

# *Neutrino* 2000



## Atmospheric Neutrinos

### MACRO

**Bari, Bologna, Boston, Caltech, Drexel, Indiana, Frascati, Gran Sasso, L'Aquila, Lecce, Michigan, Napoli, Pisa, Roma I, Texas, Torino**

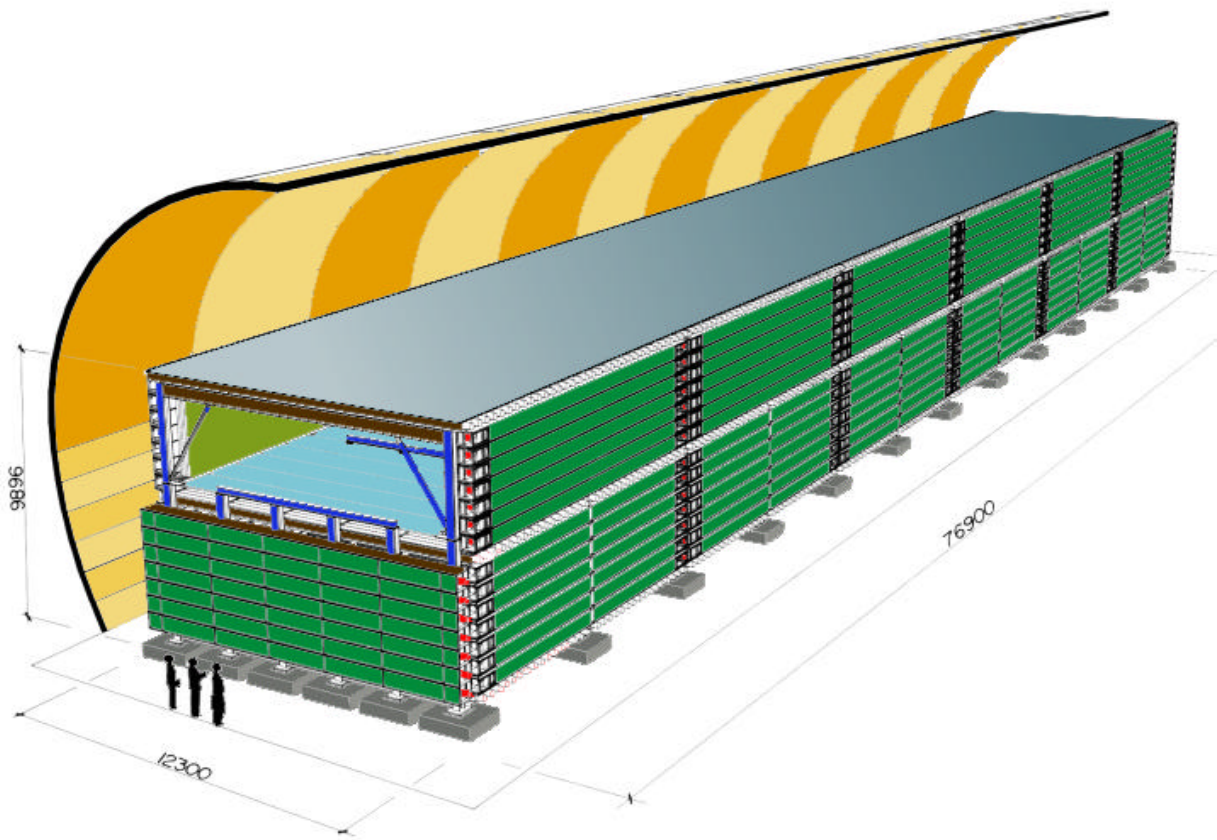
**Barry Barish**

**XIX International Conference on Neutrino Physics & Astrophysics**

Sudbury, Canada

June 16 - 21, 2000

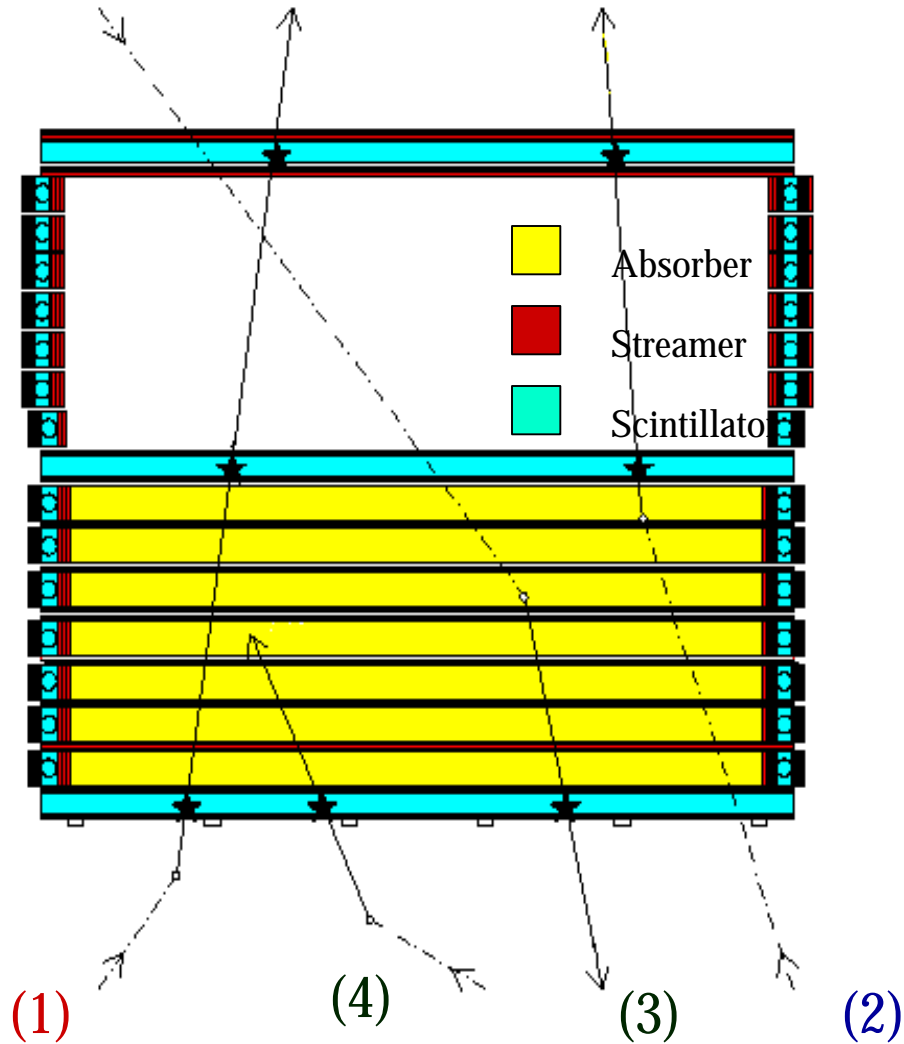
# Main features of Macro asv detector



- Large acceptance ( $\sim 10000 \text{ m}^2\text{sr}$  for an isotropic flux)
- Low downgoing  $\mu$  rate ( $\sim 10^{-6}$  of the surface rate)
- $\sim 600$  tons of liquid scintillator to measure T.O.F. (time resolution  $\sim 500\text{psec}$ )
- $\sim 20000 \text{ m}^2$  of streamer tubes (3cm cells) for tracking (angular resolution  $< 1^\circ$ )

More details in Nucl. Inst. and Meth. A324 (1993) 337.

