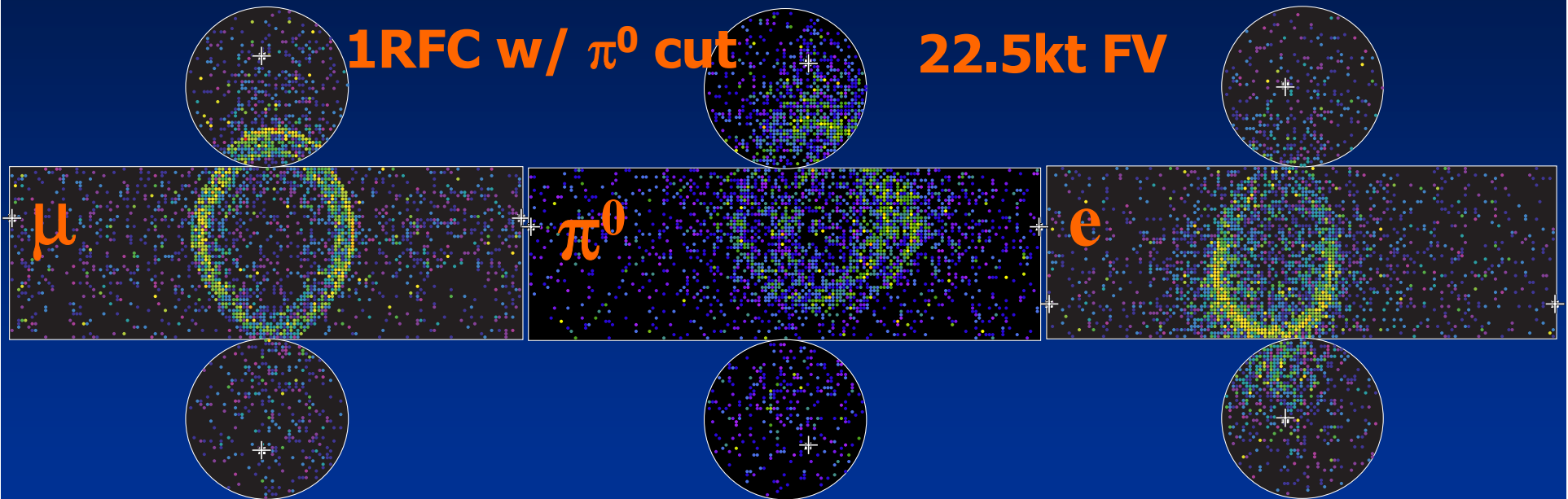
Can produce ν_e simultaneously with ν_{μ} by SPL γ beam



4. Physics Sensitivity

- Physics goals of most super ν -beam experiments are:
 - Discovery of $\nu_{\mu} \rightarrow \nu_e$ and the measurement of θ_{13}
 - CP Violation ($\nu_{\mu} \rightarrow \nu_e$ vs $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$)
 - Sign of Δm^2 by using matter effect.
 - Precision measurement of ν oscillation parameters.

ν_e appearance in JHF-Kamioka (phase 1)



