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We report the first measurement of π^0 production in neutral current (NC) interactions on argon with average neutrino energy of $\lesssim 1$ GeV. We use data from the MicroBooNE detector's 85-tonne active volume liquid argon time projection chamber situated in Fermilab's Booster Neutrino Beam and exposed to 5.89×10^{20} protons on target for this measurement. Measurements of NC π^0 events are reported for two exclusive event topologies without charged pions. Those include a topology with two photons from the decay of the π^0 and one proton and a topology with two photons and zero protons. Flux-averaged cross-sections for each exclusive topology and for their semi-inclusive combination are extracted (efficiency-correcting for two-plus proton final states), and the results are compared to predictions from the GENIE, NEUT, and NUWRO neutrino event generators. We measure cross sections of 1.243 ± 0.185 (syst) ± 0.076 (stat), $0.444 \pm 0.098 \pm 0.047$, and $0.624 \pm 0.131 \pm 0.075$ $[10^{-38} \text{cm}^2/\text{Ar}]$ for the semi-inclusive NC π^0 , exclusive NC π^0 +1p, and exclusive NC π^0 +0p processes, respectively.

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INTRODUCTION I.

69 of intense study both experimentally and within the the-70 ory community in recent years due to their role in inter-71 preting neutrino oscillation measurements and searches 72 for other rare processes in neutrino scattering [1]. While 105 of BSM models [5–9]. 73 neutrino oscillation experiments primarily rely on mea-74 suring the rate of charged current (CC) interactions, it 75 is also important that we build a solid understanding of 76 77 inclusive and exclusive neutral current (NC) neutrino interactions. 78

79 ν_e to ν_e and $\bar{\nu}_e$ measurements in the energy range of a ν_e formation to non-unitarity or light sterile neutrino oscil-81 that cannot perfectly differentiate between photon- and 82 electron-induced electromagnetic showers, and therefore $^{\ 115}$ 83 84 85 86 87 sterile neutrino oscillation searches with the upcoming 88 Short Baseline Neutrino (SBN) experimental program [2] 89 and CP violation measurements and mass hierarchy de-91 Experiment (DUNE) [3]. 92

Furthermore, NC π^{0} events can contribute as back-93 ground to searches for rare neutrino scattering processes 94 such as NC Δ resonance production followed by Δ radia-95 tive decay, or NC coherent single-photon production at ⁹⁷ energies below 1 GeV [4]. This is primarily a consequence ⁹⁸ of the limited geometric acceptance of some detectors,

⁹⁹ whereby one of the photons from a π^0 decay can escape ¹⁰⁰ the active volume of the detector. Depending on a detec-Neutrino-nucleus cross-sections have been the subject ¹⁰¹ tor's ability to resolve electromagnetic shower substruc-¹⁰² ture, NC π^0 events can further contribute as background 103 to searches for new physics beyond the Standard Model $_{104}$ (BSM), such as e^+e^- production predicted by a number

Finally, NC measurements themselves can provide a 106 ¹⁰⁷ unique channel for probing new physics. For example, ¹⁰⁸ searches for non-unitarity in the three-neutrino paradigm 109 or searches for active to sterile neutrino oscillations ¹¹⁰ are possible via NC rate disappearance measurements NC neutrino interactions are of particular importance ¹¹¹ [10, 11]. Such searches can provide complementary infew hundred MeV. This is especially true for detectors ¹¹³ lation parameters otherwise accessible only through CC 114 measurements.

Using a liquid argon time projection chamber where NC π^0 production can be misidentified as ν_e or $\bar{\nu}_e^{-116}$ (LArTPC) as its active detector, MicroBooNE [12] CC scattering. Misidentification of photons as electrons¹¹⁷ shares the same technology and neutrino target nucleus complicates the interpretation of ν_e appearance measure-¹¹⁸ as the upcoming SBN and future DUNE experiments. ments aiming to measure subtle signals. These include ¹¹⁹ MicroBooNE's 85 metric ton active volume LArTPC 120 is situated 468.5 m away from the proton beam tar-121 get in the muon-neutrino-dominated Booster Neutrino 122 Beam (BNB) at Fermilab [13] which is also used by termination with the future Deep Underground Neutrino ¹²³ SBN. The resulting neutrino beam has a mean energy $_{124} \langle E_{\nu} \rangle = 0.8 \text{ GeV}$ and is composed of 93.7% ν_{μ} , 5.8% $\bar{\nu}_{\mu}$, 125 and 0.5% $\nu_e/\bar{\nu}_e$. MicroBooNE's cross-section measure-¹²⁶ ments on argon are therefore timely and directly relevant 127 to these future (SBN and DUNE) programs.

> We present the first measurement of neutrino-induced 128 ¹²⁹ NC single- π^0 (1 π^0) production on argon with a mean ¹³⁰ neutrino energy in the 1 GeV regime, which is also the ¹³¹ highest-statistics measurement of this interaction chan-132 nel on argon to date. This measurement is relevant to ¹³³ the physics programs of experiments that operate in the ¹³⁴ few-GeV regime (SBN [2], DUNE [3], NOvA [14, 15],

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¹³⁶ argon as a target material. Additionally, this measure-¹⁹² clear final state interactions in neutrino-argon scattering ¹³⁷ ment has been used to provide an indirect constraint to ¹⁹³ models. ¹³⁸ the rate of NC $1\pi^0$ backgrounds in MicroBooNE's recent ¹³⁹ search for a single-photon excess [4]. The only previous ¹⁴⁰ results for NC $1\pi^0$ scattering on argon are from the Ar-¹⁴¹ goNeuT collaboration using the NuMI beam which has a ¹⁴² much higher mean neutrino beam energy of 9.6 GeV for ¹⁹⁵ ¹⁴³ ν_{μ} and of 3.6 GeV for $\overline{\nu}_{\mu}$ [18].

144 ¹⁴⁵ analysis are defined as

$$\nu + \mathcal{A} \to \nu + \mathcal{A}' + \pi^0 + X, \tag{1}$$

¹⁴⁶ where \mathcal{A} represents the struck (argon) nucleus, \mathcal{A}' rep- $_{147}$ resents the residual nucleus, and X represents exactly $_{202}$ MicroBooNE Collaboration. This tune specifically tar-148 one or zero protons plus any number of neutrons, but no 203 gets CC quasi-elastic (QE) and CC multi-nucleon in-149 other hadrons or leptons. The protons are identifiable in the MicroBooNE LArTPC by their distinct ionizing 150 tracks while the π^0 is identifiable through the presence 151 152 of two distinct electromagnetic showers, one for each photon from the $\pi^0 \to \gamma \gamma$ decay, with kinematic properties 154 155 variant mass.

156 ¹⁵⁷ first to perform a rate validation check and subsequently 158 in three distinct cross-section measurements. By lever-159 aging the capability of LArTPCs to detect and identify ²¹⁴ gation and re-interactions within the detector and a cus-¹⁶⁰ protons we perform the world's first exclusive $NC\pi^0 + 0p^{-215}$ tom detector response model all implemented within the ¹⁶¹ and NC π^0 +1p cross-section extractions and additionally ²¹⁶ LArSoft framework [37]. measure the cross-section for $NC\pi^0$ interactions semi-162 163 inclusively using both the one proton and zero proton 218 begins by reading out and processing the ionization samples combined. Each of these cross-section extractions utilizes a distinct signal definition. The signal 165 166 168 169 any number of protons. The signal definitions for all 224 fit to Gaussian pulses. The collection of these hits and 170 three measurements also require that there are no other 225 their characteristics such as readout time, wire chan-171 hadrons or leptons in the final state (as noted above). 226 nel number, and integrated charge are then used as in-172 173 174 lowed in the final state. Finally, the signal definitions 229 ters and matches hits across three 2D projected views 175 allow for interactions of all flavors of neutrinos that are 230 of the MicroBooNE active TPC volume to form 3D represent: ν_{μ} , $\bar{\nu}_{\mu}$, ν_{e} , and $\bar{\nu}_{e}$. 176

