## MicroBooNE and the $\nu_e$ Interpretation of the MiniBooNE Low-Energy Excess

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A new generation of neutrino experiments is testing the  $4.7\sigma$  anomalous excess of electron-like events observed in MiniBooNE. This is of huge importance for particle physics, astrophysics, and cosmology, not only because of the potential discovery of physics beyond the Standard Model, but also because the lessons we will learn about neutrino-nucleus interactions will be crucial for the worldwide neutrino program. MicroBooNE has recently released results that appear to disfavor several explanations of the MiniBooNE anomaly. Here, we show quantitatively that MicroBooNE results, while a promising start, unquestionably do not probe the full parameter space of sterile neutrino models hinted at by MiniBooNE and other data, nor do they probe the  $\nu_e$  interpretation of the MiniBooNE excess in a model-independent way.

Introduction.— A specter is haunting neutrino physics: the specter of sterile neutrinos. Over the past few decades, their existence has been postulated to explain various persistent anomalies in neutrino experiments [1], in particular the excess of electron-neutrino events in the LSND [2] experiment and the excess of electron-like events in the MiniBooNE [3] experiment. If real, sterile neutrinos would bring about a profound revolution to our understanding of early-universe cosmology [4], modify neutrino emission from astrophysical sources [5–7], and dethrone the Standard Model of particle physics. Precision measurements in particle physics, cosmology, and astrophysics will remain ambiguous until we clarify whether such sterile neutrinos exist.

Recently, the MicroBooNE collaboration has released results scrutinizing the MiniBooNE anomaly, also known as the MiniBooNE low-energy excess (MBLEE). As the MicroBooNE detector is exposed to the same neutrino beam as MiniBooNE but has superior event reconstruction capabilities, it is able to differentiate between different interpretations of the MBLEE. A first MicroBooNE analysis disfavors that the MBLEE is due to underestimated production of  $\Delta$  baryons followed by decays to photons at a significance of 94.8% C.L. [8].

Following that, three distinct and complementary analyses have been released, addressing whether the MBLEE is caused by an excess of electron-neutrinos in the beam [9–12]. These compare the MicroBooNE data to a signal template defined by the assumption that the expected spectrum of the  $\nu_e$  excess in MiniBooNE matches exactly the difference between the data and the best-fit background prediction. Assuming this nominal template, the collaboration concludes that "the results are found to be consistent with the nominal  $\nu_e$  rate expectations from the Booster Neutrino Beam and no excess of  $\nu_e$  events is observed" [9]. However, as we will show, this approach is not sufficient to exclude the  $\nu_e$  interpretation of the MBLEE in a model-independent way or even to exclude the sterile neutrino solution of the MBLEE. This is an important question that can have an impact on the future of the short-baseline neutrino program, as well as on a variety of alternative models [13–18] proposed to explain the MiniBooNE and LSND results.

Previous attempts to constrain the LSND and Mini-BooNE excesses suffered from insufficient sensitivity to cover the allowed parameter space [19–23]. It is essential to quantify the power of MicroBooNE as we search for definitive answers to this 20-year-old puzzle. This is the goal of this letter. We first analyze the constraints of MicroBooNE's latest results on  $\nu_e$  appearance in Mini-BooNE in a more model-independent way, and then we narrow down the question to the sterile neutrino scenario.

For the first stage, we follow the procedure adopted by MicroBooNE: starting from a MiniBooNE event spectrum, we translate it into an expected excess of events in MicroBooNE, and we perform a statistical analysis of the data. After verifying that we reproduce Micro-BooNE's results when using the nominal template, we repeat the analysis with a set of alternative templates that are equally successful at explaining the MBLEE. These alternative templates are allowed mostly due to the systematic uncertainties on the backgrounds. As they dominate the significance of the MBLEE, we find that they change the interpretation of the MicroBooNE results significantly.

For the second stage, we perform a fit of MicroBooNE data to a simple light sterile neutrino model. We assume both a simplified oscillation scenario with only  $\nu_{\mu} \rightarrow \nu_{e}$  appearance, as well as a fully-consistent oscillation model that accounts for oscillations in the MicroBooNE control samples and backgrounds.

