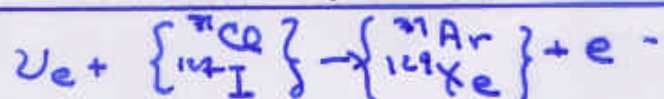


HYBRIDRADIOCHEMICAL - ELECTRONICSOLAR ν_e DETECTOR

GOAL: SEPARATELY & DIRECTLY MEASURE:

FLUX OF ν_e FROM ^8B
&

FLUX OF ν_e FROM ^7Be , pep & CNO

TECHNIQUE: SINGLE ATOM EXTRACTION &

CLASSIFICATION OF ELECTRONICALLY LABELED

ν_e INTERACTION PRODUCTS FROM

KILOTON SOLAR ν_e DETECTOR

PARTICIPANTS
FROM

UNIVERSITY OF PENNSYLVANIA

SOUTH DAKOTA SCHOOL OF MINES & TECHNOLOGY

SOUTH DAKOTA STATE UNIVERSITY

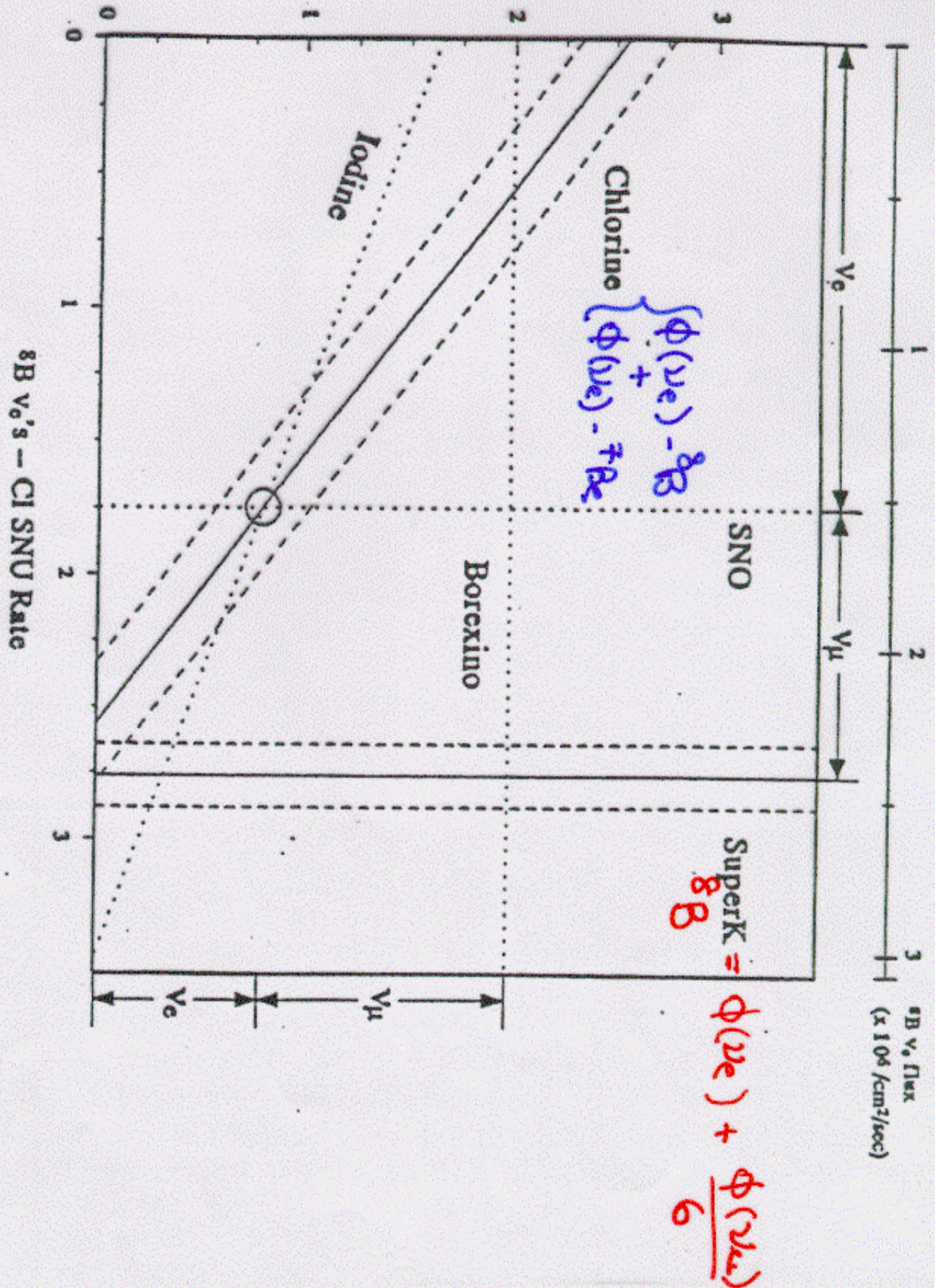
UNIVERSITY OF SOUTH DAKOTA

UNIVERSITY OF WASHINGTON

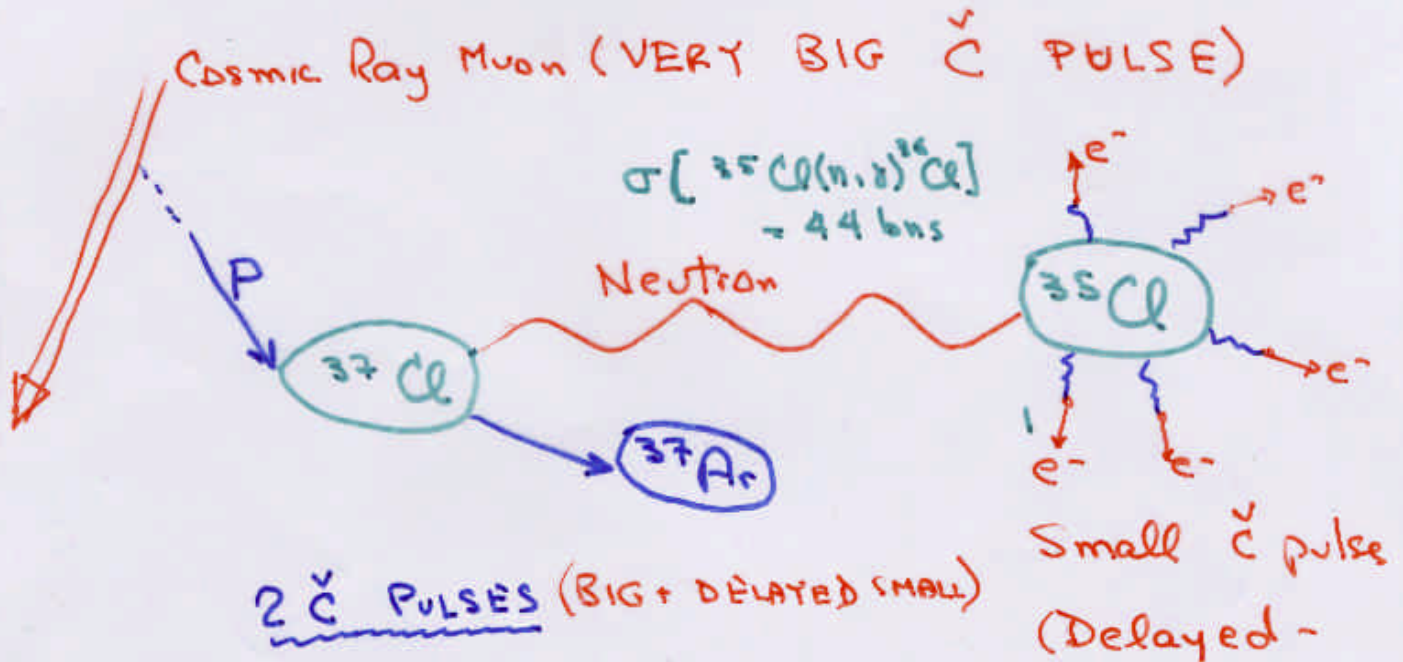
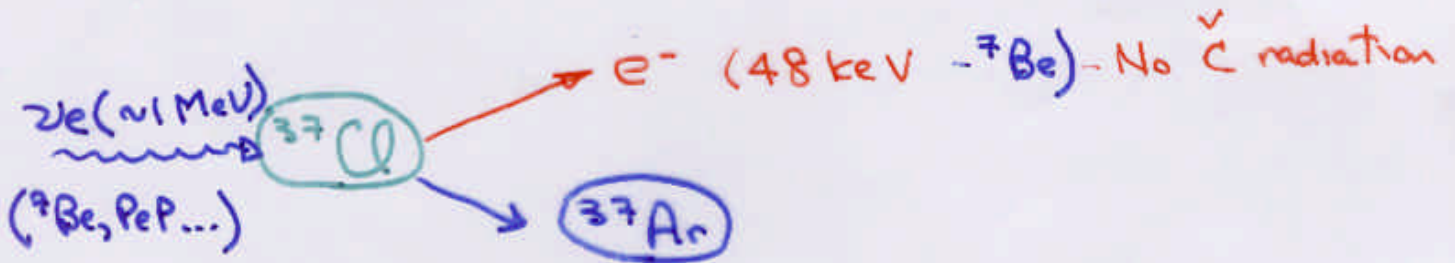
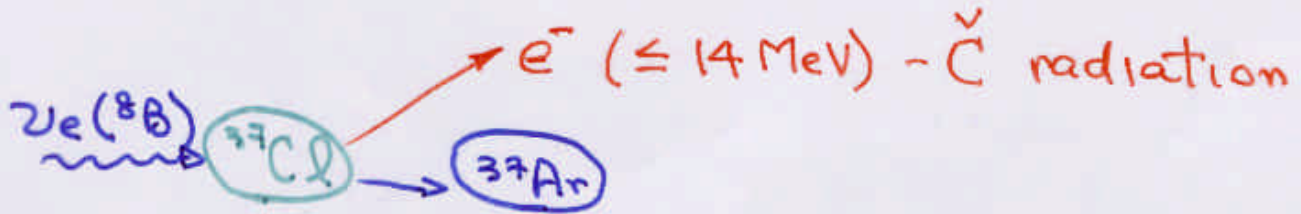
LOS ALAMOS NATIONAL LABORATORY

INSTITUTE FOR NUCLEAR RESEARCH (Moscow)