177 $_{178}$ NC π^0 measurements which typically require one and $_{233}$ fier score and aggregated into candidate neutrino inter- π^{179} only one π^0 meson and little hadronic activity in the π^{234} actions. Pandora also reconstructs a candidate neutrino 180 detector [19–27]. This differs from the more inclusive 235 interaction vertex based on the position and orientation ¹⁸¹ approach of the ArgoNeuT experiment motivated both ²³⁶ of the reconstructed tracks and showers which represents 182 by its higher energy beam as well as the need to mit- 237 the most likely position of the neutrino interaction. $_{183}$ igate the low statistics of its data sample [18]. Mak- 238 ¹⁸⁴ ing use of the MicroBooNE LArTPC's power in exam-²³⁹ constant stream of high-energy cosmic rays impinging on 185 186 187 sections extracted in this analysis extend our understand- 242 date neutrino interactions. To incorporate the effect of 188 ing of this important interaction channel. The simulta- 243 cosmic-ray contamination in the simulation, cosmic ray ¹⁸⁹ neous measurement of exclusive and semi-inclusive cross-²⁴⁴ data recorded *in situ* at MicroBooNE, when the beam ¹⁹⁰ sections in particular provides additional information for ²⁴⁵ is not present, are used as overlays (at the wire sig-

 $_{135}$ T2K [16], and Hyper-K [17]), especially those which share $_{191}$ the tuning of NC $1\pi^0$ production cross sections and nu-

II. ANALYSIS OVERVIEW

This measurement uses data corresponding to a BNB ¹⁹⁶ exposure of 5.89×10^{20} protons on target (POT), col-The interaction final states that are measured in this 197 lected during the period 2016–2018 and referred to as ¹⁹⁸ "Runs 1–3" in many of the subsequent figures. Neutrino-¹⁹⁹ argon interactions are simulated using a custom tune [28] 200 of the GENIE neutrino event generator v3.0.6 [29, 30] 201 (based on model set G18_10a_02_11a) adopted by the ²⁰⁴ teraction models and overall has very little direct effect 205 on this NC-focused analysis. GENIE v3 uses the Berger-²⁰⁶ Sehgal [31, 32] model for resonant production of π^0 and 207 includes improved agreement with an expanded data set ²⁰⁸ for the A-dependence of final state interactions (FSI), upsuch that they reconstruct to approximately the π^0 in- 209 dated form factors [33], updated diagrams for pion pro-²¹⁰ duction processes [32, 34, 35], and a new tune to neutrino-These one proton and zero proton samples are used ²¹¹ proton and neutrino-deuterium cross-section data [30]. ²¹² The MicroBooNE Monte Carlo (MC) prediction further 213 makes use of GEANT4 v4_10_3_03c [36] for particle propa-

The MicroBooNE data and MC reconstruction chain 217 219 charge signals detected on the 8,192 wires that make ²²⁰ up the three anode planes of the MicroBooNE LArTPC. definitions for the two exclusive measurements place a ²²¹ The procedure includes noise removal [38] and signal prothreshold on true proton kinetic energy of greater than ²²² cessing as described in [39] and [40]. Localized regions 50 MeV, while the semi-inclusive measurement allows for ²²³ of interest referred to as "hits" are then identified and MeV-scale photons, which may arise from nuclear de- 227 put to the Pandora pattern recognition framework for excitation processes within the struck nucleus, are al- 228 further processing [41]. The Pandora framework clus-²³¹ constructed objects. These objects are then classified as These definitions are comparable to other historical ²³² track-like or shower-like based on a multivariate classi-

Being a surface detector, MicroBooNE is subject to a ining hadronic final state multiplicities and kinematic 240 the detector that substantially outnumber the neutrino properties with high resolution, the flux-averaged cross- ²⁴¹ interactions and form the largest background to candi-

 $_{247}$ During the 2.3 ms that it takes to "drift" ionization $_{302}$ placed on the track reconstruction. ²⁴⁸ charge associated with neutrino interaction final states ²⁴⁹ across the maximum 2.56 m drift distance, $\mathcal{O}(10)$ cos-²⁵⁰ mic rays are expected to enter the detector. In order to reduce this cosmic-ray contamination, scintillation light 251 ²⁵² recorded by the MicroBooNE photo-detector system is ₃₀₄ ²⁵³ matched to candidate neutrino interactions during recon-³⁰⁵ remaining signal and background are further differenti-254 struction and is also required to occur in time with the 306 ated and separated using two tailored BDTs trained on $1.6 \ \mu s \ long \ BNB \ neutrino \ spill.$ 255

To select a high-purity sample of BNB neutrino NC 256 $_{257}$ $1\pi^0$ interactions, a series of topological, pre-selection, $_{309}$ input various reconstructed kinematic, geometric, and ²⁵⁸ and boosted decision tree (BDT)-based selections are ap- ₃₁₀ calorimetric variables both for the signal (defined as ²⁵⁹ plied. This results in two mutually exclusive final selec-260 one proton in the final state, and $2\gamma 0p$, which targets 261 262 two photons and zero protons in the final state. The ²⁶³ different selection stages are described below, along with the details of the systematic uncertainty evaluation. 264

Topological Selection and Pre-Selection Α. 265

266 267 268 270 $_{271}$ correspond to the photons expected from π^0 decay. The $_{326}$ ten reconstructed variables. Due to the existence of a $_{272}$ presence of a track corresponds to a reconstructed pro- $_{327}$ proton candidate track in the $2\gamma 1p$ sample, these ten ton exiting the nucleus while the zero-track case suggests 328 variables differ for each BDT. They are listed below. 273 either a low-energy proton that is not reconstructed or 329 274 no charged hadrons at all exiting the nucleus. 275

Once events with the desired signal topologies are iden-276 tified, a series of loose "pre-selection" requirements is ap-331 277 332 plied to reduce obvious backgrounds or mis-reconstructed 278 333 events. These pre-selection requirements include shower 279 energy thresholds of 30 MeV for the leading shower and 334 280 20 MeV for the subleading shower in both topologies. 281 336 The pre-selection also requires that the reconstructed 282 ²⁸³ neutrino interaction point be contained in a fiducial vol-²⁸⁴ ume, defined as at least 5 cm away from any TPC wall, in order to help reduce the number of selected events with 285 tracks that exit the detector. For the $2\gamma 1p$ topology, 286 287 the non-zero conversion distance of photons is explicitly used by requiring that each shower has a reconstructed ₃₄₁ 288 start point of at least 1 cm from the reconstructed neu-289 trino interaction vertex. Typically the reconstructed neu-290 trino interaction vertex is identified as the start of the re-291 constructed proton candidate track. In order to remove 292 a very small number of poorly reconstructed events in 344 293 which the candidate track is not consistent with the hy-294 ²⁹⁵ pothesis of originating from the candidate neutrino inter-²⁹⁶ action vertex, a requirement is placed to ensure the track ²⁹⁷ start point is always within 10 cm of the reconstructed ³⁴⁷ ²⁹⁸ neutrino interaction vertex. The efficiency of selecting ³⁴⁸ ²⁹⁹ NC $1\pi^0 + 0$ (1)p events using these pre-selection require- $_{300}$ ments is 82.1% (63.6%). Note that the efficiency of the 1p

²⁴⁶ nal waveform level) to simulated neutrino interactions. ³⁰¹ selection is lower because of the additional requirements

Boosted Decision Tree-Based Selection В.