# Neutrino event topologies in MACRO



Detector mass ~ 5.3 kton

Event Rate:

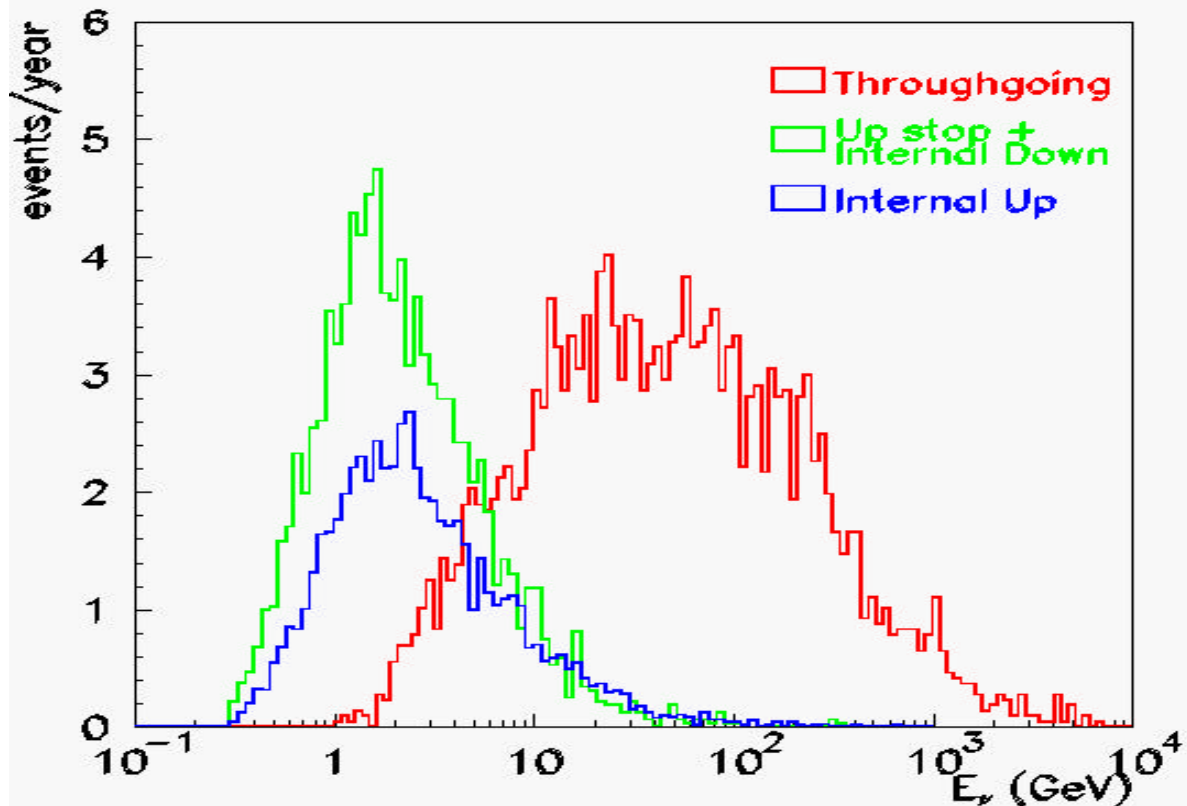
(1) up throughgoing  $\mu$  (ToF) ~ 160 /y

(2) internal upgoing  $\mu$  (ToF) ~ 50/y

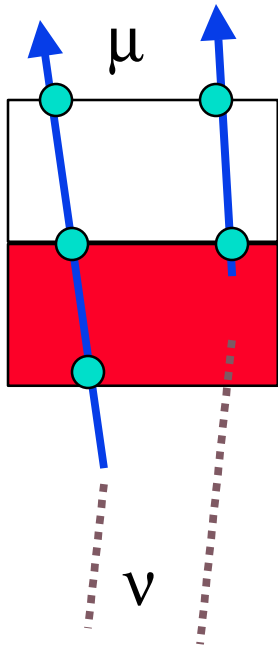
(3) internal downgoing  $\mu$  (no ToF) ~ 35/y

(4) upgoing stopping  $\mu$  (no ToF) ~ 35/y

# Energy spectra of $\nu$ events detected in MACRO

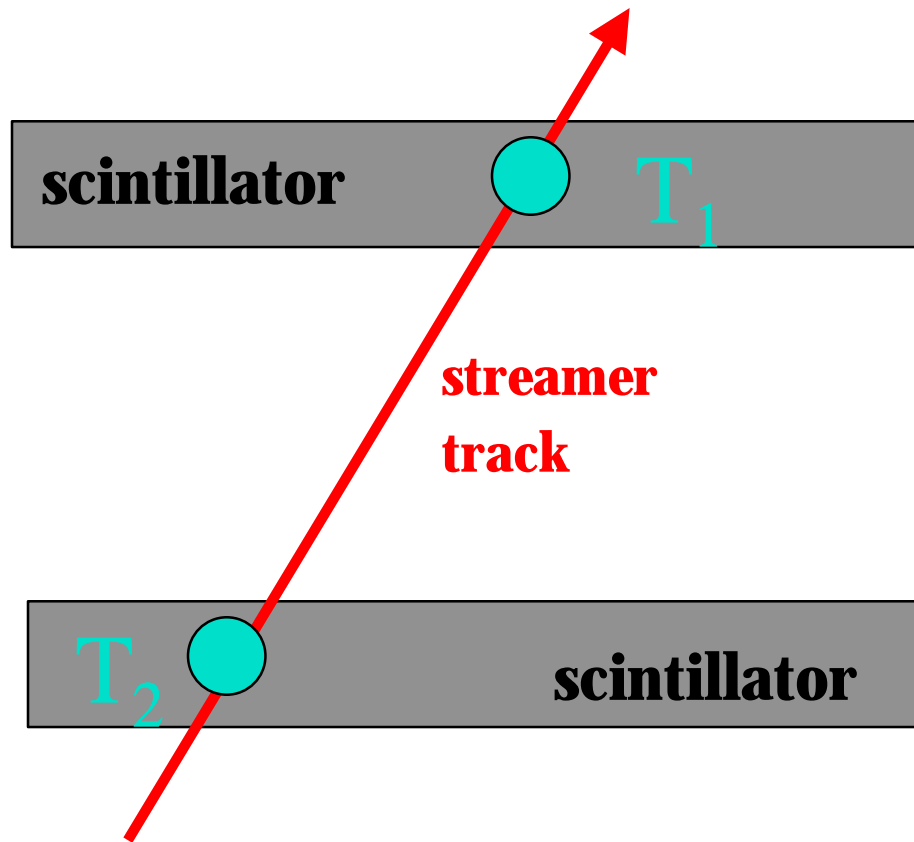


- $\langle E \rangle \sim 100$  GeV for throughgoing  $\mu$
- $\langle E \rangle \sim 5$  GeV for internal upgoing  $\mu$
- $\langle E \rangle \sim 4$  GeV for internal downgoing  $\mu$   
and for upgoing stopping  $\mu$
- Low energy events allow us to investigate the  $\nu$  oscillation parameter space independently of the throughgoing muon data

