

# New $\nu_e + \bar{\nu}_e$ appearance results from NOvA

Bruce Howard | *Indiana University*

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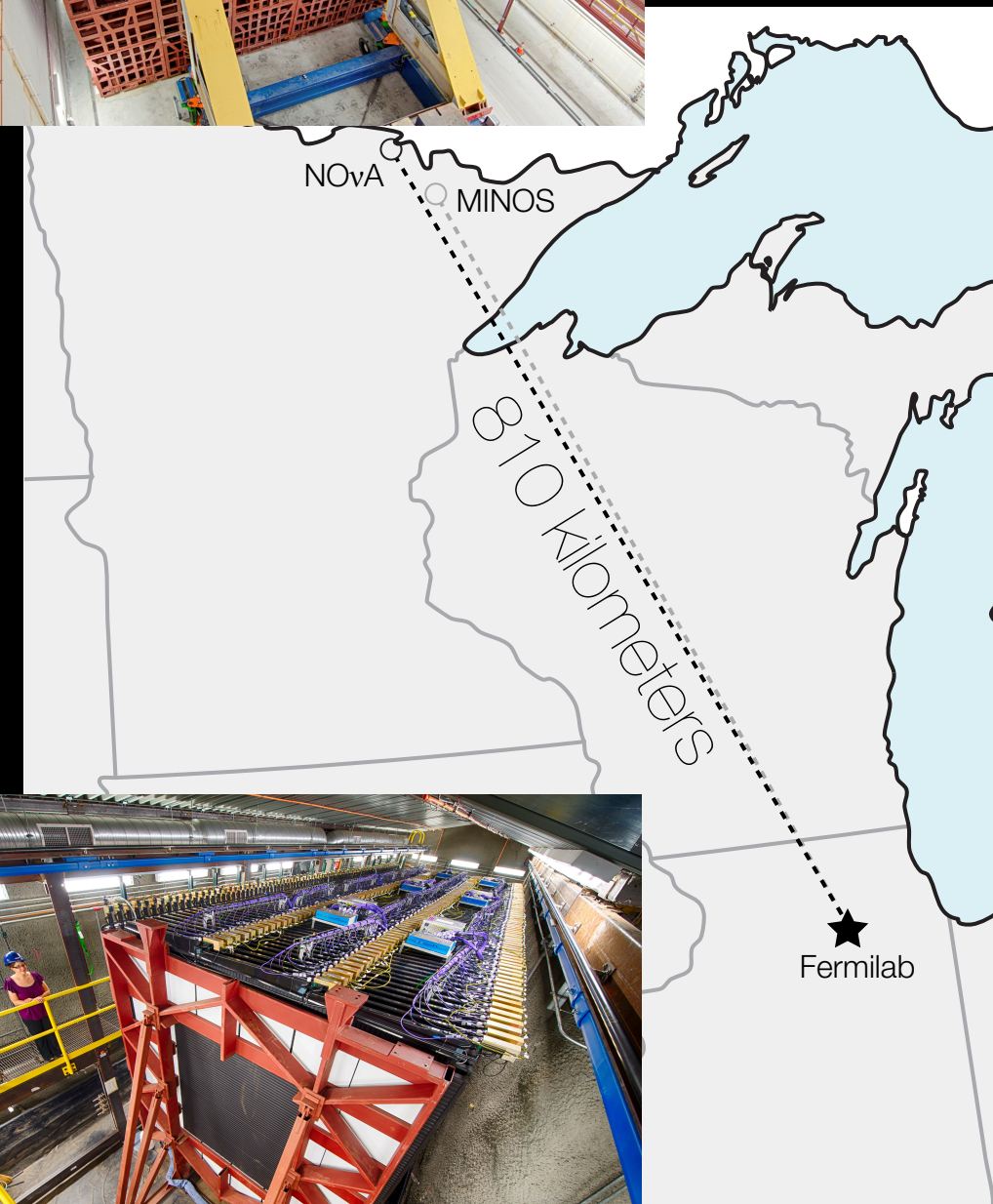
*New Perspectives 2018*  
19 June 2018

# Context

- Many details in talk from J. Vasel
- Functionally-similar near and far detectors separated by 810 km
- In addition to  $\nu_\mu$  disappearance, can study appearance  $\nu_\mu \rightarrow \nu_e$  ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ ) to probe 3-flavor neutrino oscillations
  - Mixing angles and octant sensitivity?
  - Mass splitting and hierarchy?
  - CP violation?
- Updated NOvA analysis with, for the first time, antineutrino beam data
  - Antineutrino dataset  $6.9 \times 10^{20}$  protons on target (neutrino dataset  $8.9 \times 10^{20}$  full-detector equivalent protons on target)

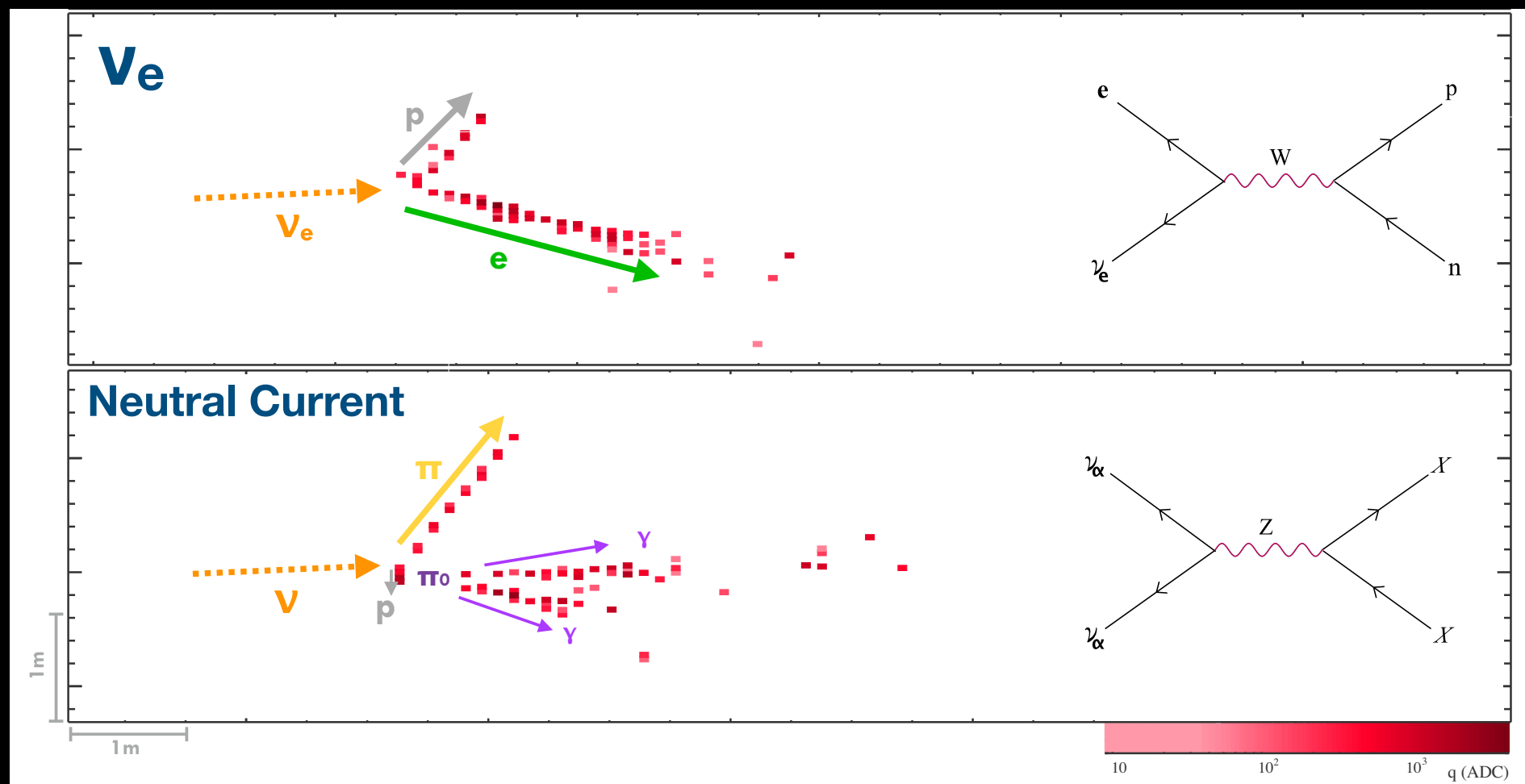


Credit: Fermilab



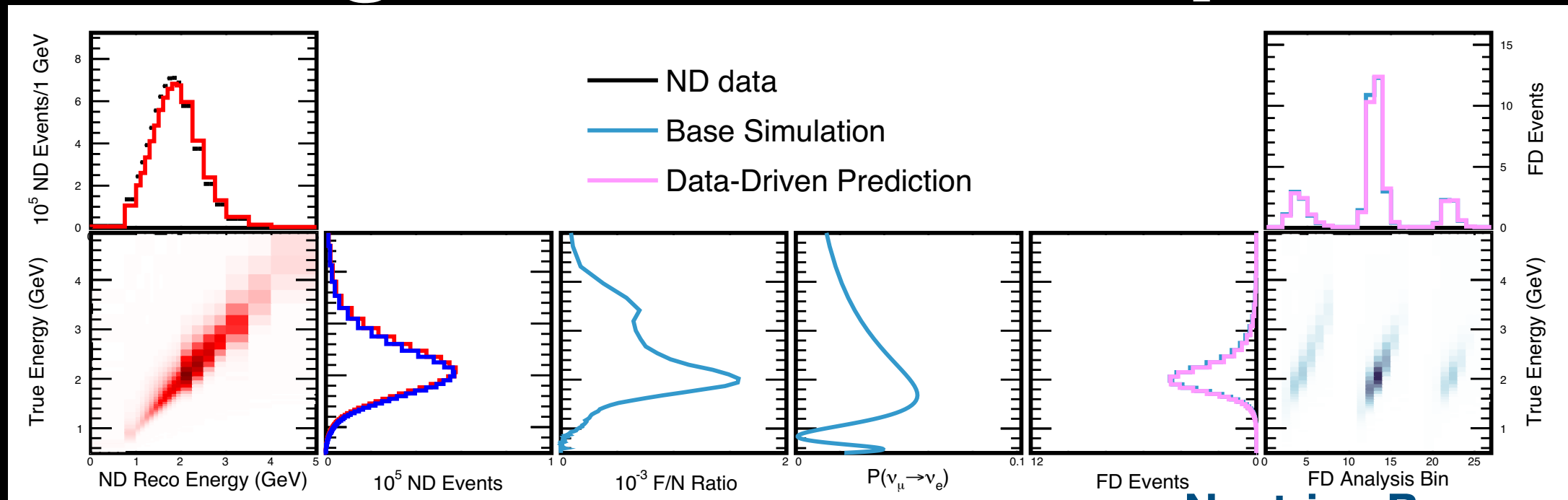
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# Event selection for $\nu_e + \bar{\nu}_e$

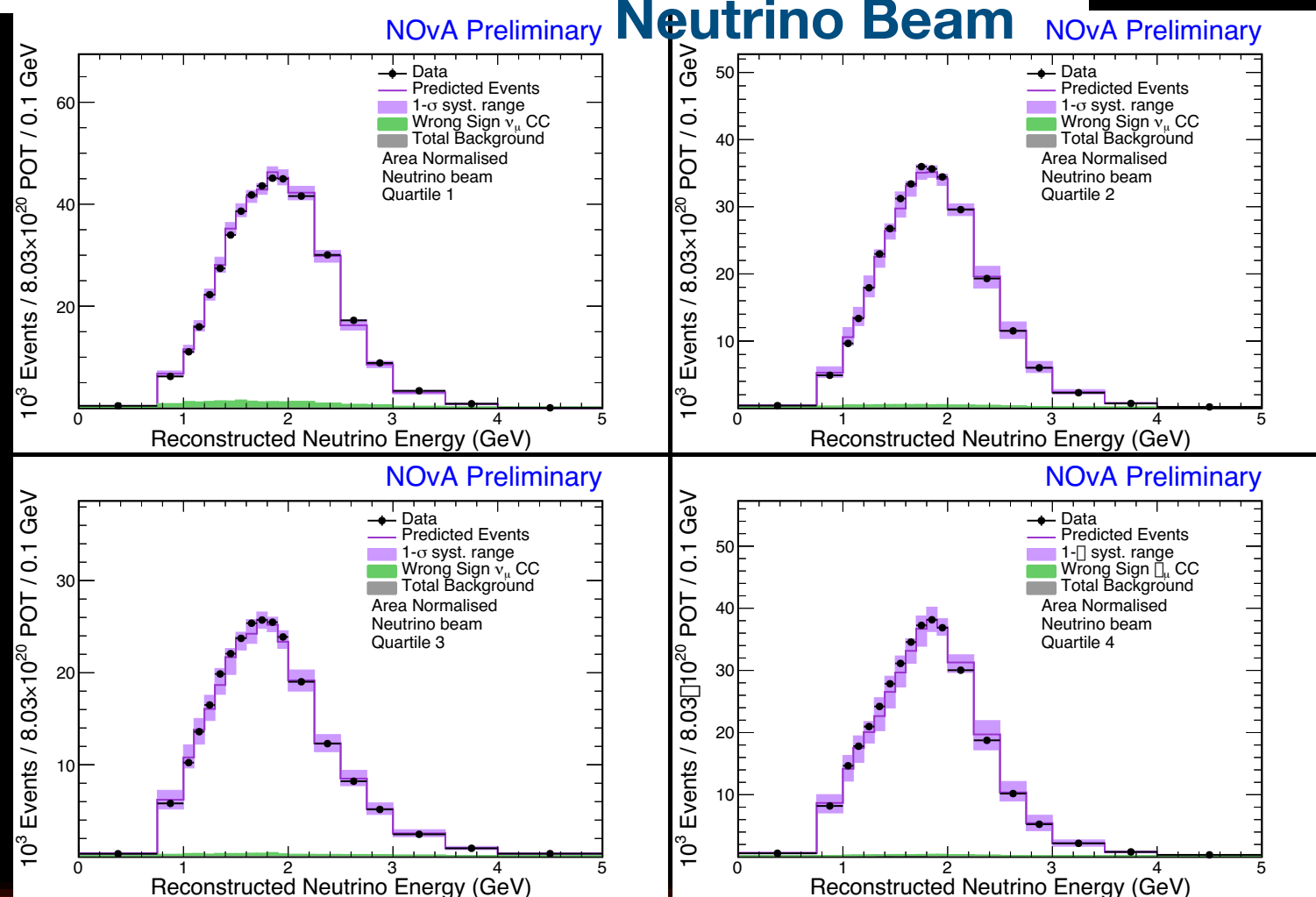


- Details of reconstruction detailed in talk by M. Groh
  - Select events w/  $\nu_e$  ( $\bar{\nu}_e$ ) classifier score & reconstructed information indicating likely  $\nu_e$  ( $\bar{\nu}_e$ ) candidate
  - NOvA uses CVN (Convolutional Visual Network): event classification based on final state topologies
  - Selections broken into two subcategories based on CVN electron score
  - Far detector: extra peripheral sample w/ larger cosmic background but recovers appearance signal
- Energy range of interest 1-4 GeV in far detector (0-4.5 GeV in near detector, peripheral)

# Predicting far detector spectra



- Functionally similar near, far detectors: extrapolation reduces overall systematics (reduction factor  $\sim 2$ )
- Near detector  $\nu_\mu$  spectrum extrapolated to far detector given oscillation parameters forms appearance signal prediction
- Expect fewer signal events in antineutrino beam than in neutrino beam: (Flux $\otimes$ Cross-section) reduced for antineutrinos and somewhat less exposure in antineutrino beam