$^7\text{Be}/\text{CNO}/\text{pep } \nu_e\text{'s} - \text{Cl SNU Rate}$



^{37}Ar PRODUCTION PROCESSES

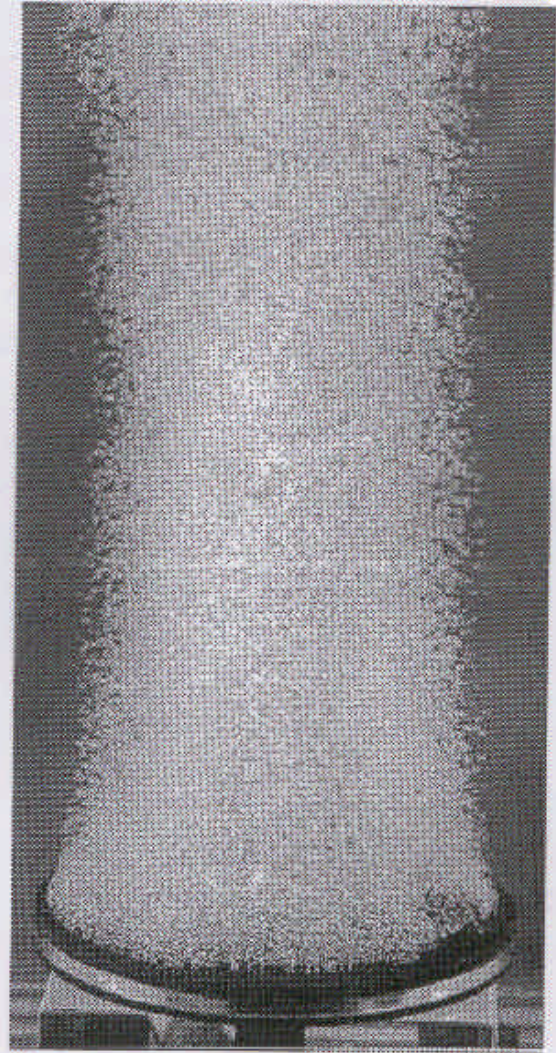
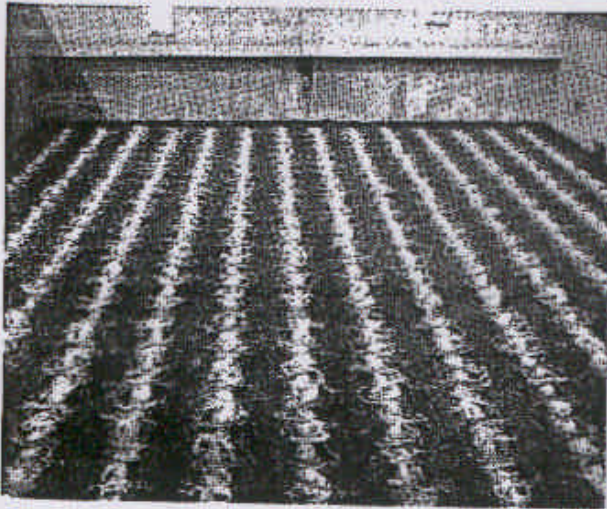
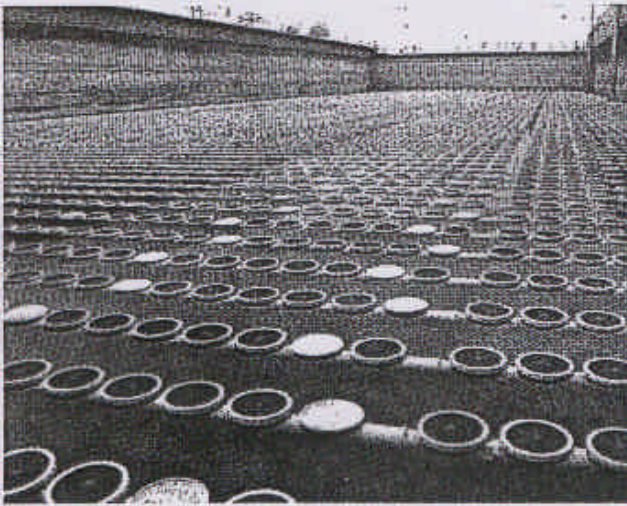


SIMILAR DIAGRAMS FOR $^{127}\text{I} \rightarrow ^{127}\text{Xe}$
 $\sigma [^{127}\text{I}(n, \gamma)^{128}\text{I}] = 6.2 \text{ bns}$

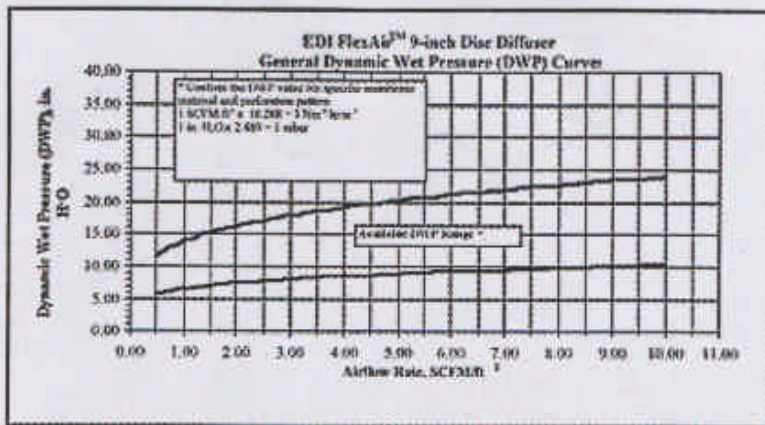
① CAN WE REMOVE EACH ^{37}Ar ATOM DUE TO γ_e FROM ^8B WHEN IT IS MADE?