Experimental Analysis. — To quantify the disagreement between MicroBooNE data and a  $\nu_e$  interpretation of the MBLEE, we proceed as follows. All our analyses start with a hypothesis for the MBLEE.<sup>1</sup> The next step is the unfolding procedure, which generates an MBLEE template before detector effects. We apply the D'Agostini iterative method [25, 26], and we have checked upon inverting the unfolding procedure we recover the original spectrum before detector effects. To obtain the expected spectrum at MicroBooNE, we rescale the unfolded spectrum to account for the differences in exposure and detector mass between MiniBooNE and MicroBooNE [9]. We then smear the events according to MicroBooNE's energy resolution [10, 11]. To infer MicroBooNE's energydependent  $\nu_e$  detection efficiency, we apply all previous steps to the intrinsic  $\nu_e$  background, and we choose the efficiency in each reconstructed energy bin such that our results match the official MicroBooNE background prediction [10, 11]. We have checked that our efficiencies are consistent with the energy-averaged efficiencies quoted in Refs. [10, 11], and that they generate an MBLEE prediction that matches the official MicroBooNE result.

For the statistical procedure, we focus on the Charged-Current (CC) Inclusive [10] and Charged-Current Quasielastic (CCQE) [11] channels, as they comprise the leading statistical power of the experiment, and we use the data releases provided by the collaboration wherever possible. For the Inclusive analysis, we have performed several statistical tests, including both Pearson- $\chi^2$  and CNP- $\chi^2$  [27], as well as calculating a test statistic with or without deriving a constraint using the conditional covariance matrix formalism [10]. For all tests, we find very good agreement with the results of Ref. [10]. For clarity, in what follows we perform all statistical tests using the CNP- $\chi^2$  formalism with the full (137, 137) covariance matrix of Ref. [10]. For the CCQE analysis, in turn, we use a Poisson likelihood, where the expectation in each bin is treated as a nuisance parameter that is constrained by the covariance matrix. Our

<sup>1</sup> To directly compare with MicroBooNE results, we base our analysis on the 2018 MiniBooNE data [24].

test statistic in this latter analysis is then determined by profiling over these nuisance parameters.

Template results.— Our first approach assesses whether MicroBooNE generically rules out a  $\nu_e$  interpretation of the MBLEE. In particular, the MiniBooNE backgrounds have large, correlated systematic uncertainties. These allow for different shapes of the MBLEE, that could affect the prediction at MicroBooNE.

First, we generate a set of LEE templates using a Markov-Chain Monte-Carlo (MCMC) [28] by independently re-scaling the normalization of the MiniBooNE backgrounds. <sup>2</sup> We group backgrounds in four classes: all intrinsic  $\nu_e$ , mis-identified  $\pi^0$ ,  $\Delta \rightarrow \gamma$ , and all others. To estimate how well each template fits the MiniBooNE data, we calculate its goodness-of-fit *p*-value using a  $\chi^2$  test statistic with the official MiniBooNE covariance matrix [29]. We follow the prescription of the MiniBooNE collaboration, that the test-statistic follows a  $\chi^2$  distribution with 8.7 degrees of freedom (while our template model has 4 degrees of freedom). This generates a set of MBLEE templates compatible with MiniBooNE data. We will refer to this analysis as the template analysis.

Figure 1 exemplifies the impact of our procedure. We show three different templates for MiniBooNE (upper panel) and the corresponding predicted excess events at MicroBooNE (lower panel): the nominal template given by the difference between the observed data and the best-fit background estimate (black); a template with significantly more events at low energies but a *p*-value of 87% (blue); a template with fewer events and a *p*-value of 87% (red); and a template corresponding to the best-fit  $\nu_{\mu} \rightarrow \nu_{e}$  two-neutrino oscillation framework assuming  $\Delta m^{2} = 0.041 \text{ eV}^{2}$  and  $\sin^{2} 2\theta = 0.92$ , which has a *p*-value of 20% (red).

The systematic uncertainties in the MiniBooNE background allow for very different shapes and rates of the MBLEE (including those coming from the sterile neutrino hypothesis) to provide a good fit to the data. These generate different predictions at MicroBooNE that will be excluded with different statistical significance. As we will show, this is crucial in constraining the interpretation of the MBLEE in terms of  $\nu_e$  events at MicroBooNE.

Let us now turn to the results of our template analysis. For concreteness, for the template analysis we focus on the Inclusive analysis, which provides the best constraints on the  $\nu_e$  interpretation of the MBLEE. We organize the templates in three categories by decreasing goodness-of-fit, p > 80%, 10%, and 1%, and classify them by their signal strength, defined as  $N/N_{\rm LEE}$  with N the number of excess events that the template predicts at MiniBooNE, and  $N_{\rm LEE} = 360$  the number of excess events.

 $<sup>^2</sup>$  Alternatively, we have also generated templates by independently varying the excess events in each bin. Both approaches produce consistent results.



