

# Search for a Sterile Neutrino in a 3+1 Framework using the Wire-Cell Inclusive Charged-Current $\nu_e$ Selection from MicroBooNE

---

June 10, 2022

The MicroBooNE collaboration  
[microboone\\_info@fnal.gov](mailto:microboone_info@fnal.gov)

## Abstract

A search for a sterile neutrino is being carried out in the MicroBooNE experiment within the 3+1 (three flavors of active neutrinos + one flavor of sterile neutrino) framework using neutrinos from the Booster Neutrino Beam (BNB) at a baseline of about 470 m with an average neutrino energy of 800 MeV. The sensitivity of this search is built upon high performance inclusive charged-current electron neutrino ( $\nu_e$ ) and muon neutrino ( $\nu_\mu$ ) event selections, which were utilized in the previous search for a low-energy excess (LEE) in the  $\nu_e$  energy spectrum at MicroBooNE. In this note, we present the results of a 3+1 oscillation fit considering both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance. Using the first three years' BNB data, no evidence of a light sterile neutrino was found and 95% C.L. exclusion limits were calculated in the 3+1 neutrino oscillation parameter space. Notably, the cancellation between  $\nu_e$  disappearance and  $\nu_e$  appearance oscillations leads to a reduced oscillation effect in the  $\nu_e$  energy spectrum resulting in a degeneracy of the oscillation parameters. Such a degeneracy is expected to be mitigated by including the neutrino events from the off-axis Neutrino from the Main Injector (NuMI) beam. The prospect of a 3+1 oscillation analysis using both BNB and NuMI is reported and it is expected to considerably improve the sensitivity to 3+1 neutrino oscillations.

## Contents

<b>1 Introduction</b>	<b>3</b>
<b>2 Analysis Approach</b>	<b>4</b>
2.1 Oscillation Model . . . . .	5
2.2 Oscillation Fit . . . . .	6
2.3 Setting an exclusion upper limit . . . . .	7
<b>3 Results from BNB</b>	<b>8</b>
<b>4 Prospect of using both BNB and NuMI data</b>	<b>19</b>
<b>5 Summary</b>	<b>22</b>
<b>Appendices</b>	<b>23</b>

## 1 INTRODUCTION

While most neutrino oscillation results are consistent with the three-neutrino framework (see Ref. [1, 2] among others), the existence of a light eV-scale sterile neutrino has been postulated to explain several experimental anomalies. These include i) the observation that calibrated  $\nu_e$  sources ( $^{51}\text{Cr}$  for GALLEX [3] and BEST [4],  $^{51}\text{Cr}$  and  $^{37}\text{Ar}$  for SAGE [5]) produced lower rates of measured  $\nu_e$  than expected in the three-neutrino framework; ii) the reactor anti-neutrino anomaly [6], which observed a deficit in the measured  $\bar{\nu}_e$  events relative to the expectation based on the reactor anti-neutrino flux calculations [7, 8]; iii) the Neutrino-4 [9] anomaly, which suggests reactor  $\bar{\nu}_e$  oscillations over distances of a few meters; and iv) the anomalous excess of  $\bar{\nu}_e$ -like events in LSND [10] and the excess of low-energy electron-like (LEE) events in MiniBooNE [11, 12]. The above experimental results could be explained by  $\nu_e/\bar{\nu}_e$  disappearance or  $\nu_e$  appearance considering light sterile neutrinos. It is worth noting that such an explanation to the reactor anti-neutrino anomaly has been disfavored by recent experimental measurements [13, 14] and improved reactor neutrino flux calculations [15, 16]. Nevertheless, there are significant challenges in explaining all available experimental results with a sterile neutrino oscillation model in a global fit [17]. It is important to clarify these experimental anomalies and the sterile neutrino, if discovered, would make a profound impact on not only particle physics, but also astrophysics and cosmology.

The MicroBooNE detector [18] is a 10.4 m long, 2.6 m wide, and 2.3 m high liquid argon time projection chamber (LArTPC), located on-axis along the Booster Neutrino Beam (BNB) at the Fermi National Accelerator Laboratory in Batavia, IL, USA. It consists of approximately 85 metric tons of liquid argon in the active volume for ionization charge detection along with an array of 32 photomultiplier tubes (PMTs) [19] for the detection of scintillation light. It sits at a distance of 468.5 m from the beryllium target of the BNB, 72.5 m upstream of the MiniBooNE detector. The recent distinct and complementary low-energy excess searches at MicroBooNE [20, 21, 22, 23], which aim to provide a definite check on the MiniBooNE anomaly, conclude that “the results are found to be consistent with the nominal  $\nu_e$  rate expectations from the Booster Neutrino Beam and no excess of  $\nu_e$  events is observed” [20], assuming a simple LEE template unfolded from the MiniBooNE excess. While these results suggest the MiniBooNE LEE has a non- $\nu_e$  origin, the current results may still be compatible with the hypothesis of a light sterile neutrino suggested by the set of experimental anomalies mentioned above.

In order to fully evaluate the possible existence of sterile neutrinos using the MicroBooNE data, a comprehensive and sensitive 3+1 (three flavors of standard model neutrinos + one

flavor of sterile neutrino) neutrino oscillation analysis is being carried out to determine the possible parameter space allowed for sterile neutrinos. Capitalizing on the high-performance neutrino selection and systematic uncertainties of the  $\nu_e$  and  $\nu_\mu$  event rate predictions in a recent MicroBooNE LEE analysis based on the Wire-Cell reconstruction [23], we present the 3+1 oscillation analysis results considering the oscillation effects from both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance for BNB neutrino events.

## 2 ANALYSIS APPROACH

This oscillation analysis utilizes the same inclusive  $\nu_e$  and  $\nu_\mu$  charged-current (CC) channels and  $\pi^0$  CC and neutral-current (NC) channels described in the published MicroBooNE Wire-Cell eLEE results [23]. The seven-channel strategy (fully contained (FC)  $\nu_e$  CC, partially contained (PC)  $\nu_e$  CC, FC  $\nu_\mu$  CC, PC  $\nu_\mu$  CC, FC  $\nu_\mu$  CC  $\pi^0$ , PC  $\nu_\mu$  CC  $\pi^0$ , and NC  $\pi^0$ ) was kept after including the oscillation effects, where the FC events are defined to be events with the reconstructed TPC activity which is fully contained within the fiducial volume and all non-FC events are defined as PC events. These seven channels are designed to be orthogonal to each other. The inclusive  $\nu_\mu$  CC channel excludes  $\nu_\mu$  CC  $\pi^0$  candidates<sup>1</sup>. This choice maintains the capability to apply data constraints across channels thereby reducing the systematic uncertainty in the 3+1 oscillation fit. The oscillation sensitivity is further improved by considering the oscillation effect event-by-event for the predicted signal and background events in all seven channels. The neutrino energy reconstruction<sup>2</sup> primarily follows a calorimetric method (sum of  $dE$ ), except where track-like particles were fully-contained and longer than 4 cm, in which case the track range was used to determine the particles' kinetic energy. The energy of an electromagnetic shower from electrons or gammas is estimated by scaling the total reconstructed charge associated with the shower with an overall energy-scale calibration factor. The details of the neutrino energy reconstruction can be found in Sec. III F of Ref. [23].

The BNB Run 1-3 data consist of  $6.4 \times 10^{20}$  protons on target (POT), which covers the first three years' data taking (about half of the full MicroBooNE data set) was used in this analysis. MicroBooNE standard overlay Monte-Carlo (MC) samples, in which the MicroBooNE GENIE tune [24] was applied and the simulation of neutrino interactions is overlaid with dedicated beam-off data, were used. The MC samples include 1) the intrinsic  $\nu_e$  and BNB  $\nu$  overlay MC samples for  $\nu_e/\nu_\mu$  disappearance events ( $\nu_e/\nu_\mu$  to  $\nu_e/\nu_\mu$  oscillation); 2) dedicated oscillation

---

<sup>1</sup>at least one  $\pi^0$  reconstructed in the final state

<sup>2</sup>For NC events, this method essentially reconstructs the energy transfer with an invisible outgoing neutrino.

samples for  $\nu_e$  appearance events ( $\nu_\mu$  to  $\nu_e$  oscillation), whose flux systematic uncertainty comes from the BNB  $\nu_\mu$  flux and the cross section and detector related systematic uncertainties from  $\nu_e$  interactions. In producing these samples the  $\nu_\mu$  appearance ( $\nu_e$  to  $\nu_\mu$  oscillation) was ignored since the intrinsic  $\nu_e$  to  $\nu_\mu$  ratio is about 0.6% and so,  $\nu_e$  to  $\nu_\mu$  oscillations would produce an insignificant difference to the total number of observed  $\nu_\mu$  events.

## 2.1 Oscillation Model

We use a 3+1 neutrino framework with the  $4 \times 4$  (4 flavors by 4 mass eigenstates) unitary PMNS matrix parameterized as below [25]

$$U = R_{34}(\theta_{34}, \delta_{34}) R_{24}(\theta_{24}, \delta_{24}) R_{14}(\theta_{14}, 0) R_{23}(\theta_{23}, 0) R_{13}(\theta_{13}, \delta_{13}) R_{12}(\theta_{12}, 0), \quad (1)$$

where  $R_{ij}(\theta_{ij}, \delta_{ij})$  denotes a counterclockwise rotation in the complex  $ij$ -plane through a mixing angle  $\theta_{ij}$  and a  $CP$  phase  $\delta_{ij}$ . Assuming the postulated fourth neutrino mass eigenstate is much heavier than the others ( $m_4 \gg m_3, m_2, m_1$ ), the short-baseline oscillation probability from  $\alpha$ -flavor to  $\beta$ -flavor neutrinos in vacuum can be expressed as

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \delta_{\alpha\beta} + (-1)^{\delta_{\alpha\beta}} \cdot \sin^2 2\theta_{\alpha\beta} \cdot \sin^2 \Delta_{41}, \quad (2)$$

where

$$\Delta_{41} \equiv \frac{\Delta m_{41}^2 L}{4E} = 1.267 \left( \frac{\Delta m_{41}^2}{\text{eV}^2} \right) \left( \frac{\text{MeV}}{E} \right) \left( \frac{L}{m} \right), \quad (3)$$

$\delta_{\alpha\beta}$  is the Kronecker delta, and  $\theta_{\alpha\beta}$  is defined as the effective mixing angle.