Backgrounds

1.8 events

9.3(*) events

11.1 events

Signal

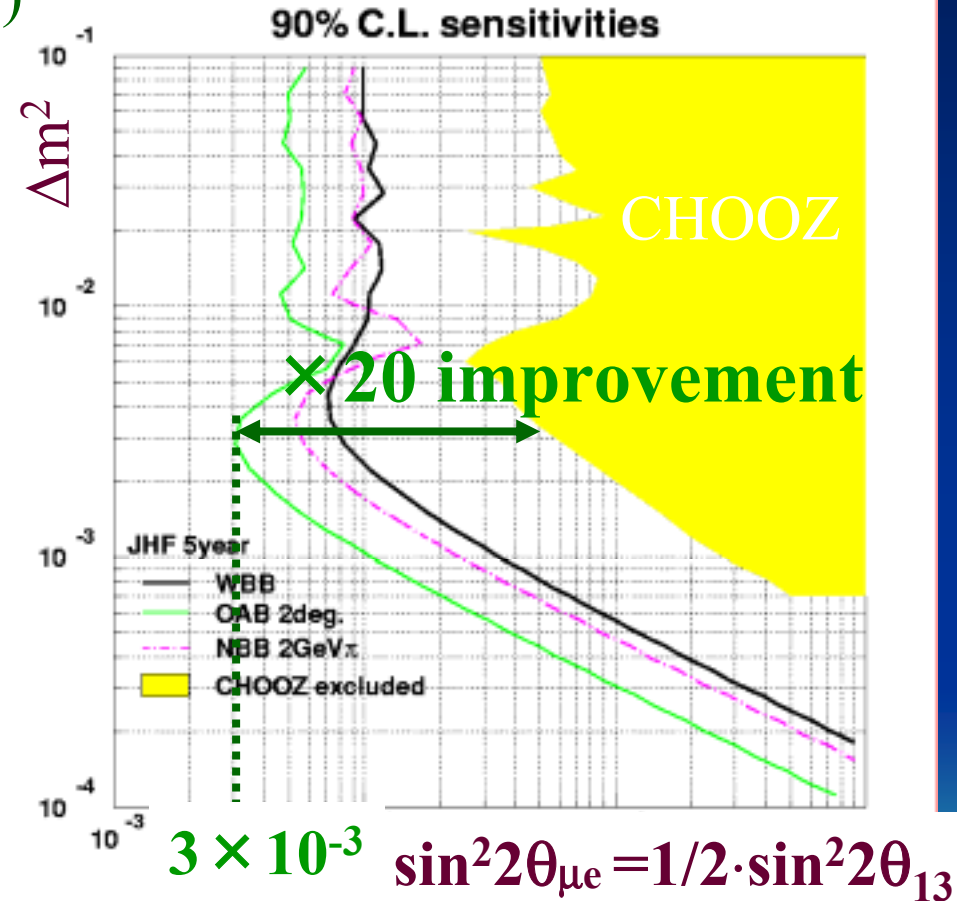
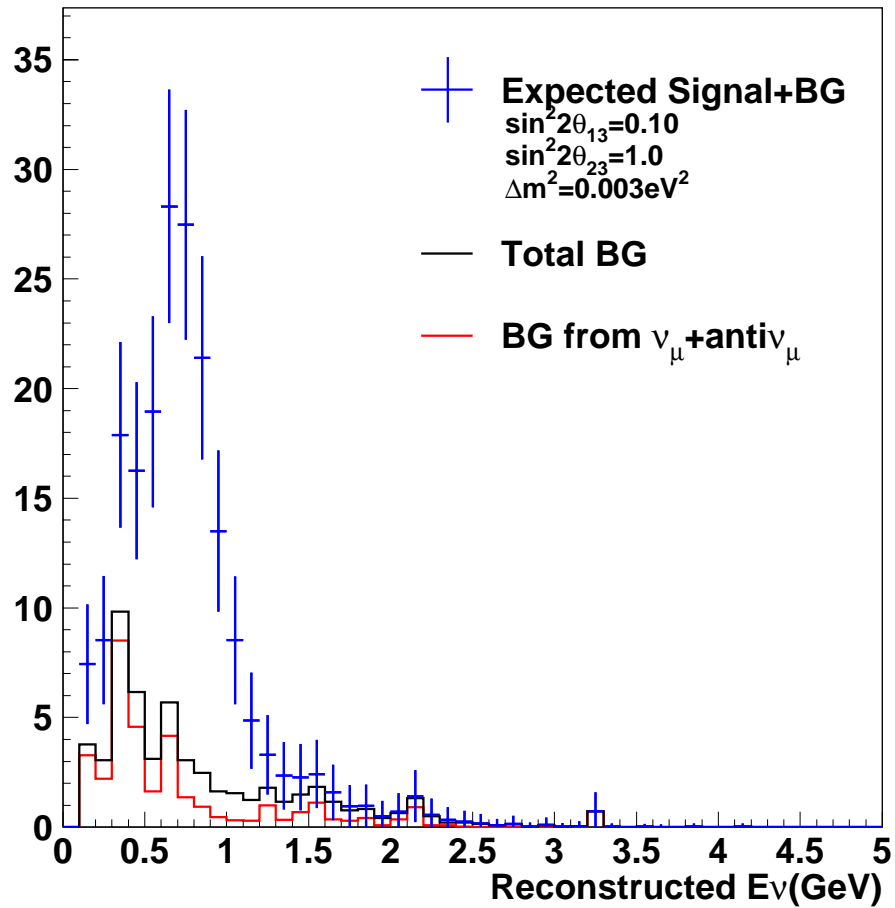
123.2 events @ $\sin^2 2\theta_{13}=0.1$, $\Delta m^2=3 \times 10^{-3} \text{eV}^2$

(*) still can be further improved

(5 years running)

ν_e appearance (continue)

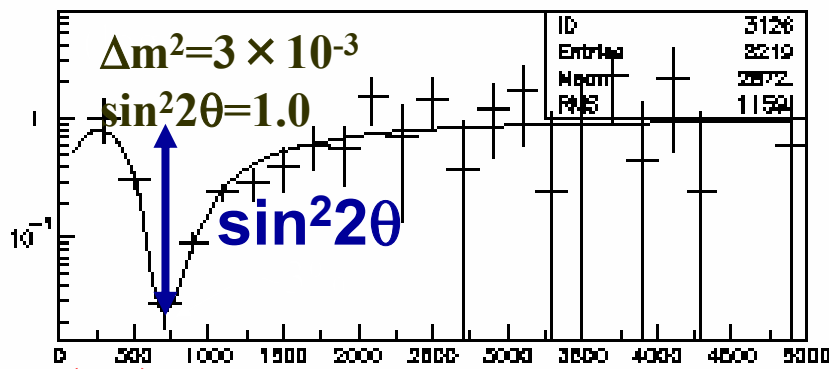
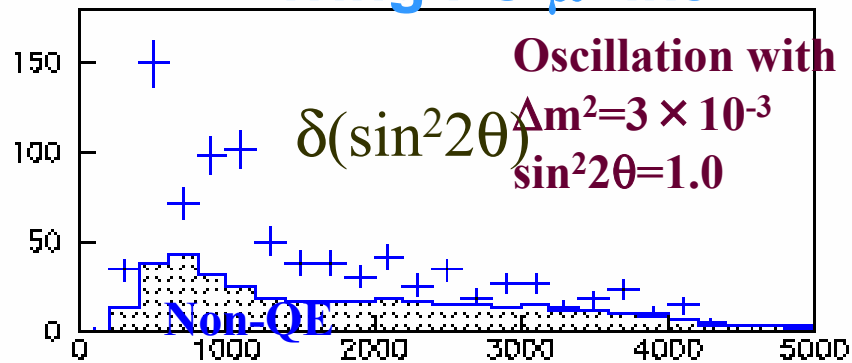
$\sin^2 2\theta_{\mu e} = 0.05$ ($\sin^2 2\theta_{\mu e} \equiv 0.5 \sin^2 2\theta_{13}$)



$\sin^2 2\theta_{13} < 0.006$ (90% C.L.)