NEED:

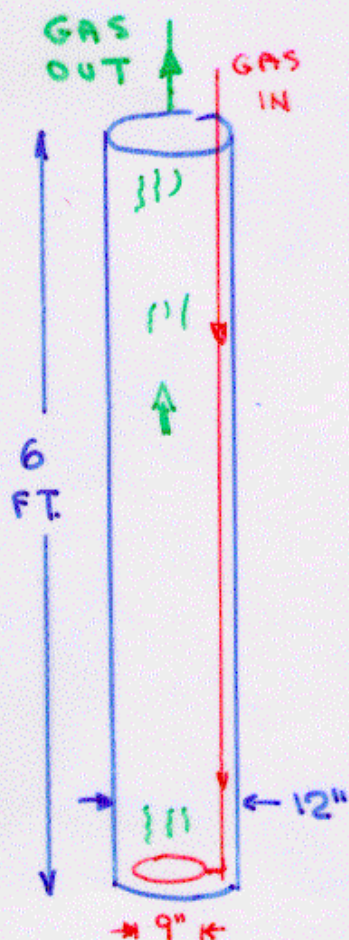
- a) FAST EXTRACTION PROCESS
- b) MULTIPLE SEGMENT DETECTOR
- c) ^8B γ_e EXTRACTION TRIGGER
- d) COSMIC RAY TRIGGER



9" Disc Dynamic Wet Pressure (DWP)



GAS DIFFUSER EXTRACTION RATE



MAKE ^{129}Xe WITH Pu-Be SOURCE

$1/e$ EXTRACTION GAS FLOW
= 30 LITERS He @ S.T.P.

$$\text{GAS COLUMN HEIGHT} = \frac{30 \text{ LITERS}}{\text{DIFFUSER AREA (410 cm}^2\text{)}} = 75 \text{ cm.}$$

SINCE DIFFUSER CAN EMIT
3 m HIGH COLUMN / min \Rightarrow

$$\tau(1/e) \sim 15 \text{ sec}$$

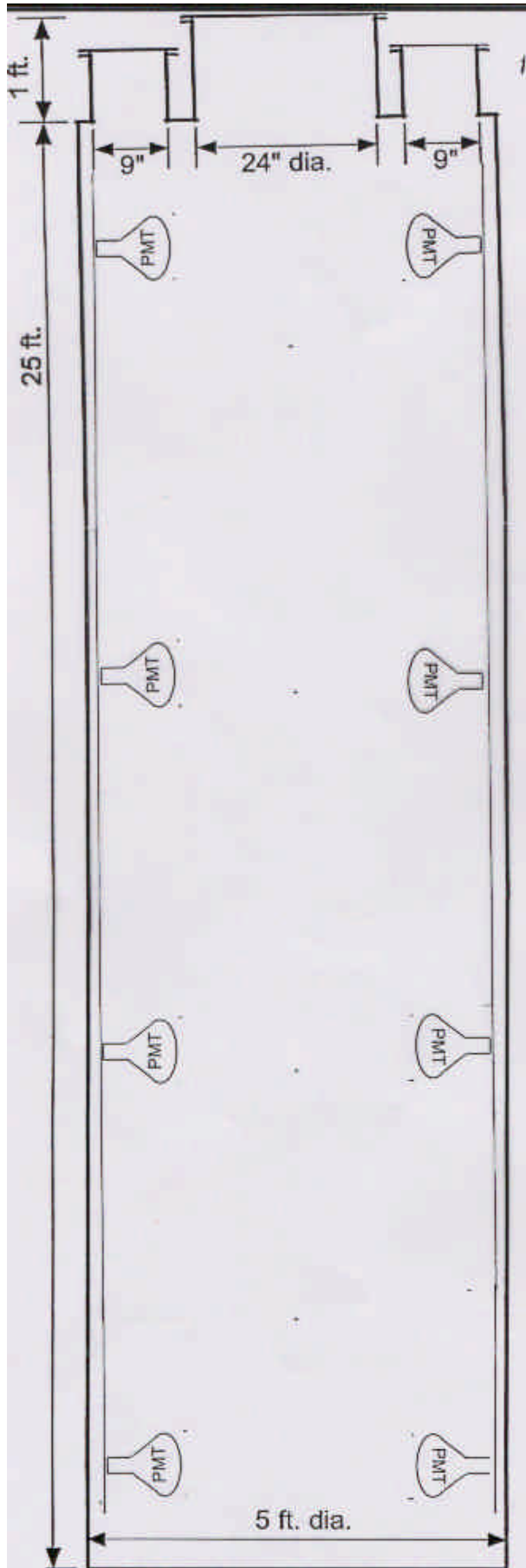
$$\& 1 - e^{-4} \text{ IN } \sim 1 \text{ MINUTE}$$