Neutrino Beam

NOvA Preliminary

NOvA Preliminary

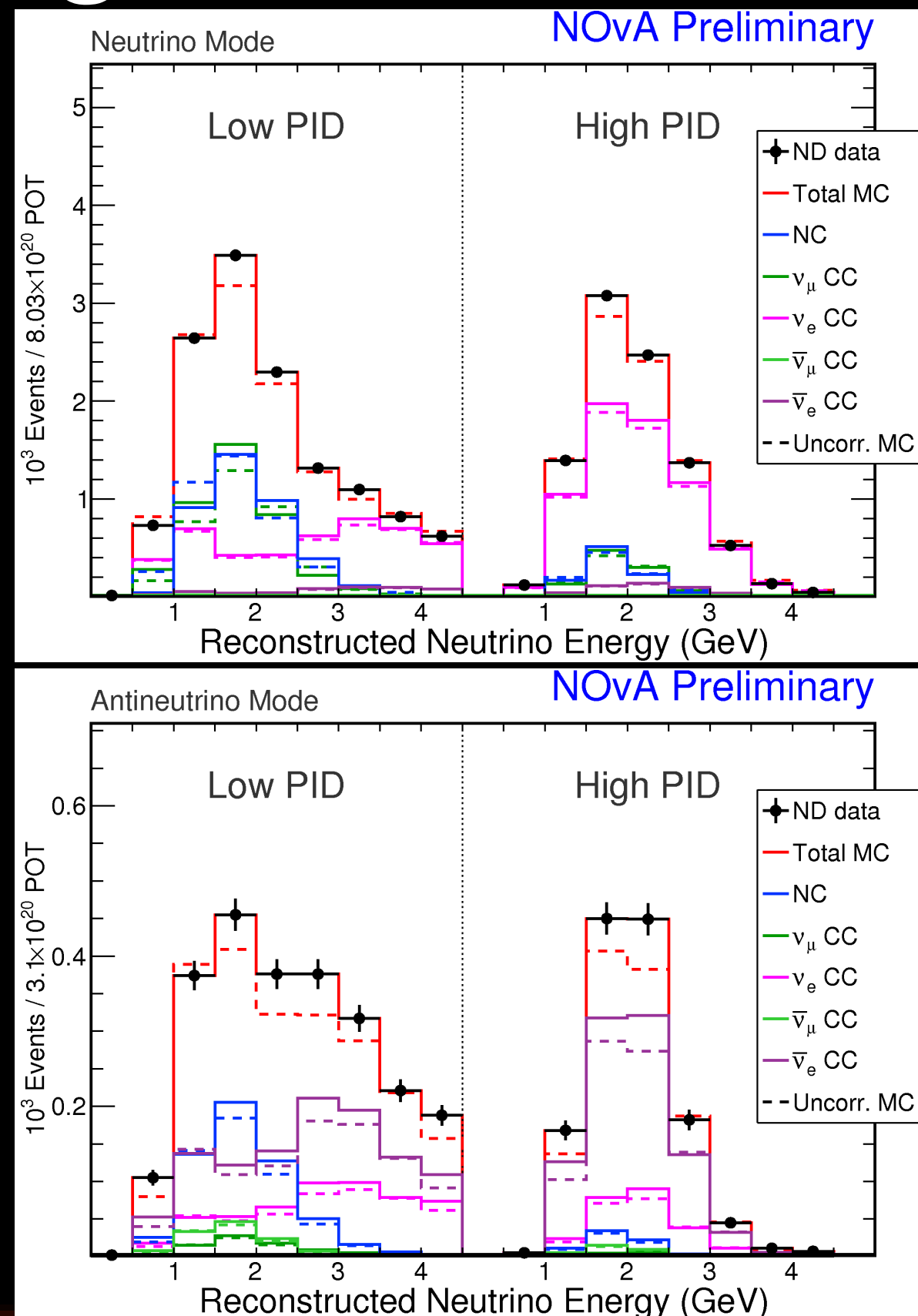
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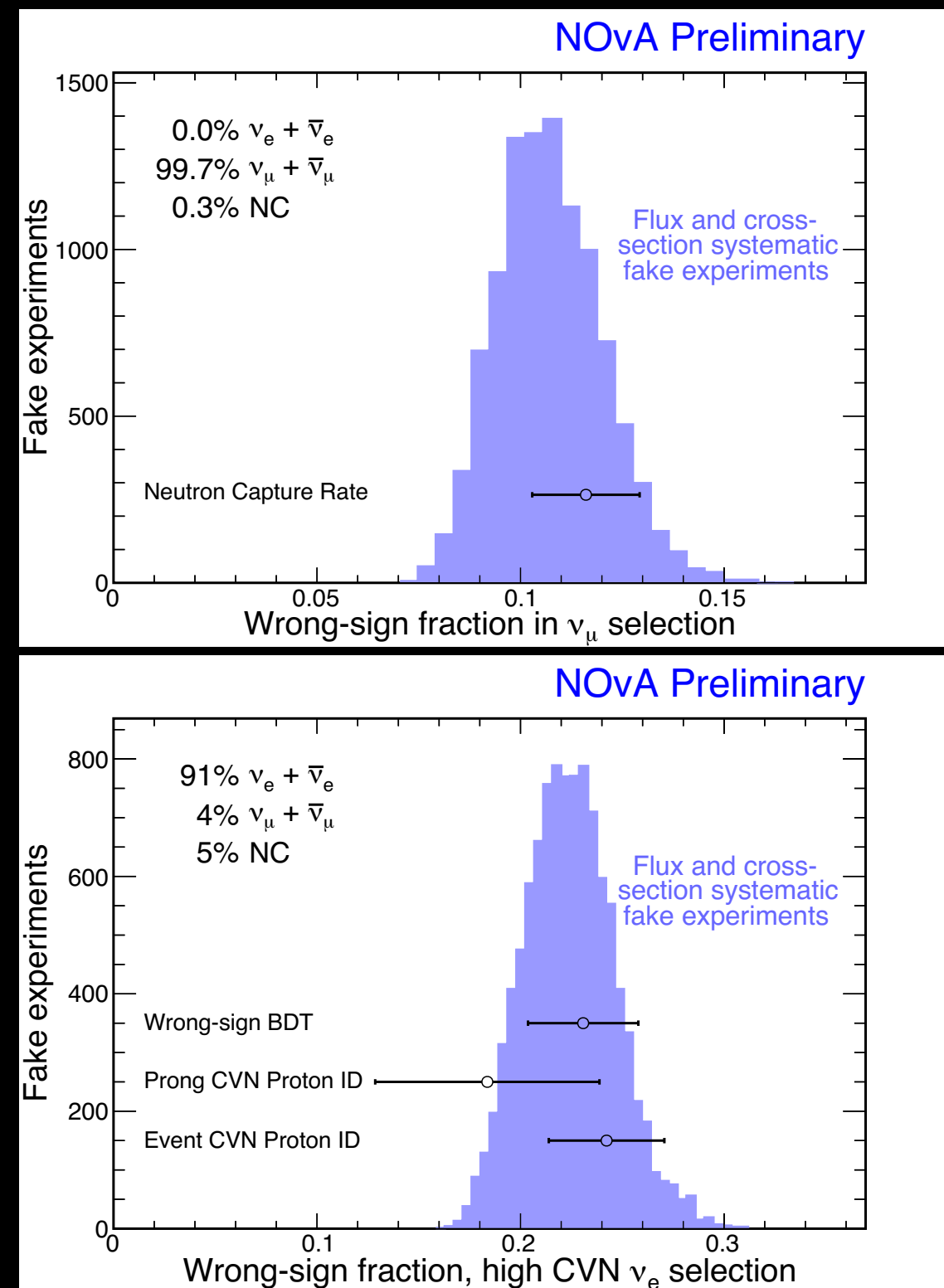
# Near Detector Background

- Selected near detector candidate spectra contain only backgrounds for appearance analysis: inform far detector background prediction via far/near ratio
- Components constrained via data-driven methods
  - In antineutrino beam, for now, data/MC differences are scaled proportionally in each energy bin to the components.
  - In neutrino beam, with higher stats, more refined method used which examines components separately.
- Decomposed ND spectra agree w/ MC by construction



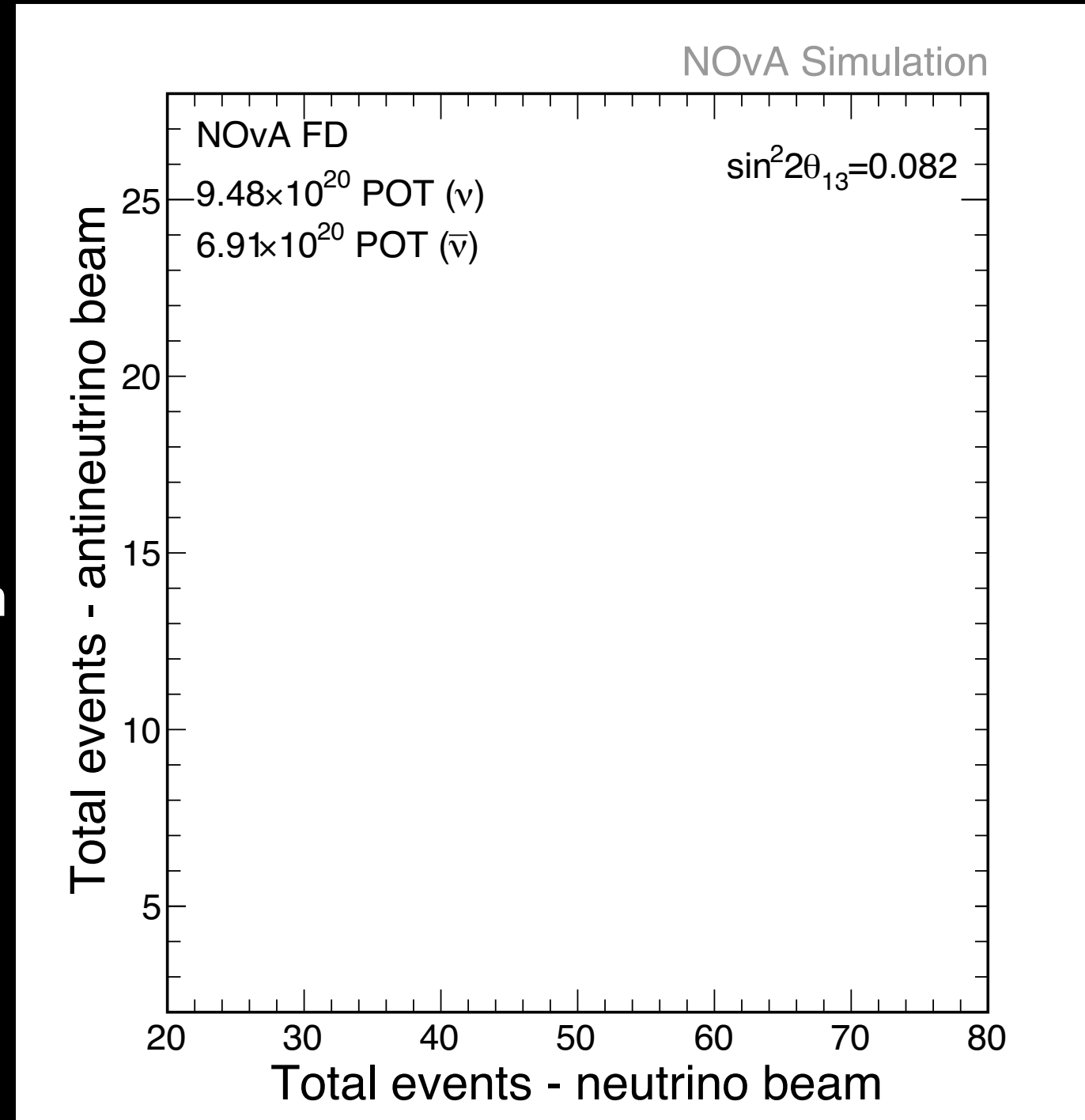
# Near Detector Background

- Wrong-sign ( $\nu$  in  $\bar{\nu}$  beam) fraction estimate in **near detector  $\bar{\nu}_\mu$  selection** 11%, checked w/ neutron capture rates
  - Oscillates to becomes appearance background**
- Wrong-sign fraction estimate in near detector **beam  $\bar{\nu}_e$  selected background** 22% for antineutrino beam in higher (purer in  $\bar{\nu}_e$ ) CVN sample. Check w/ identified protons & event kinematics



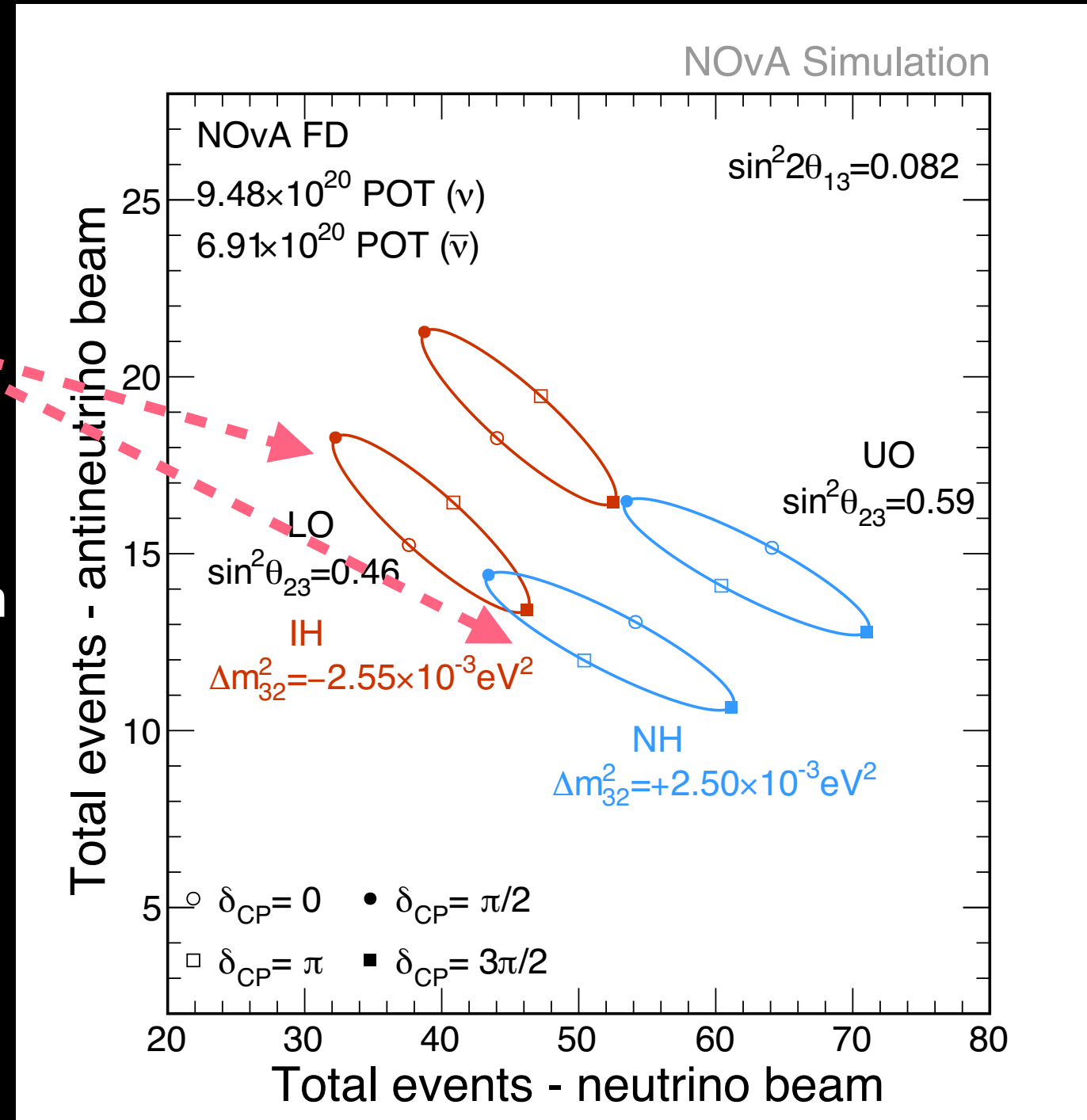
# Predicting far detector spectra

- External constraint on  $\sin^2 2\theta_{13}$  from PDG average
- Enhancement/suppression due to matter effect based on mass hierarchy. For normal hierarchy (NH), matter effect enhances  $\nu_e$ , suppresses  $\bar{\nu}_e$ 
  - Inverted Hierarchy (IH) opposite
- Oscillation parameters determine further  $\nu_e$ ,  $\bar{\nu}_e$  enhancement/suppression (overall and relative)
  - E.g. for  $\delta_{CP}=\pi/2$ , IH: suppression of  $\nu_e$  candidates relative to  $\bar{\nu}_e$
  - Alternatively, large suppression of  $\bar{\nu}_e$  relative to  $\nu_e$  for  $\delta_{CP}=3\pi/2$ , NH
- Predictions range from 10-22  $\bar{\nu}_e$  candidates, 30-75  $\nu_e$  candidates