Table 1 shows the connection between the effective mixing angles and the PMNS matrix elements. There are four independent oscillation parameters:  $\Delta m_{41}^2$ ,  $\sin^2 \theta_{14}$ ,  $\sin^2 \theta_{24}$ , and  $\sin^2 \theta_{34}$  to describe the mixing of electron ( $e$ ) neutrinos, muon ( $\mu$ ) neutrinos, tau ( $\tau$ ) neutrinos, and postulated light sterile ( $s$ ) neutrinos.

$\sin^2 2\theta_{ee}$	$= \sin^2 2\theta_{14}$	$= 4(1 -  U_{e4} ^2) U_{e4} ^2$
$\sin^2 2\theta_{\mu\mu}$	$= 4 \cos^2 \theta_{14} \sin^2 \theta_{24} (1 - \cos^2 \theta_{14} \sin^2 \theta_{24})$	$= 4(1 -  U_{\mu4} ^2) U_{\mu4} ^2$
$\sin^2 2\theta_{\mu e}$	$= \sin^2 2\theta_{14} \sin^2 \theta_{24}$	$= 4 U_{\mu4} ^2 U_{e4} ^2$
$\sin^2 2\theta_{es}$	$= \sin^2 2\theta_{14} \cos^2 \theta_{24} \cos^2 \theta_{34}$	$= 4 U_{e4} ^2 U_{s4} ^2$
$\sin^2 2\theta_{\mu s}$	$= \cos^4 \theta_{14} \sin^2 2\theta_{24} \cos^2 \theta_{34}$	$= 4 U_{\mu4} ^2 U_{s4} ^2$

**Table 1:** Definition of the 3+1 effective oscillation angles using the parameterized PMNS matrix as shown in Eq. 1. ‘s’ denotes the postulated light sterile neutrinos.

## 2.2 Oscillation Fit

In the 3+1 oscillation parameter fit, we adopt the covariance matrix formalism to construct the  $\chi^2$  test statistic:

$$\chi^2 = (M - P(\boldsymbol{\theta}))^T \times Cov_{full}^{-1}(M, P(\boldsymbol{\theta})) \times (M - P(\boldsymbol{\theta})), \quad (4)$$

where  $\boldsymbol{\theta}$  represents the oscillation parameters,  $M$  and  $P$  are vectors of the measurements and the predictions for the seven channels, respectively. The  $Cov_{full}$  is the full covariance matrix:

$$Cov_{full} = Cov_{CNP}^{stat} + Cov_{MC,stat}^{sys} + Cov_{flux}^{sys} + Cov_{xsec}^{sys} + Cov_{det}^{sys}, \quad (5)$$

where  $Cov_{CNP}^{stat}$  is the statistical covariance matrix following the Combined-Neyman-Pearson (CNP) formalism [26] with the diagonal elements corresponding to  $3 / (1/M_i + 2/P_i)$  for the  $i$ th bin.  $Cov_{MC,stat}^{sys}$  is a diagonal covariance matrix accounting for the statistical fluctuation in making the MC central value prediction. The other three covariance matrices  $Cov_{flux}^{sys}$ ,  $Cov_{xsec}^{sys}$ , and  $Cov_{det}^{sys}$  are the covariance matrices corresponding to the neutrino flux, cross section<sup>3</sup>, and detector response systematic uncertainties, respectively. Details of the estimation of these covariance matrices can be found in Sec. V of Ref. [23].

The prediction  $P(\boldsymbol{\theta})$  and the covariance matrix  $Cov(M, P(\boldsymbol{\theta}))$  both depend on the oscillation parameters,  $\boldsymbol{\theta}$ , in the  $\chi^2$  minimization. Given a set of  $\boldsymbol{\theta}$ , the oscillation effect is applied as an event weight to each event based on its baseline, true neutrino energy, initial and final neutrino flavors, and the interaction type. For NC interactions, we apply an oscillation probability  $1 - P_{\nu_e \rightarrow \nu_s}$  for  $\nu_e$  events and  $1 - P_{\nu_\mu \rightarrow \nu_s}$  for  $\nu_\mu$  events. We fix  $\theta_{34}$  at zero ( $\cos^2 \theta_{34} = 1$ ) for NC events as it has a negligible impact on this analysis. Therefore, the actual oscillation parameters in the fit are  $\Delta m_{41}^2$ ,  $\sin^2 \theta_{14}$ , and  $\sin^2 \theta_{24}$ . The  $\sin^2 2\theta_{ee}$  and  $\sin^2 2\theta_{\mu e}$  mixing parameters obtained from the 3+1 oscillation analysis are converted from the two mixing angles following the formula shown in table 1. We vary the systematic uncertainty  $Cov^{sys}(P(\boldsymbol{\theta}))$  for different oscillation parameters via the fractional covariance matrix approach “ $Cov_{ij} = P_i \cdot F_{ij} \cdot P_j$ ” where  $P$  is the prediction as a function of oscillation parameters and  $F$  is the fractional covariance matrix, which is decoupled from the total number of events in each bin.  $F$  was found to have a negligible dependence on the oscillation parameters and is assumed to be constant in our calculation.

<sup>3</sup>Includes both neutrino-argon scattering cross section and final-state hadron-argon interaction cross section.

### 2.3 Setting an exclusion upper limit

In the absence of a discovery of new physics beyond the  $3\nu$  hypothesis, the CLs method [27, 28] is used to set the limits of oscillation parameters based on the  $\chi^2$  defined in Eq. 4. The CLs method is essentially a two-hypothesis test that compares a  $4\nu$  hypothesis to the  $3\nu$  hypothesis. We use the test statistic

$$\Delta\chi^2(x) = \chi_{4\nu}^2(x) - \chi_{3\nu}^2(x), \quad (6)$$

to determine if a  $4\nu$  hypothesis can be excluded at a certain confidence level.  $\chi_{4\nu}^2(x)$  is the  $\chi^2$  value of the measurement  $x$  with the prediction from a given set of  $4\nu$  oscillation parameters, and  $\chi_{3\nu}^2(x)$  is the  $\chi^2$  value corresponding to the  $3\nu$  hypothesis. The  $\text{CL}_s$  is then defined as

$$\text{CL}_s = \frac{\text{CL}_{s+b}}{\text{CL}_b}, \quad (7)$$

$$\text{CL}_{s+b} = \text{Prob}(\Delta\chi^2 \geq \Delta\chi_{\text{obs}}^2 | 4\nu), \quad (8)$$

$$\text{CL}_b = \text{Prob}(\Delta\chi^2 \geq \Delta\chi_{\text{obs}}^2 | 3\nu), \quad (9)$$

where  $\text{CL}_{s+b}$  and  $\text{CL}_b$  are the p-values (right-tail probability) of the observed  $\Delta\chi_{\text{obs}}^2$  when the  $4\nu$  or  $3\nu$  hypothesis is true. The region with  $\text{CL}_s \leq 1 - \alpha$  is excluded at the confidence level of  $\alpha$ . By throwing pseudo-experiments corresponding to a  $3\nu$  or  $4\nu$  hypothesis, the distribution of  $\Delta\chi^2$  as defined in Eq. 6 can be obtained, as illustrated in figure 1, therefore the  $\text{CL}_{s+b}$  and  $\text{CL}_b$  can be calculated for a given  $\Delta\chi_{\text{obs}}^2$  value either from data or a pseudo-experiment. We refer to this as the "frequentist CLs method".

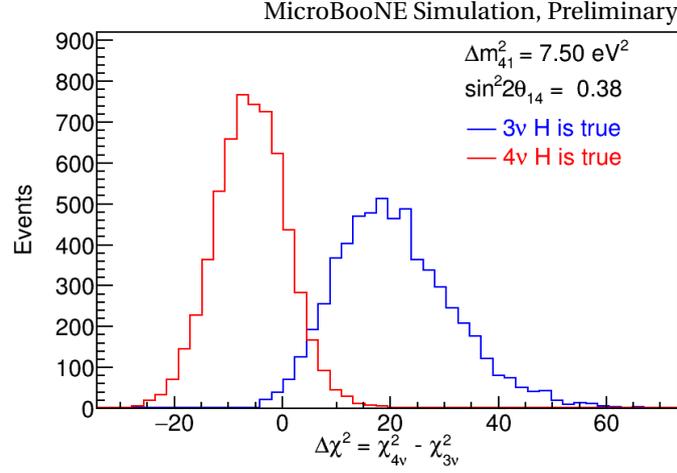
When the statistical and systematic uncertainties are small enough [29], the distribution of  $\Delta\chi^2$  as defined in Eq. 6 asymptotically follows a Gaussian distribution with  $\mu = \Delta\chi^2(x = \text{Asimov dataset})$  and  $\sigma = 2\sqrt{\mu}$ . Hence, Eq. 7 can be approximated to be

$$\text{CL}_s \approx \frac{1 + \text{Erf}\left(\frac{\Delta\chi^2(4\nu \text{ Asimov}) - \Delta\chi^2(x)}{\sqrt{8|\Delta\chi^2(4\nu \text{ Asimov})|}}\right)}{1 + \text{Erf}\left(\frac{\Delta\chi^2(3\nu \text{ Asimov}) - \Delta\chi^2(x)}{\sqrt{8|\Delta\chi^2(3\nu \text{ Asimov})|}}\right)}, \quad (10)$$

where in our case the Asimov dataset<sup>4</sup> [30] corresponds to the central value prediction for a given hypothesis without any fluctuations. We call this the "Gaussian  $\text{CL}_s$  method" [29]. The Gaussian approximation is invalid when the statistical and/or systematic uncertainty is

---

<sup>4</sup>The Asimov dataset is such that when one uses it to evaluate the estimator for all parameters, one obtains the true parameter values.