BUBBLE (1 mm dia) RISE VELOCITY $\sim 20 \text{ cm/sec}$

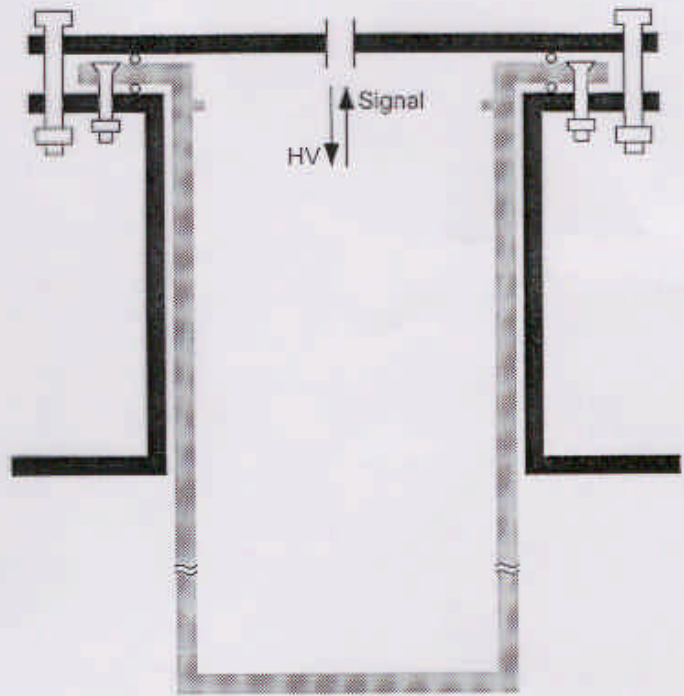
FOR 8.5 m HIGH DETECTOR - $T(\text{RISE}) = 40 \text{ sec}$