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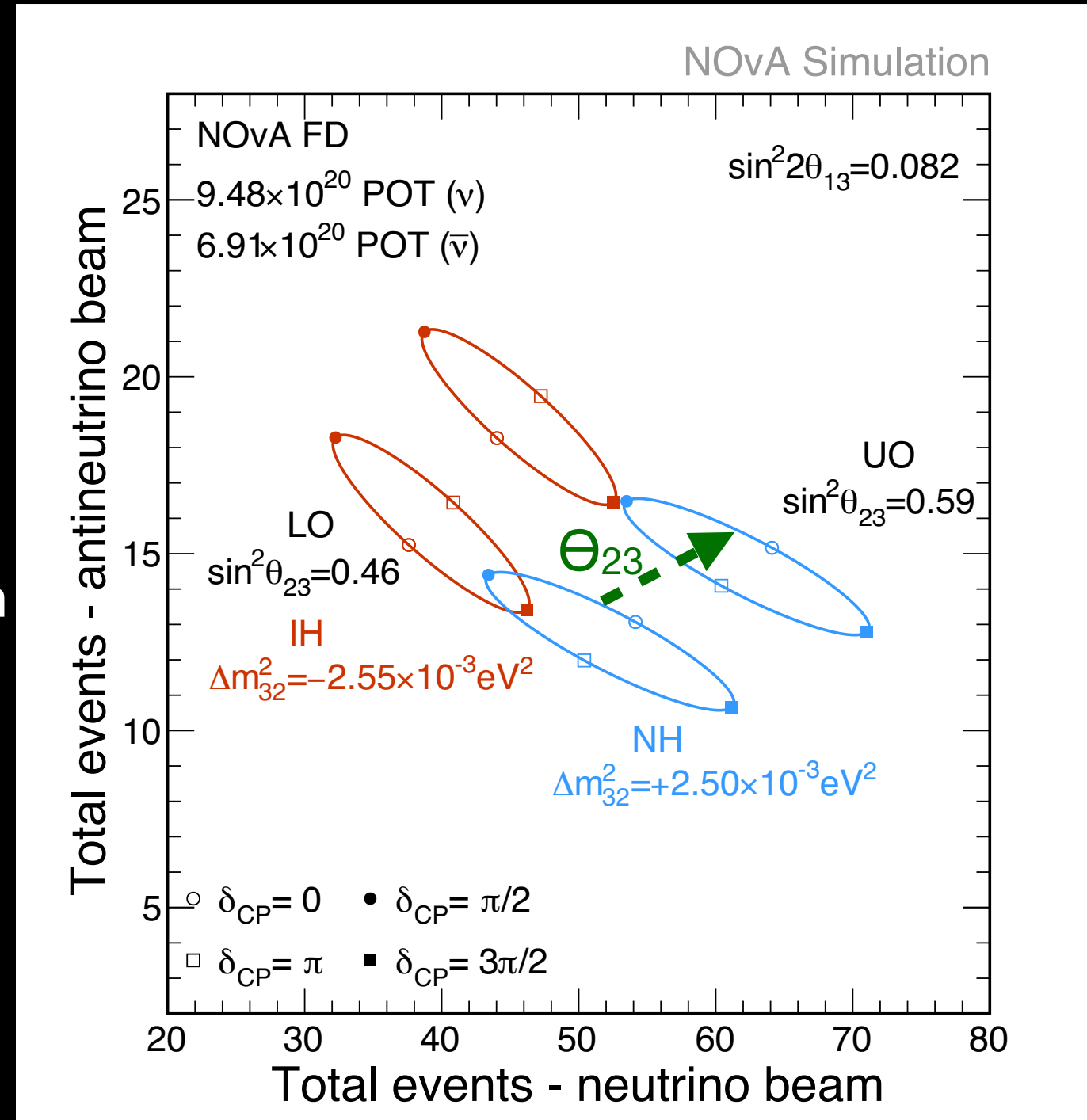
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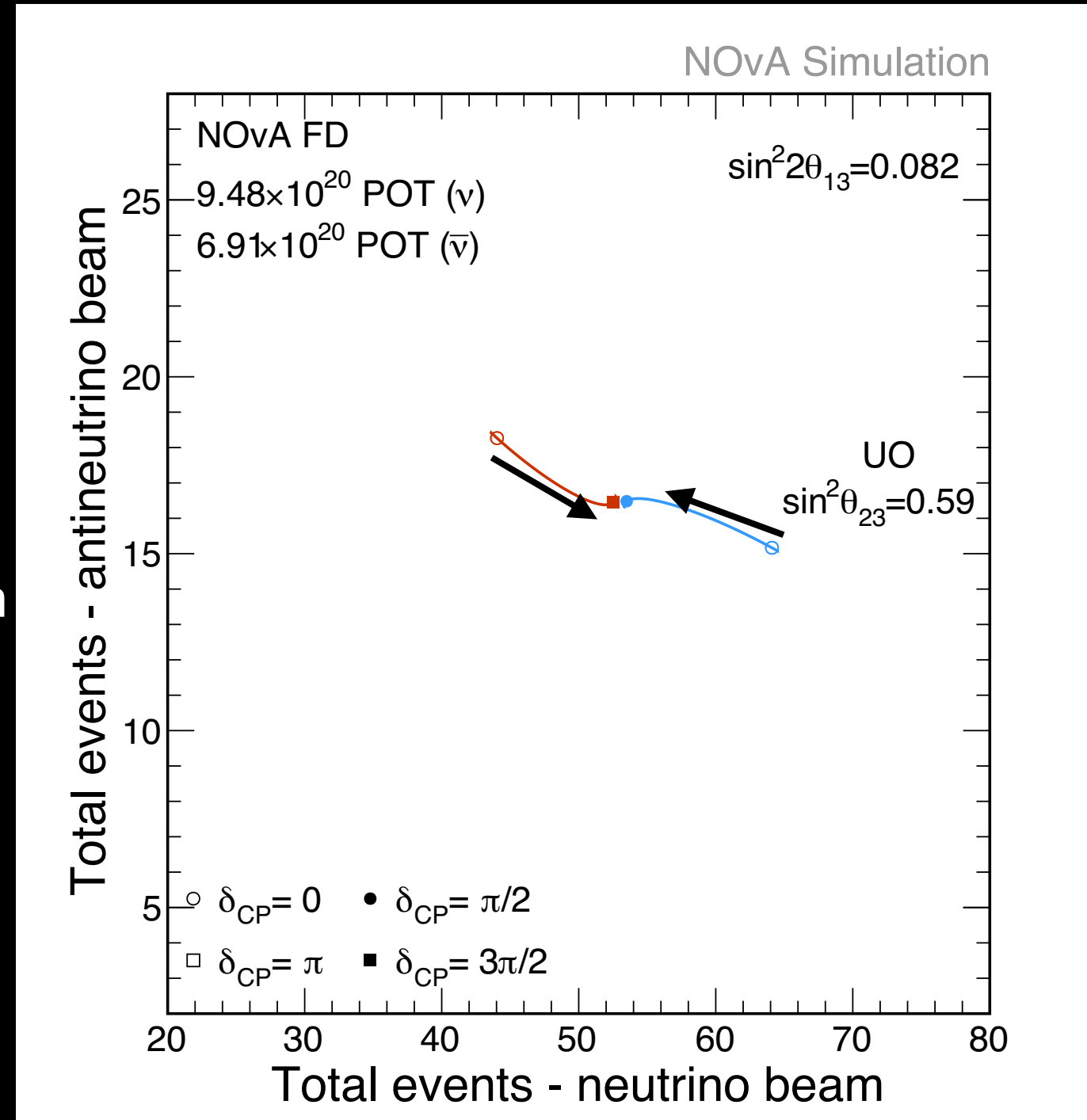
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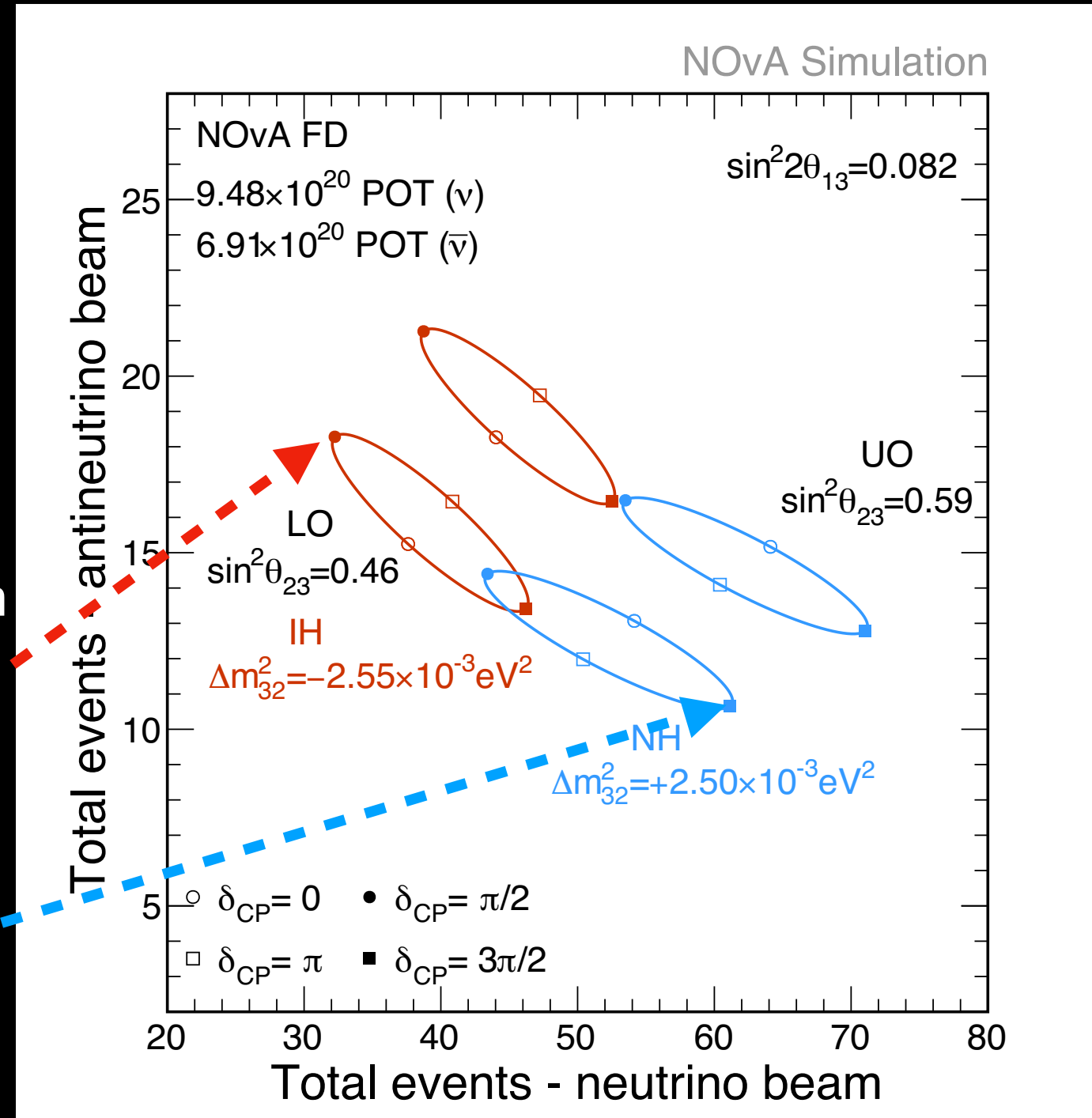
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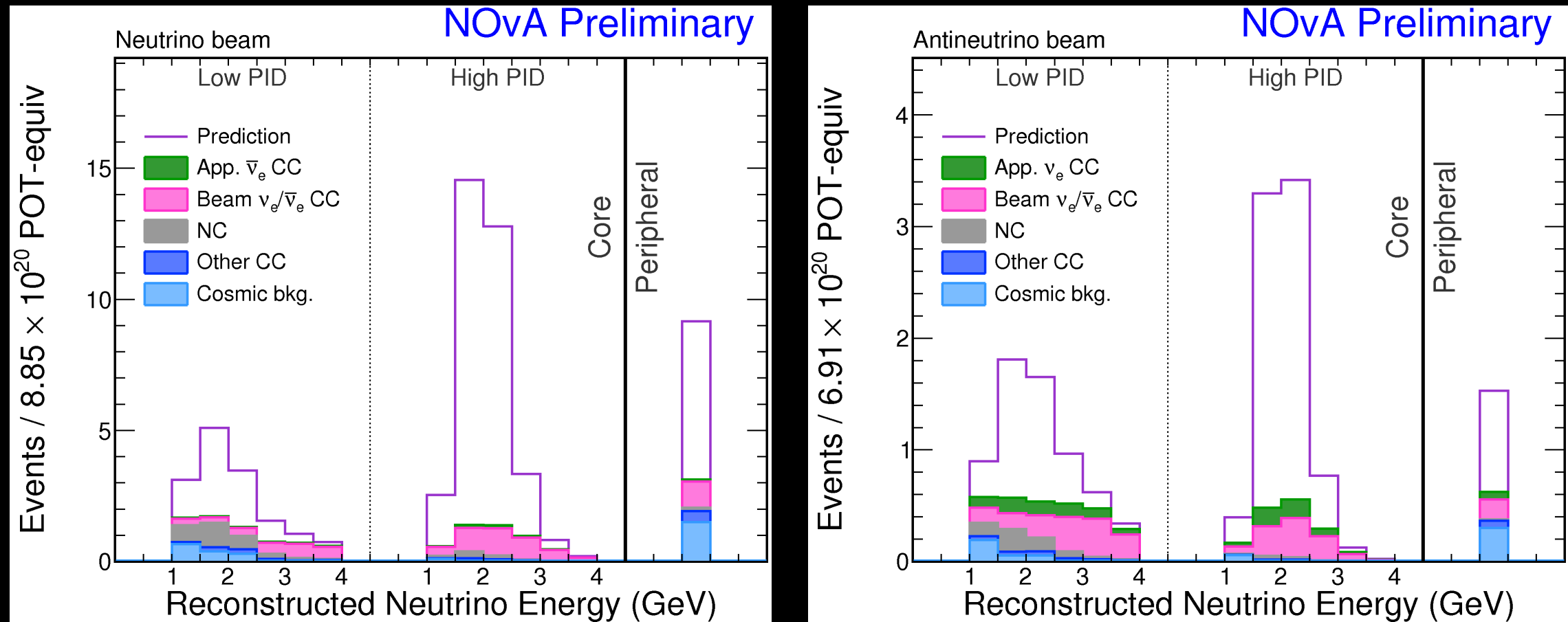


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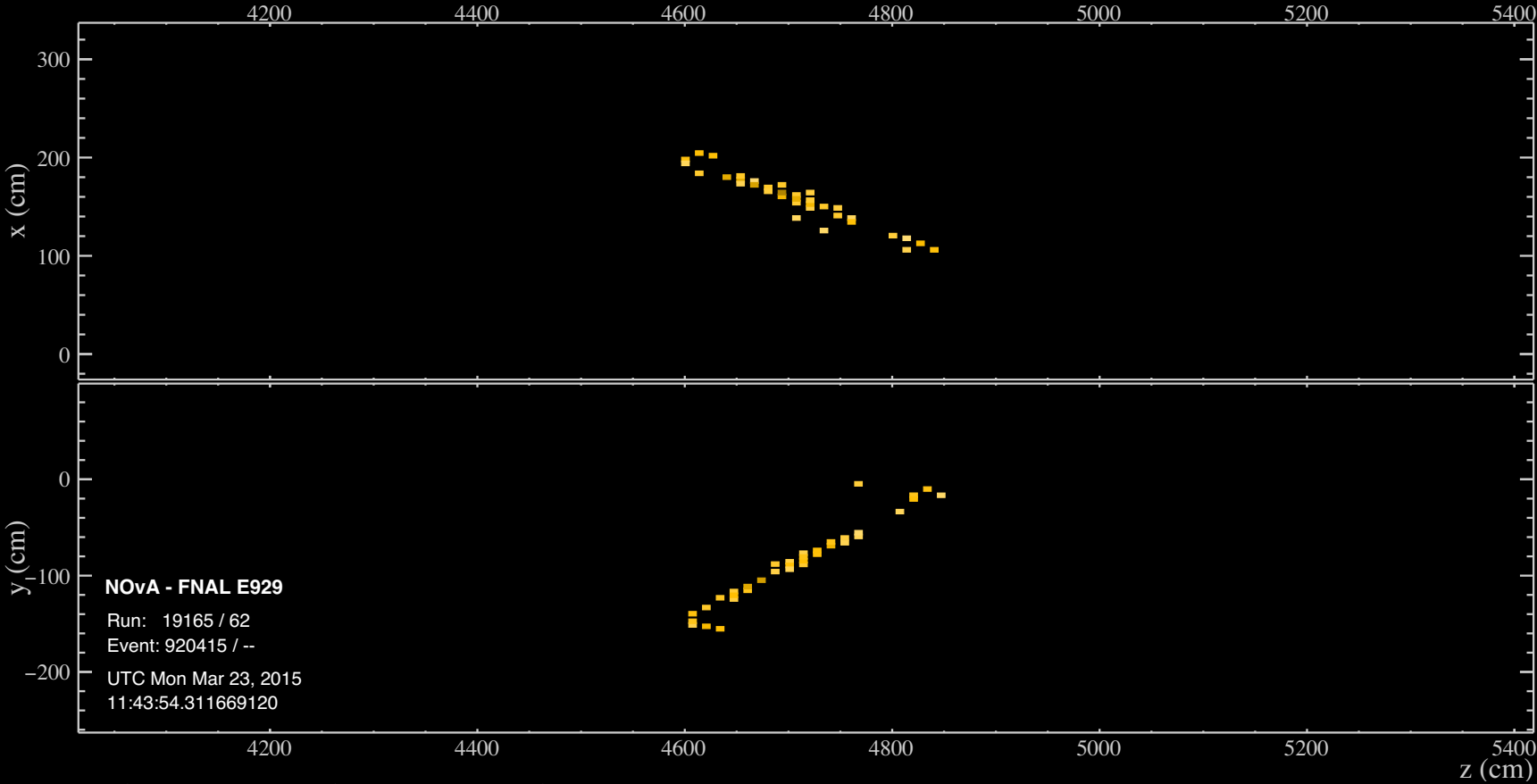


- An example for resulting prediction of candidates at far detector after constraining backgrounds with near detector and extrapolating oscillation signal prediction at a set of oscillation parameters
- Overall expectation varies based on oscillation parameters by changing the appearance signal and oscillated wrong-sign background
- Measured best-fit oscillation parameters will be the ones which give best agreement of overall expectation and the data (for *both* the  $\bar{\nu}_e$  appearance and  $\bar{\nu}_\mu$  disappearance)