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After applying the pre-selection requirements, the ³⁰⁷ simulation. The gradient boosting algorithm XGBoost [42] is used to train each of the BDTs. They take as 308 ³¹¹ an NC interaction with identically one π^0 in the final tion topologies: $2\gamma 1p$, which targets two photons and $_{312}$ state) and for the background interactions. Because the ³¹³ two tailored BDTs target different topologies, notably $_{314}$ including one track in the case of $2\gamma 1p$ and zero tracks in $_{315}$ the case of $2\gamma 0p$, the signal definitions used for the two ³¹⁶ BDTs are slightly different. NC π^0 events with exactly ³¹⁷ one proton with true kinetic energy above 20 MeV are 318 used as the training signal for the $2\gamma 1p$ BDT while ³¹⁹ NC π^0 events with no protons with true kinetic energy ₃₂₀ above 20 MeV are used as the training signal for the The event selection begins with topology-based crite- $_{321} 2\gamma 0p$ BDT. We note that the 20 MeV threshold used ria for candidate neutrino interactions identified by Pan- 322 in the BDT training is lower than the 50 MeV proton dora and targets two mutually exclusive topological def- 323 kinetic energy threshold used later during cross-section initions: (a) two showers and one track $(2\gamma 1p)$, and (b) $_{324}$ extraction, as during training we are aiming to push the two showers and zero tracks $(2\gamma 0p)$. The two showers $_{325}$ threshold as low as possible. Each BDT is trained on

³³⁰ Variables used in both $2\gamma 1p$ and $2\gamma 0p$ BDTs:

- Leading and subleading shower impact parameters: The perpendicular distance between the backprojection of the reconstructed shower and the candidate neutrino interaction point which is a metric of how well each shower "points" back to the reconstructed neutrino interaction point.
- Leading and subleading shower conversion distances: Defined as the distance between the reconstructed start of the shower and reconstructed neutrino interaction point.
- Reconstructed energy of the leading shower.

$_{342}$ Variables used in only the $2\gamma 1p$ BDT:

- Reconstructed track length.
- Reconstructed track vertical angle: Defined as the arctangent of the track direction in the vertical plane with respect to the beam axis.
- Distance from track end to TPC wall: Calculated as the shortest distance to the closest TPC wall.
- Reconstructed mean energy deposition per unit length (dE/dx) of the track.

• Ratio of dE/dx of the first half of track to that of 403 351 the second half of the track: A metric for identi- 404 352 fying stopping proton tracks that contain a Bragg 405 353 peak. 354 406

Variables used in only the $2\gamma 0p$ BDT: 355

- Reconstructed energy of the subleading shower. 356
- Leading and subleading shower geometric length 357 per unit energy: The ratio of each shower's geo- 411 358 metric length to its reconstructed energy. The ge- 412 359 ometric length is an estimate of the 3D extent of 413 360 the electromagnetic shower. 361
- 362 363 364 365 366 in origin. 367
- Reconstructed leading shower vertical angle: Direc-368 tion in the vertical plane with respect to the beam 369 axis. 370

By construction BDT scores lie on the interval of [0, 422]371 372 1]. After training, the resulting BDT score distributions, 423 diction include contributions from uncertainties in the 373 374 375 systematic and statistical uncertainties (the definition of 427 ulations and cosmic ray data). 376 these systematic uncertainties is described in detail in 428 377 378 379 380 381 382 383 384 eight categories, based on GENIE truth-level information: 436 gated effects on the final event distributions. 385

- NC $1\pi^0$: All neutral current interactions that pro-386 duce one exiting π^0 regardless of incoming neutrino 387 flavor. This is our targeted signal selection, and it 388 is further split into two sub-categories, "NC $1\pi^0$ 389 Coherent" and "NC $1\pi^0$ Non-Coherent" contribu-390 tions, based on their interaction types. 391
- NC $\Delta \rightarrow N\gamma$: Leading Standard Model source of 392 NC single-photon production below 1 GeV origi-393 nating from radiative decay of the $\Delta(1232)$ baryon. 394
- CC $\nu_{\mu} 1 \pi^0$: All ν_{μ} CC interactions that have one 395 true exiting π^0 . 396
- CC $\nu_e/\overline{\nu}_e$ Intrinsic: All CC ν_e or $\overline{\nu}_e$ interactions 397 regardless of whether or not a π^0 was emitted. 398
- BNB Other: All remaining BNB neutrino interac-399 tions that take place in the active TPC volume of 400 MicroBooNE and are not covered by the above five 401 categories. 402

• Dirt (Outside TPC): All BNB neutrino interactions that take place outside the MicroBooNE active TPC but have final states that enter and interact inside the active TPC detector. This can originate from scattering off liquid argon in the cryostat vessel outside the active TPC volume or from interactions in the concrete and "dirt" surrounding the cryostat itself.

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• Cosmic Data: Coincident cosmic ray interactions that take place during a BNB spill but without any true neutrino interaction present.

The final NC $1\pi^0$ -enriched samples are selected by 414 • Pandora "neutrino score": A multivariate classifier 415 placing a requirement on the BDT score distribution that in the Pandora reconstruction suite which scores all $_{416}$ maximizes the product of NC $1\pi^0$ signal efficiency and reconstructed neutrino candidates based on their 417 purity. This corresponds to a threshold on the BDT geometric and kinematic features as to how likely a $_{418}$ scores of > 0.854 and > 0.950 for $2\gamma 1p$ and $2\gamma 0p$, respectively. candidate is due to a neutrino interaction or cosmic 419 tively. The final distributions are provided and discussed 420 in Sec. IIE.

$\mathbf{C}.$ Systematic Uncertainty Evaluation

Systematic uncertainties on the MC simulation pretested on a statistically independent simulation and data 424 neutrino flux, the cross-section modeling, hadron reset, are shown in Fig. 1. The simulation and data points 425 interactions, detector effects, and the effect of finite agree across the full range of BDT classifier score within 426 statistics used in the background predictions (both sim-

The flux systematic uncertainties incorporate hadron Sec. II C). The bimodal distribution of the $2\gamma 1p$ BDT re- $_{429}$ production uncertainties where the Booster proton beam sponse indicates greater separation power between signal 430 hits the beryllium target, uncertainties on pion and nuand background compared to that for $2\gamma 0p$ because the 431 cleon scattering in the target and surrounding aluminum addition of the reconstructed track gives access to an en- 432 magnetic focusing horn of the BNB, and mismodeling of tirely separate handle on background rejection. For this 433 the horn current. Following [43], these are implemented and subsequent MC simulation comparisons to data, the 434 by reweighting the flux prediction according to neutrino simulation predictions are broken down into the following 435 type, parentage, and energy, and studying the propa-

> The cross-section uncertainties incorporate modeling 437 ⁴³⁸ uncertainties on the GENIE prediction [28–30], evaluated 439 by GENIE reweighting tools. The default GENIE uncer-440 tainties on NC resonant production arising from NC 441 resonant vector and axial mass parameters of $m_V =$ $_{442}$ 0.840 ± 0.084 GeV and $m_A = 1.120 \pm 0.224$ GeV, respec-443 tively, were assumed. GENIE uses an effective cascade empirical model for hadronic final-state interactions, called 444 hA2018, which allows for reweighting to estimate the ef-445 fect on final distributions. For more information on cross-446 section uncertainties in MicroBooNE, please see [28]. 447

> 448 The hadron-argon reinteraction uncertainties are associated with the propagation of hadrons through the 449 detector, as modeled in GEANT4 [36]. Both charged 450 pions and proton reinteractions during propagation 451 ⁴⁵² were considered and their impact estimated using the 453 GEANT4REWEIGHT tool [44].

> The detector modeling and response uncertainties are 454 455 evaluated using MicroBooNE's novel data-driven tech-⁴⁵⁶ nique for assessing and propagating LArTPC detector-



(b) $2\gamma 0p$

FIG. 1: The BDT classifier score for (a) $2\gamma 1p$ and (b) $2\gamma 0p$ targeted selections. Higher scores indicate more NC $1\pi^0$ signal-like events, and lower scores indicate more background-like events. The red vertical lines indicate the threshold positions, keeping all events to the right, for the final selections. The distribution above 0.95 is omitted for $2\gamma 1p$ because there are no events in this region.

⁴⁵⁷ related systematic uncertainties [45]. This approach uses ⁴⁸⁸ pected. in situ measurements of distortions in the TPC wire 458 readout waveform signals - caused by detector effects such as electron diffusion, electron drift lifetime, elec-460 tric field, and the electronics response - to parameterize 461 these effects at the TPC wire level. This provides a de-463 effects on the high level variables and, subsequently, the $_{492}$ and $2\gamma 0p$ selections, respectively. For the $2\gamma 1p$ selection, $_{466}$ corresponding to variations in the charge recombination $_{494}$ purity is 63.5% while for the $2\gamma 0p$ selection, the efficiency 467 model, the scintillation light yield, and space charge ef- $_{495}$ and purity are 54.8% and 59.6%, respectively. The $2\gamma 1p$



FIG. 2: Reconstructed shower energy vs. true shower energy for a sample of simulated true NC $1\pi^0$ events. Only showers with a reconstructed energy of at least 20 MeV are considered.