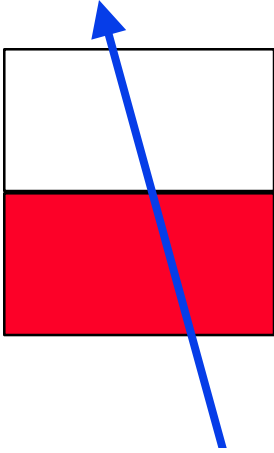


## Neutrino induced upward-going muons

*identified by the time-of-flight method*



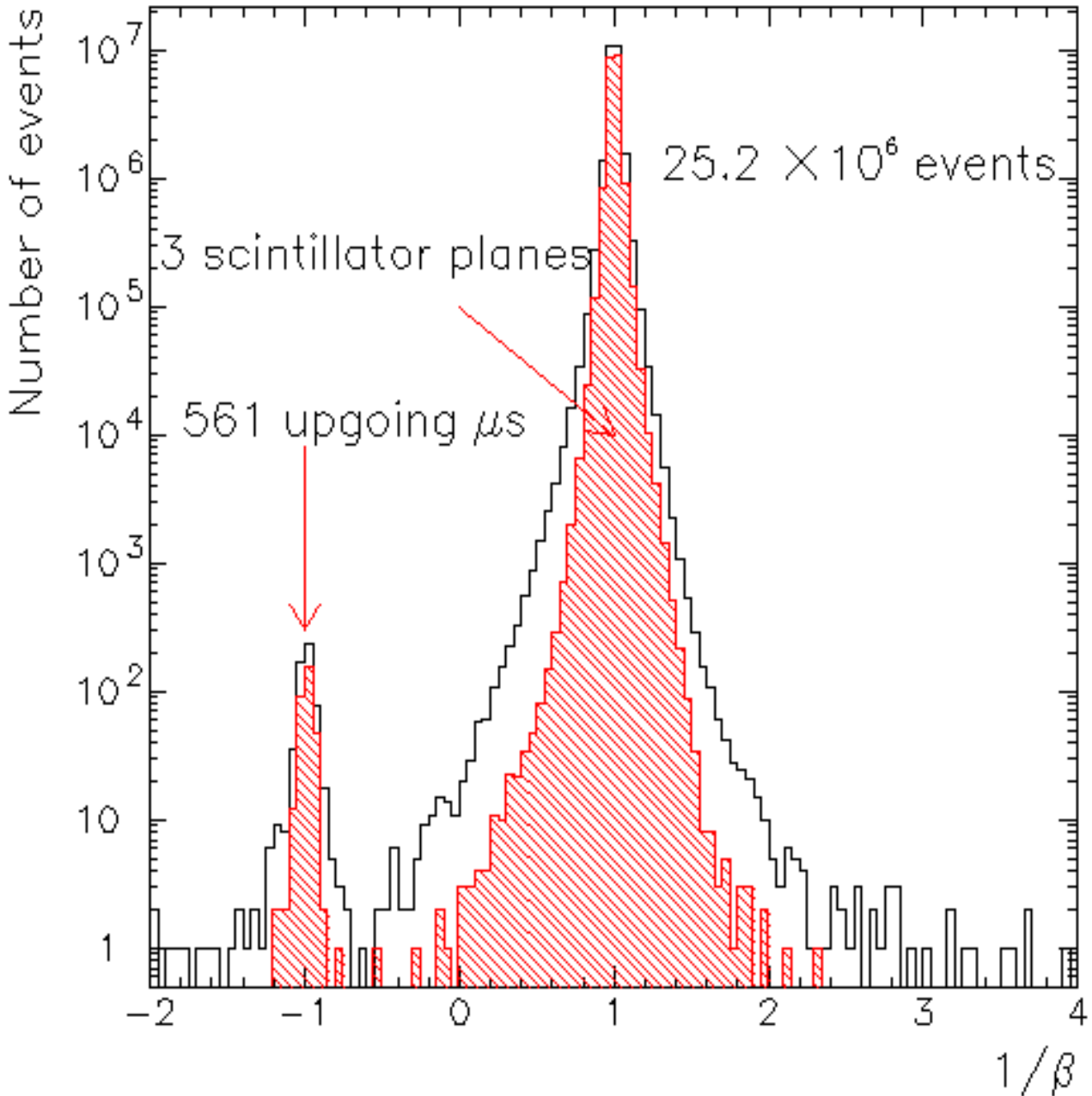
$$1/\beta = (T_2 - T_1) * c / l$$



# Upward throughgoing muons

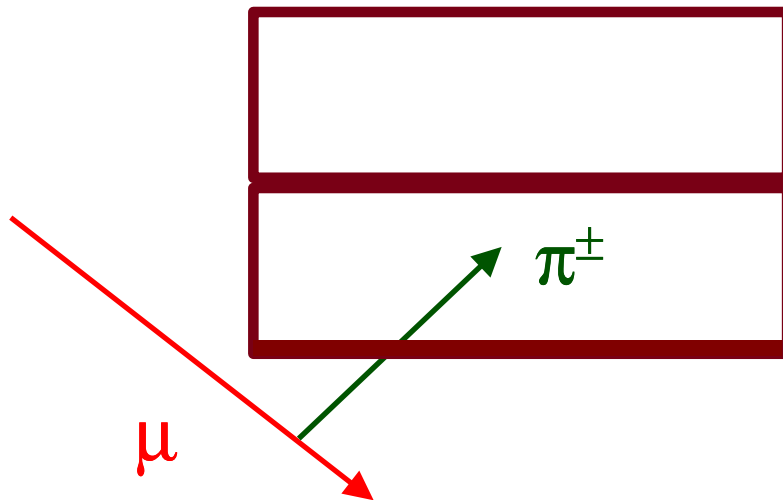
**External  $\nu$ -interactions** 768 events

upward vs downward muons



## Background source for $\nu_\mu$ events: pion production by downward $\mu$ 's

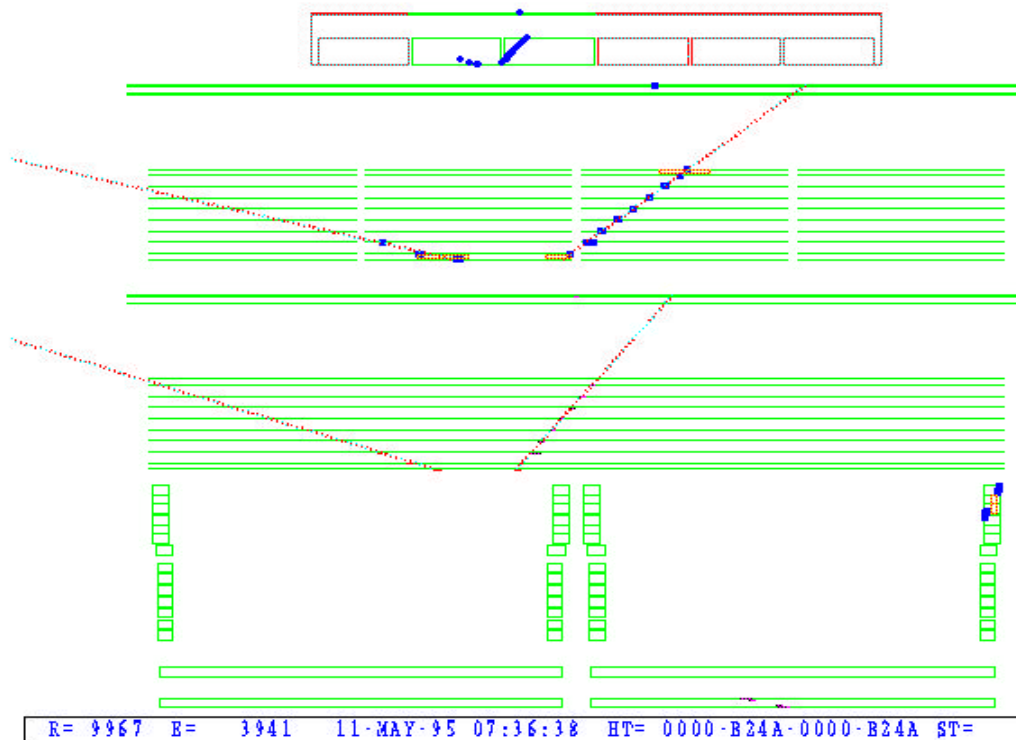
(MACRO Collab., *Astroparticle Phys.* 9('98)105)



- rate of downgoing  $\mu$ 's  $\sim 10^5$ x rate of upgoing  $\mu$ 's
- background mainly **pions produced at large angles** by  $\mu$  interactions in the rock around the detector  
( $\mu + N \rightarrow \mu + \pi^\pm + X$ )
- Estimated background in MACRO:
  - ~ 5% for the stopping muon sample  
(MACRO Coll., PL **B478** ('00) 5)
  - ~ 1% for the up-throughgoing  $\mu$  ( $>200$  g/cm<sup>2</sup> cut)  
MACRO Coll., PL **B434** ('98) 451)

# Pion production at large angle

- upgoing charged  $\pi$  produced by  $\mu$  interactions in the rock under the detector
- a background source for upward stopping and throughgoing muons
- studied 243 events with downgoing  $\mu$  plus upgoing particles



## ▪ Estimated background

- upward throughgoing  $\mu$ 's: ~1%
- low energy events: ~5%



# Total observed flux and comparison with Monte Carlo

## Data

<b>number of events:</b>	<b>768 events</b>
background (wrong $\beta$ )	18
background ( $\pi$ 's from muons)	12.5
Internal neutrino interactions	14.6

## Total

**723 events**



## Prediction (Monte Carlo)

<b>989 <math>\pm</math> 17%</b>	
Bartol neutrino Flux	( $\pm$ 14%)
GRV-LO-94 cross section	( $\pm$ 9%)
Lohmann muon energy loss	( $\pm$ 5%)