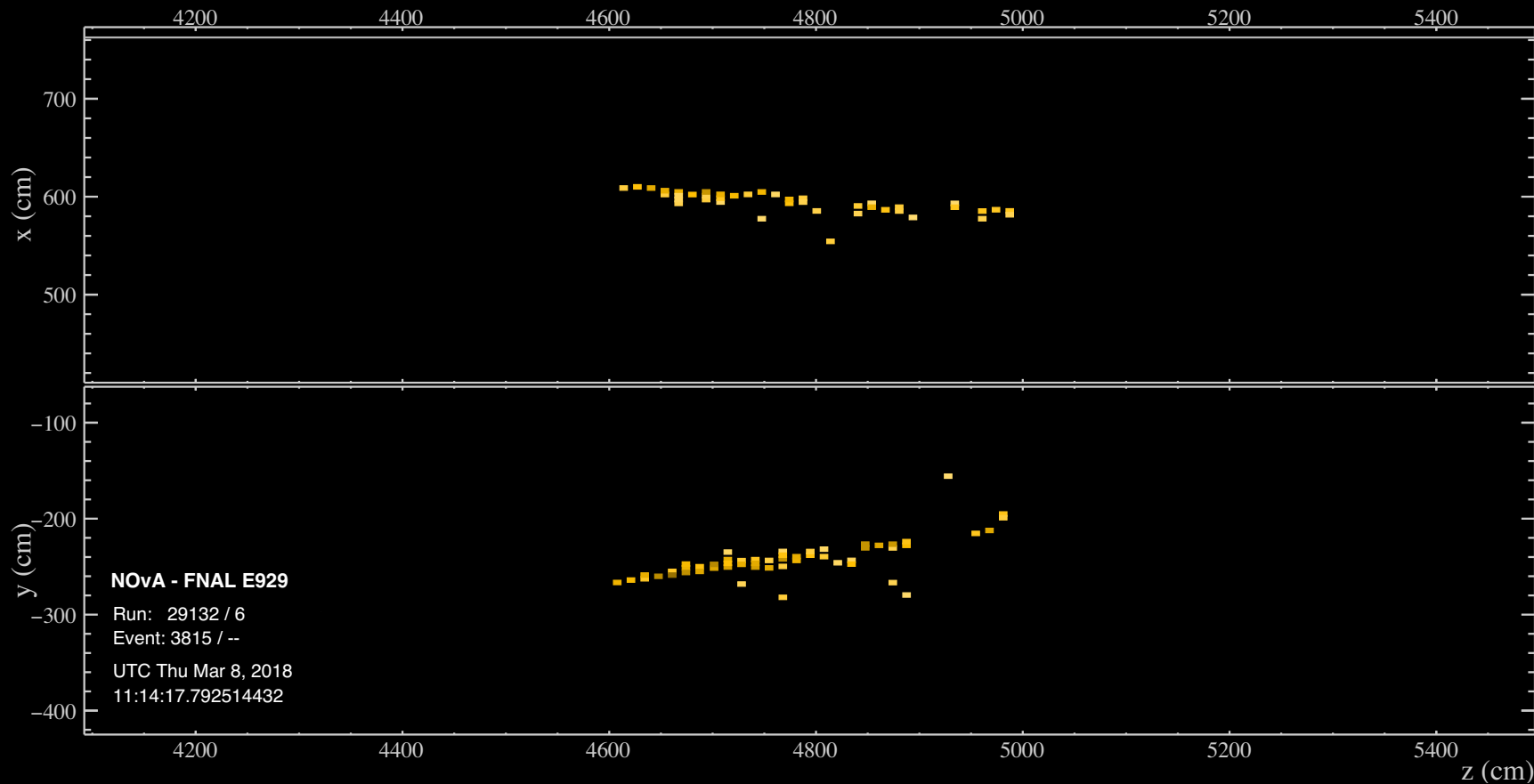


# Opening the box

*Far detector candidate,  
neutrino beam*

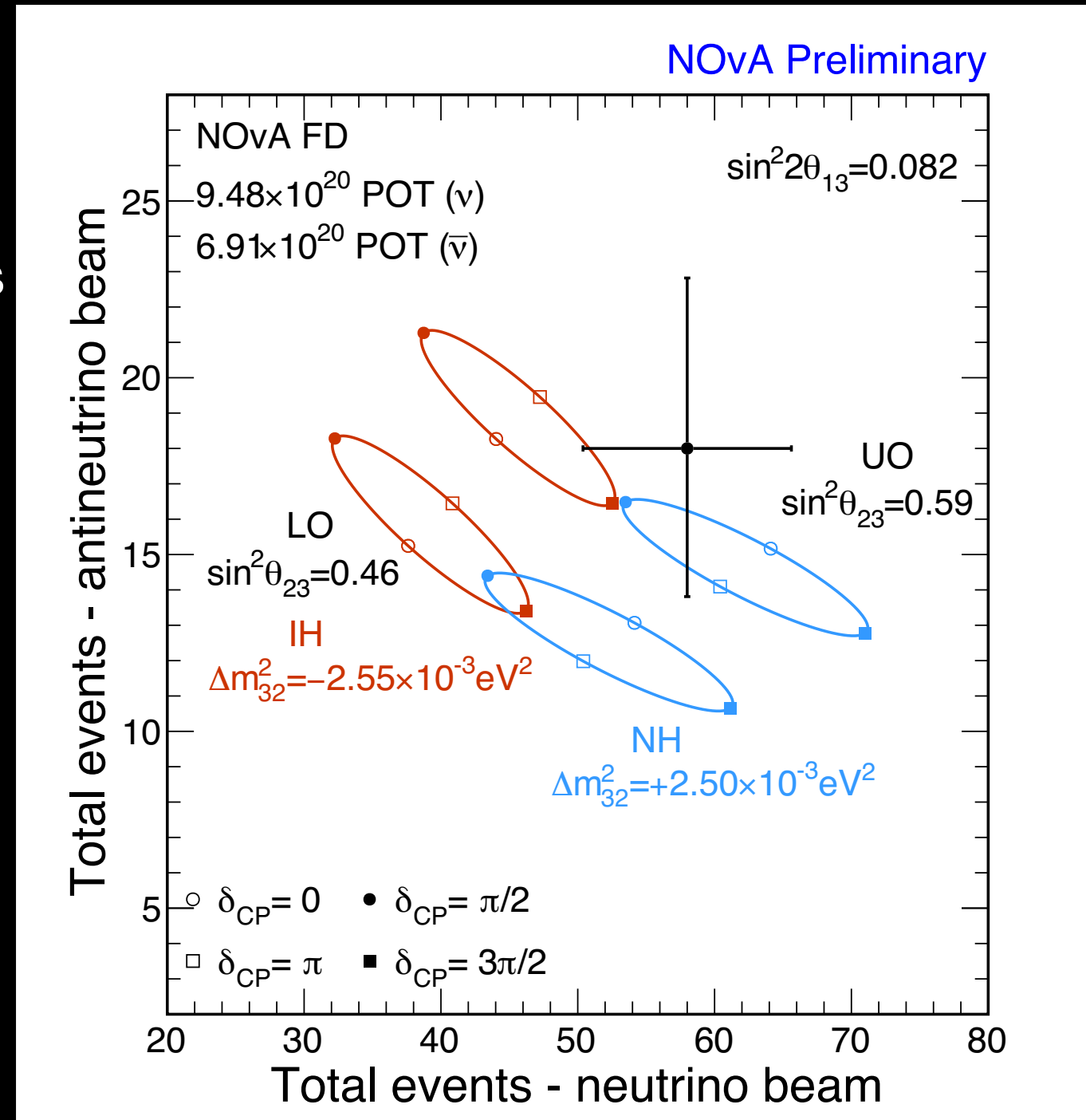


*Far detector candidate,  
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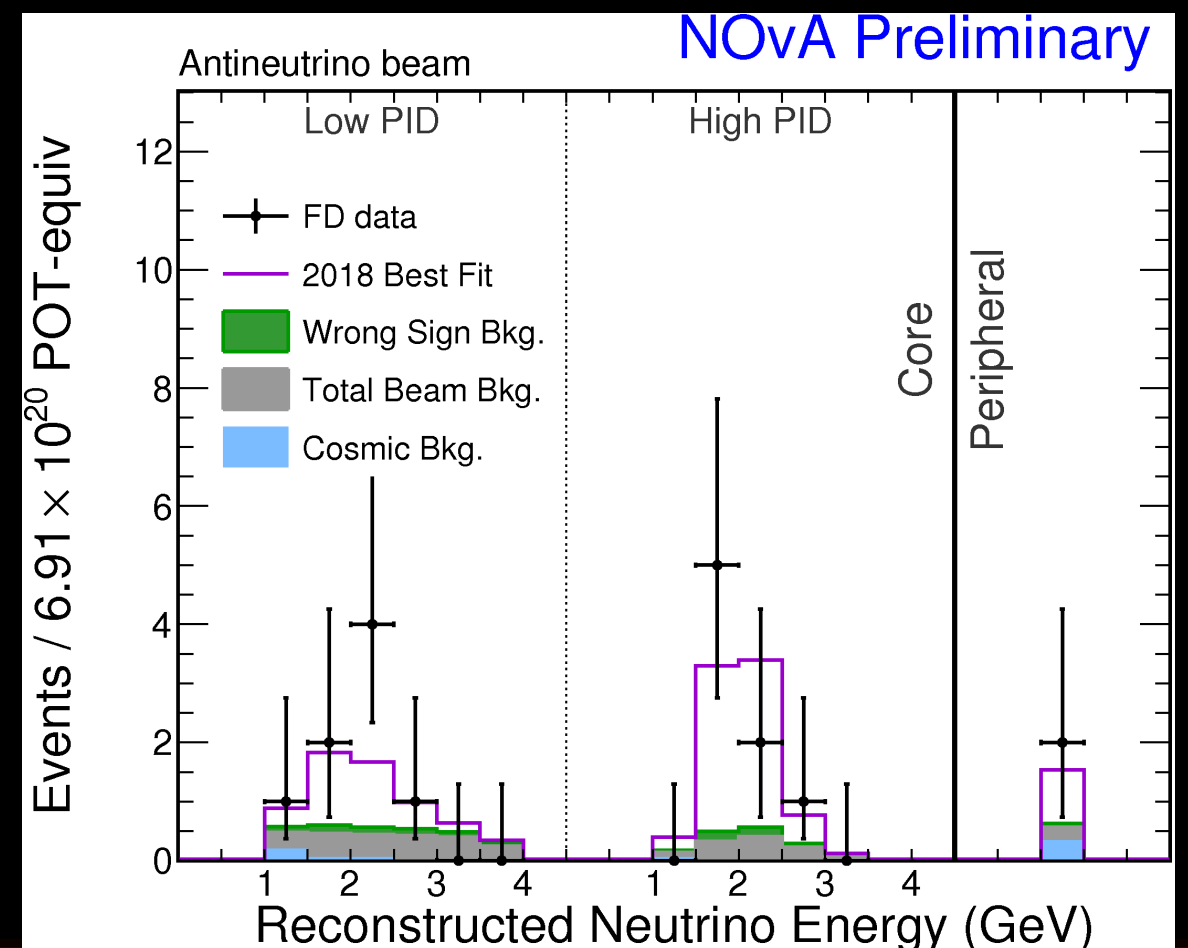
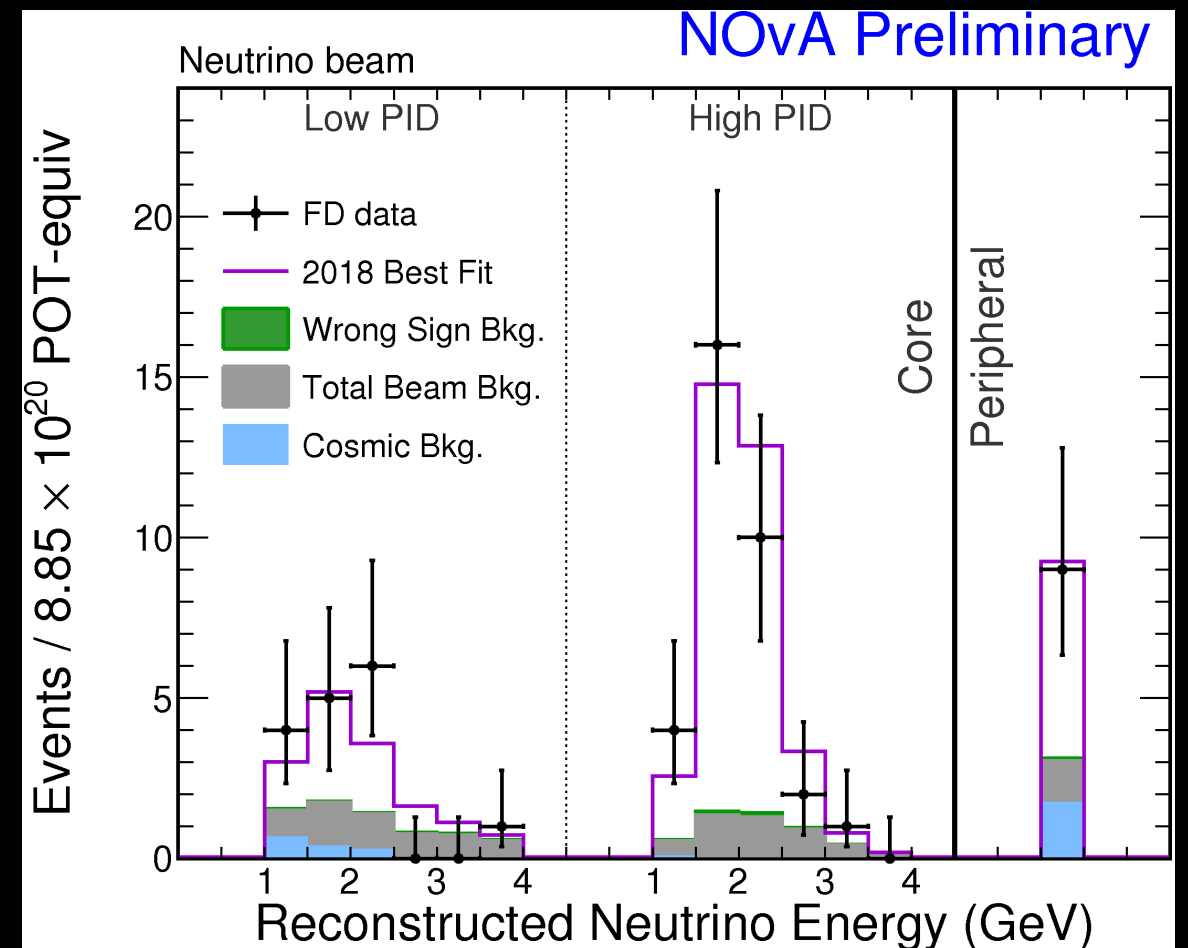
# $\nu_e, \bar{\nu}_e$ Results

- In neutrino beam, select 58  $\nu_e$  candidates
  - Background prediction: 15.1 events
  - ~12 from beam (< 1 wrong-sign oscillated) and ~3 cosmic origin
- In antineutrino beam, select 18  $\bar{\nu}_e$  candidates
  - Background prediction: 5.3 events
  - Mostly of beam origin (~1 wrong-sign oscillated), <1 cosmic
- **>4 $\sigma$  evidence for  $\bar{\nu}_e$  appearance!**



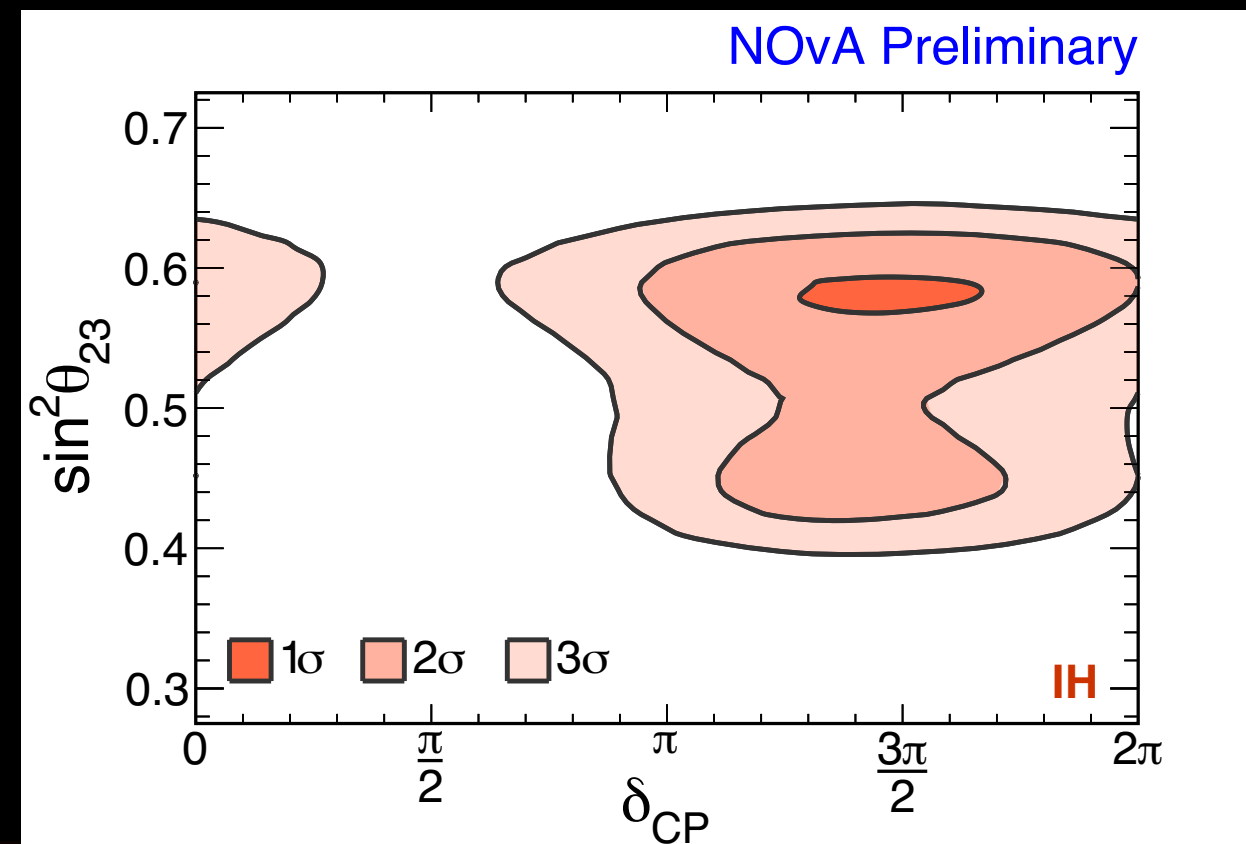
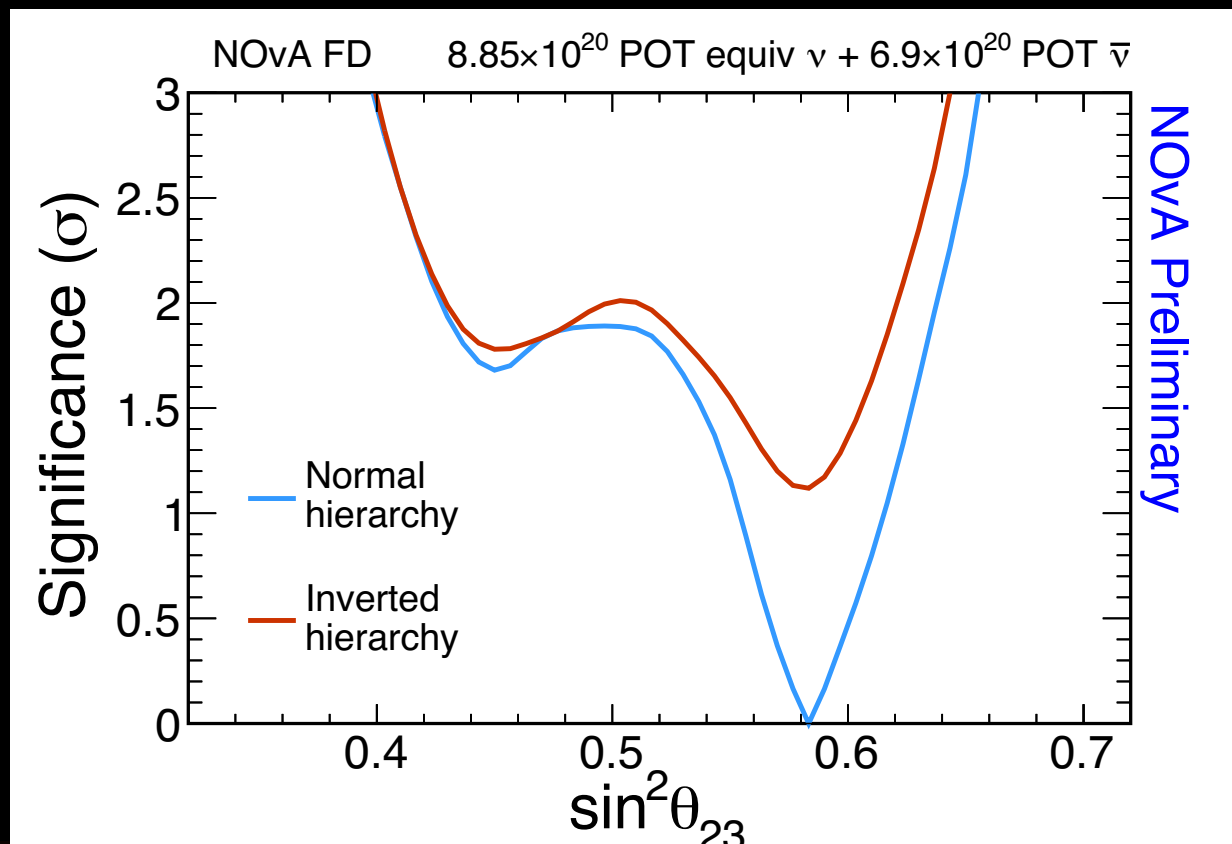
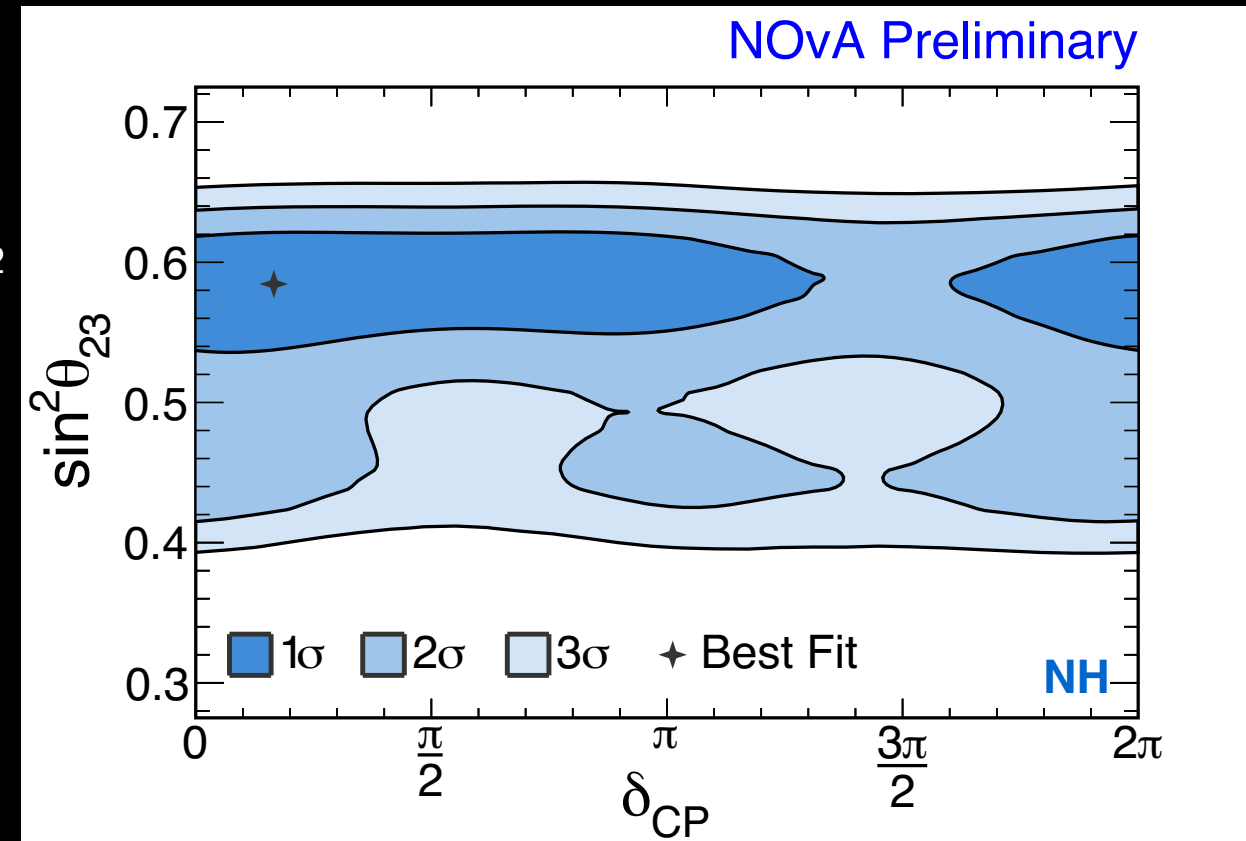
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  - Background prediction: 5.3 events
  - Mostly of beam origin (~1 wrong-sign oscillated), <1 cosmic
  - **>4 $\sigma$  evidence for  $\bar{\nu}_e$  appearance!**
- Best fit prediction is for 59.0  $\nu_e$  candidates and 15.9  $\bar{\nu}_e$  candidates*