ν_μ disappearance

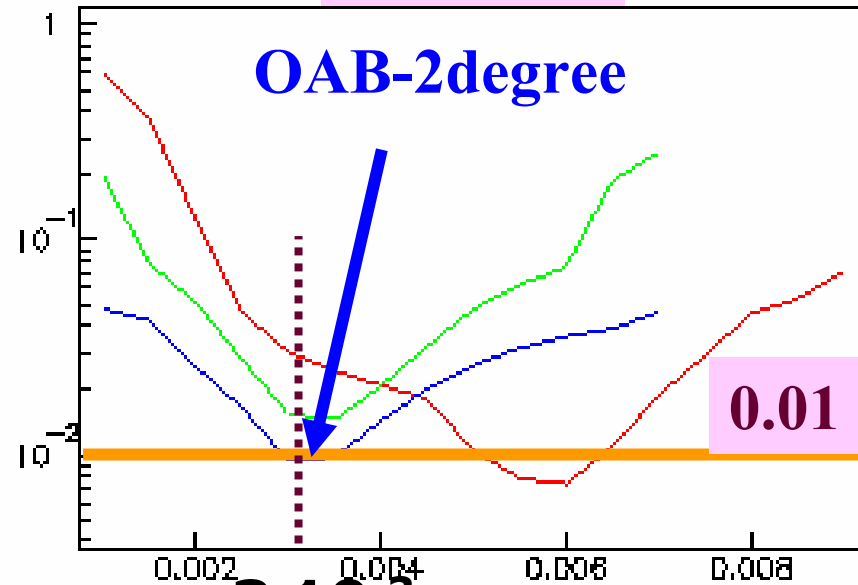
1ring FC μ -like



Δm^2 Reconstructed E_ν (MeV)

$\delta \sin^2 2\theta$

OAB-2degree



True Δm^2

$$\delta \sin^2 2\theta_{23} \sim 0.01$$

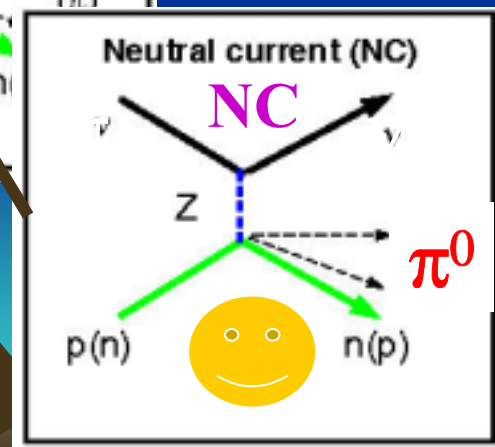
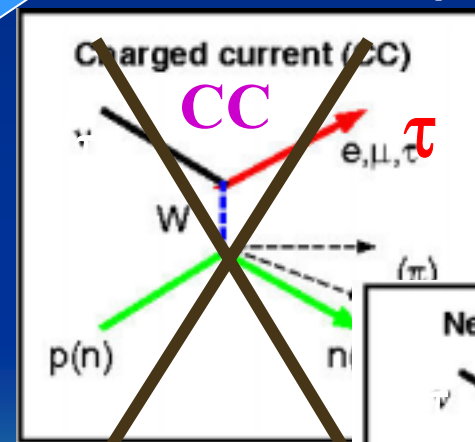
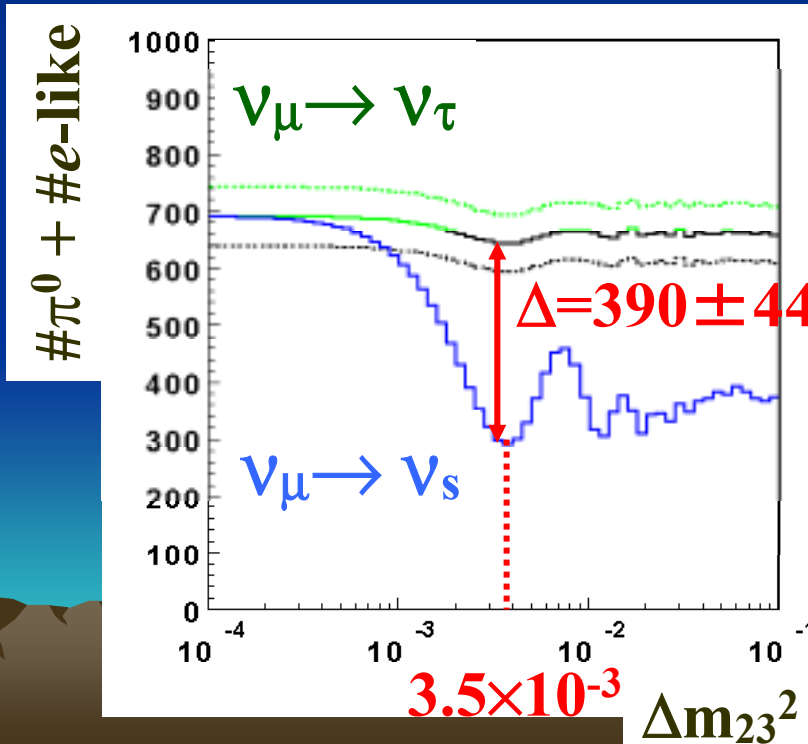
$$\delta \Delta m_{23}^2 < 1 \times 10^{-4} \text{eV}^2$$

$\nu_\mu \rightarrow \nu_\tau$ confirmation w/ NC interaction

- **NC π^0 interaction** ($\nu + N \rightarrow \nu + N + \pi^0$)

$\nu_\mu \rightarrow \nu_e$ CC + NC (~0.5CC) ~ 0 ($\sin^2 2\theta_{\mu e} \sim 0$)
 $\nu_\mu \rightarrow \nu_\mu$ CC + NC (~0.5CC) ~ 0 (maximum oscillation)
 $\nu_\mu \rightarrow \nu_\tau$ NC