FIG. 1. The event rate at MiniBooNE (top) and Micro-BooNE (bottom) in reconstructed neutrino energy. We show several MiniBooNE templates, including that of the 3+1 oscillation best-fit, as non-stacked histograms. For MicroBooNE we show the predictions based on our unfolding procedure together with the data and unconstrained background to fully contained events.

Figure 2 shows the result of the template analysis. Each point corresponds to a different template with a well-defined signal strength and  $\Delta \chi^2_{\mu B}$ . To indicate the proximity to the best-fit MiniBooNE template, the points are colored by the three goodness-of-fit categories defined above. For each template, we compute the corresponding MicroBooNE  $\chi^2_{\mu B}$ , and construct the difference with respect to the no-excess hypothesis,  $\Delta \chi^2_{\mu B}$ . We also show as a black line the result of profiling over all templates with the same signal strength. The horizontal lines then correspond to the MicroBooNE 1-, 2-, 3-, and 4- $\sigma$  exclusion limits [30, 31].

As we can see in Fig. 2, the introduction of shape and normalization uncertainties in the MBLEE template can either enhance or mitigate MicroBooNE's sensitivity. To illustrate the variability of the template shapes and normalization, we have marked with a star the two templates shown in Fig. 1, corresponding to two extreme points in the p > 80% region.

Nevertheless, several templates that are a good fit to MiniBooNE data cannot be excluded by MicroBooNE. In particular, we observe a large density of templates with good fits to MiniBooNE data, p > 80% (10%), well below the  $\Delta \chi^2_{\mu B} = 9$  (4) line. We thus conclude that, while recent MicroBooNE results indeed constrain the



FIG. 2. The  $\Delta \chi^2$  of the Inclusive  $\nu_e CC$  analysis at Micro-BooNE for various templates found by our MCMC. Each point corresponds to a specific template that provides a good fit to MiniBooNE data with a *p*-value greater than 80%, 10%, and 1%. Horizontal lines correspond to  $\Delta \chi^2_{\mu B} = 1, 4, 9, 16$ . The stars correspond to templates 1 and 2 presented in Fig. 1.

 $\nu_e$  interpretation of the MiniBooNE excess in a model-independent way, they do not completely rule it out.

Sterile Neutrino Analysis.— The analysis above does not rely on any particular particle physics model. As an example of a specific physics model that can explain the MBLEE, we now turn to light sterile neutrinos. They provide a simple scenario that could lead to  $\nu_{\mu} \rightarrow \nu_{e}$ transitions at short baselines and have been extensively studied in the literature [1, 4, 32–34].

To perform analyses including sterile neutrinos, we use Ref. [35] and first calculate, as a function of oscillation parameters, the expected excess spectrum in MiniBooNE. Using the same procedure discussed above, we map these spectra into the expected excesses in MicroBooNE's Inclusive and CCQE analyses. We note that the efficiencies for MiniBooNE and MicroBooNE CCQE analyses decrease at large energies, while in the Inclusive Micro-BooNE analysis they stay constant.

Leveraging the data released coincident with Refs. [10, 11], we can also account for oscillations of the  $\nu_{\mu}$  and  $\nu_{e}$  CC background expectations <sup>3</sup> in MicroBooNE's analyses to allow for a complete, four-neutrino oscillation analysis.

We start by discussing the results of the simplified sterile neutrino model, which assumes the backgrounds to be independent of the sterile neutrino parameters. This

<sup>&</sup>lt;sup>3</sup> More concretely, we use the data releases provided along with Ref. [10] to determine the expected  $\nu_{\mu,e}$  CC fully- and partiallycontained spectra as a function of *true* neutrino energy. Oscillations are included with respect to true energy, then the distributions are mapped into *reconstructed* neutrino energy (again, using Ref. [10]) where test statistics are calculated.



FIG. 3. MicroBooNE constraints on the sterile neutrino parameter space at  $3\sigma$  C.L. (blue, inclusive and orange, CCQE). For reference, we show the MiniBooNE 1-, 2-, and  $3-\sigma$  preferred regions in shades of grey [24], the future sensitivity of the three SBN detectors (pink [36]), and existing constraints from KARMEN (green [19]) and OPERA (gold [37]).

simplified model is parametrized by a squared-mass difference  $\Delta m_{41}^2$  and an effective mixing angle  $\sin^2 2\theta_{\mu e} \equiv 4|U_{e4}U_{\mu4}|^2$  with U the leptonic mixing matrix. Figure 3 presents the results of our analyses of MicroBooNE's inclusive and CCQE channels in blue and orange, respectively, at  $3\sigma$  C.L.