# Joint $\bar{\nu}_\mu, \bar{\nu}_e$ Results

- Joint analysis best fit prefers:
  - Normal hierarchy  $\sin^2\theta_{23} = 0.58 \pm 0.03$
  - $\delta_{CP} = 0.17\pi$   $\Delta m_{32}^2 = (2.51^{+0.12}_{-0.08}) \times 10^{-3} \text{ eV}^2$
  - Normal hierarchy preferred at  $1.8\sigma$**
  - Exclude  $\delta_{CP} = \pi/2$  in IH at  $> 3\sigma$**
  - Non-maximal mixing preferred at  $1.8\sigma$**
  - Upper octant preference at similar level**
- Full joint analysis requires statistical corrections via pseudo-experiment (Feldman-Cousins procedure)
  - The story and importance of this procedure is the subject of next talk (D. Doyle)

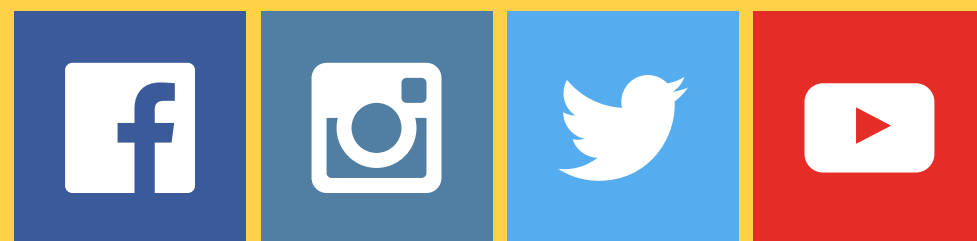




# Summary

- Updated analysis now incorporates antineutrino beam for the first time (nearly  $7 \times 10^{20}$  protons on target)
  - Includes previous data set of nearly  $9 \times 10^{20}$  (full-detector equivalent) protons on target with neutrino beam
- First strong evidence ( $>4\sigma$ ) for long-baseline  $\bar{\nu}_e$  appearance in antineutrino beam!
  - Achieved this in first  $\bar{\nu}$  analysis due to excellent beam performance
- Joint  $\bar{\nu}_\mu$  disappearance and  $\bar{\nu}_e$  appearance analysis has slight preference towards normal mass hierarchy and non-maximal  $\Theta_{23}$  [upper octant]
- Further data and analysis upgrades are expected to provide further sensitivity to key oscillation parameters in the coming years





<http://novaexperiment.fnal.gov>



# Backup

# Neutrino Oscillation

H. Nunokawa, S. Parke, J. W. F. Valle. "CP Violation and neutrino oscillations." *Prog.Part.Nucl.Phys.*, 60 (2008) 338-402

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & \sin^2\theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 \\
 & + \sin 2\theta_{23} \sin 2\theta_{13} \sin 2\theta_{12} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \frac{\sin(aL)}{(aL)} \Delta_{21} \cos(\Delta_{31} + \delta) \\
 & + \cos^2\theta_{23} \sin^2 2\theta_{12} \frac{\sin^2(aL)}{(aL)^2} \Delta_{21}^2
 \end{aligned}$$

*To first order, where*

$$\Delta_{ij} = \Delta m_{ij}^2 L / 4E$$

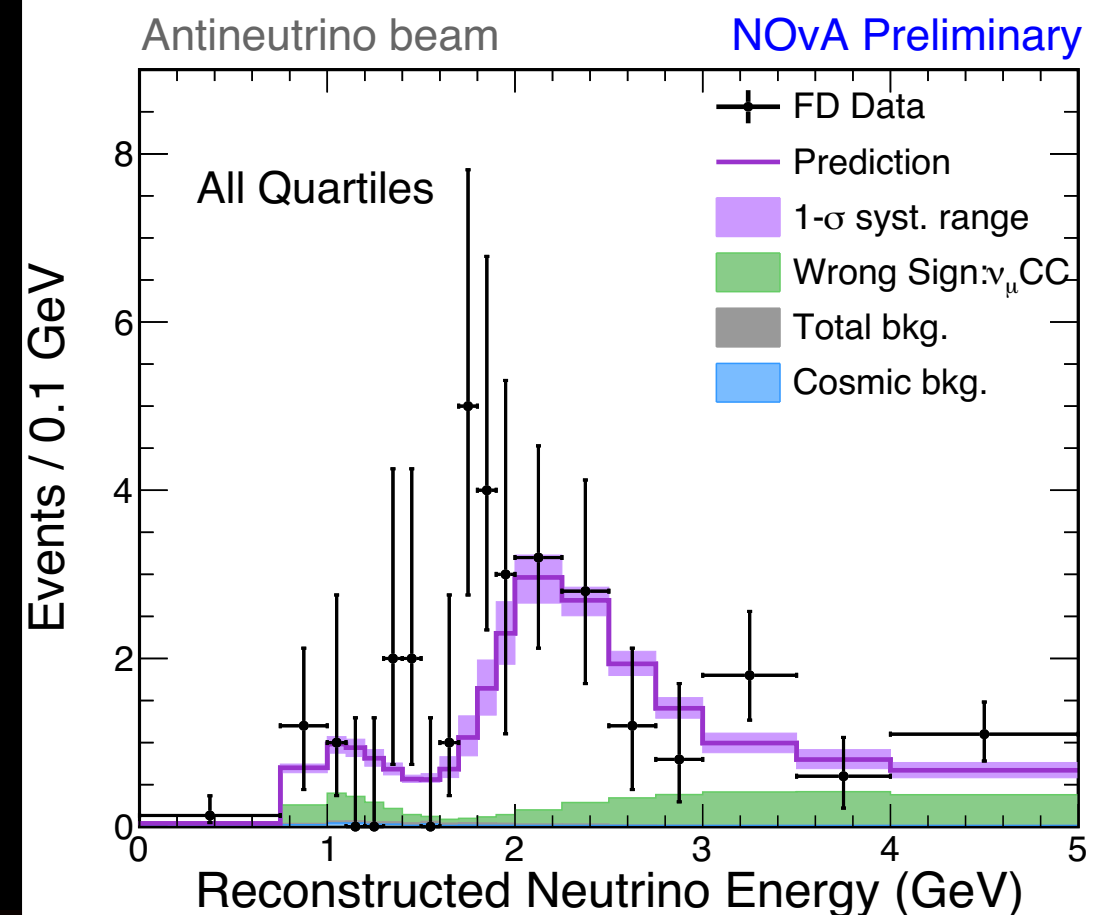
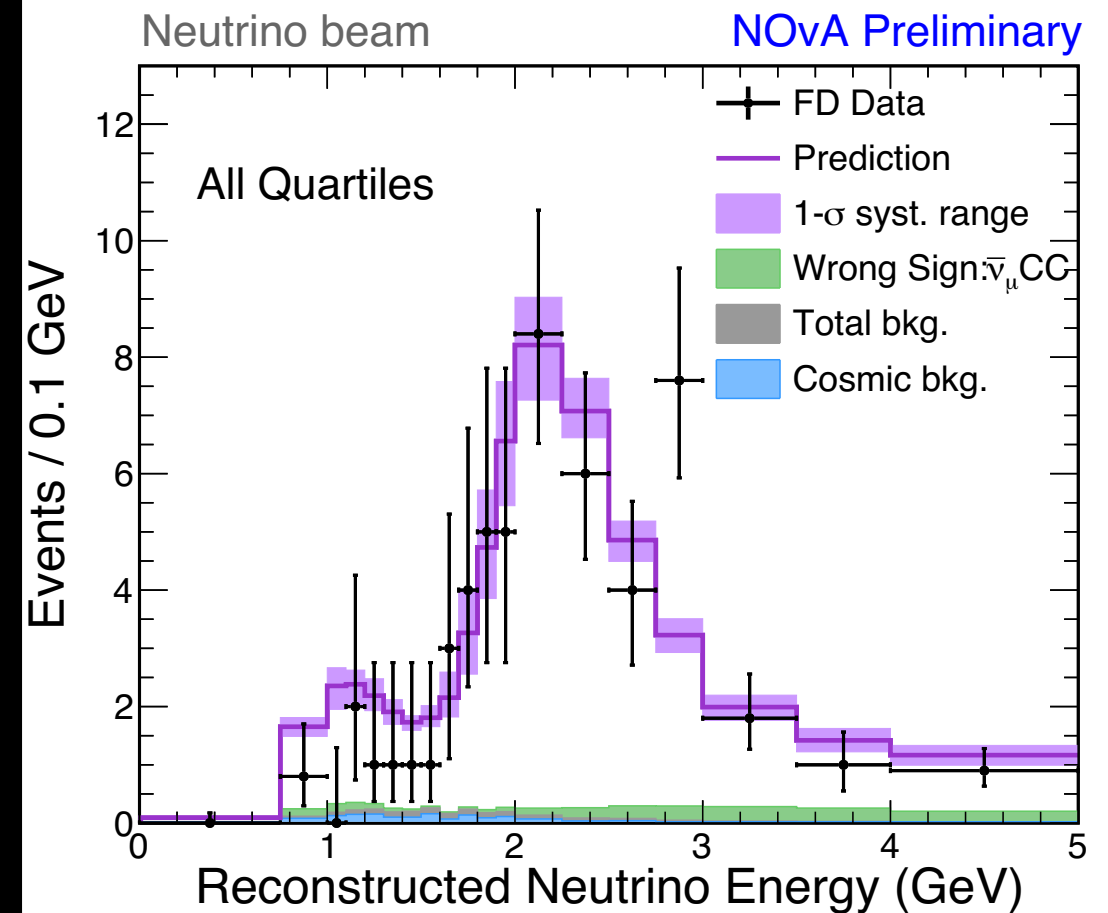
$$a = G_F N_e / \sqrt{2}$$

*And in the case of antineutrinos, the sign of the CP violating phase ( $\delta$ ) and sign of "a" flip.*



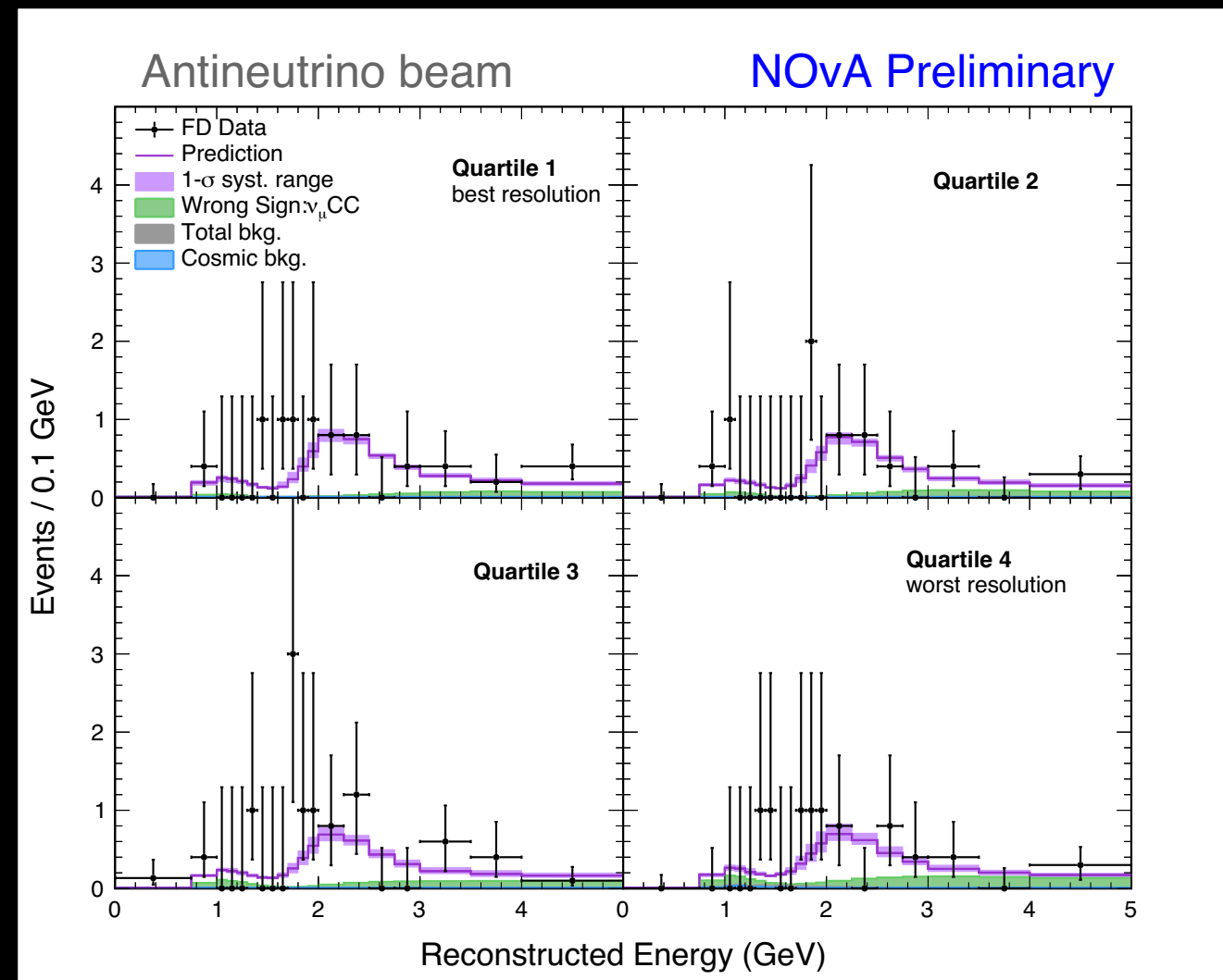
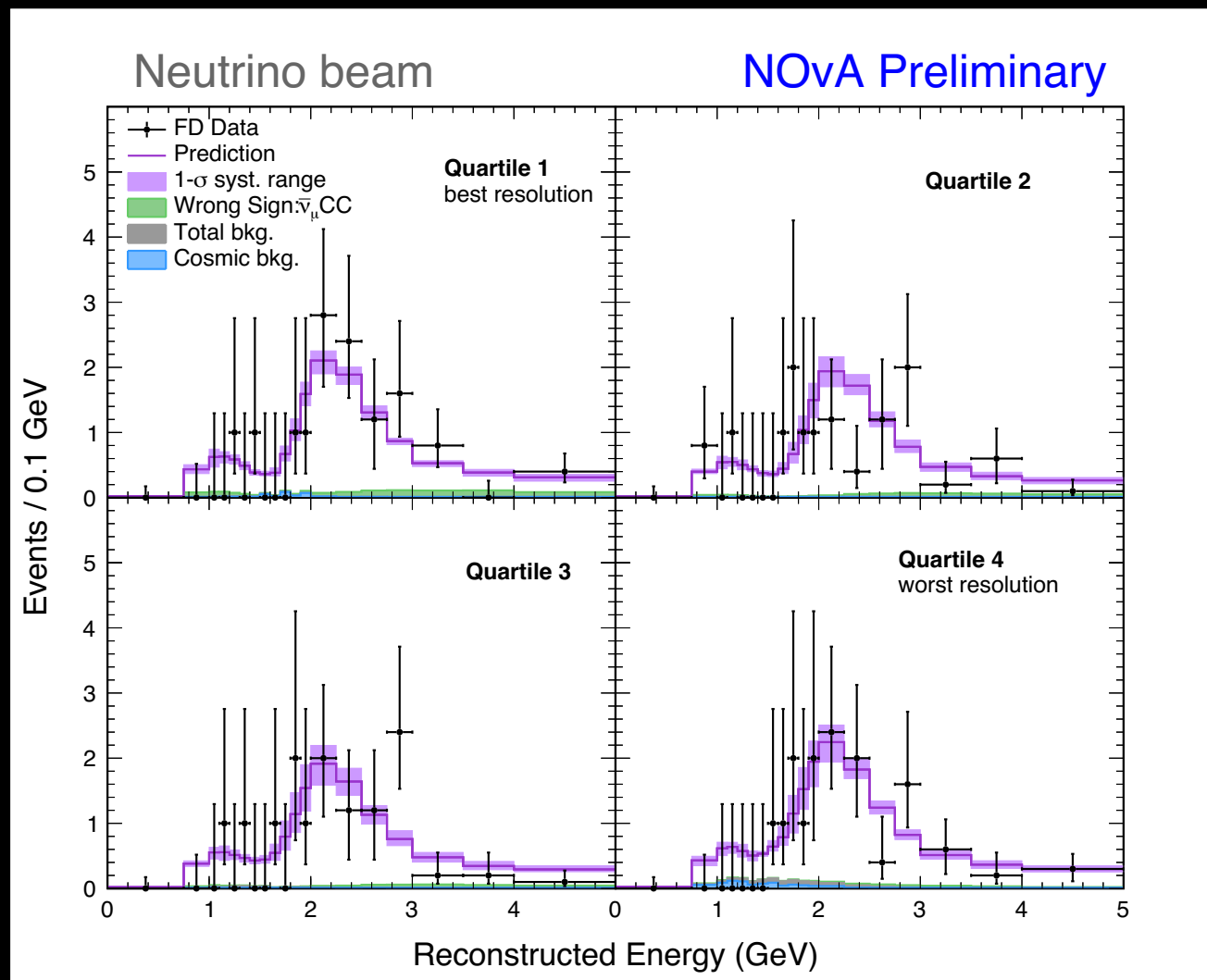
# Recall $\bar{\nu}_\mu$ Result