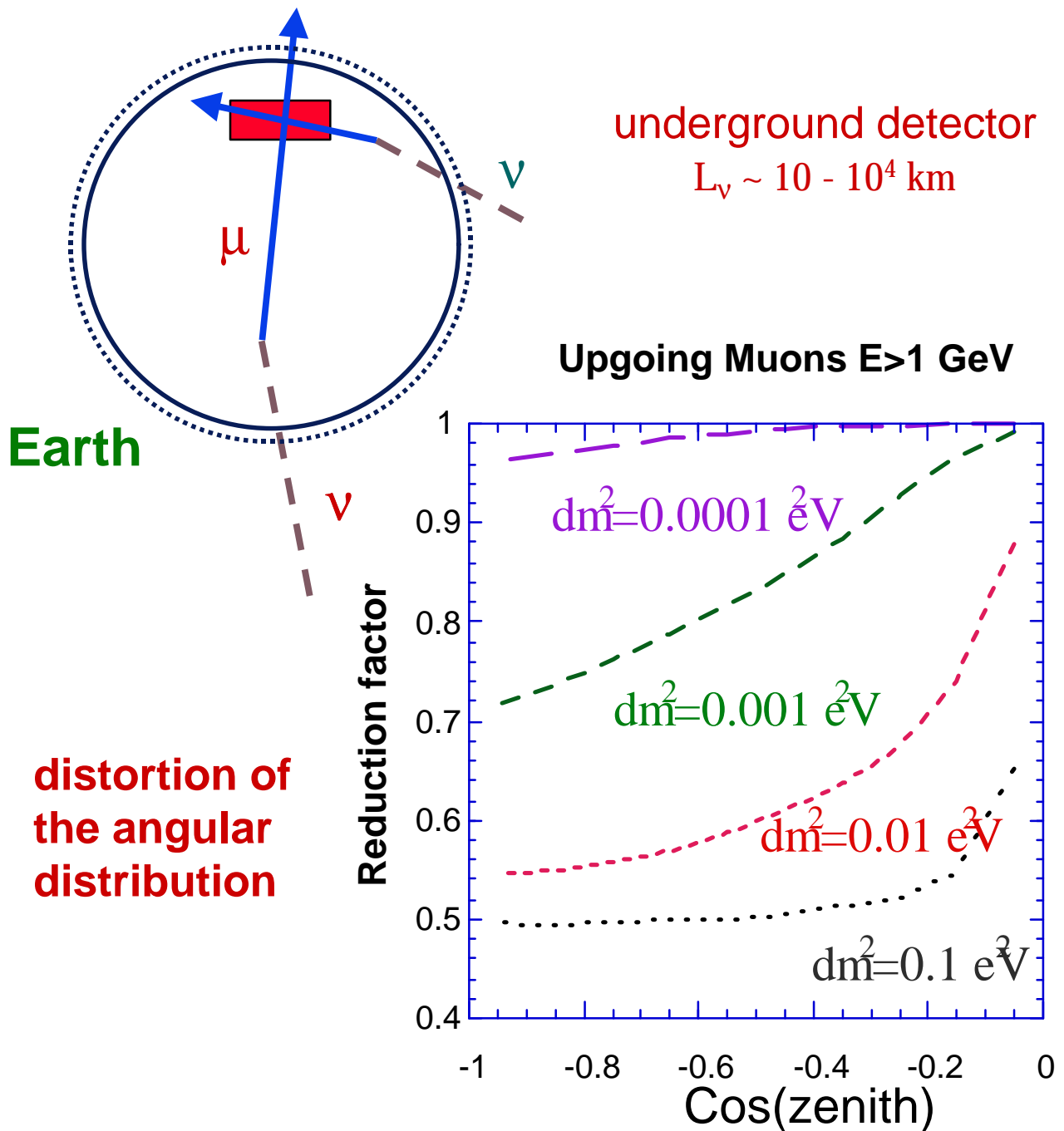
## data/prediction

**R = 0.731**

$\pm 0.028_{\text{stat}} \pm 0.044_{\text{syst}} \pm 0.124_{\text{theor}}$

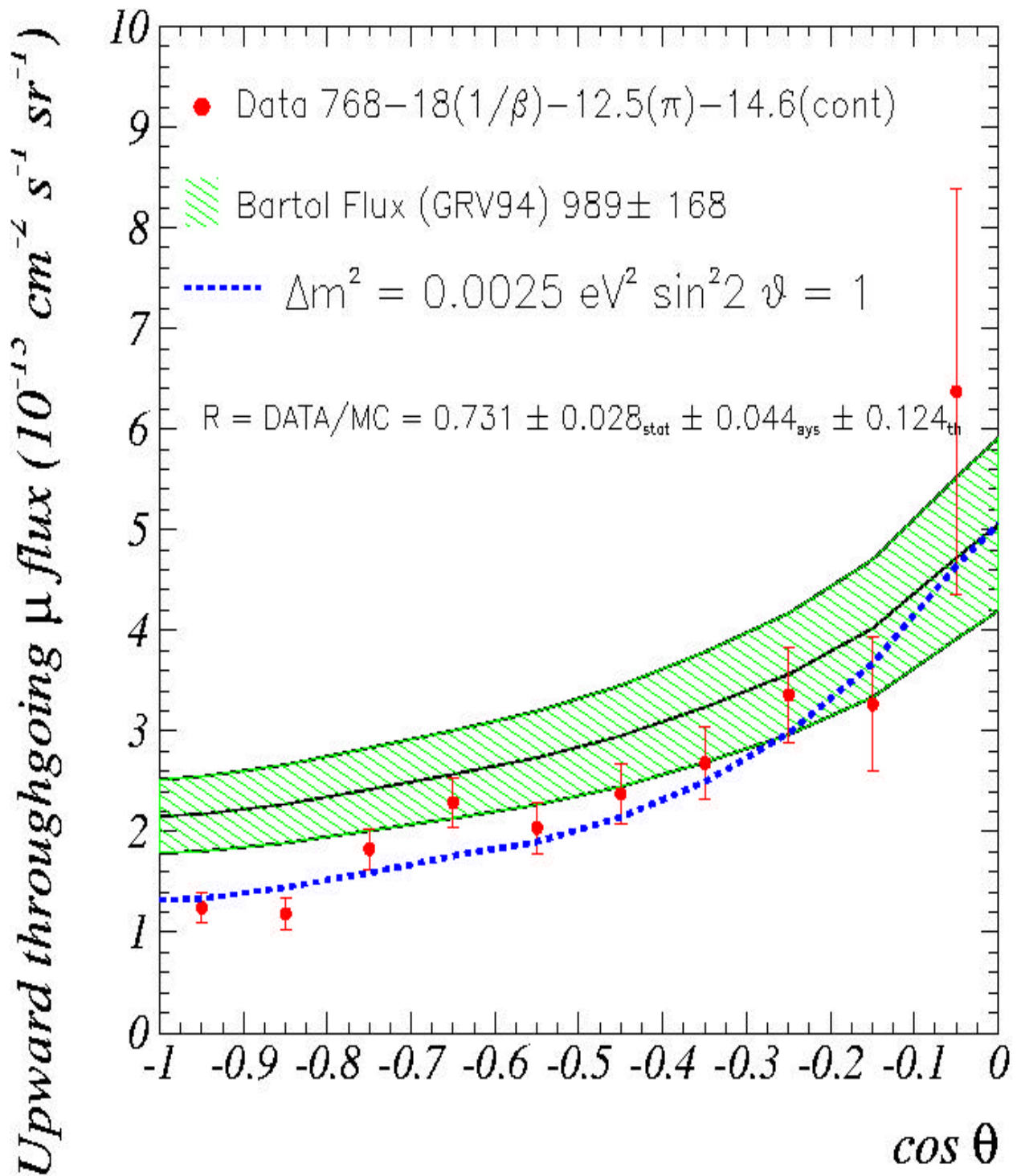
# Effects of $\nu_\mu$ oscillations on MACRO events