π^0 is sensitive to ν_τ flux. \rightarrow Limit on ν_s ($\delta f(\nu_s) \sim 0.1$)



CP Violation Study w/ Hyper-K (Phase 2)

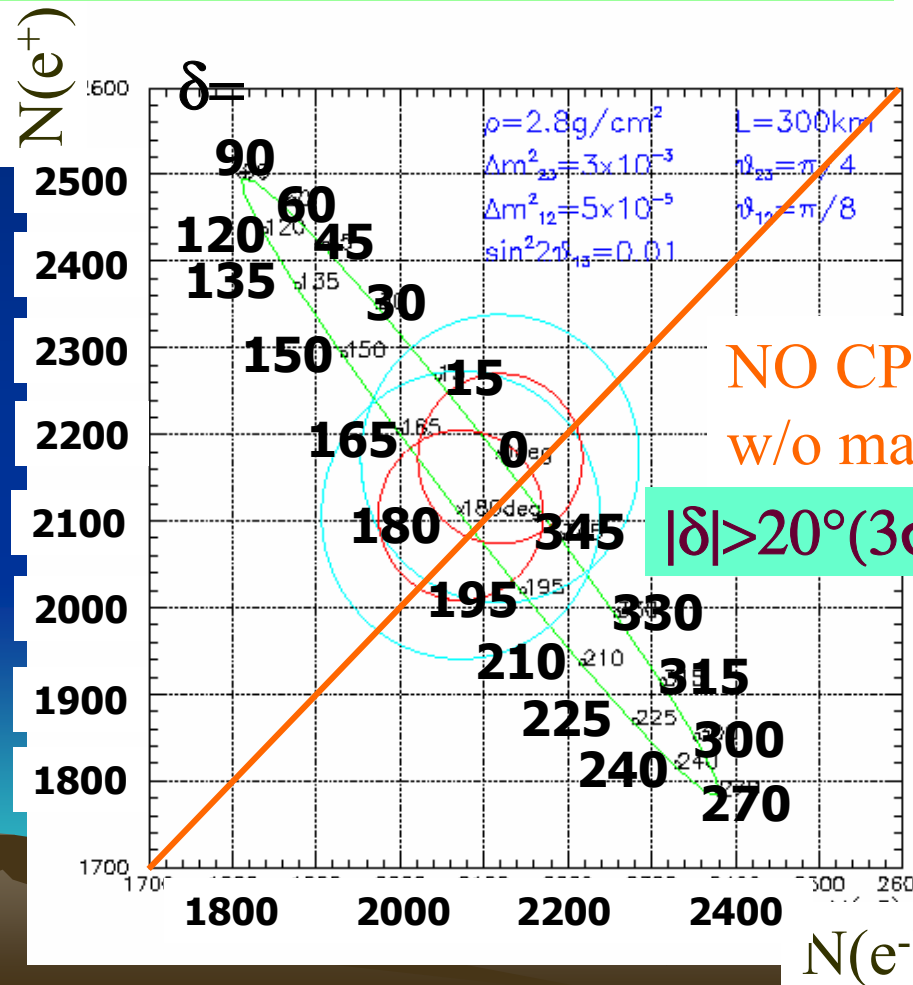
- Compare $\nu_{\mu} \rightarrow \nu_e$ (2 years) with $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ (6 years)

$$A_{CP} = \frac{P(\nu_{\mu} \rightarrow \nu_e) - P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)}{P(\nu_{\mu} \rightarrow \nu_e) + P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)} \approx \frac{\Delta m_{12}^2 L}{4E_{\nu}} \cdot \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta$$

$$\begin{aligned} \Delta m_{12}^2 &= 5 \times 10^{-5} eV^2, \\ \Delta m_{23}^2 &= 3 \times 10^{-3} eV^2 \\ \sin^2 2\theta_{13} &= 0.01 \\ \theta_{23} &= \pi/4, \theta_{12} = \pi/8 \end{aligned}$$