- Details in D. Torbunov talk
- Select 113  $\nu_\mu$  candidates in neutrino beam and 65  $\bar{\nu}_\mu$  candidates in antineutrino beam
  - Expecting  $730^{+38}_{-49}$   $\nu_\mu$  candidates in neutrino beam and  $266^{+12}_{-14}$   $\bar{\nu}_\mu$  candidates in antineutrino beam, in case of no oscillations
- Cosmic background prediction  $\sim 2$  events in neutrino beam and  $< 1$  event in antineutrino beam
- The results of the  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance and the  $\nu_e$  and  $\bar{\nu}_e$  appearance are then studied in a joint analysis and result in a joint oscillation fit in 3 neutrino paradigm

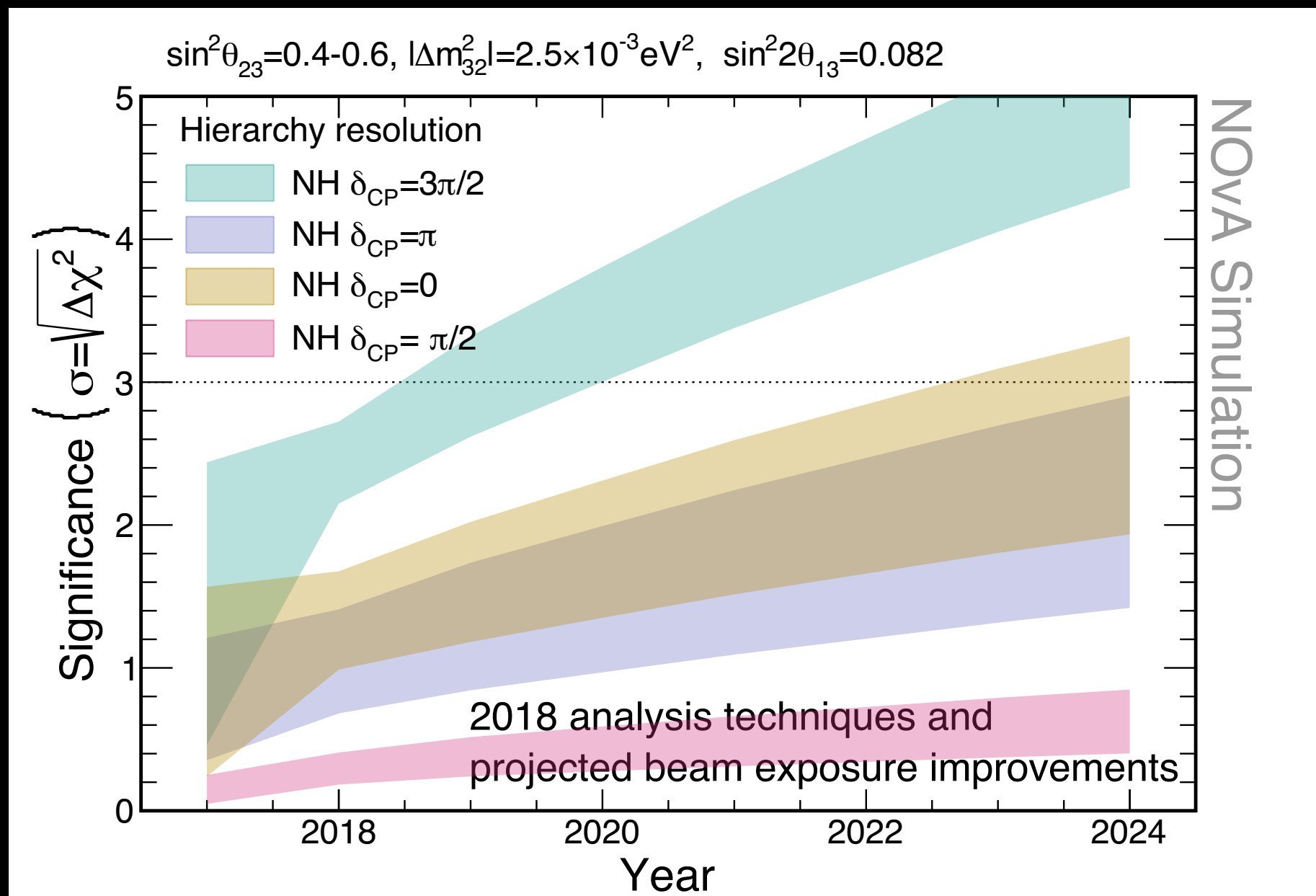


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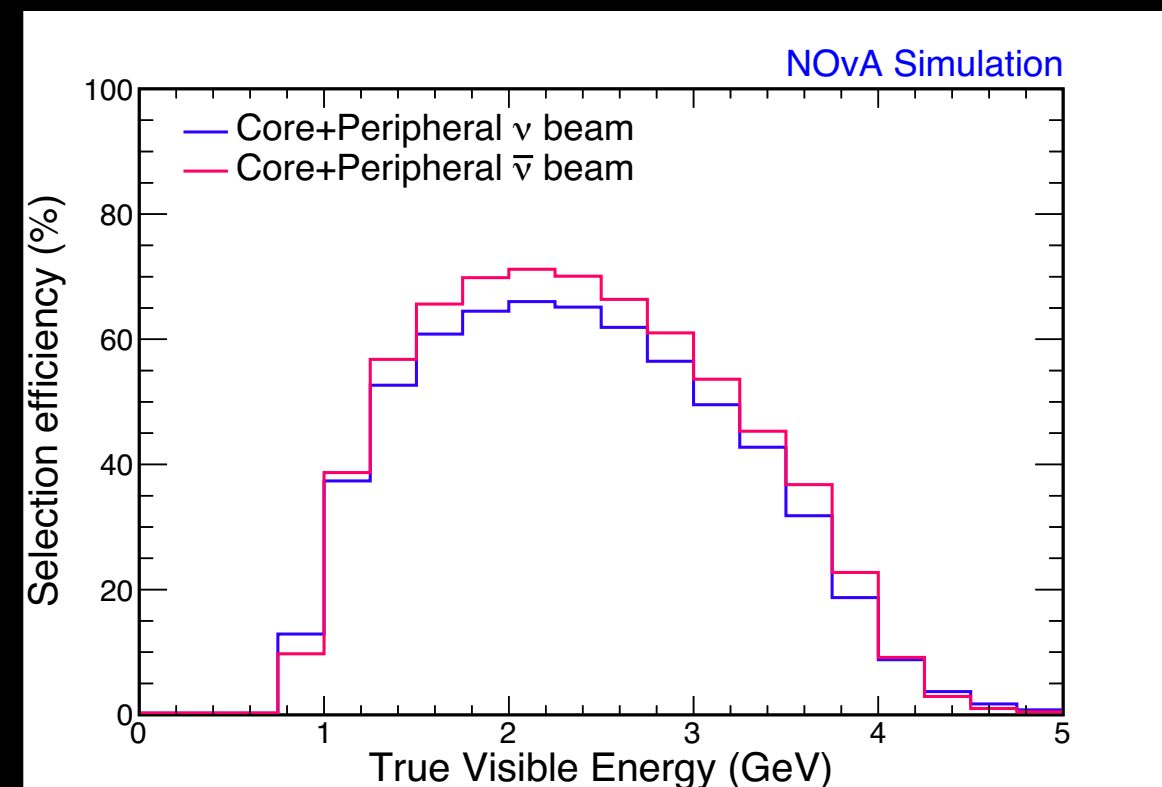
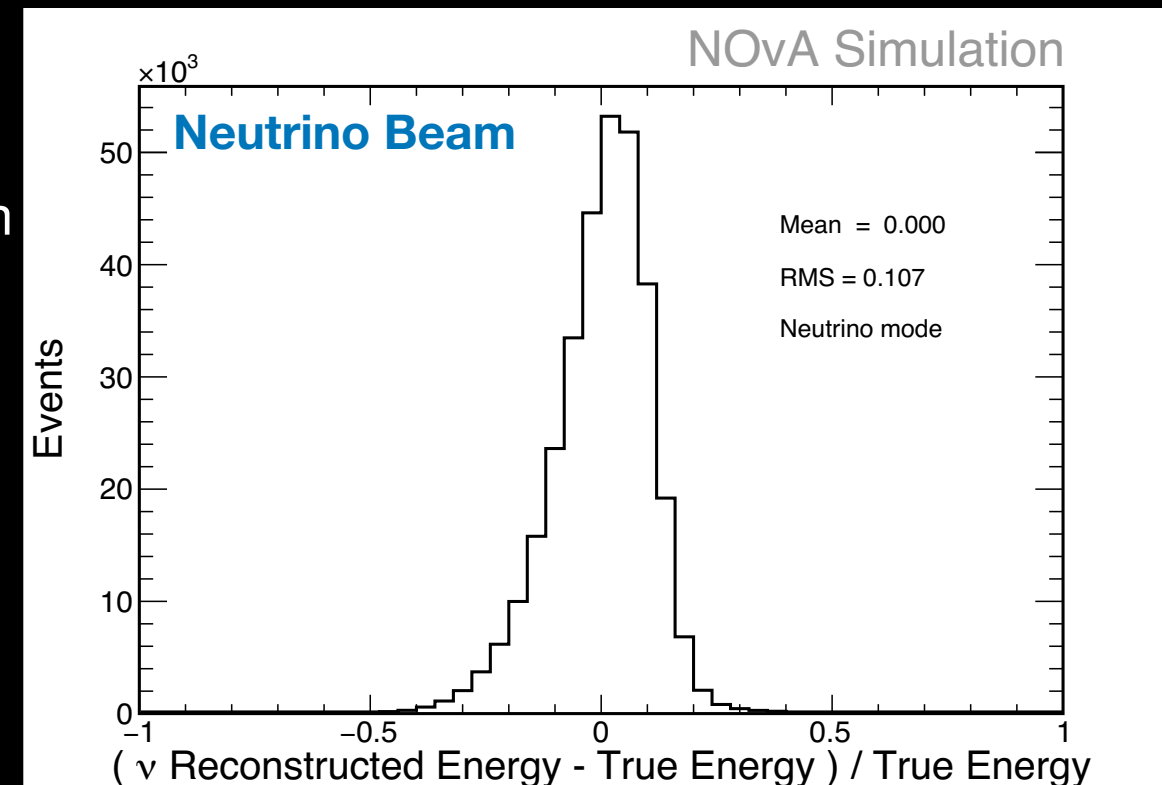
# Future sensitivity



**Mass hierarchy sensitivity  $3\sigma$  by 2020 for most favorable parameters. Can reach  $3\sigma$  by 2024 for wider range of parameters.**

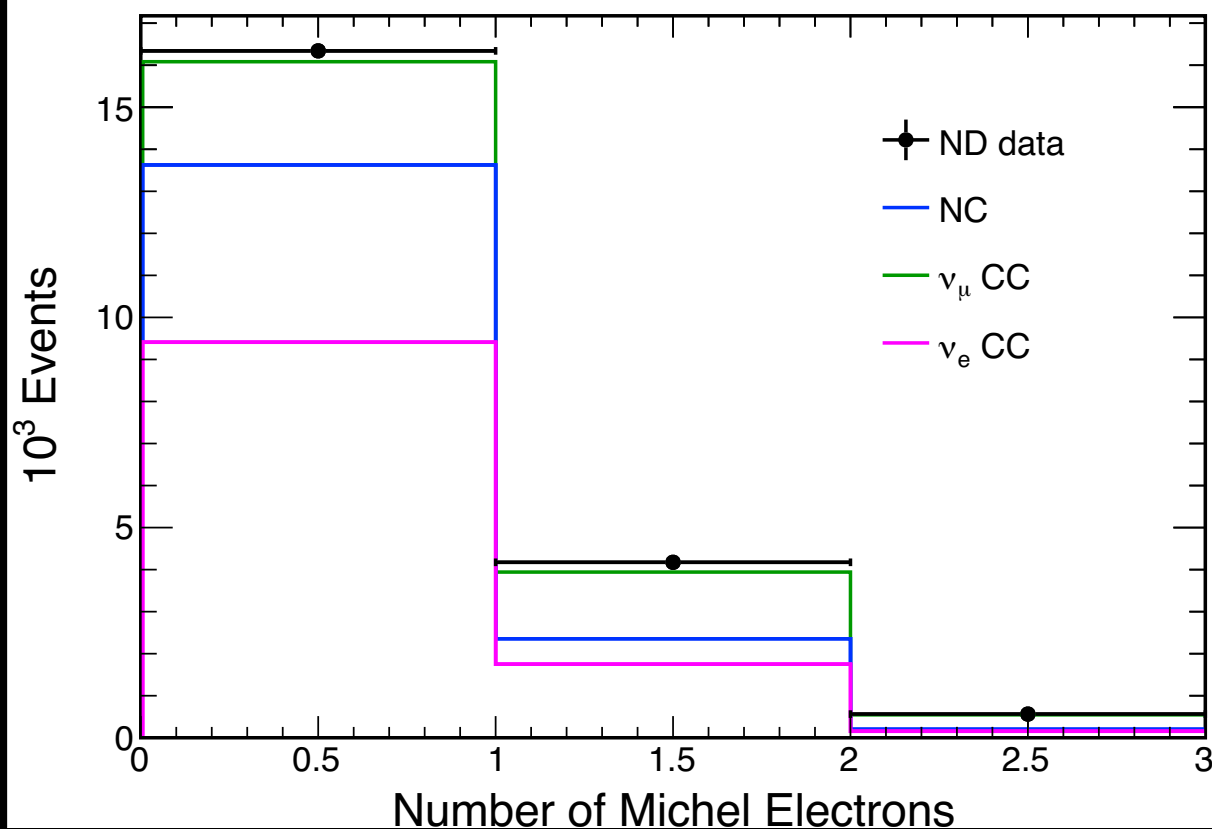
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- Select events w/  $\nu_e$  ( $\bar{\nu}_e$ ) classifier score & reconstructed information indicating likely  $\nu_e$  ( $\bar{\nu}_e$ ) candidate
  - Near detector and far detector selection broken into two subcategories based on CVN electron score.
  - Far detector has further peripheral sample with larger cosmic background but recovers appearance signal
- Energy estimator for  $\nu_e, \bar{\nu}_e$  characterizes energy deposits as coming from EM activity (or not) utilizing CVN on specific prongs. Estimator is 2nd-order polynomial in EM energy and non-EM energy
  - 11% E resolution in neutrino beam, 9% antineutrino
  - Energy range of interest 0-4.5 GeV in near detector, 1-4 GeV in far detector (0-4.5 GeV in peripheral)