→ Flux reduction depending on zenith angle  
for the upward throughgoing events



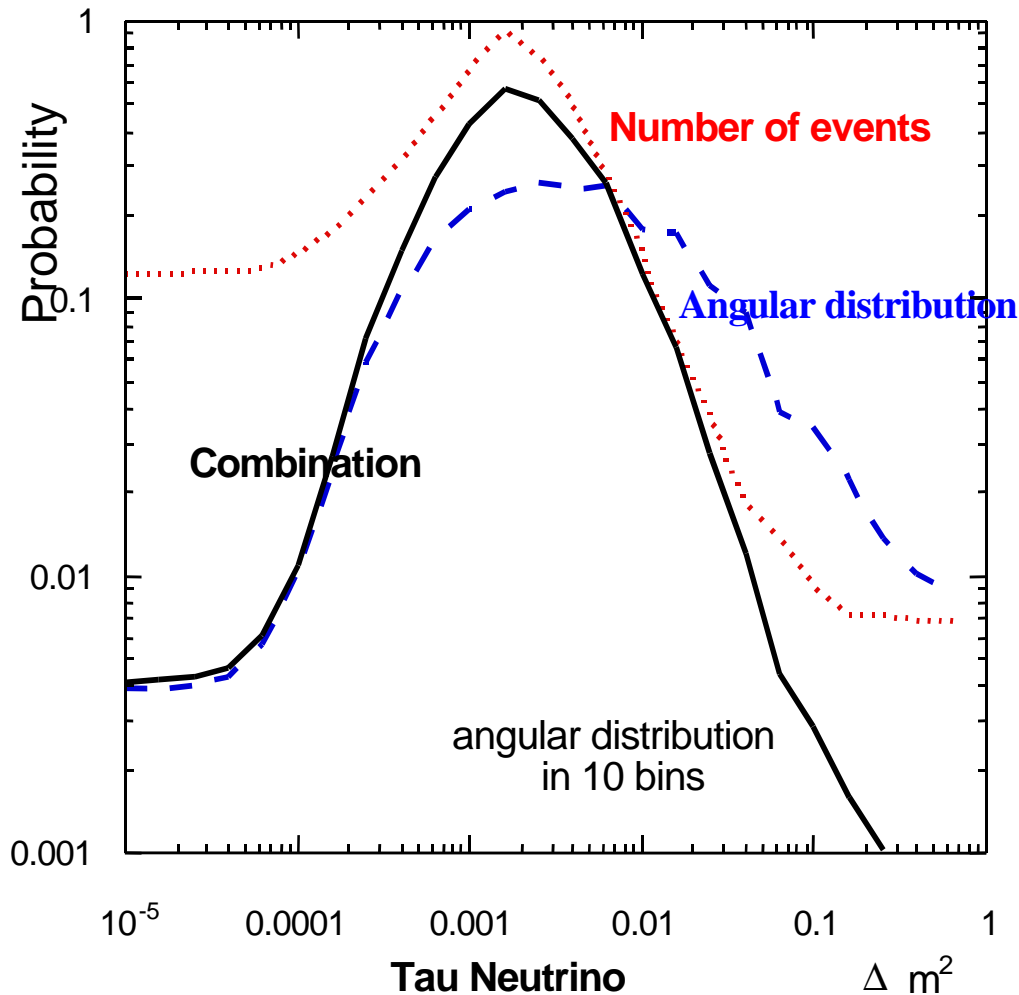
# Zenith angle distribution

$E_\mu > 1 \text{ GeV}$



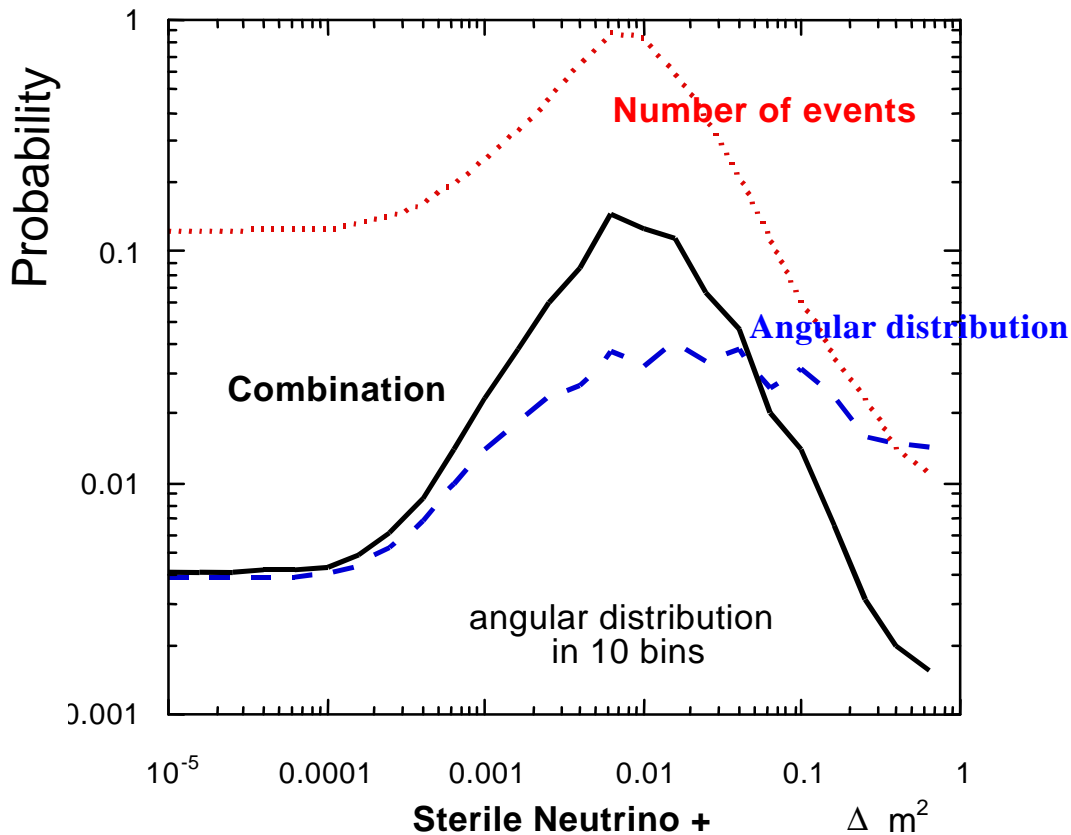
# Probabilities of $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations (for maximal mixing)

Peak probability from the angular distribution: 26%  
from the combination: 57%



- The peak probability from the angular distribution agrees with the peak probability from the total number of events
- Probability for no-oscillation:  $\sim 0.4\%$