**Figure 1:**  $\Delta\chi^2$  distribution from pseudo-experiments assuming the  $4\nu$  hypothesis (red) or the  $3\nu$  hypothesis (blue) is true. The results are from the oscillation parameters at  $(7.50 \text{ eV}^2, 0.38)$  of  $(m_{41}^2, \sin^2 2\theta_{14})$ .

relatively large<sup>5</sup>. By default we use the frequentist  $CL_s$  method to calculate the sensitivity and set the exclusion upper limit for data. To illustrate the potential improvement of sensitivity combining both BNB and NuMI data (Sec. 4), the Gaussian  $CL_s$  method with Asimov datasets is used as it is computationally inexpensive and shows reasonable consistency with the frequentist  $CL_s$  method median sensitivity (e.g. figure 6 sensitivity versus figure 12 sensitivity with  $\sin^2\theta_{24}=0.005$ ).

### 3 RESULTS FROM BNB

In a 3+1 oscillation analysis, both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance are considered. The  $\nu_e$  disappearance can cancel the appearance of  $\nu_e$  events resulting in a degeneracy of the oscillation parameters as shown in the following equation

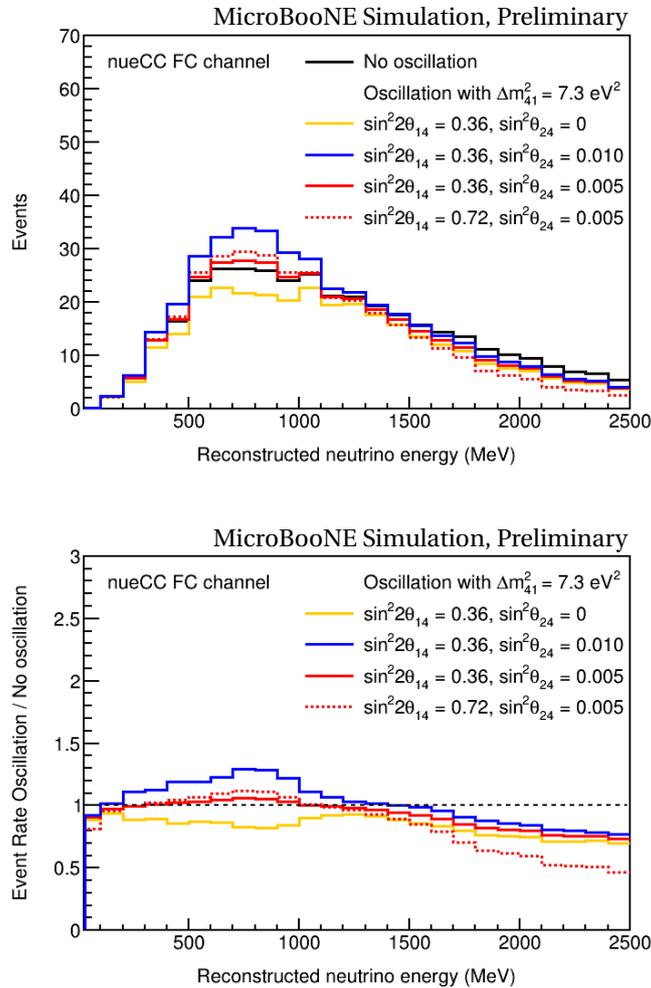
$$N_{\nu_e} = N_{\text{intrinsic } \nu_e} \cdot P_{\nu_e \rightarrow \nu_e} + N_{\text{intrinsic } \nu_\mu} \cdot P_{\nu_\mu \rightarrow \nu_e} \quad (11)$$

$$= N_{\text{intrinsic } \nu_e} \cdot \left[ 1 + (R_{\nu_\mu/\nu_e} \cdot \sin^2\theta_{24} - 1) \cdot \sin^2 2\theta_{14} \cdot \sin^2 \Delta_{41} \right] \quad (12)$$

where  $R_{\nu_\mu/\nu_e}$  is the ratio of intrinsic  $\nu_\mu$  and  $\nu_e$  events and this equation is for  $\nu_e$  of a given true neutrino energy. In the case of BNB only, the degeneracy of  $\sin^2\theta_{24}$  and  $\sin^2 2\theta_{14}$  happens when  $\sin^2\theta_{24}$  approaches  $1/\bar{R}_{\nu_\mu/\nu_e}^{\text{BNB}} \approx 0.005$ . Figure 2 illustrates the energy spectra of the  $\nu_e$  CC

<sup>5</sup>A large systematic uncertainty here means either large in magnitude or with large bin-to-bin correlations.

fully contained events at different values of the oscillation parameters corresponding to no oscillation effect and characteristic  $4\nu$  oscillation parameters referencing the Neutrino-4 best-fit values. The yellow and blue curves show obvious oscillation effects, but the red curve shows a weak oscillation effect below 1500 MeV due to the cancellation between  $\nu_e$  disappearance and appearance. Red dashed and red solid curves show that the  $\nu_e$  CC energy spectra are insensitive to oscillations when  $\sin^2\theta_{24}$  is at 0.005. Given this degeneracy around  $\sin^2\theta_{24}=0.005$ , the best-fit of  $(\Delta m_{41}^2, \sin^2\theta_{14}, \sin^2\theta_{24})$  could be especially biased when the best-fit  $\sin^2\theta_{24}$  is close to this degeneracy point which manifests as a local minimum of the  $\chi^2$  distribution in the oscillation parameter space. Meanwhile, the sensitivity around this region gets worse.



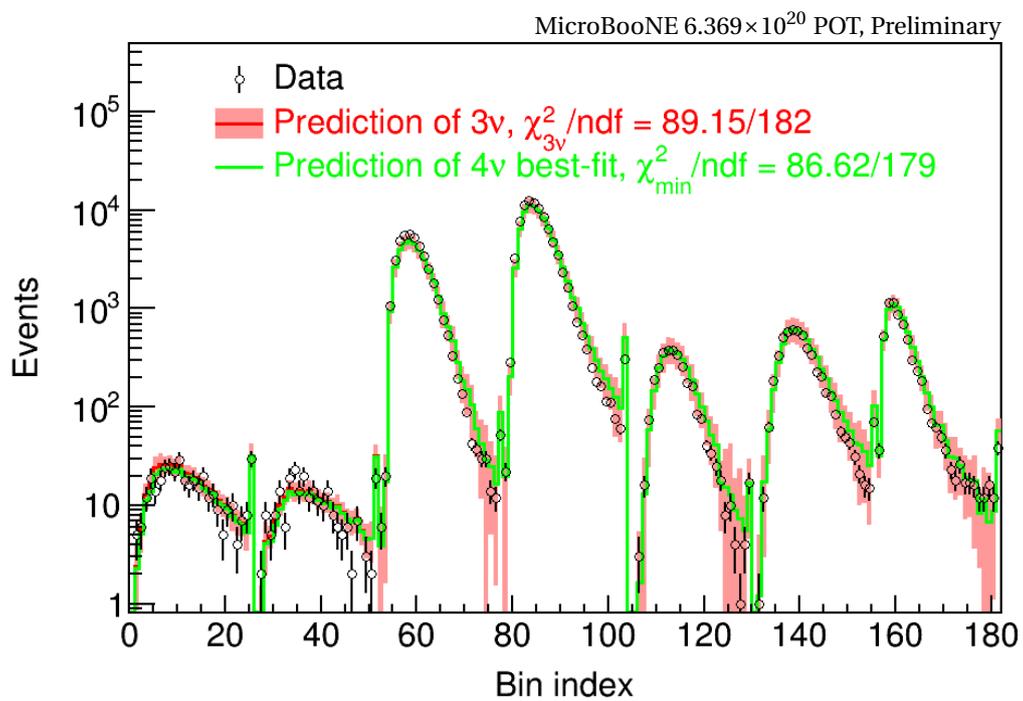
**Figure 2:** Energy spectra of the selected  $\nu_e$  CC fully contained events at different values of the oscillation parameters.

The simultaneous fit of  $(\Delta m_{41}^2, \sin^2\theta_{14}, \sin^2\theta_{24})$  using the 7 channels gives a data best-fit result of  $(1.295, 0.936, 0)$  with a  $\chi_{\min}^2(\text{data})/\text{ndf}=86.62/179$ . The best-fit value of  $\sin^2\theta_{14}$  corresponds to  $\sin^2 2\theta_{14} = 0.240$ . In comparison, assuming the  $3\nu$  hypothesis is true, the  $\chi_{3\nu}^2(\text{data})/\text{ndf}$  is equal to  $89.15/182$ . In this fit, the initial values of the oscillation parameters are obtained by scanning a 3-D grid of  $60 \times 60 \times 60$  points that are evenly distributed in a log-scale from  $0.1\text{-}100 \text{ eV}^2$ ,  $0.0001\text{-}1.0$ , and  $0.0001\text{-}1.0$  for  $\Delta m_{41}^2$ ,  $\sin^2\theta_{14}$ , and  $\sin^2\theta_{24}$ , respectively. The initial values correspond to the grid point which gives the minimal  $\chi^2$  relative to the others. The final best-fit values are obtained by float parameter minimization using MINUIT. Figure 3 and figure 4 show the energy spectra of the data, the  $3\nu$  prediction (after constraints from  $\nu_{\mu}$  CC, CC  $\pi^0$  and NC  $\pi^0$  channels), and the best-fit  $4\nu$  prediction for the 7 channels and the  $\nu_e$  CC channels, respectively. The compatibility between the data and the  $3\nu$  hypothesis is quantitatively calculated using the Feldman-Cousins approach [31]. The p-value of the data for the  $3\nu$  hypothesis is found to be 0.426 by comparing  $\chi_{3\nu}^2(\text{data}) - \chi_{\min}^2(\text{data}) = 2.53$  with the distribution of  $\Delta\chi_{\text{FC}}^2 (= \chi_{3\nu}^2 - \chi_{\min}^2)$ . Figure 5 shows the  $\Delta\chi_{\text{FC}}^2$  distribution from two thousands of pseudo-experiments assuming the  $3\nu$  hypothesis is true. The data result ( $\Delta\chi^2(\text{data}) = 2.53$ ) is also shown in the same figure. The p-value indicates a good compatibility between the data and the  $3\nu$  hypothesis, which are consistent within  $1\sigma$ . We'll report the  $\text{CL}_s$  exclusion upper limits from the data.