# Near detector background

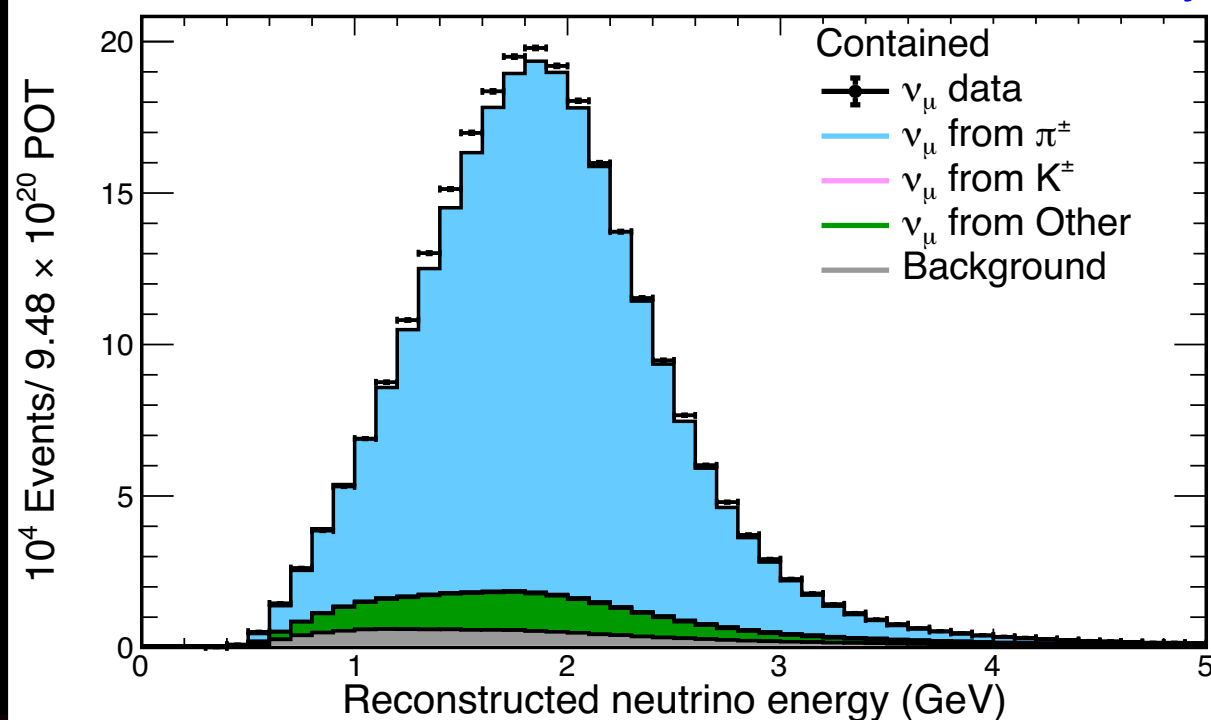
$\nu_\mu$  constraint



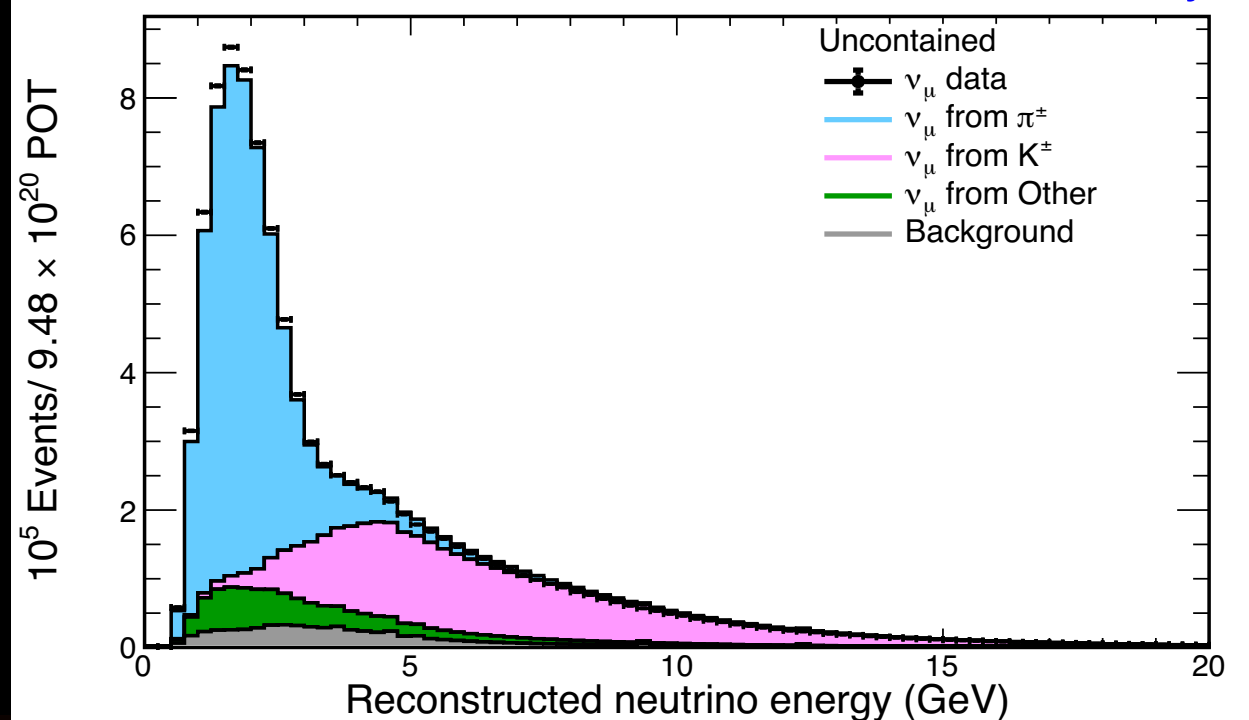
- Muon neutrino spectrum examined to constrain flux by parent hadron, which informs beam  $\nu_e$  component
- Michel electrons used to constrain  $\nu_\mu$  CC

$\nu_e$  constraint

NOvA Preliminary

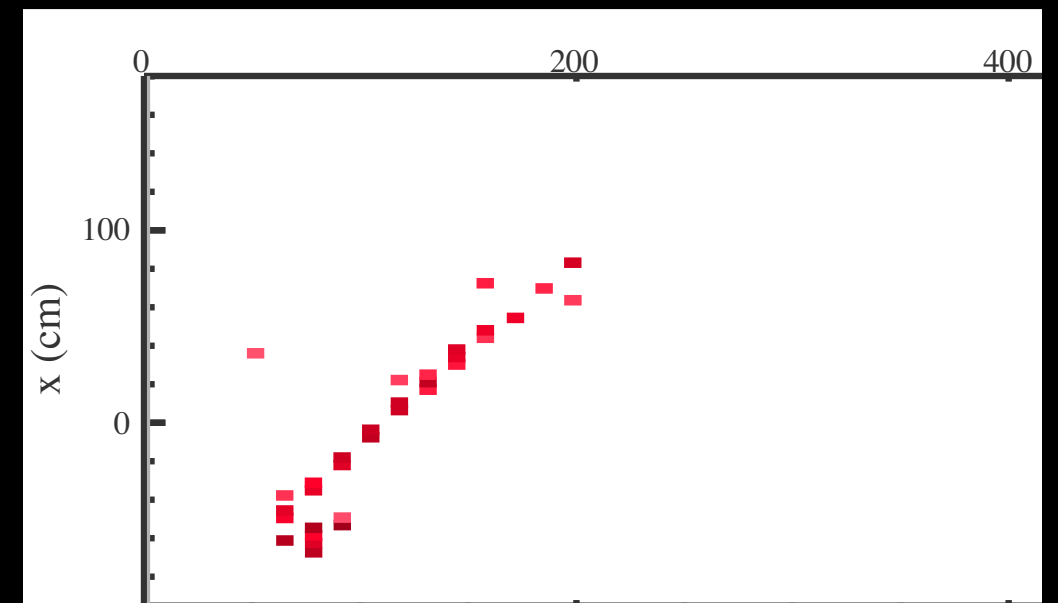
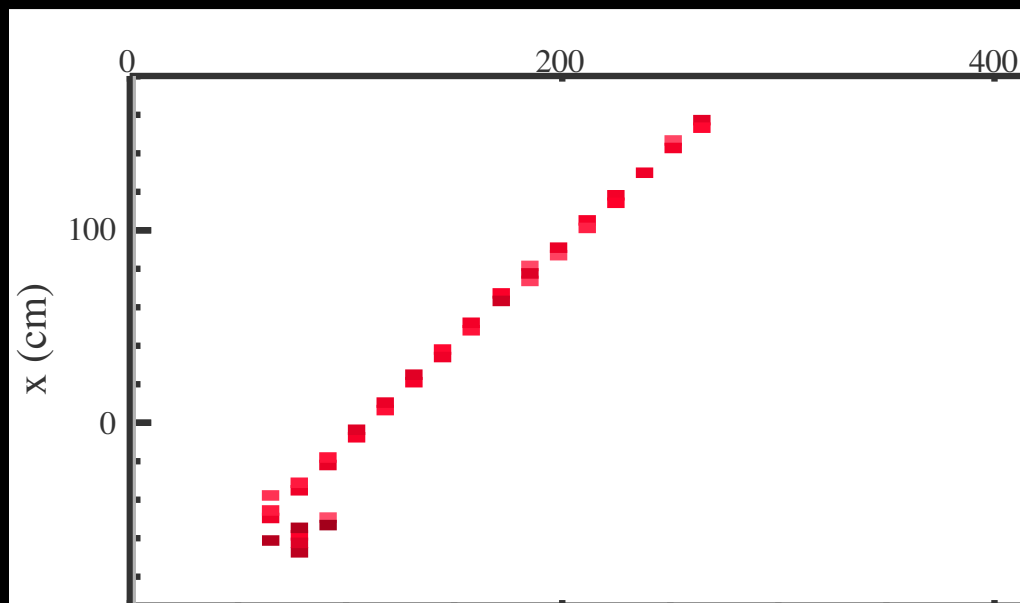


NOvA Preliminary



# Event selection Cross-checks

- Muon-removed electron (MRE) events replace muons in selected near detector  $\nu_\mu$  data candidates with electron showers of same momentum
  - Takes advantage of hadronic components directly from data
  - Electron shower simulation
  - Probes understanding of selection efficiency for  $\nu_e$  events



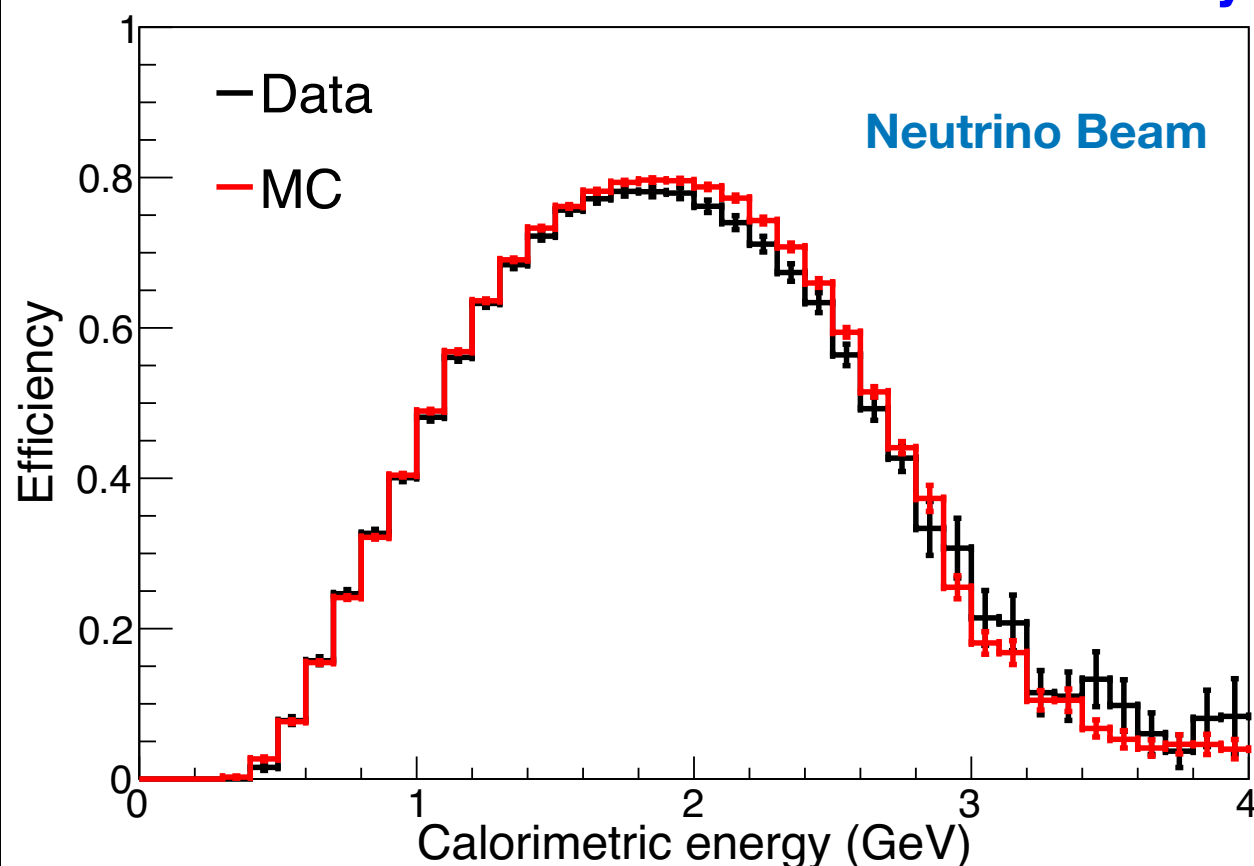
*Example*



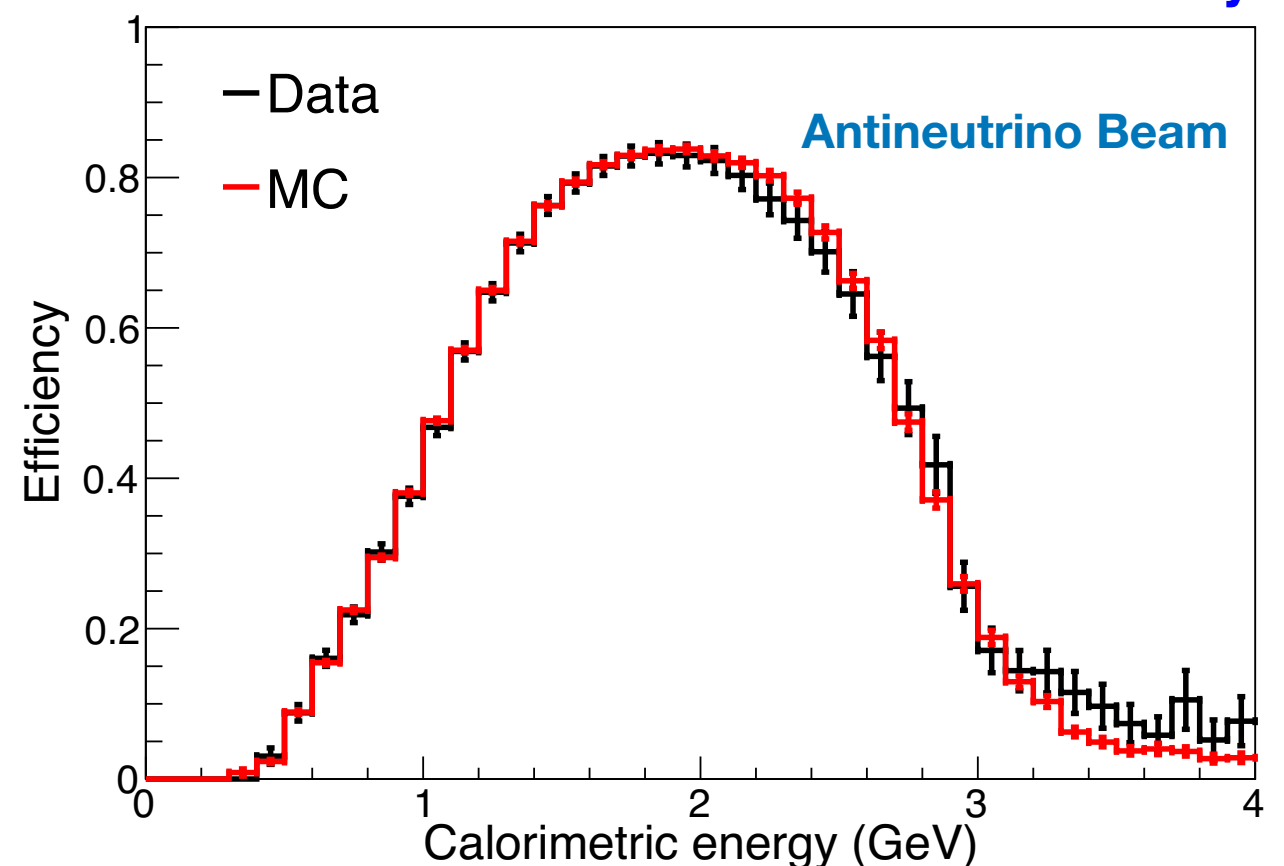
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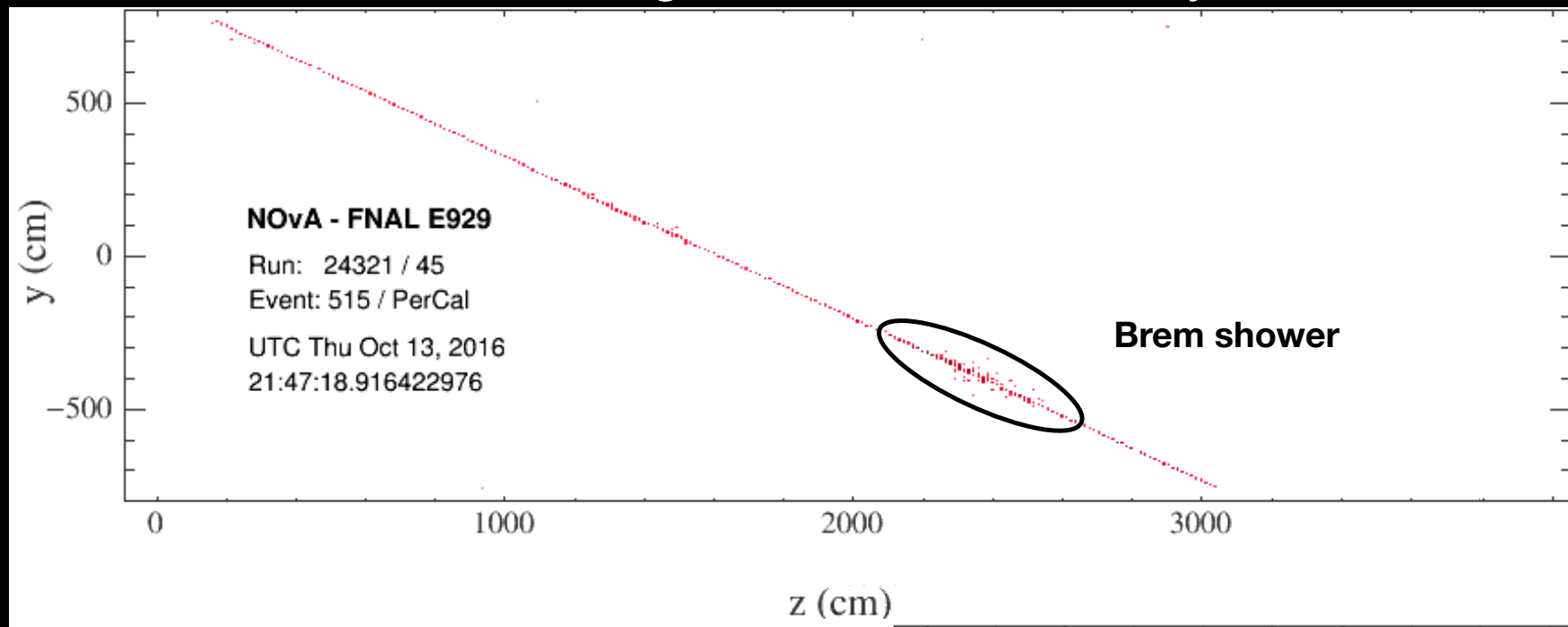