# Probabilities for sterile neutrino oscillations with maximum mixing

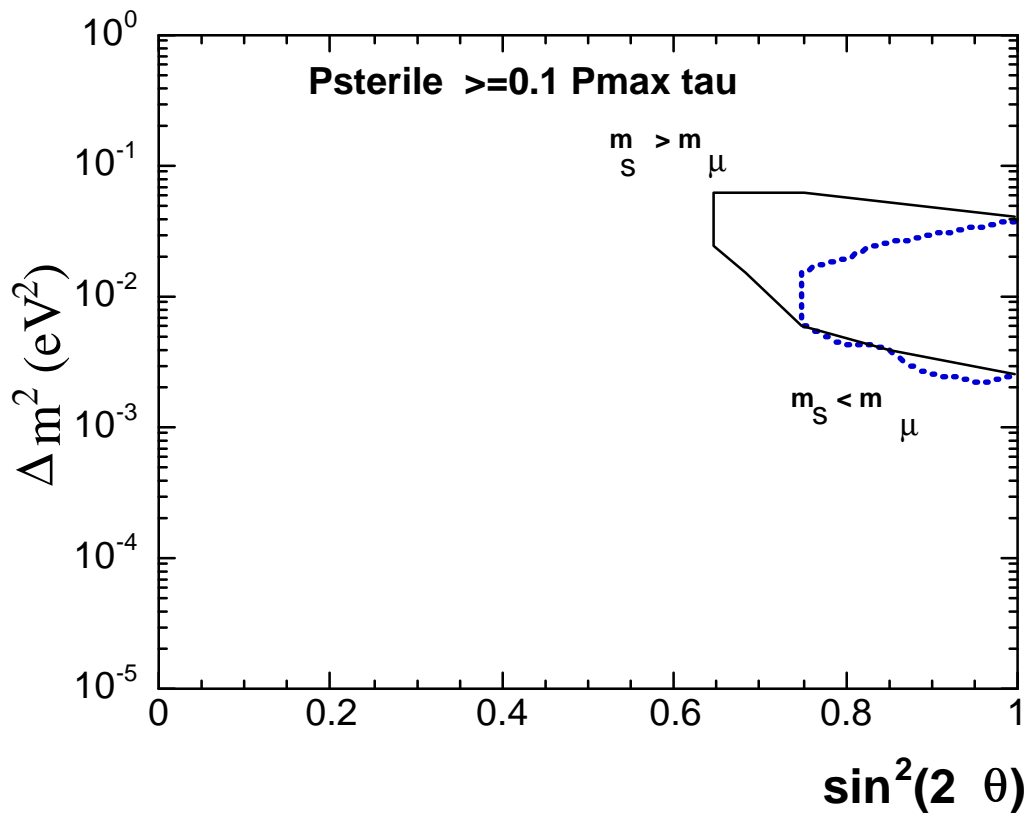


**Peak probabilities lower than that for tau neutrinos:**

- from the angular distribution: **4.1 %**
- from combination: **14.5 %**

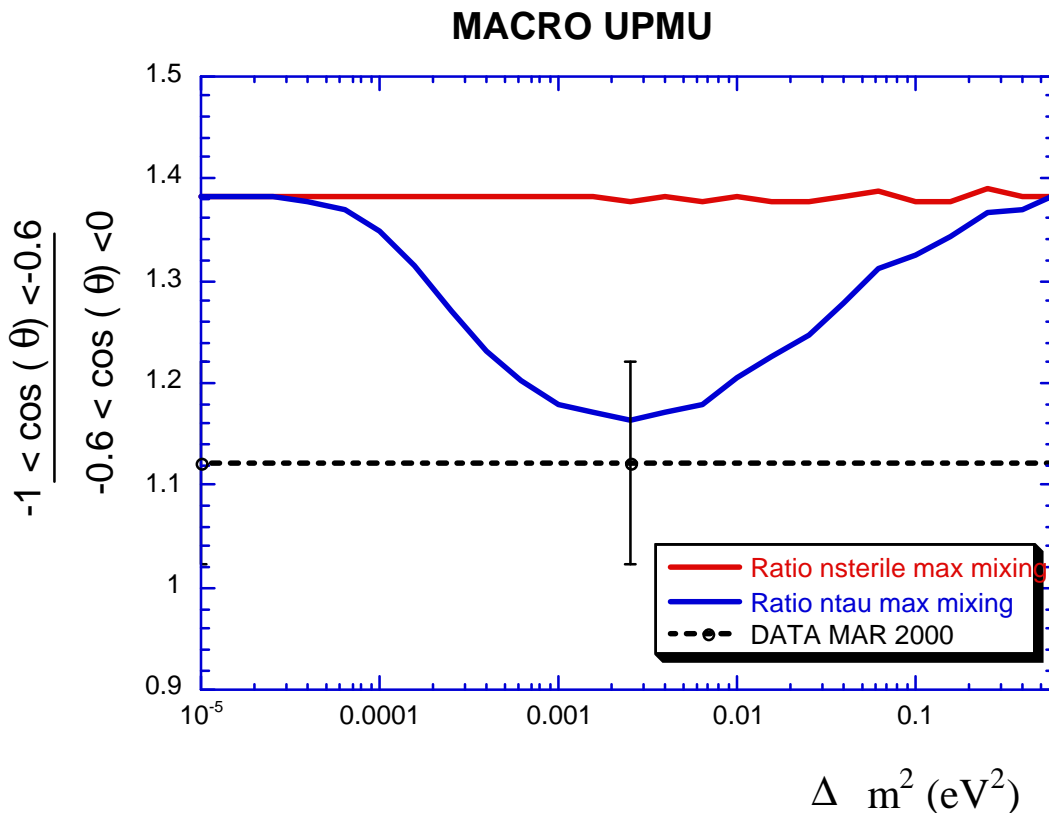
# Probabilities for sterile neutrino oscillations

ANGULAR DISTRIBUTION + Normalization



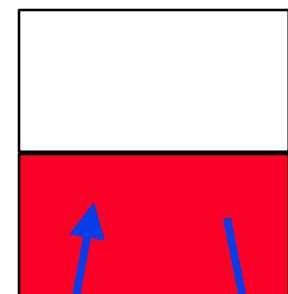
# Test of oscillations with the ratio vertical / horizontal

- Ratio (*Lipari- Lusignoli, Phys Rev D57 1998*) can be statistically more powerful than a  $\chi^2$  test:
  - 1) the ratio is sensitive to the sign of the deviation
  - 2) there is gain in statistical significance
- Disadvantage: the structure in the angular distribution of data can be lost.



P best Tau/ P best Sterile =  
15 (5% systematic in each bin)

**Measured value  $\sim 2 \sigma$  (and one sided) from the expected value for sterile neutrino**



Upward going  
stopping

Internal  
downgoing

## Internal Downgoing $\mu$ 's and Upward Going Stopping $\mu$ 's

from MC simulation:

- $E_\nu \sim 4$  GeV
- mixture
  - upward going stopping  $\mu$        $\sim 50\%$
  - internal downgoing  $\mu$        $\sim 50\%$
- $\sim 87\%$  from  $\nu_\mu$  –charged current events

**DATA - Background = 229 events**

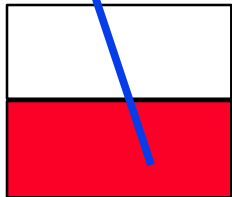


# Data vs Monte Carlo Predictions

## Low Energy Neutrino Events

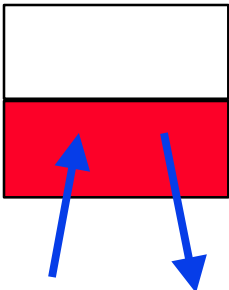
- $\Phi_\nu$ : Bartol  $\nu$  flux with geomagnetic cutoffs  
(error ~ 20%)
- $\sigma_\nu = \text{Q.E.} + 1\pi$  (Lipari et al., PRL74 (1995) 4384)  
+ DIS (GRV-LO-94 PDF) (error ~ 15%)
- $\epsilon(E_\mu, \theta_{\text{zenith}})$ : detector response and acceptance  
(systematic error ~ 10 %)

Internal upward going  $\mu$ 's

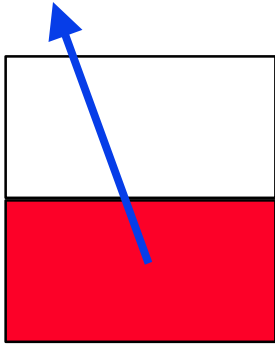


MC:  $247 \pm 25_{\text{sys}} \pm 62_{\text{theo}}$   
**DATA:**  $135 \pm 12_{\text{stat}}$

Internal down + Upgoing stop

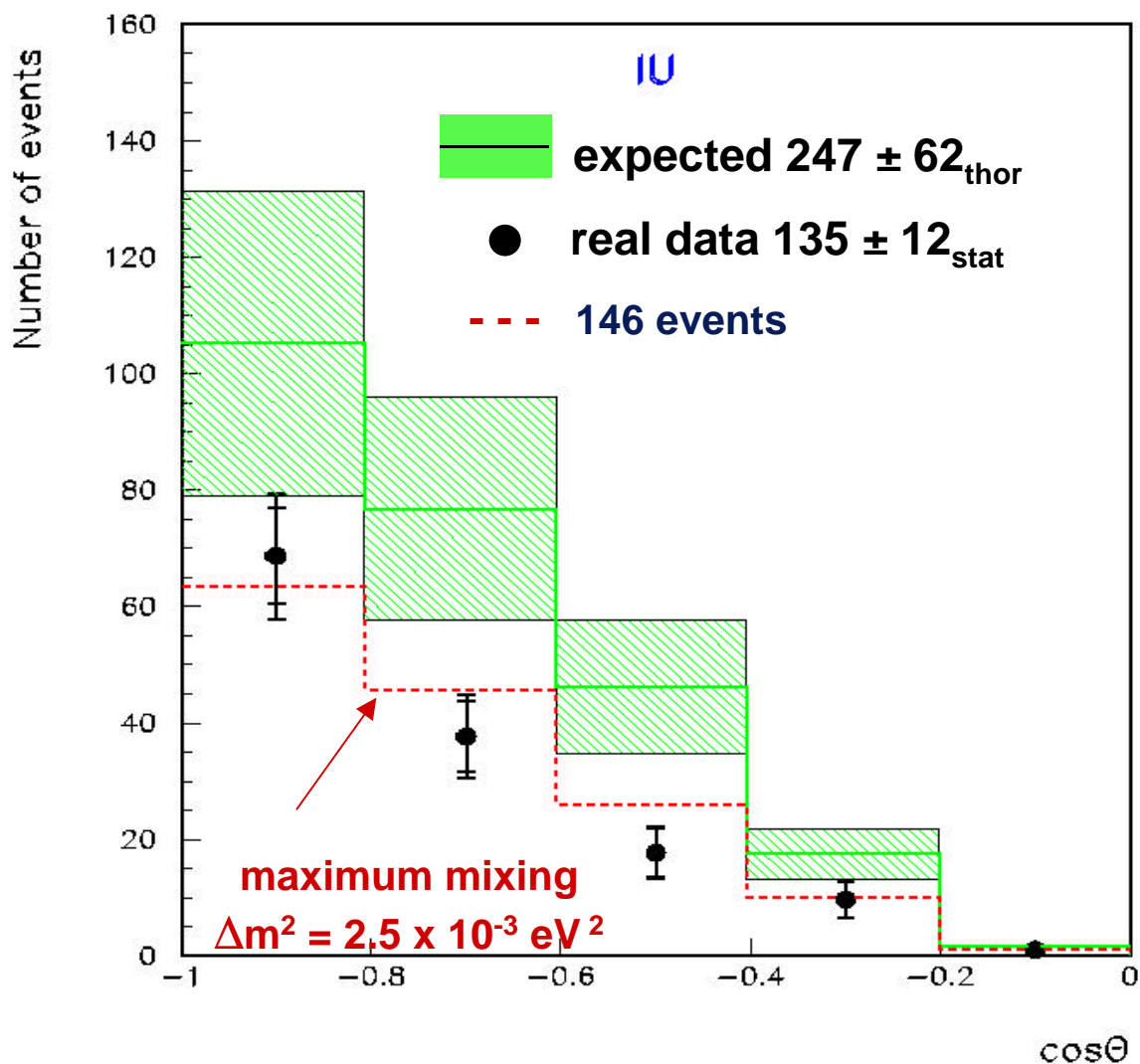


MC:  $329 \pm 33_{\text{sys}} \pm 82_{\text{theo}}$   
**DATA:**  $229 \pm 15_{\text{stat}}$