Shower Energy Calibration D.

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Electromagnetic shower reconstruction in LArTPCs 470 is known to be a lossy process primarily due to mis-471 clustering and thresholding effects. Current reconstruc-472 473 tion algorithms often miss small, low-energy hits in an electromagnetic shower when clustering objects, and 474 some of the hits that are reconstructed may fall below the 475 energy threshold. On average, these effects are expected to yield shower energy losses of approximately 20% [48]. 477 This can be seen in Fig. 2 where the reconstructed shower ⁴⁷⁹ energy falls systematically below the true shower energy 480 in simulation. By performing a linear fit to the most probable values of reconstructed shower energy in bins 482 of true shower energy, shown as the pink straight line in Fig. 2, a correction factor is extracted which brings the reconstructed values closer to expectation. This fit 484 results in an energy correction that is applied to all re-485 constructed showers, 486

$$E_{\rm corr} = (1.21 \pm 0.03)E_{\rm reco} + (9.88 \pm 4.86) \,\,{\rm MeV},$$
 (2)

⁴⁸⁷ and represents a correction of approximately 20%, as ex-

E. **Final Selected Spectra**

After applying the BDT requirements, 1130 selected 490 tector model-agnostic way to study and evaluate their $_{491}$ data events remain with 634 and 496 falling into the $2\gamma 1p$ final event distributions. Additional detector systematics 493 the BDT score requirement efficiency is 69.9% and the fects [46, 47] are separately evaluated and also included. $_{496}$ and $2\gamma 0p$ BDT selection efficiencies and purities are both ⁴⁹⁷ calculated relative to a NC $1\pi^0$ final state, allowing for ⁴⁹⁸ any number of protons. The efficiencies at each stage



(a) π^0 momentum dependence



(b) Proton kinetic energy dependence

FIG. 3: Efficiencies of the final $2\gamma 1p$, $2\gamma 0p$ and combined $2\gamma (0 + 1)p$ selections as a function of (a) true π^0 momentum and (b) true leading exiting proton kinetic energy. Events in which there are no exiting protons are included in the first bin. As can be seen, a threshold of ≈ 50 MeV proton kinetic energy is where events start to shift between the $2\gamma 0p$ and $2\gamma 1p$ selections.

⁴⁹⁹ of the analysis are provided in Table I, and the total ef-⁵⁰⁰ ficiency for each selection is shown as a function of (a) 501 true π^0 momentum and (b) true proton kinetic energy $_{502}$ in Fig. 3. Overall, the 1*p* selection is more efficient and $_{503}$ of higher signal purity relative to the 0p selection due 504 to the existence of a reconstructed particle track which greatly helps to tag the neutrino interaction point and 505 reject backgrounds. This track information, particularly 506 track calorimetry, provides an additional handle on the 507 neutrino interaction mode; a proton-like track is highly 508 indicative of an NC $1\pi^0$ interaction whereas CC interac-509 tions generally have a muon track in the final state. 510

⁵¹¹ The resulting distributions as a function of the recon-⁵¹² structed two-photon invariant mass are shown in Fig. 4.



(b) $2\gamma 0p$

FIG. 4: The reconstructed diphoton invariant mass for both the (a) $2\gamma 1p$ and (b) $2\gamma 0p$ final selected data. The result of fitting a Gaussian plus linear function to the data is shown in cyan.

TABLE I: NC $1\pi^0$ efficiencies for the $2\gamma 1p$ and $2\gamma 0p$ selections. The topological and combined efficiencies are evaluated relative to all true NC $1\pi^0$ events inside the active TPC. The pre-selection and BDT selection efficiencies are evaluated relative to their respective preceding selection stage. The final efficiencies are the combined total efficiency for each selection.

Selection Stage	$2\gamma 1p$ eff.	$2\gamma 0p$ eff.
Topological	10.5%	6.60%
Pre-selection	59.4%	77.3%
BDT Selection	69.9%	54.8%
Final Efficiencies	4.36%	2.81%

⁵¹³ The invariant mass is reconstructed from the energy and

514 direction of the two photon candidate showers as

$$M_{\gamma\gamma}^2 = 2E_{\gamma 1}E_{\gamma 2}(1 - \cos\theta_{\gamma\gamma}), \qquad (3)$$

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⁵¹⁵ where $\cos \theta_{\gamma\gamma}$ is the opening angle between the two show-⁵¹⁶ ers. For the $2\gamma 1p$ case where a track has been identified ⁵¹⁷ as a candidate proton, the directions of the showers and ⁵¹⁸ thus the opening angle between them are calculated by ⁵¹⁹ constructing the direction between the candidate neu-⁵²⁰ trino interaction point and the shower start point. For ⁵²¹ the $2\gamma 0p$ selection, however, no such candidate track ex-⁵²² ists. Instead, the shower direction and opening angle are ⁵²³ entirely estimated from the geometric shape of the show-⁵²⁴ ers themselves.

A Gaussian-plus-linear fit is performed to each ob-525 526 served distribution in data to extract the reconstructed 527 π^0 invariant mass while taking into account the non- $_{528} \pi^0$ background contamination. For the $2\gamma 1p$ event sam-₅₂₉ ple, this fit gives a Gaussian mean of $138.9\pm2.1 \text{ MeV}/c^2$ so with a width of 31.7 ± 2.4 MeV/ c^2 . For the $2\gamma0p$ event ⁵³¹ sample, the corresponding fit gives a Gaussian mean of $_{532}$ 143.3 \pm 3.2 MeV/ c^2 with a width of 47.9 \pm 4.9 MeV/ c^2 . As 533 a goodness-of-fit test, the resulting χ^2 per degree of free- $_{534}$ dom is 1.20 and 1.45 for the $2\gamma 1p$ and $2\gamma 0p$ fits, respec-⁵³⁵ tively. These both show agreement with the expected invariant mass of the π^0 of 134.9770 ± 0.0005 MeV/ c^2 536 [49] giving confidence and validation of the calorimetric 537 energy reconstruction of the showers. Additional distri-538 butions showing the reconstructed π^0 momentum as well as the reconstructed angle of the outgoing π^0 with respect 540 to the incoming neutrino beam are provided in Fig. 5. 541

Two additional reconstructed distributions of interest are highlighted. First, the reconstructed cosine of the center-of-mass (CM) decay angle is shown in Fig. 6. This is defined as the angle between the lab-frame π^0 momentum direction and the decay axis of the two daughter photons in the CM frame,

$$\cos \theta_{\rm CM} = \frac{E_{\pi^0}}{|P_{\pi^0}|} \frac{|E_{\gamma 1} - E_{\gamma 2}|}{E_{\gamma 1} + E_{\gamma 2}}.$$
 (4)

⁵⁴⁸ This quantity should be an isotropic flat distribution for ⁵⁴⁹ true $\pi^0 \rightarrow \gamma \gamma$ signal events, and any deviation from ⁵⁵⁰ this can highlight regions of inefficiency in reconstruc-⁵⁵¹ tion or selection. As can be seen in Fig. 6, for both ⁵⁵² $2\gamma 1p$ and $2\gamma 0p$ selections, the distributions taper off at ⁵⁵³ high $\cos \theta_{\rm CM}$ which corresponds to increasingly asymmet-⁵⁵⁴ ric π^0 decays. When reconstructing asymmetric π^0 de-⁵⁵⁵ cay events, it is more likely that the subleading photon ⁵⁵⁶ shower is missed due to its low energy. Note, however, ⁵⁵⁷ that the observed data show the same trend as the sim-⁵⁵⁸ ulation within uncertainty.