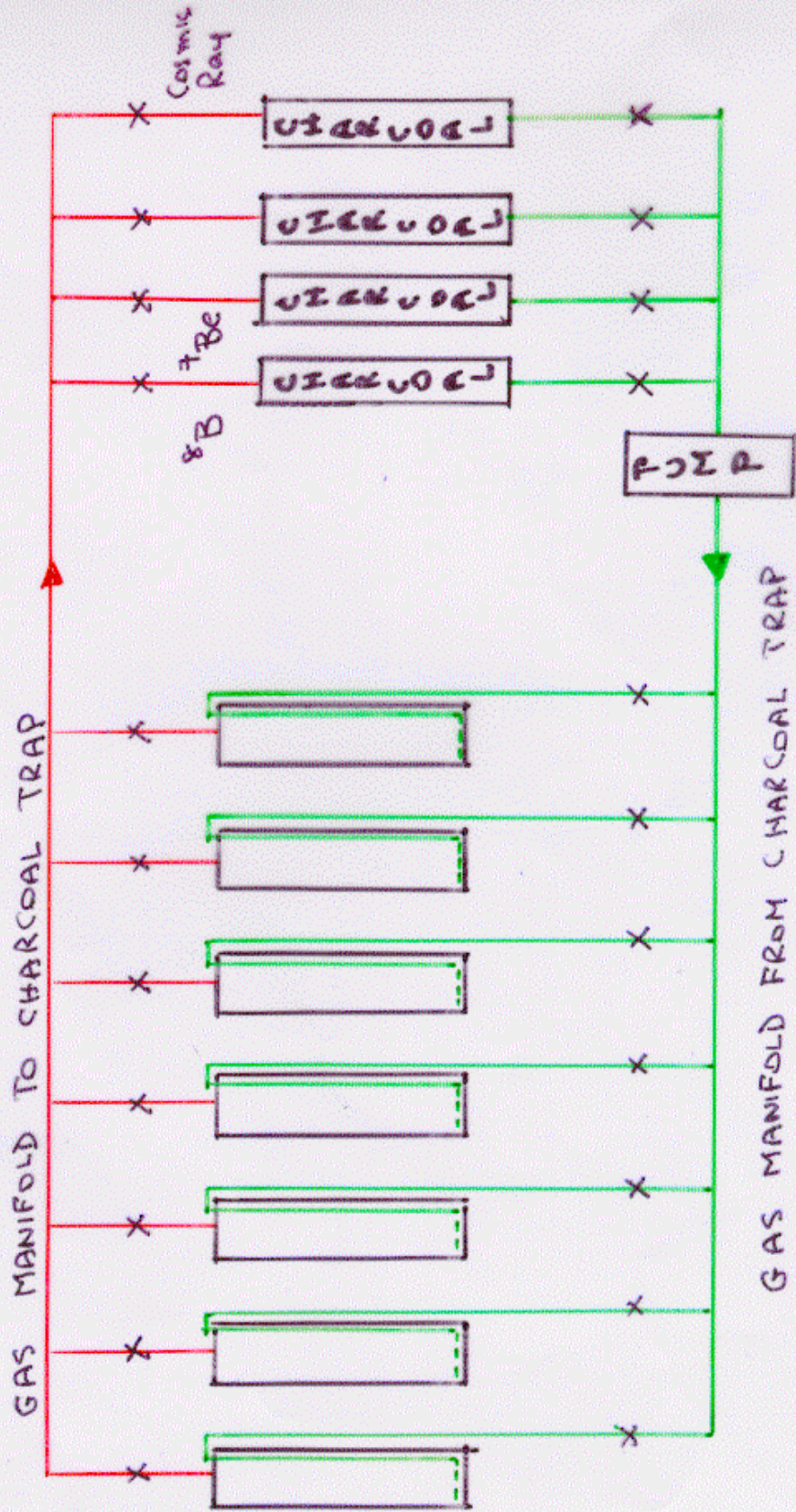
TOTAL EXTRACTION TIME = 1 min + 40 sec

$$\underline{\underline{\sim 2 \text{ min}}}$$



Seal Detail





ANOTHER POSSIBILITY

ASSUME SK SIGNAL CONSISTS OF

$$\left. \begin{aligned} \phi(\nu_e) &= 1.90 \times 10^6 \\ \phi(\nu_\mu) &= 3.25 \times 10^6 \end{aligned} \right\} \Sigma = \frac{\sigma(\nu_e)}{\sigma(\nu_\mu)} \text{ from BP 98}$$

$$\text{SINCE } \sigma(\nu_\mu + e^- \rightarrow) = \frac{1}{6} \sigma(\nu_e + e^- \rightarrow)$$

THE OBSERVED INTERACTION RATE

WILL BE

$$\begin{array}{cc} (\nu_e) & (\nu_\mu) \end{array}$$

$$1.90 + \frac{3.25}{6} = 1.90 + 0.54 = \underline{\underline{2.44}}$$

$$\phi(\nu_e) = 1.90 \times 10^6 \Rightarrow 2.17 \text{ SNU}(\nu_e)$$

DETECTION RATES

CASE ① $\phi(\nu_e)$ from ${}^7\text{Be} = 0 \Rightarrow$ all signal all ${}^8\text{B}$

SK - $\phi(\nu_e) = 2.44 \pm 0.09 \times 10^6 / \text{cm}^2 \text{sec}$

	C_2Cl_4	${}^{37}\text{Cl}$	$\text{NaI} + \text{H}_2\text{O}$	${}^{127}\text{I}$
1 MODULE	24 TONS	5.3 TONS	24 TONS	9 TONS
# TARGET ATOMS	8.7×10^{28}	${}^{37}\text{Cl}$		5×10^{28} ${}^{127}\text{I}$
${}^8\text{B}$ - ν_e int/day-module	.02			.04
TOTAL RATE IN 50 MODULES		1/DAY		2/DAY (500 TONS ${}^{127}\text{I}$)

CASE ② SK $\left\{ \begin{array}{l} \phi(\nu_e) = 1.9 \times 10^6 \\ \phi(\nu_\mu) = 3.25 \times 10^6 \end{array} \right. \left(1.9 + \frac{3.25}{6} = 2.44 \right)$

Now Cl has $2.16 \text{ SNU } {}^8\text{B} + 0.4 \text{ SNU } {}^7\text{Be}$

${}^8\text{B}$ RATE	0.84/DAY (300/YR)	1.7/DAY (620/YR)
${}^7\text{Be}$ RATE	0.15/DAY 55/YR	0.41/DAY 150/YR

FOR 3 YEARS OF DATA

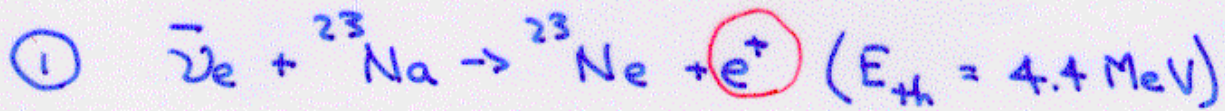
& ASSUMING THAT 60% OF PRODUCED ATOMS
ARE COUNTED

	^{37}Cl	^{127}I
^8B - Case ①	657	1314
^8B - case ②	552	1117
Δ	$= 105 \pm 35 (3\sigma)$	$197 \pm 49 (4\sigma)$
- - - - -		
^7Be	$= 100 \pm 10 (\text{stat})$	$270 \pm 16 (\text{stat})$
Cosmic Ray + Neutron Abund	$\sim 60-70$	~ 5

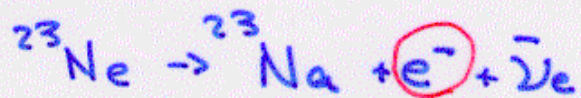
$\bar{\nu}_e$ DETECTION IN IODINE DETECTOR

(NaI + H₂O)

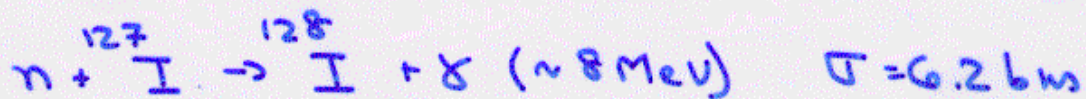
DOUBLE PULSE PATTERNS



followed by ($\tau_{1/2} = 37.6 \text{ sec}$)



followed by ($\tau_{\text{mod}} \sim 300 \mu\text{sec}$)



$\textcircled{} \equiv \text{CERENKOV RADIATORS}$

HAXTON (1988)