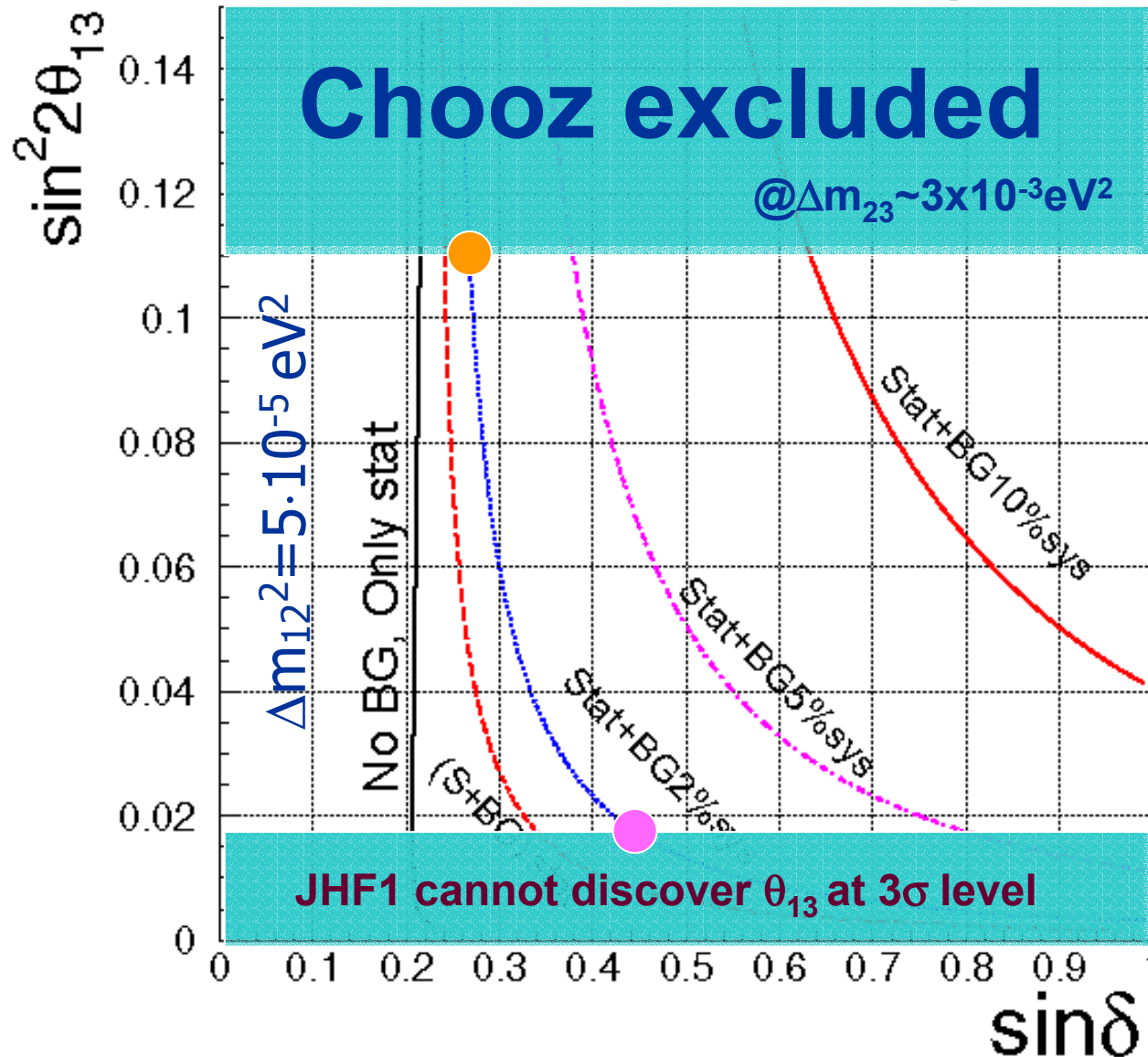
Asymmetry

$$\begin{aligned} &\equiv \frac{N(e^+) - N(e^-)}{N(e^+) + N(e^-)} \\ &\approx \frac{\sim 2000 - \sim 2000}{\sim 2000 + \sim 2000} \\ &\approx 0.02 \end{aligned}$$



CP Sensitivity (3σ)

JHF-HK CPV Sensitivity



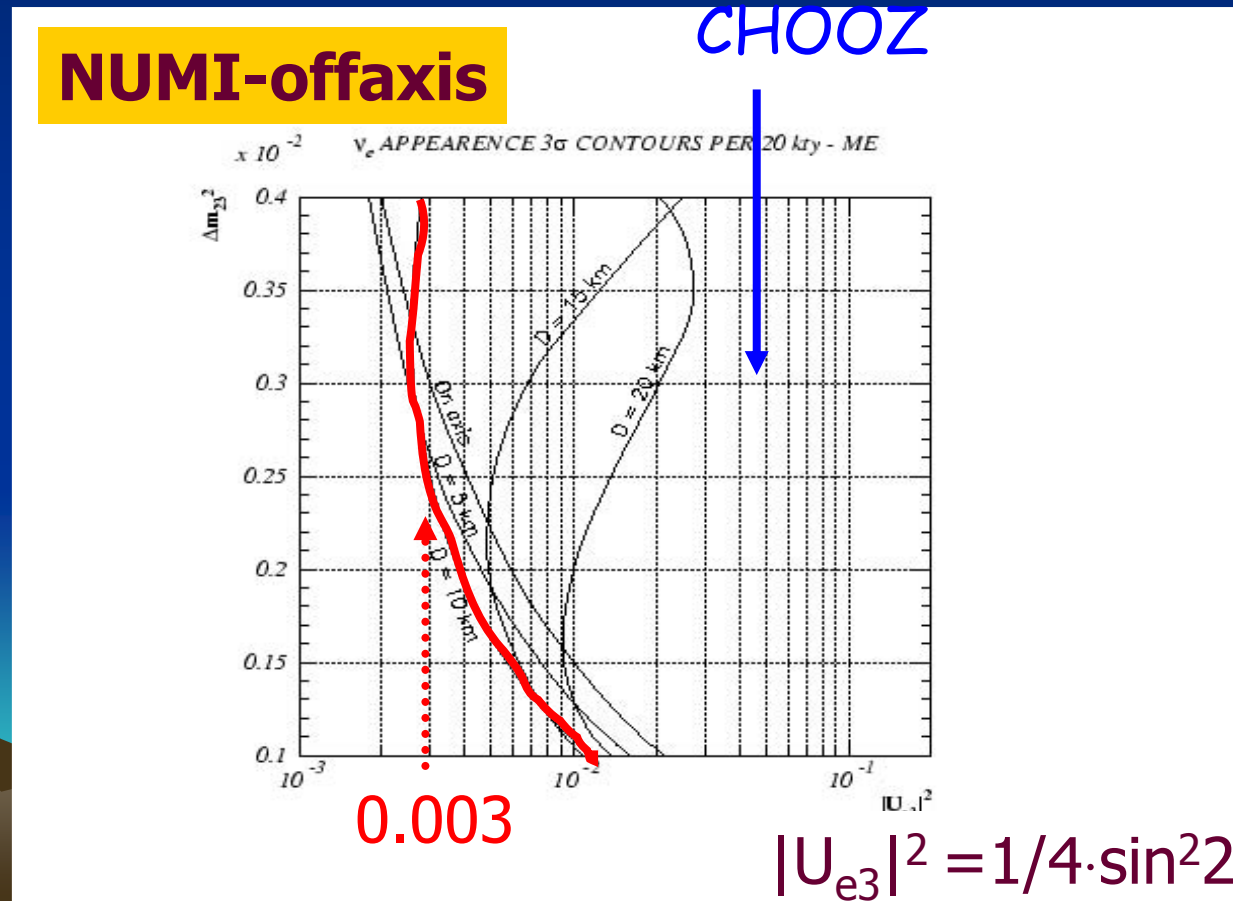
assuming 2% BG uncertainty

$\sin^2 2\theta_{13} = 0.01$
 $\rightarrow \sin \delta > 0.44$
 (26deg)