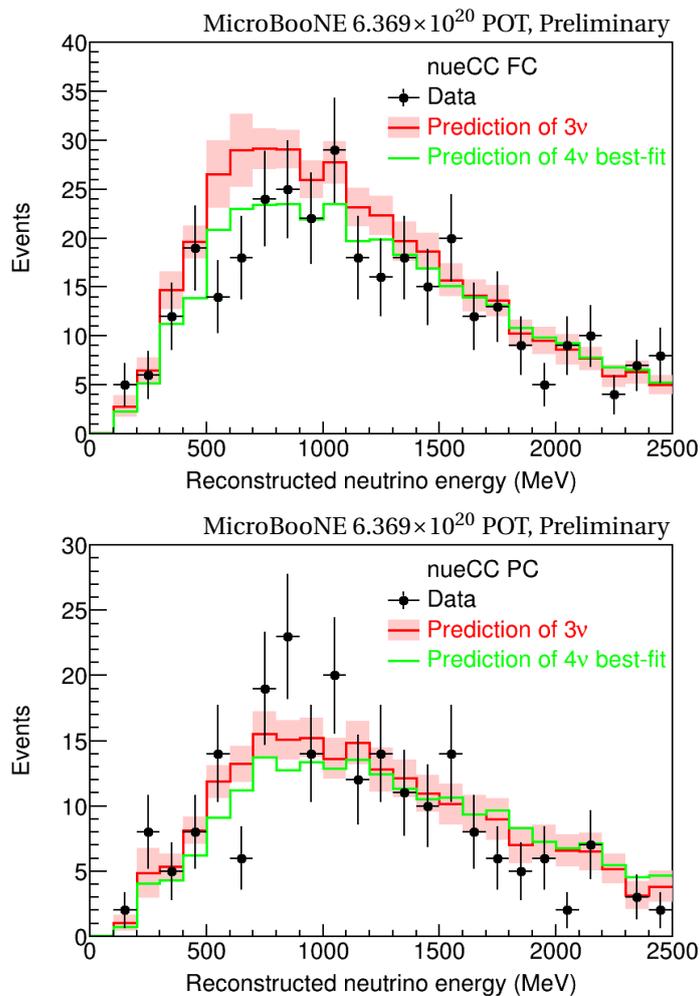
Bin index	Channel	Energy range (GeV)	Bin width (GeV)	Kinematic type
0-24	FC $\nu_e$ CC	0.0-2.5	0.1	neutrino energy
25	FC $\nu_e$ CC	$\geq 2.5$	-	neutrino energy
26-50	PC $\nu_e$ CC	0.0-2.5	0.1	neutrino energy
51	PC $\nu_e$ CC	$\geq 2.5$	-	neutrino energy
52-76	FC $\nu_{\mu}$ CC	0.0-2.5	0.1	neutrino energy
77	FC $\nu_{\mu}$ CC	$\geq 2.5$	-	neutrino energy
78-102	PC $\nu_{\mu}$ CC	0.0-2.5	0.1	neutrino energy
103	PC $\nu_{\mu}$ CC	$\geq 2.5$	-	neutrino energy
104-128	FC $\nu_{\mu}$ CC $\pi^0$	0.0-2.5	0.1	neutrino energy
129	FC $\nu_{\mu}$ CC $\pi^0$	$\geq 2.5$	-	neutrino energy
130-154	PC $\nu_{\mu}$ CC $\pi^0$	0.0-2.5	0.1	neutrino energy
155	PC $\nu_{\mu}$ CC $\pi^0$	$\geq 2.5$	-	neutrino energy
156-180	NC $\pi^0$	0.0-2.5	0.1	energy transfer
181	NC $\pi^0$	$\geq 2.5$	-	energy transfer

**Table 2:** Definition of bin indices in figure 3.

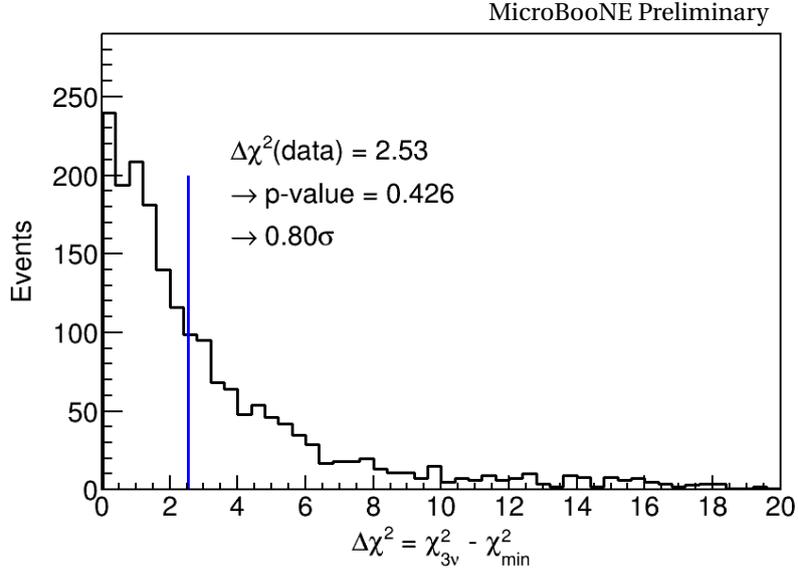
The results considering a full 3 active neutrinos + 1 sterile neutrino oscillation scenario are shown below. Figure 6 and figure 7 show the 95% C.L. exclusion limits in the 2D parameter



**Figure 3:** Energy spectra of the 7 channels: data (black circle), prediction of  $3\nu$  hypothesis (red curve), and prediction from the best fit of  $4\nu$  hypothesis (green curve). The error bar on each data point represents the statistical uncertainty. The red band on the prediction represents the systematic uncertainty. Definition of bin indices can be found in table 2.



**Figure 4:** Energy spectra of the  $\nu_e$  CC FC channel (top) and PC channel (bottom): data (black circle), prediction of the  $3\nu$  hypothesis (red curve) after constraints from  $\nu_\mu$  CC, CC  $\pi^0$  and NC  $\pi^0$  channels, and the prediction of the best fit  $4\nu$  hypothesis (green curve). The error bar on each data point represents the statistical uncertainty. The red band on the prediction represents the systematic uncertainty. These two figures are zoom-in figures from figure 3.



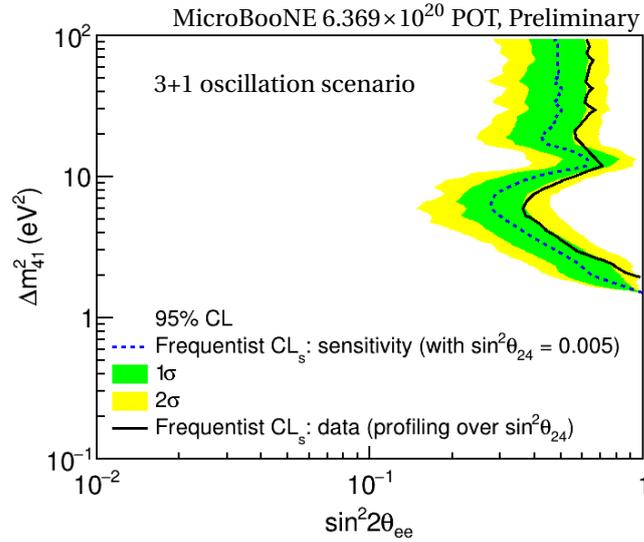
**Figure 5:** Distribution of  $\Delta\chi^2$  assuming  $3\nu$  hypothesis is true. The data  $\Delta\chi^2(\text{data})$  value is indicated by the blue vertical line. The data measurement has a p-value of 0.426 for the test statistic from the  $3\nu$  hypothesis pseudo-experiments, corresponding to only a  $0.80\sigma$  discrepancy.

space of  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{ee}$  and  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$ . Both sensitivity and data results are from the frequentist  $\text{CL}_s$  method. No obvious inconsistency was found in data compared to the sensitivity. In figure 6 the data result allows slightly more oscillation phase space due to the overall deficit we observed in the  $\nu_e$  CC channel. In figure 7 the sensitivity curve ends at  $\sin^2 2\theta_{\mu e} = 0.005$  because  $\sin^2 2\theta_{\mu e}$  is equal or less than  $\sin^2 \theta_{24}$  which is fixed at 0.005. The data exclusion limit is a 2D profiled result obtained by minimizing over  $\sin^2 \theta_{24}$  at each point of the 2D parameter space. The sensitivity result ( $1\sigma$  and  $2\sigma$  bands) corresponds to  $\sin^2 \theta_{24}$  fixed at 0.005, for which we expect to have the least sensitive 2D projected sensitivity<sup>6</sup> because of maximum parameter degeneracy at this  $\sin^2 \theta_{24}$  value.

Figure 8 shows the comparisons of our result with the allowed regions of Neutrino-4 [9] and gallium anomaly [4] results in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{ee}$  parameter space. The Neutrino-4 and gallium experiments have no  $\nu_\mu$  events from their neutrino sources, and their results correspond to a full 3+1 oscillation scenario that effectively rule out disappearance only measurements. For MicroBooNE, especially in this study, both  $\nu_e$  and  $\nu_\mu$  events from the BNB beam are used and the result (solid red curve; 2D profiling) corresponds to a full 3+1 oscillation analysis considering both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance. In comparison, the MicroBooNE  $\nu_e$  disappearance only result (dashed green curve) corresponding to  $\sin^2 \theta_{24} = 0$  is

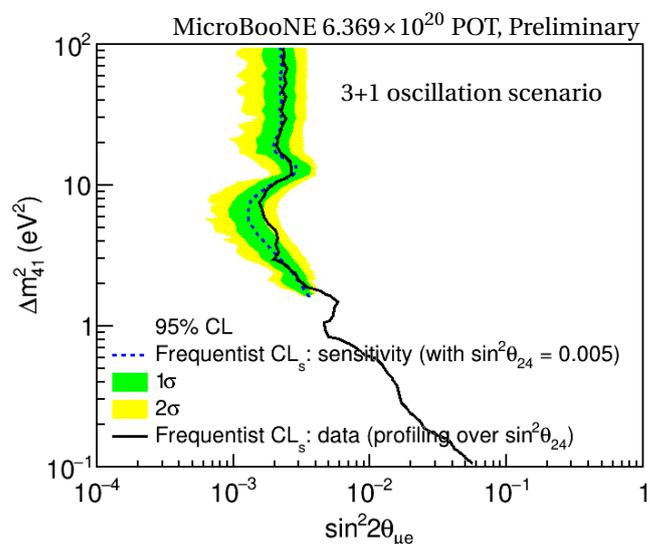
<sup>6</sup>The least sensitive result is more suitable to be compared to the 2D profiled result.

also presented, which provides a more stringent limit than the full 3+1 oscillation result, as expected. A fraction of the Neutrino-4 and gallium anomaly allowed regions are excluded by this analysis result at 95% CL. With the addition of NuMI data (see Sec. 4), the physics sensitivity can be extended covering the full Neutrino-4 and gallium anomaly  $2\sigma$  allowed region as indicated by the dashed magenta curve in figure 8. This BNB+NuMI sensitivity was calculated based on about 50% and 20% of the full MicroBooNE BNB and NuMI datasets, in the full 3+1 oscillation scenario with  $\sin^2\theta_{24}$  at 0.005.