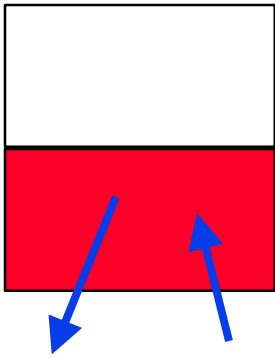


## Internal upward events

$$\frac{Data}{MC_{(IU)}} = 0.55 \pm 0.04_{stat} \pm 0.06_{sys} \pm 0.14_{th}$$

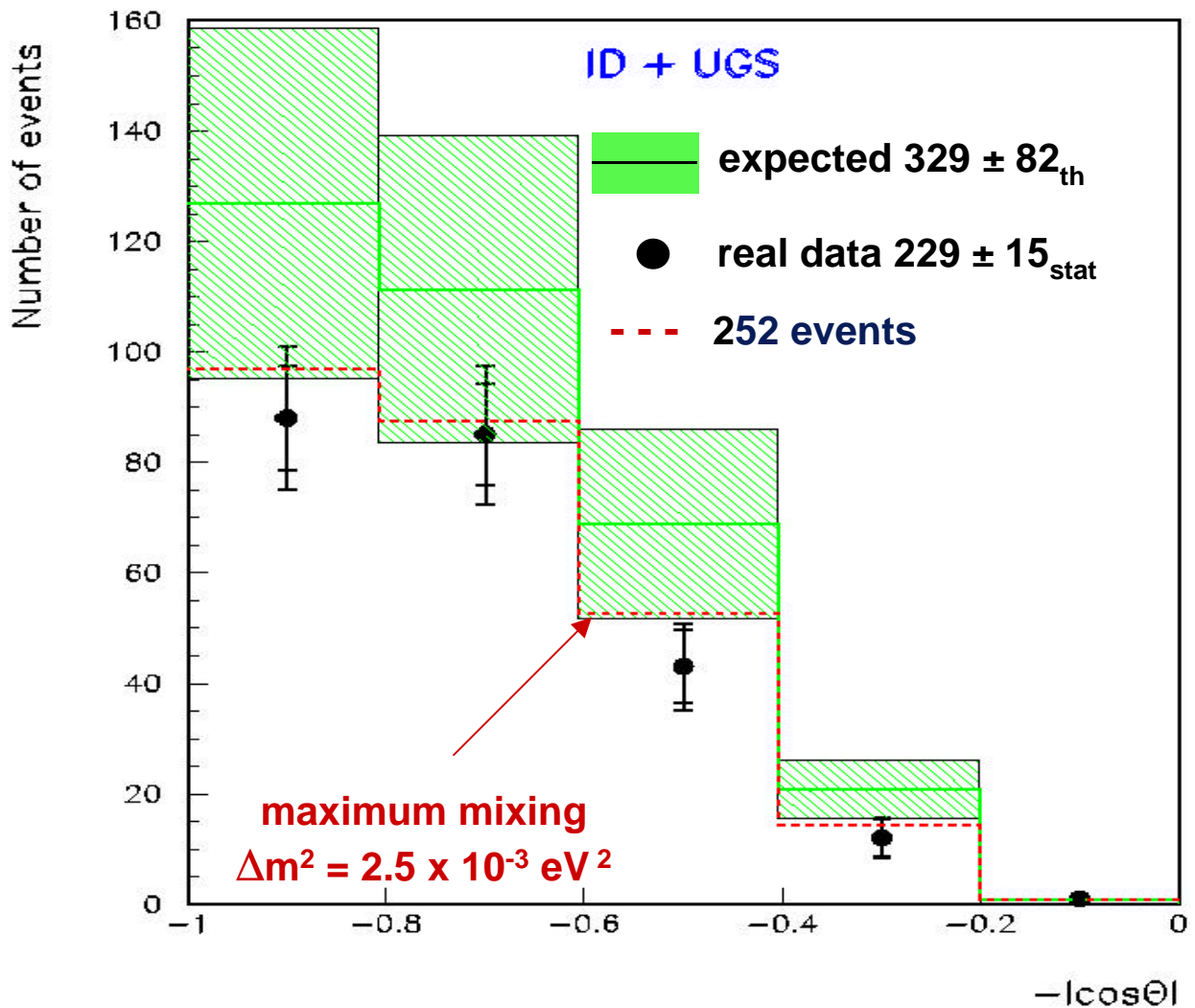


- Data consistent with a **constant deficit** in all **zenith angle** bins ( $\chi^2/\text{d.o.f.} = 3.1/4$  on shape)
- Probability for NO oscillations:  $\sim 4.3\%$  (one side)



# Internal downward + Upward stopping events

$$\frac{Data}{MC_{(ID+UGS)}} = 0.70 \pm 0.04_{stat} \pm 0.07_{sys} \pm 0.18_{th}$$



- Data consistent with a **constant deficit** in all **zenith angle** bins
- Probability for NO oscillations: ~ 12% (one side)

# Ratio of event types at low energy

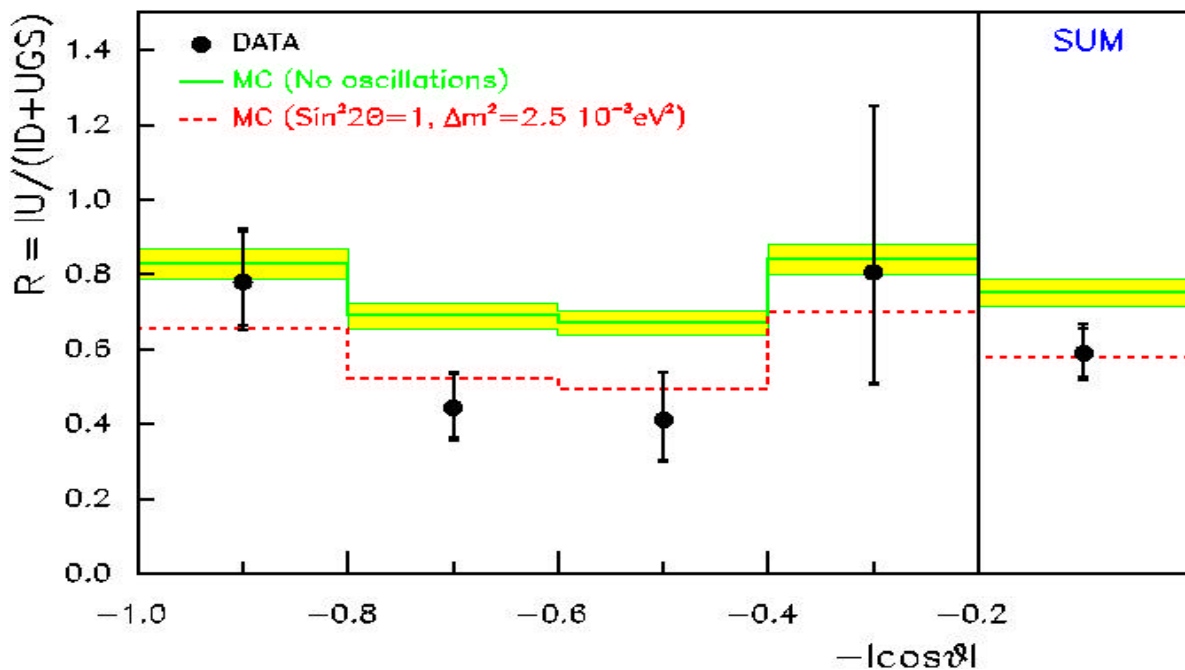
(Internal upward) / (Internal downward + Upward stopping)

- Most of the theor. uncertainties canceled (<5%)
- Systematic errors reduced (~6%)