Lande -

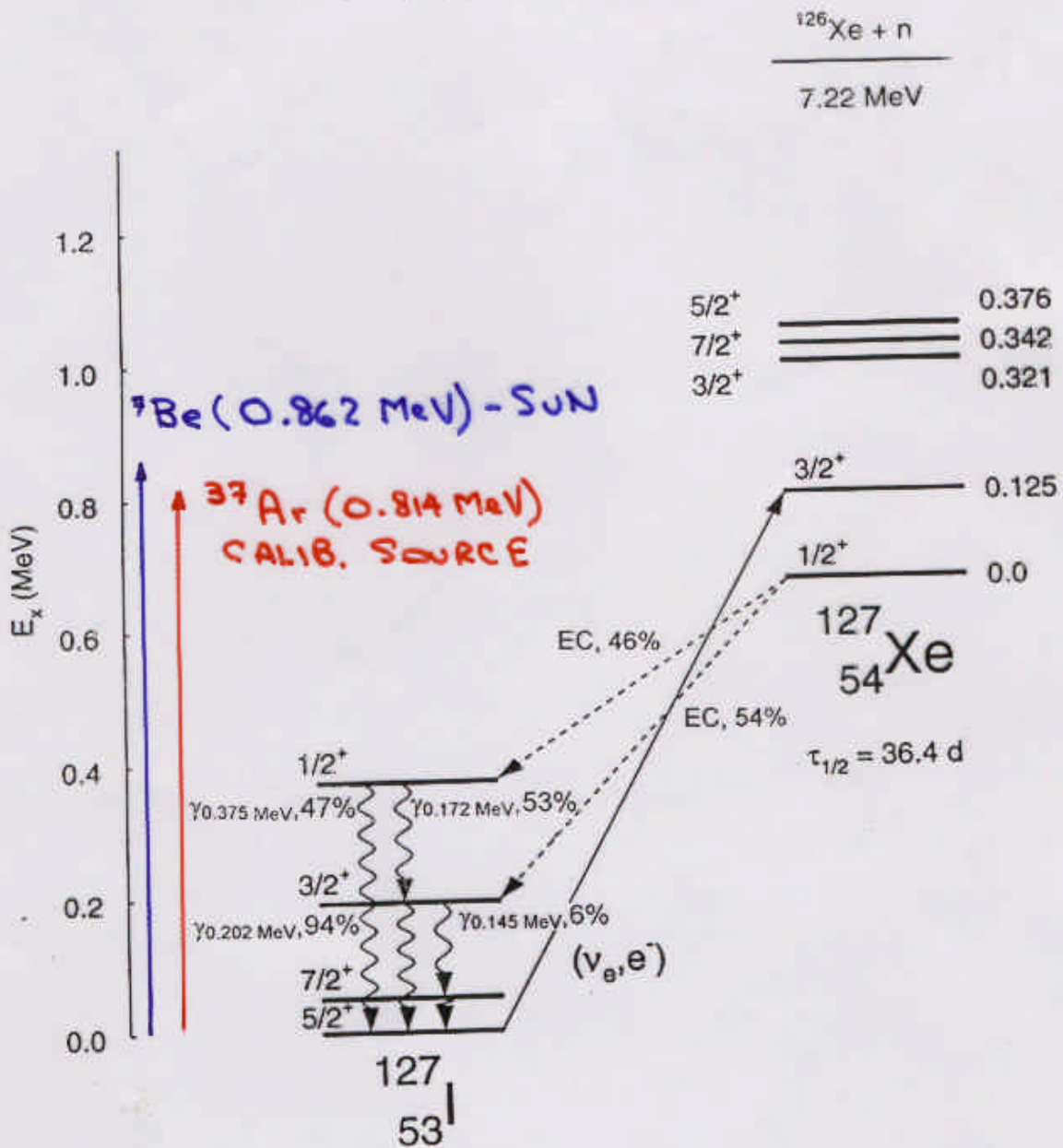
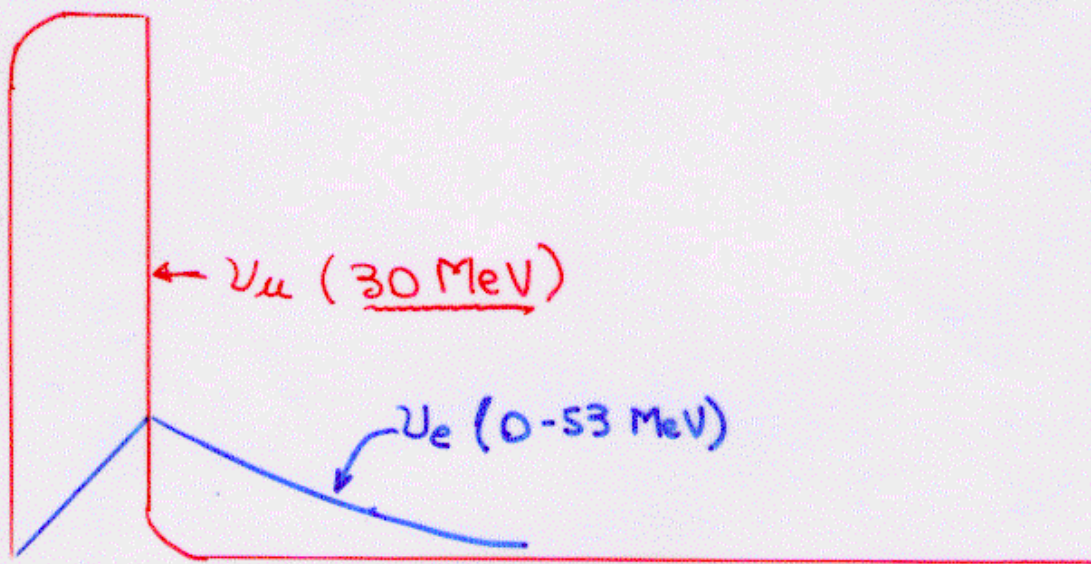


Figure 1.4 Decay scheme for $^{127}\text{Xe} \rightarrow ^{127}\text{I}$. The decay proceeds via electron capture. The absence of a ground state to ground state transition means that ^{127}Xe decays to an excited state of ^{127}I . Deexcitation of the ^{127}I nucleus results in nuclear gamma rays in the proportions and energy shown.