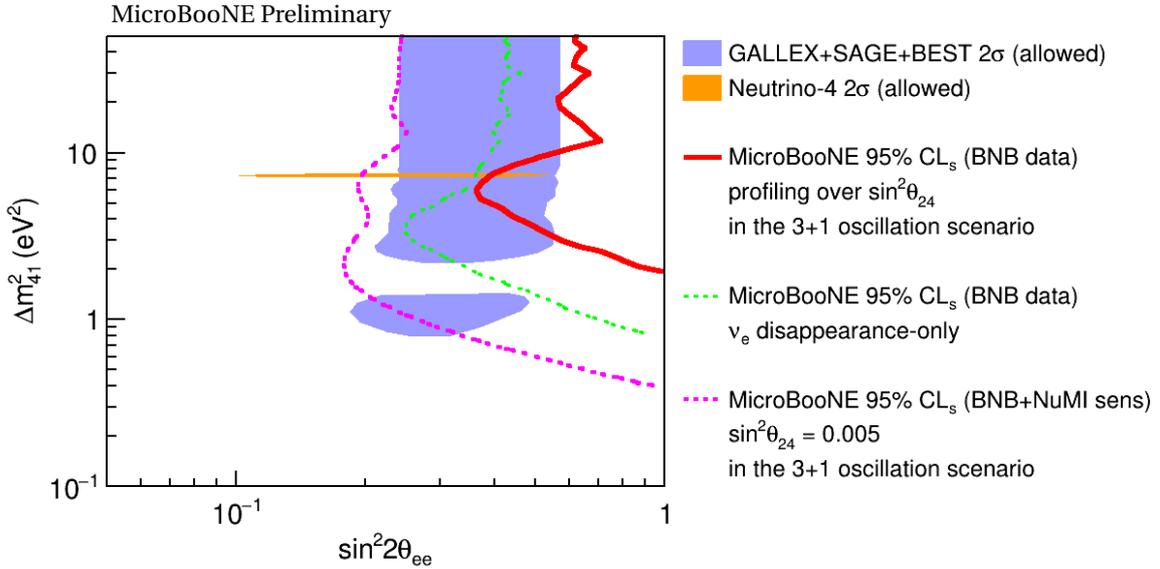


**Figure 6:** MicroBooNE 95% C.L. exclusion limit in the 2D parameter space of  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{ee}$ . Both sensitivity (fixed  $\sin^2\theta_{24}$  at 0.005) and data (2D profiling by minimizing over  $\sin^2\theta_{24}$ ) results are from the frequentist  $CL_s$  method in the full 3+1 neutrino oscillation scenario. For each  $\Delta m_{41}^2$ , the frequentist  $CL_s$  sensitivity curve corresponds to the median (50% quantile) value and the green and yellow bands correspond to  $(50 \pm 68.3/2)\%$  and  $(50 \pm 95.5/2)\%$  quantiles in the distribution of 95% C.L. exclusion limits from  $3\nu$  pseudo-experiments.

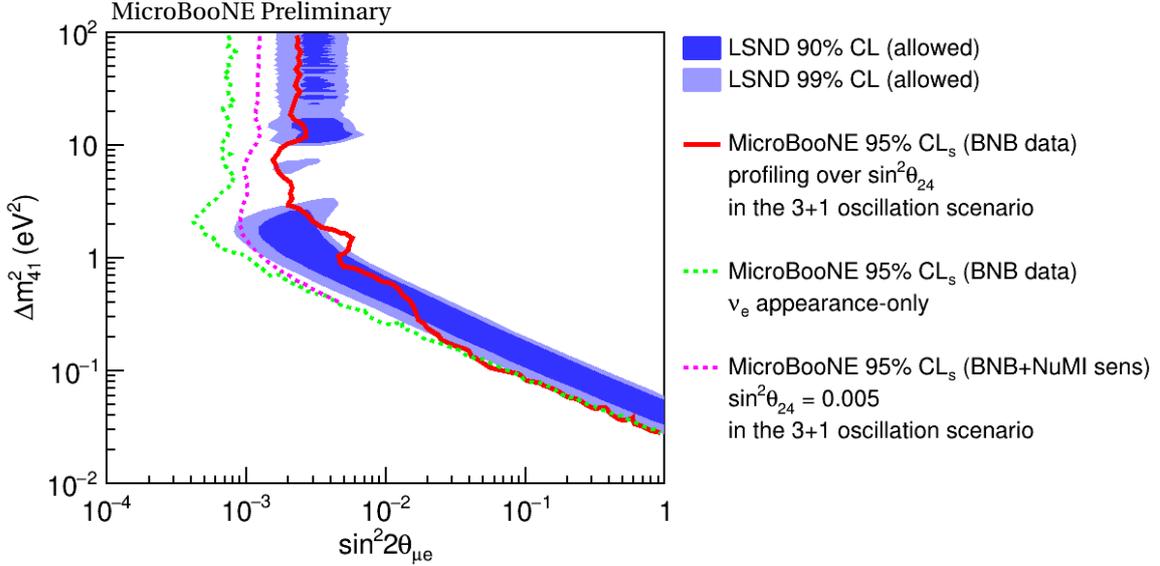
Figure 9 shows the comparisons of our result with the allowed region of LSND [10] in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$  parameter space. The LSND result is for a  $\bar{\nu}_e$  appearance only oscillation analysis with low intrinsic  $\bar{\nu}_e$  background, on which the impact of  $\bar{\nu}_e$  disappearance is insignificant. The MicroBooNE full 3+1 neutrino oscillation result (solid red curve) considering both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance are reported. In comparison, the  $\nu_e$  appearance only result (dashed green curve), which is not allowed in the 3+1 neutrino oscillation framework, is also presented. Part of the LSND allowed region is excluded by the full 3+1 oscillation analysis result at 95% CL. With the addition of NuMI data (see Sec. 4), the physics sensitivity can be extended covering the full LSND 99% CL allowed region as indicated by the dashed magenta curve in figure 9. This BNB+NuMI sensitivity was calculated based on about 50%



**Figure 7:** MicroBooNE 95% C.L. exclusion limit in the 2D parameter space of  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$ . Both sensitivity (fixed  $\sin^2 \theta_{24}$  at 0.005) and data (2D profiling by minimizing over  $\sin^2 \theta_{24}$ ) results are from the frequentist  $CL_s$  method in the full 3+1 neutrino oscillation scenario. For each  $\Delta m_{41}^2$ , the frequentist  $CL_s$  sensitivity curve corresponds to the median (50% quantile) value and the green and yellow bands correspond to  $(50 \pm 68.3/2)\%$  and  $(50 \pm 95.5/2)\%$  quantiles in the distribution of 95% C.L. exclusion limits from 3 $\nu$  pseudo-experiments. The sensitivity curve ends at  $\sin^2 2\theta_{\mu e} = 0.005$  because  $\sin^2 2\theta_{\mu e}$  is equal or less than  $\sin^2 \theta_{24}$  which is fixed at 0.005 in the sensitivity calculation.



**Figure 8:** MicroBooNE 95% confidence level frequentist CL<sub>s</sub> limits in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{ee}$  parameter space. The GALLEX+SAGE+BEST [4] and Neutrino-4 [9] 2 $\sigma$  allowed regions are shown in the shaded areas. The red solid curve represents the MicroBooNE 95% CL<sub>s</sub> exclusion limit (2D profiling by minimizing over  $\sin^2\theta_{24}$ ) in the full 3+1 neutrino oscillation scenario (both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance) using BNB Run 1-3 data. The green dashed curve represents the data exclusion limit in the  $\nu_e$  disappearance-only scenario as opposed to the full 3+1 oscillation result. The magenta dashed curve represents the MicroBooNE 95% CL<sub>s</sub> sensitivity in the full 3+1 oscillation scenario with a fixed  $\sin^2\theta_{24}$  at 0.005 when both BNB and NuMI data are combined. This sensitivity is calculated based on the Gaussian approximation CL<sub>s</sub> method which is computationally inexpensive and the result is similar to the frequentist CL<sub>s</sub> median sensitivity. BNB data correspond to the Run 1-3 data-taking period with a data exposure of  $6.369 \times 10^{20}$  POT, and NuMI data correspond to the Run 1 data-taking period with a data exposure of  $1.917 \times 10^{20}$  POT.



**Figure 9:** MicroBooNE 95% confidence level frequentist CL<sub>s</sub> limits in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$  parameter space. The LSND 90% and 99% CL allowed regions [10] are shown in the shaded areas. The LSND result is for a  $\bar{\nu}_e$  appearance only oscillation analysis with low intrinsic  $\bar{\nu}_e$  background, on which the impact of  $\bar{\nu}_e$  disappearance is insignificant. The red solid curve represents the MicroBooNE 95% CL<sub>s</sub> exclusion limit (2D profiling by minimizing over  $\sin^2 \theta_{24}$ ) in the full 3+1 neutrino oscillation scenario (both  $\nu_e$  appearance and  $\nu_e / \nu_\mu$  disappearance) using BNB Run 1-3 data. The green dashed curve represents the data exclusion limit in the  $\nu_e$  appearance only scenario (physically not allowed by the 3+1 oscillation framework) as opposed to the full 3+1 oscillation result. The magenta dashed curve represents the MicroBooNE 95% CL<sub>s</sub> sensitivity in the full 3+1 oscillation scenario of a fixed  $\sin^2 \theta_{24}$  at 0.005 when both BNB and NuMI data are combined. This sensitivity is calculated based on the Gaussian approximation CL<sub>s</sub> method which is computationally inexpensive and the result is similar to the frequentist CL<sub>s</sub> median sensitivity. BNB data correspond to the Run 1-3 data-taking period with a data exposure of  $6.369 \times 10^{20}$  POT, and NuMI data correspond to the Run 1 data-taking period with a data exposure of  $1.917 \times 10^{20}$  POT.