While MicroBooNE data disfavors part of the region preferred by MiniBooNE, we find that there is still a large viable fraction of the parameter space, even within the  $1\sigma$ preferred region of MiniBooNE. We find it unlikely that future MicroBooNE results will significantly improve on this, even though MicroBooNE has only analyzed about half of their data set, because of a deficit in their Inclusive data that generates more sensitivity than expected (*c.f.* Fig. 1, this could be due to an underfluctuation in the data or to background mismodelling). This highlights the importance of searching for sterile neutrinos with the three SBN detectors — SBND, MicroBooNE, and ICARUS — which will probe the full  $2\sigma$  region preferred by MiniBooNE with less dependence on the neutrino cross section and flux.

Finally, we stress that a fully-consistent four-neutrino analysis should also consider oscillations of the backgrounds. This is relevant at MiniBooNE [16, 38], and even more for the MicroBooNE Inclusive analysis: while the former has large non-neutrino induced backgrounds, the dominant background in the latter is beam- $\nu_e$  contamination. Moreover, since other neutrino samples (particularly CC  $\nu_{\mu}$ ) are used to constrain systematics and



FIG. 4. MicroBoone constraints in the planes  $\Delta m_{41}^2$  versus  $\sin^2(2\theta_{ee})$  (left) and  $\sin^2(2\theta_{\mu\mu})$  (right). In each of these panels, we have either fixed (solid lines) or profiled over (dashed) the unseen mixing angle. For comparison, we show existing constraints and preferred regions (see Refs. [39–53]).

backgrounds, oscillations should also be considered for these samples. In this regard, the CCQE analysis is complementary and distinct from the Inclusive one, as it focuses on lower neutrino energies where the neutrino contribution to the background is small. To first order, the CCQE analysis performs an appearance search, while the Inclusive simultaneously searches for appearance and disappearance.

Figure 4 presents our results in a consistent fourneutrino approach, considering oscillations of all  $\nu_e$  and  $\nu_{\mu}$  samples. We show the MicroBooNE-Inclusive constraints in the planes  $\Delta m_{41}^2 vs. \sin^2 (2\theta_{ee}) \equiv 4|U_{e4}|^2 (1 - |U_{e4}|)^2$  (left panel) and  $\sin^2 (2\theta_{\mu\mu}) \equiv 4|U_{\mu4}|^2 (1 - |U_{\mu4}|)^2$  (right panel). As one can see, even in the absence of neutrino appearance, *i.e.*,  $U_{e4}$  or  $U_{\mu4}$  equal zero, Micro-BooNE can still set a limit on neutrino disappearance. For muon neutrinos, the disappearance sensitivity comes from the large  $\nu_{\mu}$  data sample. For electron neutrinos, on the other hand, the sensitivity derives from the large  $\nu_e$  background.

In the two panels of Fig. 4 we perform two analyses (both in blue): solid lines present the sensitivity on a mixing angle when the other is fixed to zero, whereas the dashed lines present the sensitivity when the other mixing angle has been profiled over. The disappearance prospects for  $\nu_e$  are compared against hints of sterile neutrinos in Gallium experiments [39–43] and constraints from solar [44], and reactor antineutrino [45–51] experiments. The bottom panel, showing MicroBooNE  $\nu_{\mu}$  disappearance constraints, is contrasted against constraints from MINOS/MINOS+ [52] and results from Ice-Cube [53], including a 90% C.L. preferred region and best-fit point.

As one final remark on the importance of the complete four-neutrino analysis, Fig. 5 shows the constraints from MicroBooNE-Inclusive as a function of  $\sin^2 (2\theta_{\mu e})$ after profiling over  $\sin^2 (2\theta_{\mu\mu})$  in comparison with preferred regions of MiniBooNE under the same set of assumptions [54]. Even more than in Fig. 3, we see that allowed MiniBooNE parameter space persists despite the MicroBooNE-Inclusive constraints.

Our results emphasize that, while the signal-oscillation only analysis is simple and intuitive, accounting for oscillations in *all* samples is the only fully-consistent approach and can affect the interpretation of the results. This will be even more relevant for the full SBN program, particularly due to different oscillation effects among the three detectors as well as due to the increased analysis sensitivity — we *strongly* advocate for the adoption of this standard moving forward with short-baseline searches for anomalous neutrino (dis)appearance.