Figure 7 additionally highlights the reconstructed photon conversion distance for all showers in the final $2\gamma 1p$ selection. Well-reconstructed showers with conversion distances as far as 100 cm from the candidate neutrino interaction are observed. This helps validate the assumption that the reconstructed showers are indeed likely to be true photons as $\mathcal{O}(100)$ MeV photons are expected to

⁵⁶⁶ have a mean free path in argon of ≈ 20 cm. Note that, as ⁵⁶⁷ the $2\gamma 0p$ selection does not have any visible hadronic ac-⁵⁶⁸ tivity for tagging the interaction point, the corresponding ⁵⁶⁹ conversion distance is significantly harder to estimate.

⁵⁷⁰ Finally, Fig. 8 shows two example event displays of se-⁵⁷¹ lected events in data for both the $2\gamma 1p$ and $2\gamma 0p$ topolo-⁵⁷² gies. Each event shows two well-reconstructed showers ⁵⁷³ pointing back to a common interaction point with prop-⁵⁷⁴ erties consistent with those being photons from a π^0 de-⁵⁷⁵ cay.

III. NC π^0 RATE VALIDATION

NC 1 π^0 events contribute as a dominant background to NC single-photon production measurements carried out radiative decay [4], NC coherent single-photon production, or more rare e^+e^- pair production motivated in BSM theories. In addition to using these selected events as a calibration sample for understanding and validating shower reconstruction performance, they are also used to validate the observed overall rate of this process as curvalidate the observed overall rate of this sample rate from a sufficient description of the observed data, this sample can and has been used to provide an *in situ* constraint on NC 1 π^0 mis-identified backgrounds, *e.g.* as in [4]. Alternatively, these measurements can be used to motivate GENIE tuning.

592 As shown in Fig. 5, both the $2\gamma 1p$ and the $2\gamma 0p$ se-⁵⁹³ lections see an overall deficit in data relative to the MC ⁵⁹⁴ prediction. This is more pronounced in the $2\gamma 1p$ selection 595 where the ratio of the number of selected data events to ⁵⁹⁶ the number of selected simulated events is 0.79. As it is ⁵⁹⁷ also known that the GENIE branching fraction of coherent NC $1\pi^0$ production on argon is significantly lower than ⁵⁹⁹ expectation extrapolated from MiniBooNE's π^0 measure-600 ment on mineral oil [23], the possibility of a correction 601 to GENIE predictions on both non-coherent and coher-₆₀₂ ent NC $1\pi^0$ production is explicitly examined. The MC ⁶⁰³ predictions are fitted to data allowing both coherent and non-coherent NC $1\pi^0$ rates to vary. Both normalization-604 605 only and normalization plus shape variations to the coherent and non-coherent rates are explored; all yield sim-607 ilar conclusions. This section describes the normalization ⁶⁰⁸ plus shape variation fit in detail.

The normalization plus shape variation fit is performed as a function of reconstructed π^0 momentum for both 610 as a function of reconstructed π^0 momentum for both 611 2 γ 1p and 2 γ 0p selections, using [0, 0.075, 0.15, 0.225, 612 0.3, 0.375, 0.45, 0.525, 0.6, 0.675, and 0.9] GeV/c bin lim- 613 its. In the fit, MC predicted coherent NC 1 π^0 events are scaled by a normalization factor N_{coh} , and MC predicted 615 non-coherent NC 1 π^0 events are scaled on an event-by- 616 event basis depending on their corresponding true π^0 mo- 617 mentum according to $(a + b|\vec{p}_{\pi^0}^{true}|)$, where the true π^0 618 momentum is given in [GeV/c]. This form of scaling is 619 chosen for non-coherent NC 1 π^0 as the simplest correc- 620 tion to the decline in data-to-MC ratio as reconstructed



(c) $2\gamma 1p$: Reconstructed cosine of π^0 angle



FIG. 5: The reconstructed π^0 momentum ((a) and (b)) and reconstructed π^0 angle with respect to the neutrino beam ((c) and (d)) for both the $2\gamma 1p$ ((a) and (c)) and $2\gamma 0p$ ((b) and (d)) final selected data. The prediction shows agreement with the observed data within assigned uncertainties for the ranges shown, although a systematic deficit is observed in the total event rates as is discussed in Sec. III.

 $_{621}$ 1 π^0 momentum increases, seen in Fig. 5a. This linear $_{639}$ the cross-section normalization uncertainties of coherent $_{622}$ scaling as a function of π^0 momentum was chosen be- $_{640}$ and non-coherent NC $1\pi^0$ are only removed for the pur-623 cause it was the simplest implementation that was con- 641 poses of this fit and not for the cross section extraction ⁶²⁴ sistent with the observed data-to-MC deficit, as observed ⁶⁴² described in the following section. 625 in Fig. 5a.

626 627 628 630 631 633 634 including bin-to-bin systematic correlations. As the goal 652 ally large uncertainties, the momentum-dependent shift $_{635}$ of the fit is to extract the normalization and scaling pa- $_{653}$ is preferred over the GENIE CV at the 1.43σ level. The $_{636}$ rameters of the coherent and non-coherent NC $1\pi^0$ rates, $_{654}$ 1D marginalized $\Delta\chi^2$ distributions in Fig. 9 also confirm $_{638}$ and non-coherent NC $1\pi^0$ are not included. Note that $_{656}$ uncertainty. The data and MC comparisons of the recon-

in Fig. 5a. At each set of fitting parameters (N_{coh}, a, b) , a χ^2 ⁶⁴³ The data-extracted best-fit parameters correspond to is evaluated between the scaled prediction for this pa-⁶⁴⁵ and a = 0.98 and b = -1.0 [c/GeV] for the scaling pa-⁶⁴⁵ between the scaled prediction for this pa-⁶⁴⁵ between the scaled prediction for the scale prediction for rameter set and the observed data using the Combined- $_{646}$ rameters of the non-coherent NC $1\pi^0$ events with a χ^2 per Neyman-Pearson χ^2 [50]. The χ^2 calculation makes $_{647}$ degree of freedom (dof) of 8.46/17. The χ^2/dof at the use of a covariance matrix including statistical and sys- $_{648}$ GENIE central value (CV) prediction is 13.74/20 yield-tematic uncertainties and correlations corresponding to $_{649}$ ing a $\Delta\chi^2$ between the GENIE CV and the best-fit point the scaled prediction. Flux, cross section, detector and $_{650}$ of 5.28 for 3 dof. Although the goodness-of-fit χ^2/dof GEANT4 systematic uncertainties are included in the fit 651 values for both scenarios are acceptable due to the generthe cross-section normalization uncertainties of coherent 655 that the GENIE CV prediction agrees with data within



(b) $2\gamma 0p$

FIG. 6: The reconstructed cosine of the center-of-mass angle for the (a) $2\gamma 1p$ and (b) $2\gamma 0p$ final selected data.

⁶⁵⁷ structed π^0 momentum distributions scaled to the best-⁶⁵⁸ fit parameters are provided in Fig. 10 and, compared to ⁶⁵⁹ those corresponding to the GENIE CV, show better agree-⁶⁶⁰ ment with data after the fit.

While the data suggest that GENIE may over-estimate 661 ₆₆₂ NC non-coherent $1\pi^0$ production and under-estimate NC coherent $1\pi^0$ production, the results demonstrate that 663 the GENIE prediction of NC $1\pi^0$ s is accurate within un-664 certainty. This validates the approach of using the measured NC $1\pi^0$ event rate as a powerful in situ constraint 666 of GENIE-predicted NC $1\pi^0$ backgrounds as in [4]. On the 667 other hand, it is natural to extract a data-driven NC $1\pi^0$ 669 cross-section on argon using these selections, and com-670 pare to a number of neutrino event generators, including 671 GENIE. This is described below.



Events

35

300

250

20

150 100

50

0.5

Data/Prediction

674



FIG. 7: The reconstructed conversion distance of both photons in the $2\gamma 1p$ final selection. There are well-reconstructed showers with conversion distances as far as 100 cm from the candidate neutrino interaction.