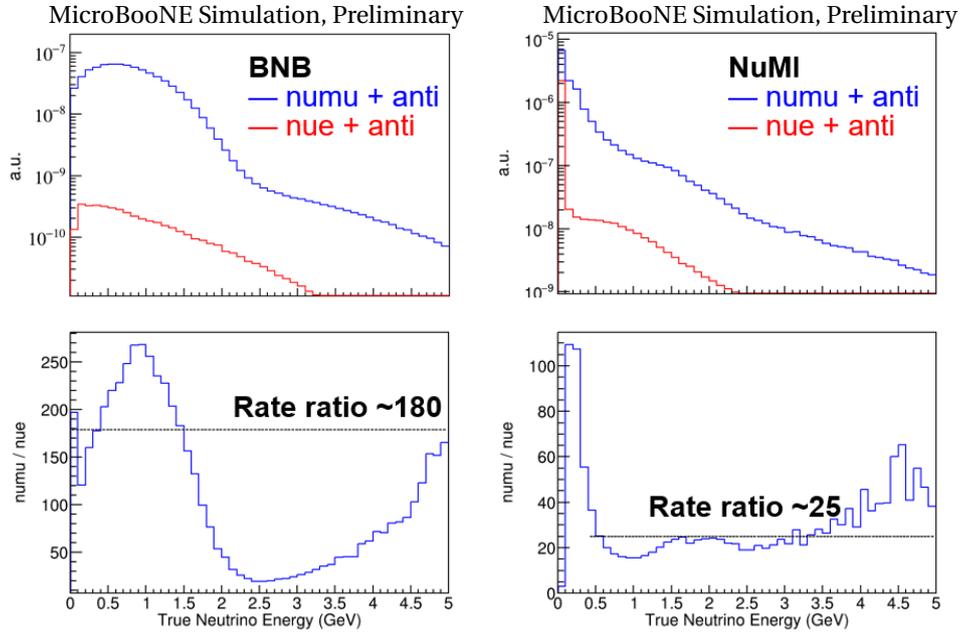
and 20% of the full MicroBooNE BNB and NuMI datasets, in the full 3+1 oscillation scenario with  $\sin^2\theta_{24}$  at 0.005. With the same caveats with respect to the  $\nu_e$  appearance only versus full 3+1 oscillation analyses, the MicroBooNE 3+1 results also exclude parts of the MiniBooNE allowed regions [32] which assumes the origin of the low-energy excess is solely due to  $\nu_e$ .

Although the main results in this note are for the full 3+1 oscillation analysis, results for the simplified scenarios of  $\nu_e$  disappearance only,  $\nu_e$  appearance only, and  $\nu_\mu$  disappearance only can also be found in the appendix.

## 4 PROSPECT OF USING BOTH BNB AND NUMI DATA

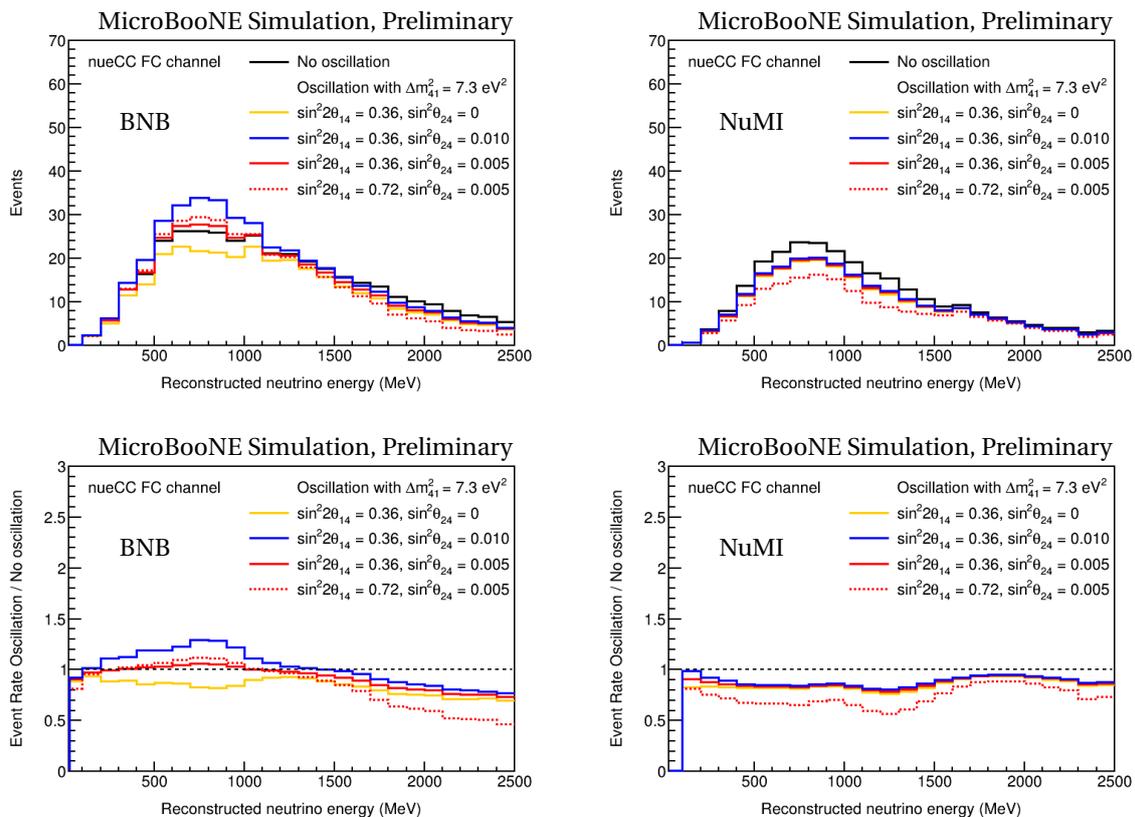
In this section, the prospect of the addition of NuMI (Neutrino at the Main Injector) data in this 3+1 oscillation analysis is presented.

Figure 10 shows that the intrinsic flux and  $\nu_\mu$  to  $\nu_e$  ratio in NuMI is quite different from that in BNB. The addition of NuMI events in the 3+1 oscillation helps to break the degeneracy of the oscillation parameters. Figure 11 shows the energy spectra of the selected  $\nu_e$  CC fully contained events at different values of the oscillation parameters for BNB and NuMI, respectively.



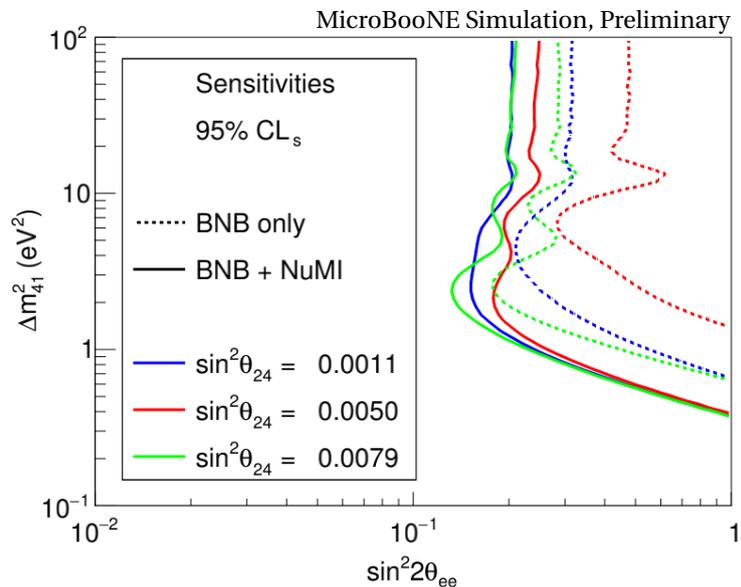
**Figure 10:** BNB and NuMI intrinsic  $\nu_e$  and  $\nu_\mu$  flux and their ratios as a function of true neutrino energy.

Because the NuMI degeneracy point is around  $\sin^2\theta_{24} \sim 0.04$  which is far away from the BNB degeneracy point, the addition of NuMI data can significantly reduce the BNB degeneracy. The BNB+NuMI sensitivity compared to the BNB only result is expected to have a weak dependence on  $\sin^2\theta_{24}$  rather than a local minimum (which leads to worse sensitivity) around  $\sin^2\theta_{24} \sim 0.005$ . The NuMI data can also provide more statistics of  $\nu$  events, particularly the  $\nu_e$  events. For example, the NuMI Run 1 sample provides a similar amount of  $\nu_e$  CC events as BNB Run 1-3. In addition, the NuMI neutrinos also extend the range of L/E values, which would increase sensitivity to a broader parameter space. Figure 12 shows the comparison of BNB-only (Run 1-3) and BNB+NuMI (BNB Run 1-3 and NuMI Run 1) sensitivity results for  $\sin^22\theta_{14}$  for different  $\sin^2\theta_{24}$  values. The BNB+NuMI sensitivity is significantly improved and



**Figure 11:** Energy spectra of the selected events from BNB (left) and NuMI (right) in the  $\nu_e$  CC FC channel at different values of oscillation parameters.

has largely reduced degeneracy around  $\sin^2\theta_{24} \sim 0.005$  compared to the BNB only result.



**Figure 12:** Sensitivities (Gaussian approximation CLs method) at 95% confidence level for the full 3+1 oscillation scenario ( $\nu_e$  appearance +  $\nu_e/\nu_\mu$  disappearance) at different  $\sin^2\theta_{24}$  values. Dashed curves correspond to BNB only results. BNB data correspond to the Run 1-3 data-taking period with a data exposure of  $6.369 \times 10^{20}$  POT, and NuMI data correspond to the Run 1 data-taking period with a data exposure of  $1.917 \times 10^{20}$  POT.