SNS NEUTRINO

PROTONS ON TARGET



8 CHARCOAL TRAPS

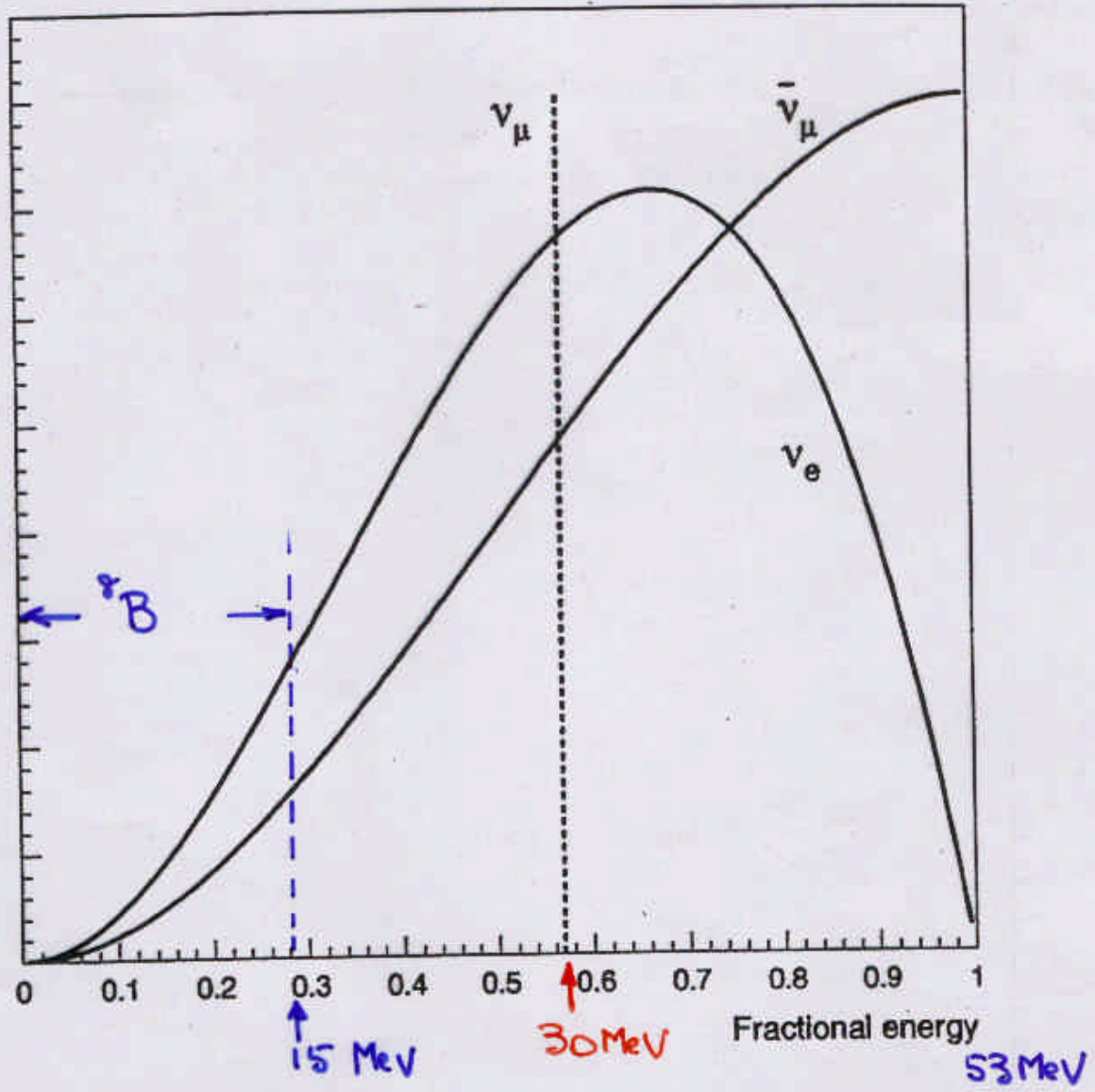
ENERGY	τ_1 (0-650ns)	τ_2 (650ns-10μs)	τ_3 (10μs-1ms)	τ_4 (1ms-16ms)
53 15 MeV	$\nu_\mu \rightarrow \nu_e$ TEST REGION	$\sigma(\nu_e - \text{SUPER-NOVA})$		
15 0 MeV	EARLY ν_e - COMPARE TO TOP	$\sigma[\nu_e + {}^8\text{B}]$	N bkgnd	C.R. bkgnd

ν_μ

ν_e

NEUTRONS

COSMIC RAYS



POSSIBLE LAYOUT OF HYBRID DETECTOR AT SNS