672 IV. INCLUSIVE AND EXCLUSIVE NC $1\pi^0$ 673 CROSS-SECTIONS ON ARGON

A. Methodology

The prescription for calculating the cross section is provided in Eq. (5) where the components are defined as follows: $N_{NC1\pi^0}^{\text{obs}}$, N_{cosmic} , and N_{bkg} denote the number of selected data events, the number of background events arising from cosmic rays traversing the detector, and the number of expected beam-correlated background events, respectively; $\epsilon_{NC1\pi^0}$ denotes the efficiency of selecting NC1 π^0 events; Φ denotes the integrated flux; and N_{targets} denotes the number of argon atoms in the fiducial volume of the analysis.

$$\sigma_{NC1\pi^0} = \frac{N_{NC1\pi^0}^{\text{obs}} - N_{\text{cosmic}} - N_{bkg}}{\epsilon_{NC1\pi^0} \Phi N_{\text{targets}}}.$$
 (5)

This calculation is performed independently using each 686 of the $2\gamma 1p$ and $2\gamma 0p$ selections to measure an exclusive 687 cross section. These measurements are denoted as the 688 NC π^0 +1p and NC π^0 +0p cross sections, respectively; in 689 each case one or zero protons is explicitly required in the signal definition (described in detail below). Addi-690 tionally, the calculation is performed using the combined 691 $2\gamma(0+1)p$ selection to measure a semi-inclusive cross section, $NC\pi^0$, with no requirement on the number of protons in the signal definition. Note that this semi-604 ⁶⁹⁵ inclusive measurement is efficiency-corrected to include 2+ proton final states that are not included in the final ⁶⁹⁷ selected events. As noted in Section IIC, the simulation is run multiple times to encompass the effect of varying 698 ⁶⁹⁹ underlying sources of systematic uncertainty. The cal-⁷⁰⁰ culation of each cross section is performed separately in





FIG. 8: Event displays of candidate NC $1\pi^0$ events found in the MicroBooNE data using (a) the $2\gamma 1p$ selection and (b) the $2\gamma 0p$ selection, on the MicroBooNE TPC collection plane. The horizontal axis here corresponds to the increasing wires, with an associated distance in cm. The vertical axis represents the TPC drift time. The aspect ratio of this plot is set such that the length scale shown for the horizontal axis is the same for the vertical axis.

701 each of these systematic "universes" to guarantee that 702 all correlations between components of the cross section 703 are handled correctly. This is done using tools from the MINERvA Analysis Toolkit [51]. 704

Both selections, as well as their combination, corre-705 706 707 708 709 $_{711}$ by the ratio between the POT of the $2\gamma 1p$ data sample $_{741}$ ing the total number of simulated events truly satisfying $_{712}$ (smaller POT) and the POT of the $2\gamma 0p$ data sample $_{742}$ the corresponding signal definition. In each of the exclu-⁷¹³ and then are added to the $2\gamma 1p$ distributions. This op-⁷⁴³ sive measurements, the signal definition is taken to be ⁷¹⁴ eration is performed for $N_{NC1\pi^0}^{\text{obs}}$, N_{cosmic} , N_{bkg} , and the ⁷⁴⁴ NC1 π^0 with exactly zero or one final-state proton with a 715 numerator of the efficiency.

716 717 fore there is no systematic uncertainty attributed to 747 notably allowing for any number of protons in the fi- $_{718}$ them. These values are reported in Table II. N_{bkg} is $_{748}$ nal state. The efficiency for each analysis is reported in 719 extracted from the simulation, and we note that many 749 Table II. 720 of the key backgrounds in this analysis are shared with 750 The integrated flux is calculated separately for all four



FIG. 9: The distribution of marginalized $\Delta \chi^2$, as a function of flat normalization factor for (a) coherent NC $1\pi^0$ momentum-independent scaling factor, (b) non-coherent NC $1\pi^0$ momentum-independent scaling factor, and (c) coefficient of momentum-dependent scaling factor for NC non-coherent $1\pi^0$, marginalized over the other two parameters. The red arrows indicate parameter values expected for the GENIE central value prediction. The 1σ , 90% and 99% C.L. lines are based on the assumption that the distribution follows a χ^2 distribution with 1 degree of freedom.

722 dominant contributions to the uncertainty on the back-723 ground event rate for each analysis are from FSI re-724 lated to inelastic nucleon scattering, pion and nucleon 725 absorption, and pion charge-exchange. The axial and ⁷²⁶ vector mass parameters, m_A and m_V , respectively, in the 727 charged current resonant form factors are also sources of 728 significant uncertainties; this is consistent with expec-729 tation because of the large background due to charged-⁷³⁰ current interactions in which a π^0 is produced.

731 The efficiency of the selection is constructed using as 732 the numerator the number of signal events passing all re-733 construction cuts and analysis BDTs in simulation and 734 as the denominator the total number of signal events pre-⁷³⁵ ceding the application of any cuts or analysis BDTs. The spond to approximately, but not identically, the same 736 difference in signal definition between the semi-inclusive POT, provided in Table II (due to differences in the com- 737 measurement and each of the two exclusive measureputational processing of the two samples). To extract the 738 ments is contained in the efficiency denominator. The semi-inclusive cross section from the combined $2\gamma(0+1)p_{739}$ exclusive measurements and the semi-inclusive measureselection, the relevant $2\gamma 0p$ distributions are scaled down ₇₄₀ ment each use a distinct efficiency denominator, reflect-745 kinetic energy above 50 MeV. In the semi-inclusive mea- $N_{NC1\pi^0}^{\text{obs}}$ and N_{cosmic} are measured in data and there-

⁷²¹ MicroBooNE's search for NC Δ radiative decay [4]. The τ_{51} neutrino species $(\nu_{\mu}, \bar{\nu}_{\mu}, \nu_{e}, \bar{\nu}_{e})$, and the sum of these





FIG. 10: The data-MC comparison for (a) $2\gamma 1p$ and (b) $2\gamma 0p$ selections, as a function of reconstructed π^0 momentum. Monte Carlo predictions at the central value and at the best-fit point ($N_{coh} = 2.6, a = 0.98$, b = -1.0 [c/GeV]) are both shown, with prediction and corresponding systematic error evaluated at the GE-NIE central value in salmon, and at the best-fit in blue. Note that the systematic uncertainties on the plot include MC intrinsic statistical error and all the systematic errors (flux, cross-section and detector), with the exception of cross section normalization uncertainties on coherent and non-coherent NC $1\pi^0$.

752 integrated fluxes is used to normalize each cross section 753 measurement. This choice was made because of the inability to identify the species of the incident neutrino 754 based on the neutral current final state. The integrated 755 flux is varied within each flux systematic "universe", and 756 the correlations between each varied flux and the corre-757 sponding variations in the predicted background and ef-758 ficiency are taken into account when extracting the cross 759 760 sections.

The number of argon atoms used is calculated as 761

 $_{^{762}}N_{\rm targets} = \rho V N_A/M_{Ar},$ where $V = 5.64 \times 10^7 {\rm cm}^3$ is the ⁷⁶³ fiducial volume of the analysis, $\rho = 1.3954$ g/cm³ is the 764 density of argon at the temperature in the cryostat, and $_{765}$ $M_{Ar} = 39.948$ g/mol is the molar mass of argon. A 1% ⁷⁶⁶ uncertainty is assigned to the number of targets to reflect 767 variation in the argon density through temperature and 768 pressure fluctuations.

Results and Interpretation в.

769

The calculation of each cross section from its compo-770 ⁷⁷¹ nents follows from Eq. (5) and is summarized in Table II. The resulting cross sections are shown in Fig. 11, com-772 pared to the simulated cross sections from several neutrino event generators including GENIE, NUWRO [52], ⁷⁷⁵ and NEUT [53]. The GENIE curve shown is generated ⁷⁷⁶ using the MicroBooNE cross-section "tune" [28], which does not modify the GENIE v3.0.6 central value predic-777 tion (because the tune did not adjust the NC interaction 778 779 model), but does define the uncertainty on the predic-780 tion. We observe a consistent deficit in data compared to GENIE for the combined semi-inclusive measurement 781 and for each of the individual NC π^0 +1p and NC π^0 +0p 782 exclusive measurements. Overall, the NEUT predictions 783 most closely match the reported measurements across 784 semi-inclusive and exclusive final states. Additionally we note that while NUWRO is generally consistent with 786 the other generators in its semi-inclusive and exclusive 787 ⁷⁸⁸ 1p predictions, its exclusive 0p prediction is higher com-789 pared to NEUT and GENIE predictions. The extracted semi-inclusive NC π^0 cross section is 1.24 \pm 0.19 (syst) \pm 790 0.08 (stat) $[10^{-38} \text{cm}^2/\text{Ar}]$ which is 26% lower than the 791 GENIE prediction of $1.68 [10^{-38} \text{cm}^2/\text{Ar}]$. 792

The corresponding breakdown of uncertainty for each 793 of the measurement channels is shown in Fig. 12. In all 794 cases the flux, GENIE, and statistical uncertainties are 795 dominant. The dominant contributions to the GENIE un-796 certainties enter into the cross section via the background subtraction and, as noted above, arise from the modeling 798 of final-state interactions and the axial and vector mass 799 ⁸⁰⁰ parameters governing CC resonant pion production.