large $\sin^2 2\theta_{13}$
 $\rightarrow \sin \delta > 0.25$
 (14deg)

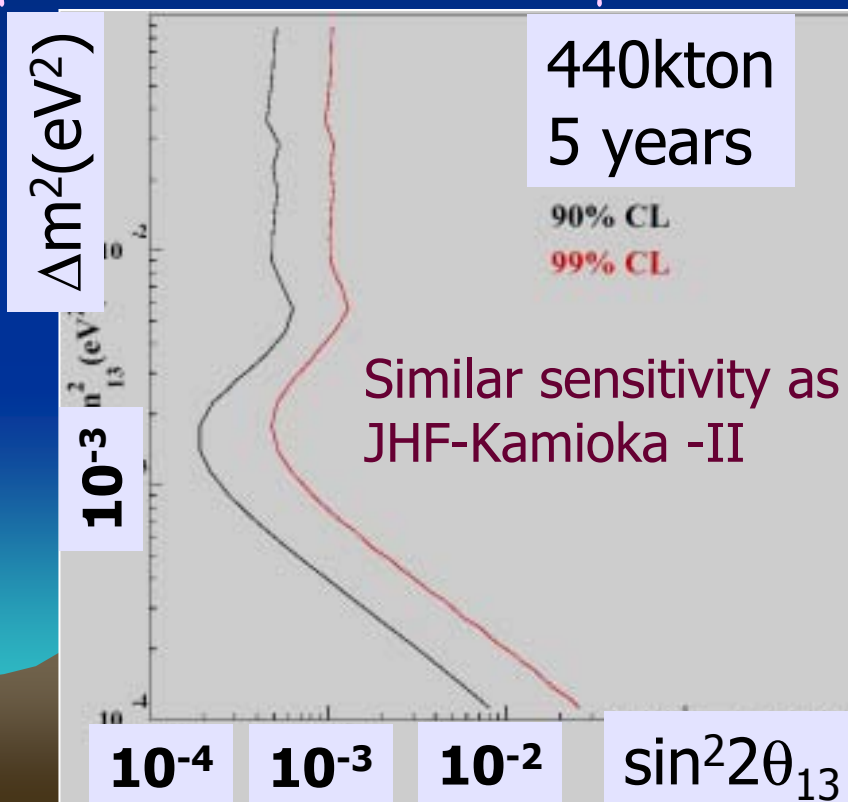
US Super ν beam

- They are studying the physics potential of several options, which are competitive to JHF-Kamioka project.