In this complete picture, we find results consistent with no oscillations, with a best-fit point at  $\Delta m_{41}^2 = 1.38 \text{ eV}^2$ ,  $\sin^2 (2\theta_{ee}) = 0.2$ ,  $\sin^2 (2\theta_{\mu\mu}) = 0$  with a significance of  $0.95\sigma$ . Our results do not agree with Ref. [55]. We believe that this stems from the treatment of systematic uncertainties, as we consider correlated systematics by using the official MicroBooNE covariance matrix; and due to the fact that we account for oscillations in the partially-contained  $\nu_e$  sample, which is used to obtain the constrained fully-contained  $\nu_e$  sample. We also implement oscillations as a function of true neutrino energy.

Conclusions. — Does MicroBooNE rule out the  $\nu_e$  interpretation of the MiniBooNE low-energy excess? And does it disfavor the sterile neutrino explanation of the excess? While current MicroBooNE analyses give us invalu-



FIG. 5. MicroBooNE constraint as a function of  $\Delta m_{41}^2$ and  $\sin^2(2\theta_{\mu e})$  after profiling over the unseen mixing angle in a consistent four-neutrino analysis. Preferred MiniBooNE regions [54] are shown in grey. Other constraints/projections, faded for clarity, are identical to those in Fig. 3.

able insights on the MiniBooNE anomaly, we find that they still do not provide definitive answers to either of these two questions. Uncertainties on MiniBooNE backgrounds significantly impact MicroBooNE's reach, and consequently, the MiniBooNE puzzle remains wide open. To demonstrate this quantitatively, we have developed a model-independent analysis and have carried out a fullyconsistent sterile neutrino fit of MicroBooNE data in the context of the MiniBooNE excess. In the first analysis, we find MiniBooNE excess templates with goodness-of-fit better than 10% that are allowed by MicroBooNE data at  $< 2\sigma$ . In the sterile neutrino analysis, we find that MicroBooNE's  $3\sigma$  exclusion does not cover the entire  $1\sigma$ preferred region by MiniBooNE.

Our findings highlight the importance of running the full SBN program, and of complementing it with the worldwide efforts to search for light sterile neutrinos in reactor [56–60], radioactive source [39], accelerator [61–64], solar [44, 65], and atmospheric neutrino [53, 66–70] experiments. Together, these experiments will have sufficient sensitivity to answer this decades-old puzzle once and for all.

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## Supplemental Material

## Comparison of Results to Asimov Sensitivity Expectations

In the Sterile Neutrino Analysis section, see Fig. 3, we presented the  $3\sigma$  derived constraints on  $\sin^2 (2\theta_{\mu e})$  vs.  $\Delta m_{41}^2$  from the MicroBooNE Inclusive and CCQE analyses, and contrasted these constraints against existing and proposed future results. For completeness, in this appendix, we provide a comparison between the results from MicroBooNE data and the expected Asimov sensitivity, i.e., replacing the observed data with the expected background in our test statistics.



SUPPL. FIG. 1. Comparison of MicroBooNE constraints (solid lines) with the expected Asimov sensitivity (dashed) of each analysis, both at  $3\sigma$  C.L. Blue lines correspond to the inclusive analysis, where orange correspond to the CCQE one. Constraints/preferred regions in this panel, dimmed for clarity, are identical to those shown and labeled in Fig. 3.

Suppl. Fig. 1 presents this comparison. As discussed in the *Template Analysis* section, the deficit of observed events relative to the expected background rate leads to a more powerful constraint than expected in both analyses, but especially so in the Inclusive one (see e.g. Fig. 1).

Suppl. Fig. 2 repeats this data/Asimov comparison for the results of Fig. 4, comparing data results in blue with Asimov expectations in purple. Here we focus, as in the main text, on the MicroBooNE Inclusive analysis. In both panels, as in Fig. 4, solid lines correspond to the analysis with the other mixing angle fixed to zero, whereas the dashed lines include profiling over the unseen angle. We see in the right panel that the data result exceeds the sensitivity greatly – this is driven by the observed  $\nu_{\mu}$  CC excess observed in both the Fully-Contained and Partially-Contained samples. These samples also modify the expectation of  $\nu_e$  CC event rates due to large correlations, which has a non-negligible impact on the sensitivity to  $\sin^2(2\theta_{ee})$  in the left panel.



SUPPL. FIG. 2. Comparison of MicroBooNE constraints (blue) with the Asimov sensitivity expectation (purple) at 95% C.L. as a function of  $\sin^2 (2\theta_{ee})$  and  $\sin^2 (2\theta_{\mu\mu})$  vs.  $\Delta m_{41}^2$ . Faded lines are identical to those in Fig. 4 and are only dimmed for clarity of comparison. In each panel, the solid (dashed) lines correspond to fixing (profiling over) the unseen mixing angle.