## 5 SUMMARY

In this note, we present the 3+1 sterile neutrino oscillation analysis results using the BNB seven-channel selections which were developed and applied in the Wire-Cell eLEE analysis [23]. We report the best-fit result for the three oscillation parameters ( $\Delta m_{41}^2$ ,  $\sin^2\theta_{14}$ ,  $\sin^2\theta_{24}$ ) that dictate short-baseline 3+1 neutrino oscillations. Considering the full 3+1 oscillation, the data result is consistent with the  $3\nu$  hypothesis within  $1\sigma$  following the Feldman-Cousins approach. Exclusion upper limits of the several effective mixing angles and the mass-squared difference for a postulated light sterile neutrino are calculated and compared to the sensitivity results. Data exclusion upper limits and sensitivity results in the simplified  $\nu_e$  disappearance or  $\nu_e$  appearance scenarios are shown as well which are generally more stringent than the full 3+1 oscillation results. In summary,

- The MicroBooNE data are consistent with the  $3\nu$  hypothesis and provide no evidence for a sterile neutrino.
- The MicroBooNE exclusion limits cover a large fraction of sterile neutrino parameter space allowed by results from other experiments.

In the future, we plan to add the NuMI events and do a combined BNB+NuMI oscillation analysis to mitigate the degeneracy of oscillation parameters. More data from the BNB and/or NuMI will also be included with other advancements in neutrino flavor identification and energy reconstruction to further improve the physics sensitivity to 3+1 neutrino oscillations.

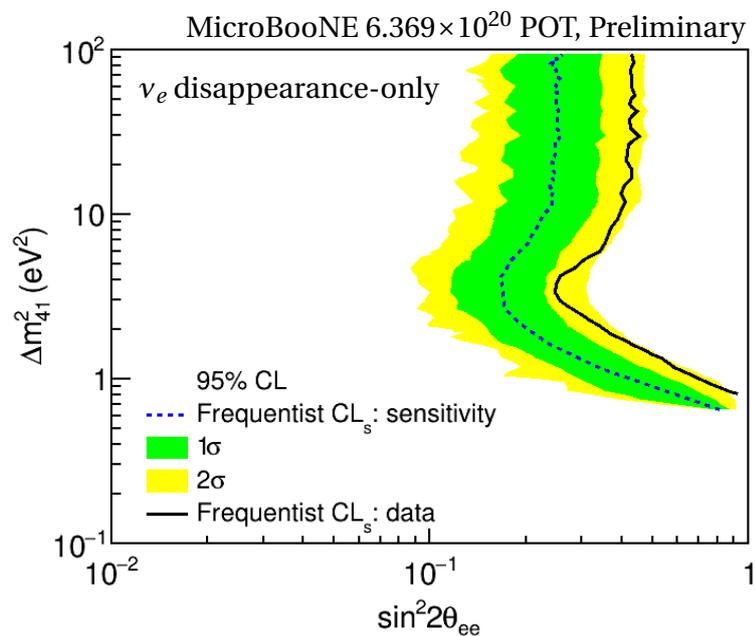
# Appendices

## RESULTS FROM SIMPLIFIED OSCILLATION SCENARIOS

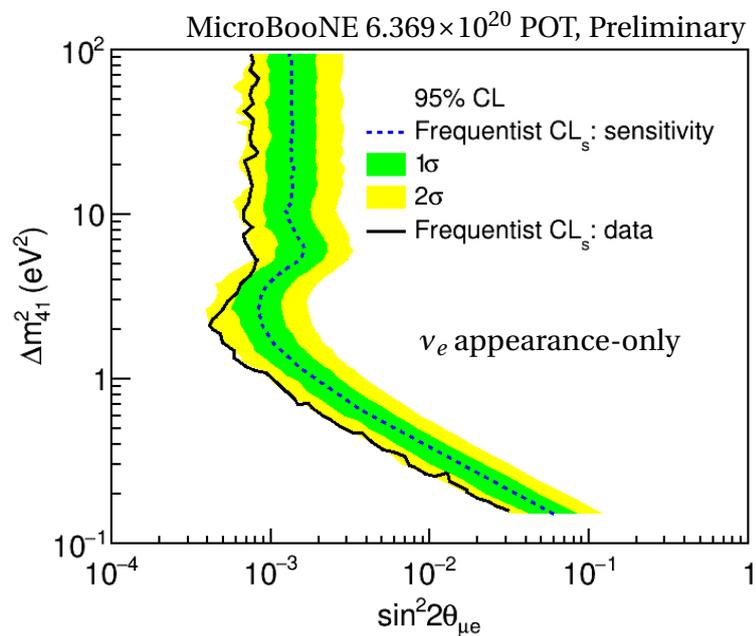
In this appendix, the  $\nu_e$  disappearance only ( $\nu_e$  to  $\nu_e$  oscillation),  $\nu_e$  appearance only ( $\nu_\mu$  to  $\nu_e$  oscillation), and  $\nu_\mu$  disappearance only ( $\nu_\mu$  to  $\nu_\mu$  oscillation) results are reported. The results can be compared to those in Ref. [33] that were calculated based on the MicroBooNE LEE data release. To first order, these results should be consistent ignoring the impact of a precise  $L/E$  determination and energy smearing for each MC event. Moreover, we set the exclusion upper limit using a frequentist approach to properly take into account low statistics and large systematic uncertainties in the neutrino selections.

Figure 13 shows the exclusion and sensitivity contours in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{ee}$  parameter space for the scenario of  $\nu_e$  disappearance only. Figure 14 shows the exclusion and sensitivity contours in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$  parameter space for the scenario of  $\nu_e$  appearance only. Because there is an overall deficit (less than  $2\sigma$ ) in the data for  $\nu_e$  CC channels compared to the prediction of the  $3\nu$  hypothesis, the  $\nu_e$  disappearance data exclusion limit in figure 13 is weaker than the sensitivity and the  $\nu_e$  appearance data exclusion limit in figure 14 is more stringent than the sensitivity. Figure 15 shows the comparisons of our result with the allowed region of MiniBooNE [32] in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$  parameter space. The MiniBooNE result is for a  $\nu_e$  appearance only oscillation analysis.

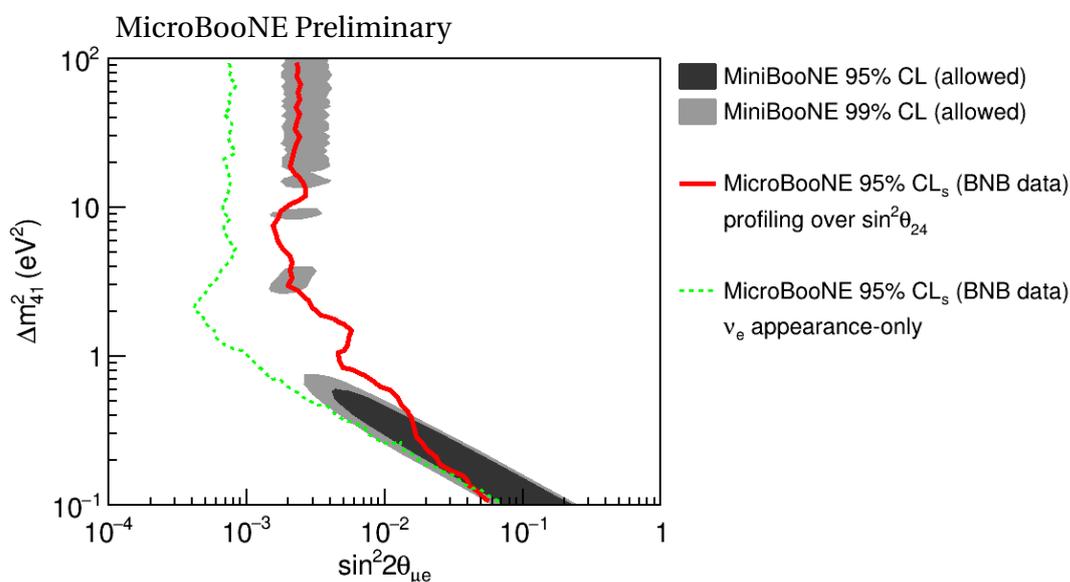
Figure 16 shows the exclusion and sensitivity contours in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu\mu}$  parameter space for the scenario of  $\nu_\mu$  disappearance only. Because of the relatively large systematic uncertainty (weak constraints from the other channels), the frequentist method produces a significantly different result from that using the Wilks' theorem or other approximations. The data-to-MC shape difference in the inclusive  $\nu_\mu$  CC channel can allow for more  $4\nu$  oscillation parameter space leading to a much weaker data exclusion limit than the median sensitivity.



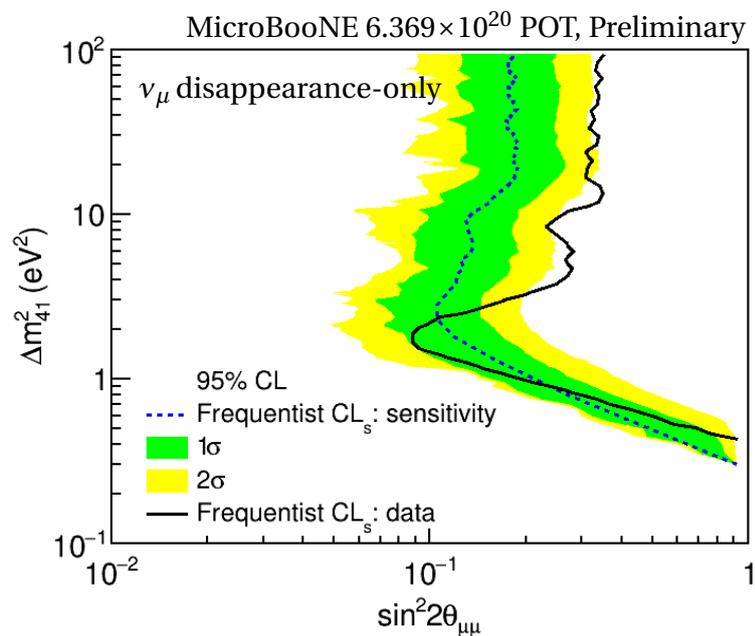
**Figure 13:** MicroBooNE 95% CL<sub>s</sub> sensitivity and data exclusion curves for  $\nu_e$  disappearance-only. The solid black curve represents the exclusion contour. The blue dashed curve represents the sensitivity contour. Seven channels are used in the oscillation fit while only the  $\nu_e$  disappearance oscillation effect is considered. The frequentist CL<sub>s</sub> sensitivity curve corresponds to the median value for each  $\Delta m_{41}^2$  and the green/yellow bands correspond to 68.3% ( $1\sigma$ ) and 95.5% ( $2\sigma$ ) confidence levels.