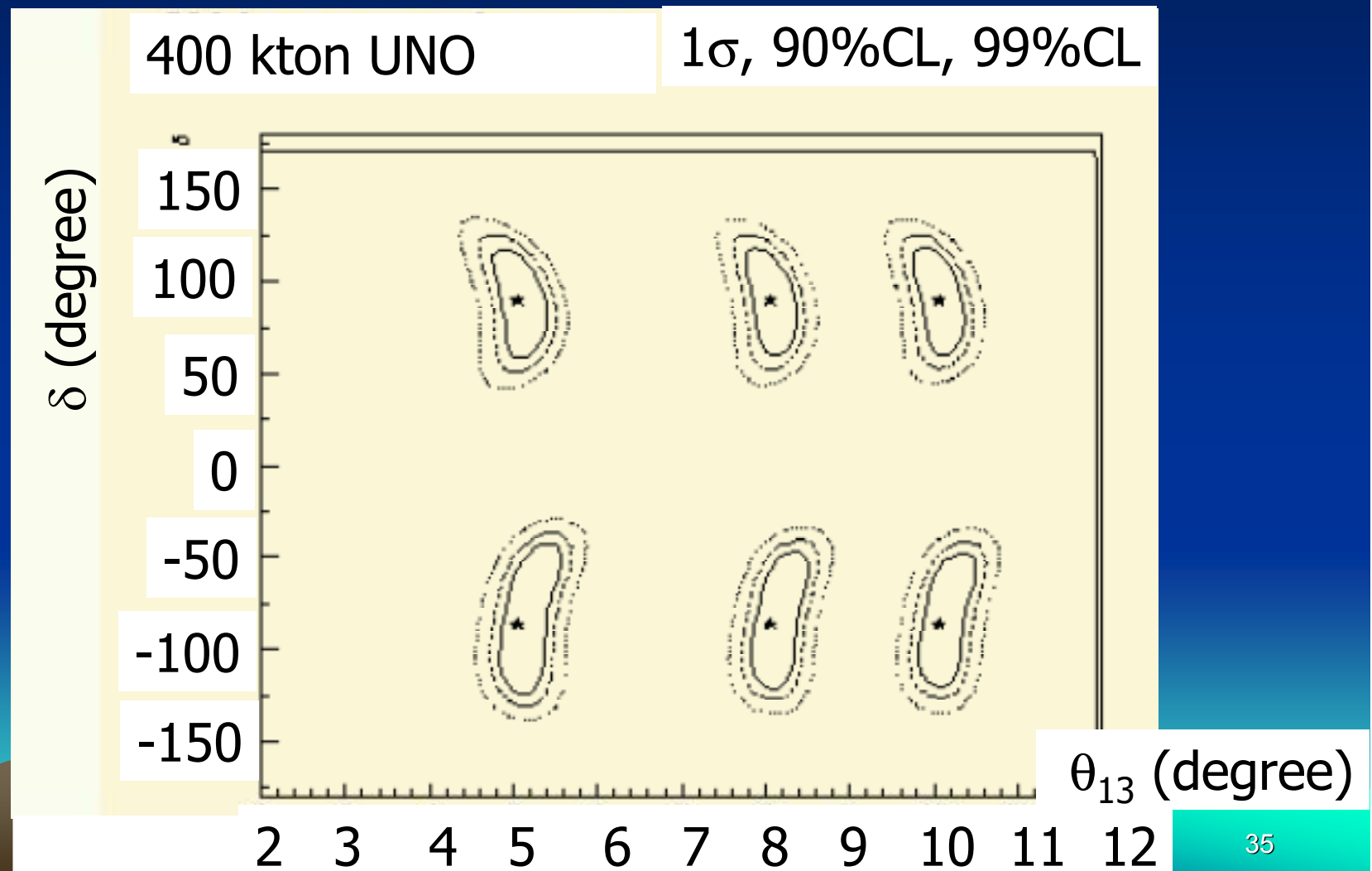
CERN: SPL+ β beams to UNO@Frejus

- Search for $\nu_{\mu} \rightarrow \nu_e$ w/ super ν beam
- Search for $\nu_e \rightarrow \nu_{\mu}$ w/ β beam
- They can also study T symmetry in ν oscillation between $\nu_{\mu} \rightarrow \nu_e$ and $\nu_e \rightarrow \nu_{\mu}$.

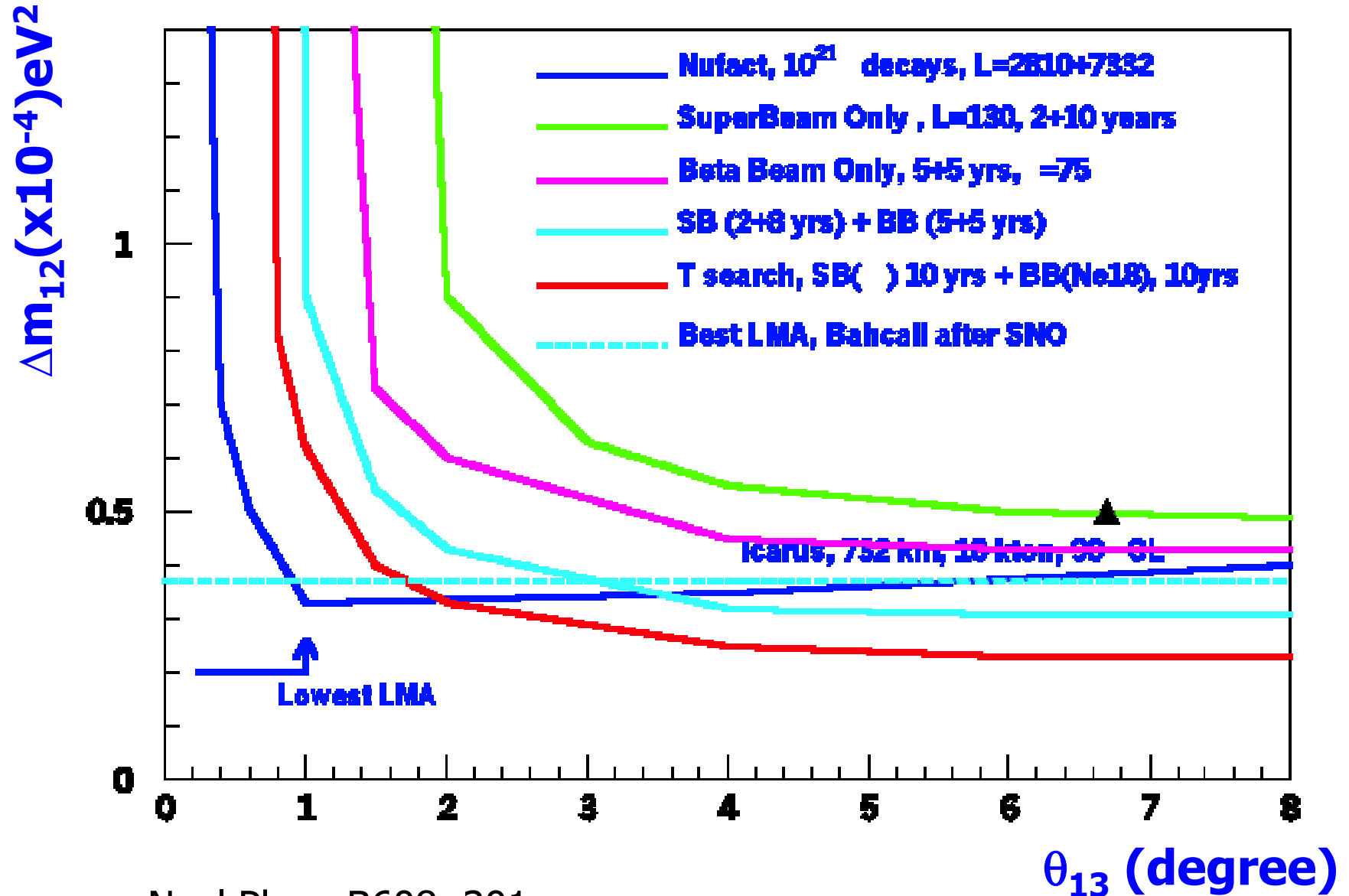


CP sensitivity (δ)