To further understand this measurement, it is instruc-801 tive to compare it to previous experimental measure-802 ments of $NC\pi^0$ production. We compare our measure-⁸⁰⁴ ment to that performed by MiniBooNE which operated ⁸⁰⁵ in the same beamline as MicroBooNE but which utilized 806 a different detector material (mineral oil, CH_2) as the $_{\rm 807}$ neutrino scattering target. In MiniBooNE's NC π^0 anal-⁸⁰⁸ ysis, they measured NC interactions wherein only one $_{1009}$ π^0 and no additional mesons exited the target nucleus ^{\$10} (no requirement on the number or identity of outgoing ⁸¹¹ nucleons was made). A final flux-averaged cross section $_{\rm 812}$ of 4.76 \pm 0.76 \pm 0.05 $[10^{-40} {\rm cm}^2/{\rm nucleon}]$ was reported [25]. We can compare this result to our semi-inclusive re-813 sult by comparing each to the same neutrino generator. ⁸¹⁵ This is shown in Fig. 13 where we compare both to the ⁸¹⁶ default GENIE v3.0.6 on argon and mineral oil respec-

TABLE II: Summary table of all inputs to the cross section calculation, reported as $\sigma \pm sys \pm stat$ uncertainty. Note that while the individual errors on the components are given here, the full uncertainty on the cross section is calculated properly assuming full correlations.

	$\mathbf{NC}\pi^0$ (semi-inclusive)	$\mathbf{N}\mathbf{C}\pi^0 + 1p$ (exclusive)	NC π^0 + 0 <i>p</i> (exclusive)
Samples Used	$2\gamma(0+1)p$ Selection	$2\gamma 1p$ Selection	$2\gamma 0p$ Selection
$N_{targets} [10^{30} \text{ Ar atoms}]$		$1.187 \pm 0.119 \pm 0.00$	
Flux $[10^{-10} \ \nu/\text{POT/cm}^2]$		$7.876 \pm 0.902 \pm 0.00$	
POT of sample $[10^{20} \text{ POT}]$	$5.84 \pm 0.12 \pm 0.00$	$5.84 \pm 0.12 \pm 0.0$	$5.89 \pm 0.12 \pm 0.00$
Efficiency	$0.089 \pm 0.003 \pm 0.001$	$0.107 \pm 0.006 \pm 0.002$	$0.060 \pm 0.003 \pm 0.001$
Selected data [evts]	$1125.9 \pm 0.0 \pm 33.5$	$634.0 \pm 0.0 \pm 25.2$	$496.0 \pm 0.0 \pm 22.3$
Cosmic data [evts]	$177.0 \pm 0.0 \pm 8.9$	$96.1 \pm 0.0 \pm 6.5$	$81.5 \pm 0.0 \pm 6.1$
Background [evts]	$345.8 \pm 51.1 \pm 9.0$	$279.6 \pm 43.5 \pm 7.2$	$208.3 \pm 33.5 \pm 7.0$
Background-subtracted rate [evts]	$603.2 \pm 51.1 \pm 35.8$	$258.3 \pm 43.5 \pm 27.0$	$206.1 \pm 33.5 \pm 24.1$
$\sigma_{NC1\pi^0} \ [10^{-38} {\rm cm}^2/{\rm Ar}]$	$1.243 \pm 0.185 \pm 0.076$	$0.444 \pm 0.098 \pm 0.047$	$0.624 \pm 0.131 \pm 0.075$



FIG. 11: Measured semi-inclusive $NC\pi^0$, exclusive NC π^0 +1p, and exclusive NC π^0 +0p cross sections, each compared to the corresponding GENIE v3 (G18_10a_02_11) cross section and its uncertainty (shaded red bands) as well as other contemporary neutrino generators.

^{\$18} slightly below the expected central value, both our re- ^{\$33} rently analyzed MicroBooNE data statistics, the nominal ⁸¹⁹ sult and MiniBooNE's agree with GENIE v3.0.6 within ⁸³⁴ GENIE neutrino event generator used for MicroBooNE 820 assigned uncertainties.

821

v. SUMMARY

In summary, we report the highest statistics measure- 840 822 823 production on argon, including the first exclusive mea-824 $_{826}$ cross sections are measured using the MicroBooNE detec- $_{844}$ 0.098 ±0.047, and 0.624 ±0.131 ±0.075 [10^{-38} cm²/Ar] for $_{227}$ tor exposed to the Fermilab Booster Neutrino Beamline, $_{245}$ the semi-inclusive NC π^0 , exclusive NC π^0 +1p, and excluwhich has $\langle E_{\nu} \rangle < 1$ GeV. As presented within this paper, ⁸⁴⁶ sive NC π^0 +0p processes compared to 1.678, 0.722, and



FIG. 12: Error budget for the semi-inclusive $NC\pi^0$, exclusive NC π^0 +1p, and exclusive NC π^0 +0p cross section measurements.

 $_{s29}$ kinematic distributions of the π^0 momentum and angle ⁸³⁰ relative to the beam direction provide some sensitivity 831 to contributions to this process from coherent and non-^{\$17} tively. We observe that while this result on argon lies ^{\$32} coherent pion production and suggest that, given the cur-835 Monte Carlo modeling describes the observed distribu-⁸³⁶ tions within uncertainties. This has provided an impor-⁸³⁷ tant validation check justifying the use of this sample as ⁸³⁸ a powerful constraint for backgrounds to single-photon ⁸³⁹ searches in MicroBooNE, *e.q.* in [4].

Using a total of 1,130 observed NC π^0 events, a fluxment to date of neutrino neutral current single pion 841 averaged cross section has been extracted for neutrinos ⁸⁴² with a mean energy of 804 MeV and has been found to surements of this process ever made in argon. These 843 correspond to 1.243 ± 0.185 (syst) ± 0.076 (stat), 0.444 ± 0.185 (syst) ± 0.076 (stat), 0.044 ± 0.185 (syst) (syst) ± 0.076 (stat), 0.444 ± 0.185 (syst) (sys



FIG. 13: Comparison of this semi-inclusive result on argon (left) as well as that from MiniBooNE on mineral oil CH_2 (right), to the same GENIE v3.0.6. While the published MiniBooNE result was originally compared to a prediction made using the NUANCE v3 generator [54], we have instead generated a prediction using GENIE to aid in a comparison between the two experimental results. MiniBooNE's statistical uncertainty is small and only the systematic error bar is visible. Shaded error bands show GENIE uncertainty only.

⁸⁴⁹ ing NEUT and NUWRO show reasonable agreement with ⁸⁷⁶ not have been possible.

850 the NEUT predictions found to be slightly more consistent with the MicroBooNE data-extracted cross-section 851 ⁸⁵² for all three exclusive and semi-inclusive processes.