**Figure 14:** MicroBooNE 95% CL<sub>s</sub> sensitivity and data exclusion curves for  $\nu_e$  appearance-only. The solid black curve represents the exclusion contour. The blue dashed curve represents the sensitivity contour. Seven channels are used in oscillation fit while only the  $\nu_e$  appearance oscillation effect is considered. The frequentist CLs sensitivity curve corresponds to the median value for each  $\Delta m_{41}^2$  and the green/yellow bands correspond to 68.3% ( $1\sigma$ ) and 95.5% ( $2\sigma$ ) confidence levels.



**Figure 15:** MicroBooNE 95% confidence level frequentist  $CL_s$  limits in the  $\Delta m_{41}^2$  vs.  $\sin^2 2\theta_{\mu e}$  parameter space. The MiniBooNE 95% and 99% CL allowed regions [32] are shown in the shaded areas, which were estimated from the  $\nu_e$  appearance only scenario. The red solid curve represents the MicroBooNE 95%  $CL_s$  exclusion limit (2D profiling by minimizing over  $\sin^2 \theta_{24}$ ) in the full 3+1 neutrino oscillation scenario (both  $\nu_e$  appearance and  $\nu_e/\nu_\mu$  disappearance) using BNB Run 1-3 data. The green dashed curve represents the data exclusion limit in the  $\nu_e$  appearance only scenario (physically not allowed by the 3+1 oscillation framework) as opposed to the full 3+1 oscillation result.



**Figure 16:** MicroBooNE 95%  $\text{CL}_s$  sensitivity and data exclusion curves for  $\nu_\mu$  disappearance-only. The solid black curve represents the exclusion contour. The blue dashed curve represents the sensitivity contour. Seven channels are used in the oscillation fit while only the  $\nu_\mu$  disappearance oscillation effect is considered. The frequentist  $\text{CL}_s$  sensitivity curve corresponds to the median value for each  $\Delta m_{41}^2$  and the green/yellow bands correspond to 68.3% ( $1\sigma$ ) and 95.5% ( $2\sigma$ ) confidence levels.

## REFERENCES

- [1] P. Adamson et al. Improved Constraints on Sterile Neutrino Mixing from Disappearance Searches in the MINOS, MINOS+, Daya Bay, and Bugey-3 Experiments. *Phys. Rev. Lett.*, 125(7):071801, 2020, 2002.00301.
- [2] M. Andriamirado et al. Improved short-baseline neutrino oscillation search and energy spectrum measurement with the PROSPECT experiment at HFIR. *Phys. Rev. D*, 103(3):032001, 2021, 2006.11210.
- [3] F. Kaether, W. Hampel, G. Heusser, J. Kiko, and T. Kirsten. Reanalysis of the GALLEX solar neutrino flux and source experiments. *Phys. Lett. B*, 685:47–54, 2010, 1001.2731.
- [4] V. V. Barinov et al. Results from the Baksan Experiment on Sterile Transitions (BEST). 9 2021, 2109.11482.
- [5] J. N. Abdurashitov et al. Measurement of the solar neutrino capture rate with gallium metal. III: Results for the 2002–2007 data-taking period. *Phys. Rev. C*, 80:015807, 2009, 0901.2200.
- [6] G. Mention, M. Fechner, Th. Lasserre, Th. A. Mueller, D. Lhuillier, M. Cribier, and A. Loutourneau. The Reactor Antineutrino Anomaly. *Phys. Rev. D*, 83:073006, 2011, 1101.2755.
- [7] Patrick Huber. On the determination of anti-neutrino spectra from nuclear reactors. *Phys. Rev. C*, 84:024617, 2011, 1106.0687. [Erratum: *Phys.Rev.C* 85, 029901 (2012)].
- [8] Th. A. Mueller et al. Improved Predictions of Reactor Antineutrino Spectra. *Phys. Rev. C*, 83:054615, 2011, 1101.2663.
- [9] A. P. Serebrov et al. Search for sterile neutrinos with the Neutrino-4 experiment and measurement results. *Phys. Rev. D*, 104(3):032003, 2021, 2005.05301.
- [10] A. Aguilar-Arevalo et al. Evidence for neutrino oscillations from the observation of  $\bar{\nu}_e$  appearance in a  $\bar{\nu}_\mu$  beam. *Phys. Rev. D*, 64:112007, 2001, hep-ex/0104049.
- [11] A. A. Aguilar-Arevalo et al. Improved Search for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  Oscillations in the MiniBooNE Experiment. *Phys. Rev. Lett.*, 110:161801, 2013, 1303.2588.
- [12] A. A. Aguilar-Arevalo et al. Updated MiniBooNE neutrino oscillation results with increased data and new background studies. *Phys. Rev. D*, 103(5):052002, 2021, 2006.16883.
- [13] F. P. An et al. Evolution of the Reactor Antineutrino Flux and Spectrum at Daya Bay. *Phys. Rev. Lett.*, 118(25):251801, 2017, 1704.01082.

- [14] F. P. An et al. Antineutrino energy spectrum unfolding based on the Daya Bay measurement and its applications. *Chin. Phys. C*, 45(7):073001, 2021, 2102.04614.
- [15] M. Estienne et al. Updated Summation Model: An Improved Agreement with the Daya Bay Antineutrino Fluxes. *Phys. Rev. Lett.*, 123(2):022502, 2019, 1904.09358.
- [16] C. Giunti, Y. F. Li, C. A. Ternes, and Z. Xin. Reactor antineutrino anomaly in light of recent flux model refinements, 10 2021, 2110.06820.
- [17] Carlo Giunti and T. Lasserre. eV-scale Sterile Neutrinos. *Ann. Rev. Nucl. Part. Sci.*, 69:163–190, 2019, 1901.08330.
- [18] R. Acciarri et al. Design and Construction of the MicroBooNE Detector. *JINST*, 12(02):P02017, 2017.
- [19] T. Brieser et al. Testing of Cryogenic Photomultiplier Tubes for the MicroBooNE Experiment. *JINST*, 8:T07005, 2013.
- [20] P. Abratenko et al. Search for an Excess of Electron Neutrino Interactions in MicroBooNE Using Multiple Final State Topologies. 10 2021, 2110.14054.
- [21] P. Abratenko et al. Search for an anomalous excess of charged-current quasi-elastic  $\nu_e$  interactions with the MicroBooNE experiment using Deep-Learning-based reconstruction. 10 2021, 2110.14080.
- [22] P. Abratenko et al. Search for an anomalous excess of charged-current  $\nu_e$  interactions without pions in the final state with the MicroBooNE experiment. 10 2021, 2110.14065.
- [23] P. Abratenko et al. Search for an anomalous excess of inclusive charged-current  $\nu_e$  interactions in the MicroBooNE experiment using Wire-Cell reconstruction. 10 2021, 2110.13978.
- [24] P. Abratenko et al. New  $CC0\pi$  GENIE model tune for MicroBooNE. *Phys. Rev. D*, 105(7):072001, 2022, 2110.14028.
- [25] Baobiao Yue, Wei Li, Jiajie Ling, and Fanrong Xu. A compact analytical approximation for a light sterile neutrino oscillation in matter. *Chin. Phys. C*, 44(10):103001, 2020, 1906.03781.
- [26] Xiangpan Ji, Wenqiang Gu, Xin Qian, Hanyu Wei, and Chao Zhang. Combined Neyman–Pearson chi-square: An improved approximation to the Poisson-likelihood chi-square. *Nucl. Instrum. Meth. A*, 961:163677, 2020, 1903.07185.

- [27] Alexander L. Read. Modified frequentist analysis of search results (The CL(s) method). In *Workshop on Confidence Limits*, pages 81–101, 8 2000.
- [28] Thomas Junk. Confidence level computation for combining searches with small statistics. *Nucl. Instrum. Meth. A*, 434:435–443, 1999, hep-ex/9902006.
- [29] X. Qian, A. Tan, J. J. Ling, Y. Nakajima, and C. Zhang. The Gaussian CL<sub>s</sub> method for searches of new physics. *Nucl. Instrum. Meth. A*, 827:63–78, 2016, 1407.5052.
- [30] Glen Cowan, Kyle Cranmer, Eilam Gross, and Ofer Vitells. Asymptotic formulae for likelihood-based tests of new physics. *Eur. Phys. J. C*, 71:1554, 2011, 1007.1727. [Erratum: *Eur.Phys.J.C* 73, 2501 (2013)].
- [31] Gary J. Feldman and Robert D. Cousins. A Unified approach to the classical statistical analysis of small signals. *Phys. Rev. D*, 57:3873–3889, 1998, physics/9711021.
- [32] A. A. Aguilar-Arevalo, B. C. Brown, J. M. Conrad, R. Dharmapalan, A. Diaz, Z. Djurcic, D. A. Finley, R. Ford, G. T. Garvey, S. Gollapinni, A. Hourlier, E.-C. Huang, N. W. Kamp, G. Karagiorgi, T. Katori, T. Kobilarcik, K. Lin, W. C. Louis, C. Mariani, W. Marsh, G. B. Mills, J. Mirabal-Martinez, C. D. Moore, R. H. Nelson, J. Nowak, I. Parmaksiz, Z. Pavlovic, H. Ray, B. P. Roe, A. D. Russell, A. Schneider, M. H. Shaevitz, H. Siegel, J. Spitz, I. Stancu, R. Tayloe, R. T. Thornton, M. Tzanov, R. G. Van de Water, D. H. White, and E. D. Zimmerman. Updated miniboone neutrino oscillation results with increased data and new background studies. *Phys. Rev. D*, 103:052002, Mar 2021.
- [33] C. A. Argüelles, I. Esteban, M. Hostert, Kevin J. Kelly, J. Kopp, P. A. N. Machado, I. Martinez-Soler, and Y. F. Perez-Gonzalez. MicroBooNE and the  $\nu_e$  Interpretation of the MiniBooNE Low-Energy Excess. 11 2021, 2111.10359.