- θ_{13} and δ measurement



CERN: SPL+ β beams to UNO@Frejus



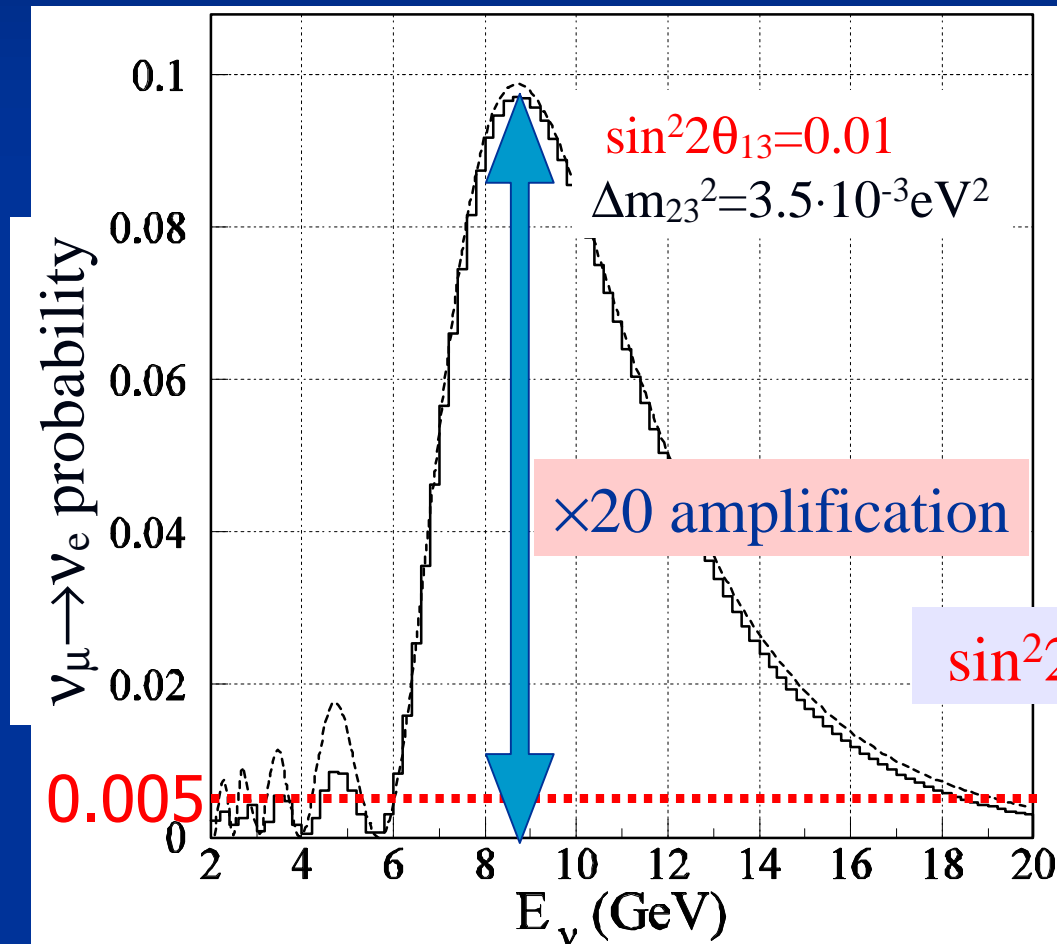
More Possibilities (1)

- Japan to Korea or China (hep-ph/0112338)
 - CP violation and the sign of Δm_{23}^2 with JHF ν beam.



More Possibilities (2)

- Fermilab to Super-Kamiokande (hep-ex/0203005)
 - Matter enhanced ν oscillation ($L=9300\text{km}$)



Remark: Near to Far extrapolation

A next generation LBL ν experiment requires the precise knowledge of ν spectrum at the far site based on the measurements at the near site.

- Hadron Production Experiments
 - CERN HARP(-III), FNAL E907
- A detector at the medium distance
 - JHF-Kamioka project proposes to build the near neutrino detector at 2km away from the target, which is far enough to see the neutrino production volume as a point source.

Non Oscillation Physics

- A near site detector with the super ν beam can probe LSND effect if demanded.
- Precise ν interaction and cross section study.
 - Quasi-elastic scattering; etc..
- $\nu_{\mu}e^{-}$ elastic scattering at low Q^2 ; the neutrino magnetic moment.

5. Summary and Conclusion

- In the next 5 years, several LBL ν experiments will provide more information on ν oscillation.

- The Super ν beam experiments will follow the trend, and further explore the physics.

JHF-Kamioka experiment will start in 2007 earliest.

- $\sin^2 2\theta_{13}$ sensitivity w/ super ν -beams
 - $0.01 \sim 0.001$ w/ 20~50 kton detectors
>10 times more sensitive than CHOOZ.
 - < 0.001 w/ 400~1000 kton detectors

5. Summary and Conclusion (Continue)

– CP Violation sensitivity

- Need a larger detector (400~1000 kton) when $\nu_{\mu} \rightarrow \nu_e$ oscillation is discovered.
- The sensitivity of CP violation phase δ is down to 10~20 degrees if $\Delta m_{12}^2 = 5 \cdot 10^{-5} \text{eV}^2$ (LMA).

– The sign of Δm_{23}^2 and matter effect

- The longer baseline ($\sim 1000\text{km}$ or more) is necessary to study it. JHF and CERN SPL has no design of the ν beam line for this purpose.

– There may be a surprise behind the ν oscillation phenomena. We should proceed the experiments with the super ν beams and the larger detectors.