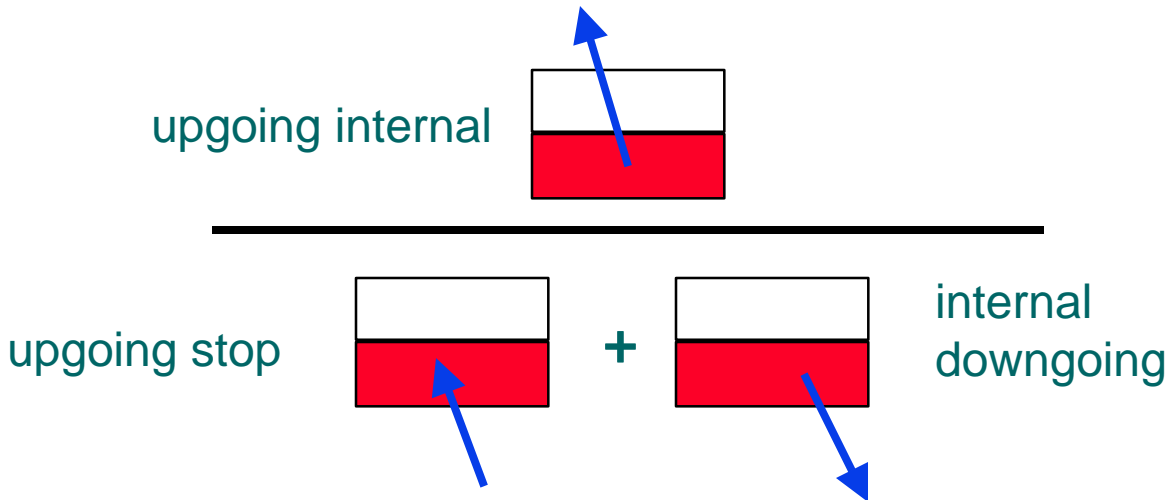
**Data:**  $R = 0.59 \pm 0.07_{\text{stat}}$

Expected (No oscillations):  $R = 0.75 \pm 0.04_{\text{sys}} \pm 0.04_{\text{th}}$   
**compatibility ~ 2.7%**

Expected  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations  $R = 0.58 \pm 0.03_{\text{sys}} \pm 0.03_{\text{th}}$   
(maximal mixing and  $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$ )



# Ratio of event types at low energy



- Most of the theor. uncertainties canceled < 5%
- Systematic errors reduced ~ 6%

**Data:**

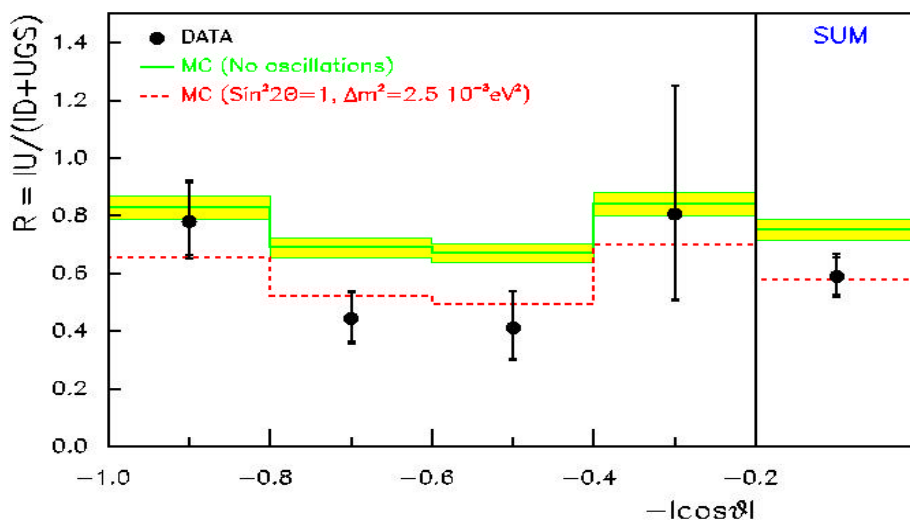
$$R = 0.59 \pm 0.07_{\text{stat}}$$

Expected (No oscillations):

$$R = 0.75 \pm 0.04_{\text{sys}} \pm 0.04_{\text{th}}$$

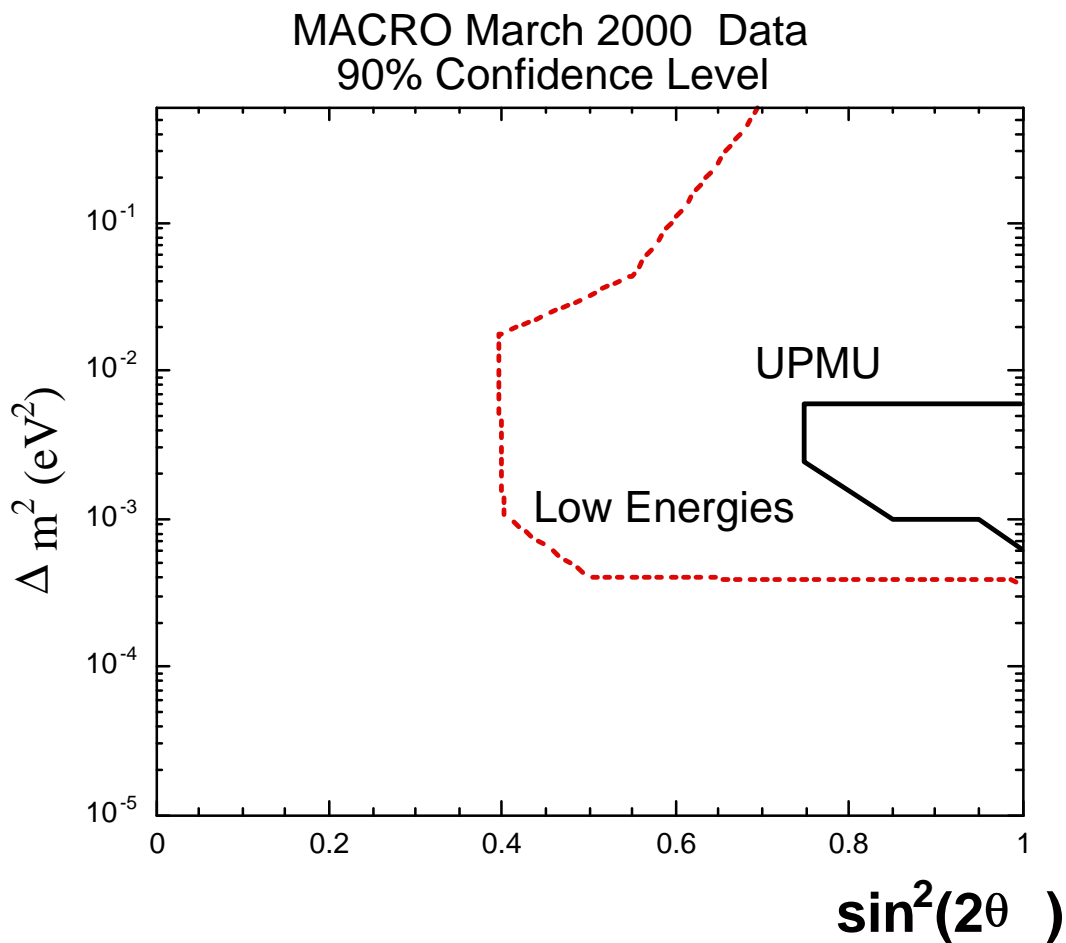
Expected  $\nu_\mu \rightarrow \nu_\tau$  oscillations

$$R = 0.58 \pm 0.03_{\text{sys}} \pm 0.03_{\text{th}}$$



**maximal mixing**  
 $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$

**MACRO  $\nu$  events:**  
**confidence level regions for**  
 **$\nu_{\mu} \rightarrow \nu_{\tau}$  oscillation**



# Summary and Conclusions

## ▪ High energy events:

- $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillation favoured with large mixing angle and  $\Delta m^2 \sim 2.5 \times 10^{-3} \text{ eV}^2$
- sterile  $\nu$  disfavoured at  $\sim 2 \sigma$  level

## ▪ Low energy events:

- event deficit with respect to expectation different for Internal upward, and internal downward + Upward going stopping
- no zenith angle distortion
- compatibility between result and expectation (no oscillations) with a probability  $< 3 \%$
- agreement with hypothesis of **oscillation** with large mixing angle and  **$10^{-3} < \Delta m^2 < 2 \times 10^{-2} \text{ eV}^2$**

## ▪ A consistent scenario arises from both high and low energy MACRO $\nu$ events