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- [1] O. Benhar, P. Huber, C. Mariani, and D. Meloni, Phys. 877 Rept. 700, 1 (2017), arXiv:1501.06448 [nucl-th]. 878
- M. Antonello et al. (MicroBooNE, LAr1-ND, ICARUS-[2]879 WA104), (2015), arXiv:1503.01520 [physics.ins-det]. 880
- B. Abi et al. (DUNE), (2020), arXiv:2002.03005 [hep-ex]. 881 [3]
- [4] P. Abratenko et al. (MicroBooNE), Phys. Rev. Lett. 128, 882
- 111801 (2022), arXiv:2110.00409 [hep-ex]. 883 and 910 E. Bertuzzo, S. Jana, P. A. N. Machado, 884 $\left| 5 \right|$
- R. Zukanovich Funchal, Phys. Rev. Lett. 121, 241801 911 885 (2018), arXiv:1807.09877 [hep-ph]. 886
- [6] P. Ballett, S. Pascoli, and M. Ross-Lonergan, Phys. Rev. 887 D 99, 071701 (2019), arXiv:1808.02915 [hep-ph]. 888
- A. Abdullahi, M. Hostert, and S. Pascoli, Phys. Lett. B [7]889 820, 136531 (2021), arXiv:2007.11813 [hep-ph]. 890
- [8] B. Dutta, S. Ghosh, and T. Li, Phys. Rev. D 102, 055017 891 (2020), arXiv:2006.01319 [hep-ph]. 892
- W. Abdallah, R. Gandhi, and S. Roy, Phys. Rev. D 104, 893 [9] 055028 (2021), arXiv:2010.06159 [hep-ph]. 894
- [10] D. Cianci, A. Furmanski, G. Karagiorgi, and 921 895 M. Ross-Lonergan, Phys. Rev. D 96, 055001 (2017), 896 arXiv:1702.01758 [hep-ph]. 897
- [11] A. P. Furmanski and C. Hilgenberg, Phys. Rev. D 103, 898 112011 (2021), arXiv:2012.09788 [hep-ex]. 800
- R. Acciarri et al. (MicroBooNE), JINST 12, P02017 [12]900 (2017), arXiv:1612.05824 [physics.ins-det]. 901
- 902 [13] A. A. Aguilar-Arevalo et al. (MiniBooNE), Phys. Rev. D 928 [27] S. Nakayama et al. (K2K), Phys. Lett. B 619, 255 (2005),

79, 072002 (2009).

903

- L. Aliaga et al. (MINERvA), Phys. Rev. D 94, 092005 [14]904 (2016), arXiv:1607.00704 [hep-ex]. 905
- 906 [15] M. A. Acero et al. (NOvA), Phys. Rev. D 102, 012004 (2020), arXiv:1902.00558 [hep-ex]. 907
- 908 [16]K. Abe *et al.* (T2K), Phys. Rev. D 87, 012001 (2013), arXiv:1211.0469 [hep-ex]. 909
 - K. Abe et al. (Hyper-Kamiokande Proto-Collaboration), 17 PTEP **2015**, 053C02 (2015), arXiv:1502.05199 [hep-ex].
- 912 18 R. Acciarri et al. (ArgoNeuT), Phys. Rev. D 96, 012006 (2017), arXiv:1511.00941 [hep-ex]. 913
- S. J. Barish et al., Phys. Rev. Lett. 33, 448 (1974). [19]914
- M. Derrick et al., Phys. Rev. D 23, 569 (1981). [20]915
- 916 [21]W.-Y. Lee et al., Phys. Rev. Lett. 38, 202 (1977).
- 917 [22] W. Krenz et al. (Gargamelle Neutrino Propane, Aachen-Brussels-CERN-Ecole Poly-Orsay-Padua), Nucl. Phys. B 918 **135**, 45 (1978). 919
- A. A. Aguilar-Arevalo et al. (MiniBooNE), Phys. Lett. B 920 23 664, 41 (2008), arXiv:0803.3423 [hep-ex].
- A. A. Aguilar-Arevalo et al. (MiniBooNE), Phys. Rev. D 922 |24|83, 052009 (2011). 923
- A. A. Aguilar-Arevalo et al. (MiniBooNE), Phys. Rev. D 924 [25] 81, 013005 (2010), arXiv:0911.2063 [hep-ex]. 925
- 926 [26] Y. Kurimoto et al. (SciBooNE), Phys. Rev. D 81, 033004 (2010).927

- arXiv:hep-ex/0408134. 929
- P. Abratenko et al. (MicroBooNE), (accepted by Phys. 930 [28]959 Rev. D) (2021), arXiv:2110.14028 [hep-ex]. 931
- C. Andreopoulos et al., Nucl. Instr. and Meth. A 614, 87 [29]932 961 (2010), arXiv:0905.2517 [hep-ph]. 933
- [30]J. Tena-Vidal et al. (GENIE), Phys. Rev. D 104, 072009 934 (2021), arXiv:2104.09179 [hep-ph]. 935
- D. Rein and L. M. Sehgal, Annals Phys. 133, 79 (1981). [31] 936
- C. Berger and L. M. Sehgal, Phys. Rev. D 76, 113004 [32]966 937 (2007), arXiv:0709.4378 [hep-ph]. 938
- K. M. Graczyk and J. T. Sobczyk, Physical Review D 77 [33] 939 (2008), 10.1103/physrevd.77.053003. 940
- K. S. Kuzmin, V. V. Lyubushkin, and V. A. Naumov, [34]941 Modern Physics Letters A 19, 2815–2829 (2004). 942
- J. A. Nowak, AIP Conference Proceedings [35] (2009).972 943 10.1063/1.3274164. 944
- [36]S. Agostinelli et al. (GEANT4), Nucl. Instrum. Meth. A 945 **506**, 250 (2003). 946
- E. Snider and G. Petrillo, Journal of Physics: Conference 947 Series 898, 042057 (2017). 948
- R. Acciarri et al. (MicroBooNE), JINST 12, P08003 949 [38](2017), arXiv:1705.07341 [physics.ins-det] 950
- C. Adams et al. (MicroBooNE), JINST 13, P07006 [39]980 951 (2018), arXiv:1802.08709 [physics.ins-det]. 952 981
- [40] C. Adams et al. (MicroBooNE), JINST 13, P07007 953 (2018), arXiv:1804.02583 [physics.ins-det]. 954
- [41] J. S. Marshall and M. A. Thomson, Eur. Phys. J. C75, 955 439 (2015). 956
- 957 [42] T. Chen and C. Guestrin, in Proceedings of the 22nd

ACM SIGKDD International Conference on Knowledge Discovery and Data Mining, KDD '16 (ACM, New York, NY, USA, 2016) pp. 785-794.

958

960

962

979

982

- [43] P. Abratenko et al. (MicroBooNE), Phys. Rev. Lett. 123, 131801 (2019), arXiv:1905.09694 [hep-ex].
- [44]J. Calcutt, C. Thorpe, K. Mahn, and L. Fields, JINST 963 16, P08042 (2021), arXiv:2105.01744 [physics.data-an]. 964
- Abratenko *et al.* (MicroBooNE), (2021),[45]Ρ. 965 arXiv:2111.03556 [hep-ex].
- [46]C. Adams et al. (MicroBooNE), JINST 15, P07010 967 (2020), arXiv:1910.01430 [physics.ins-det]. 968
- [47]P. Abratenko et al. (MicroBooNE), JINST 15, P12037 969 (2020), arXiv:2008.09765 [physics.ins-det]. 970
- D. Caratelli, Study of Electromagnetic Interactions in 971 [48] the MicroBooNE Liquid Argon Time Projection Chamber, Ph.D. thesis, Columbia University (2018). 973
- [49]P. A. Zyla et al. (Particle Data Group), PTEP 2020, 974 083C01 (2020). 975
- X. Ji, W. Gu, X. Qian, H. Wei, and C. Zhang, Nucl. [50]976 Instr. and Meth. A 961, 163677 (2020), arXiv:1903.07185 977 978 [physics.data-an].
 - B. Messerly et al. (MINERvA), EPJ Web Conf. 251, [51]03046 (2021), arXiv:2103.08677 [hep-ex].
 - [52]T. Golan, J. T. Sobczyk, and J. Zmuda, Nucl. Phys. B Proc. Suppl. 229-232, 499 (2012).
- Y. Hayato, Nucl. Phys. B Proc. Suppl. 112, 171 (2002). 983 53
- [54] D. Casper, Nucl. Phys. B Proc. Suppl. 112, 161 (2002), 984 arXiv:hep-ph/0208030. 985