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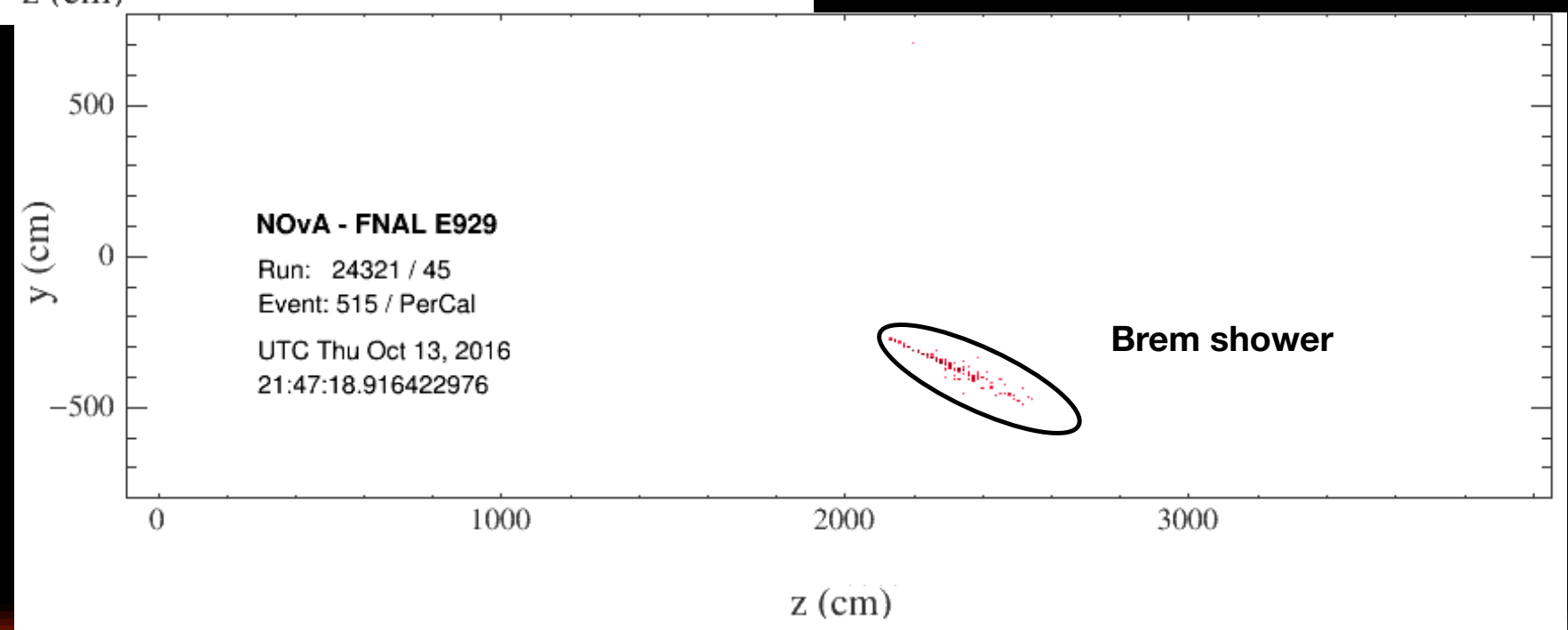


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- Muon-removed bremsstrahlung (MRBrem) removes muons from far detector cosmic data and keeps the bremsstrahlung electron shower
- Probes understanding of selection efficiency for  $\nu_e$  events

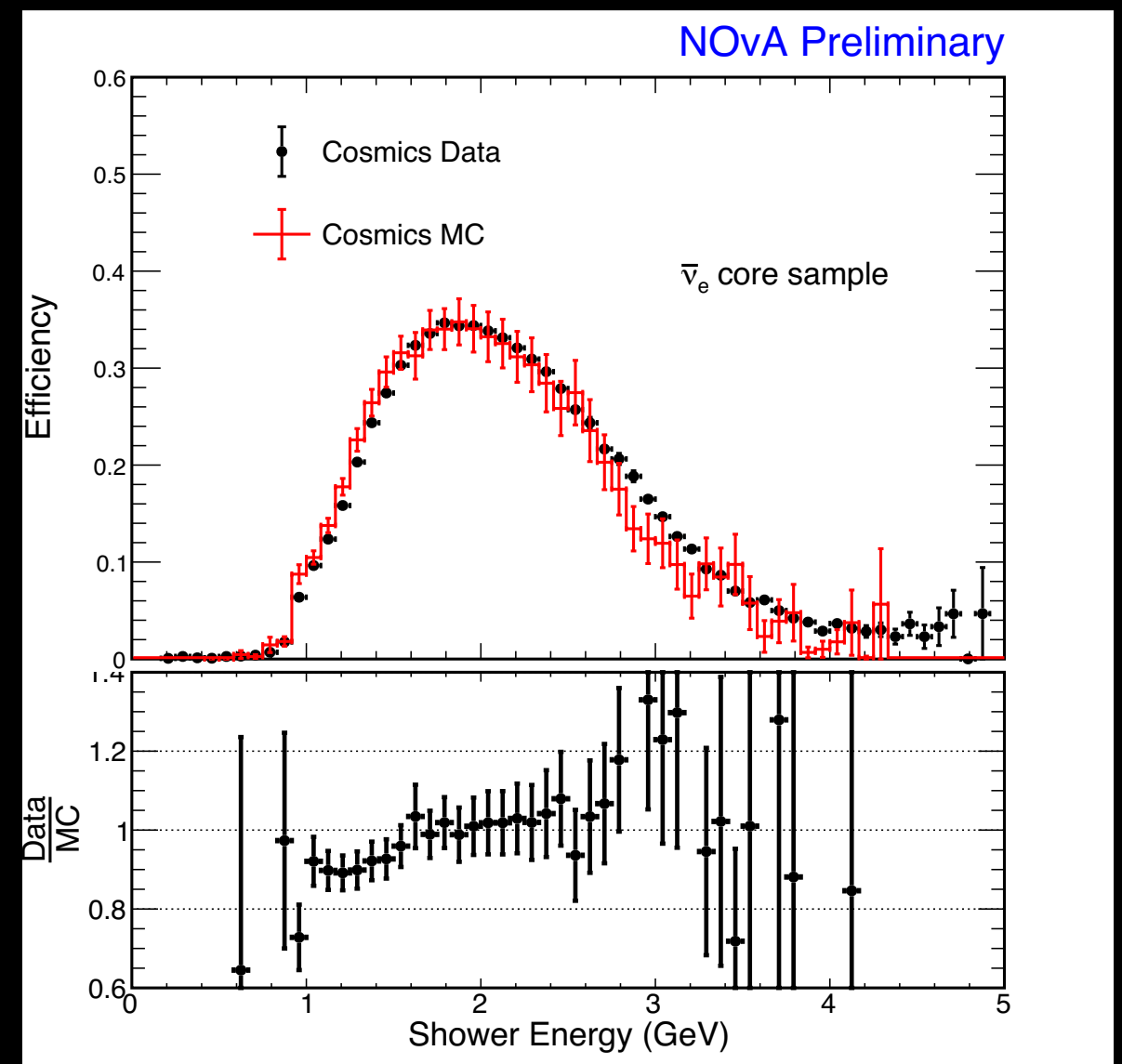
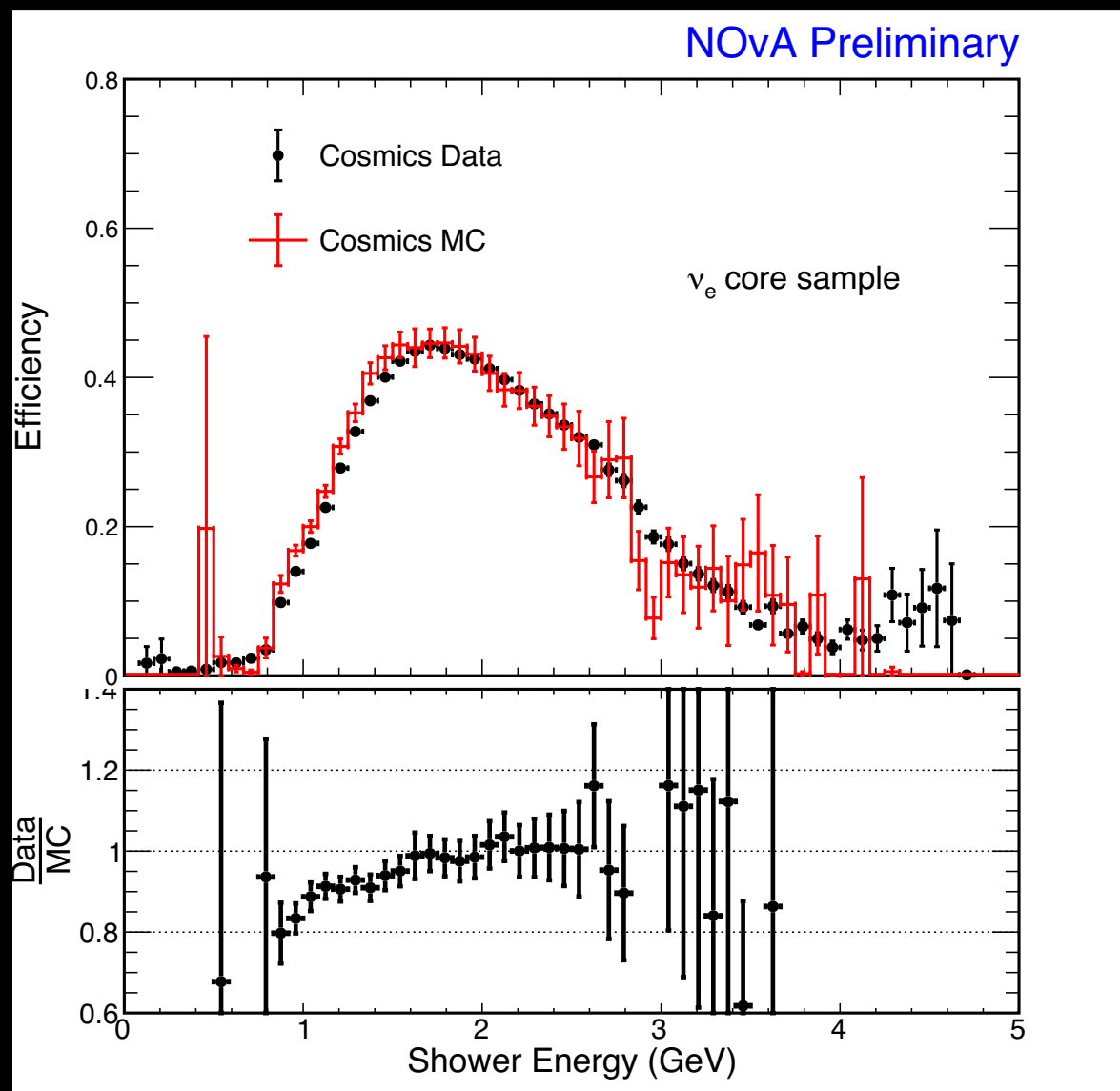


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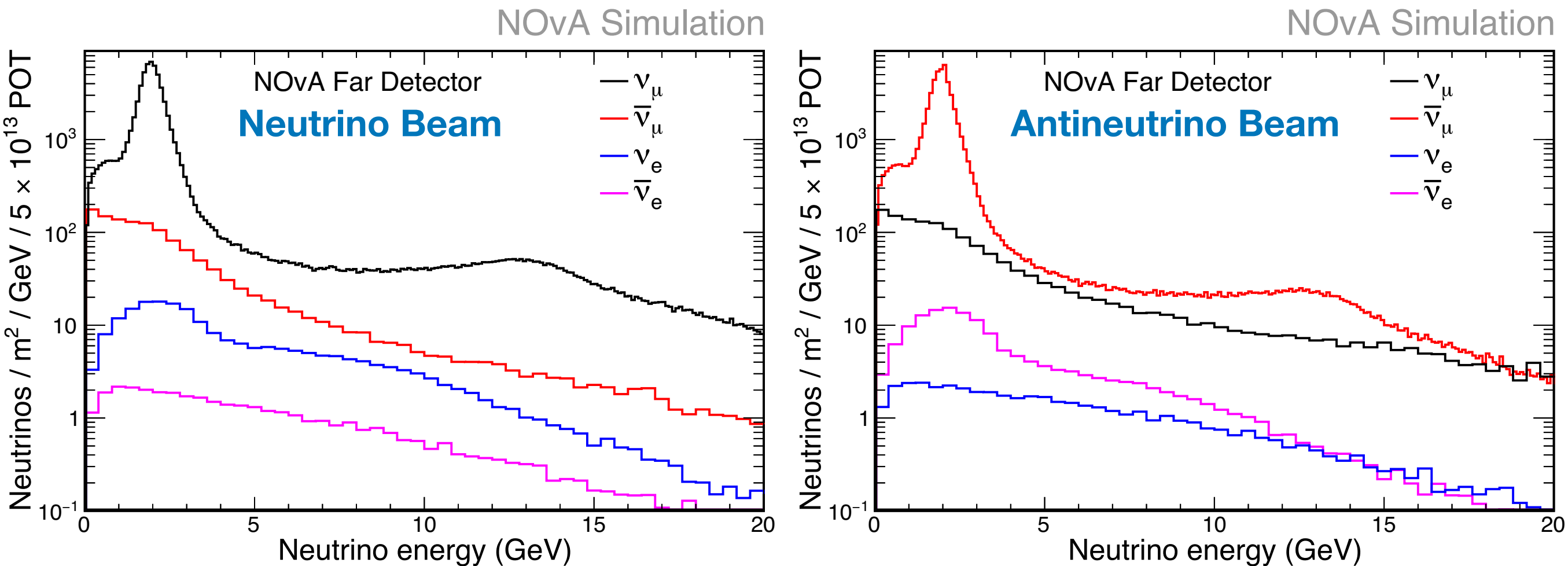
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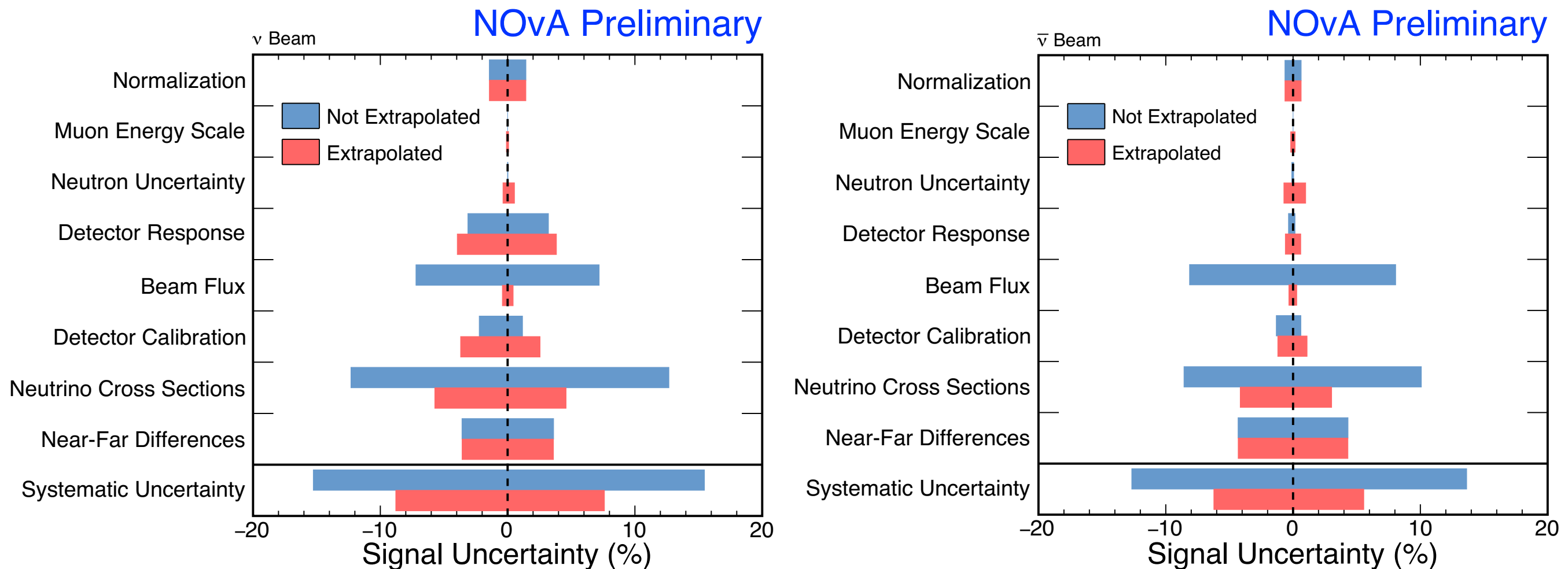
# NOvA flux

- For  $\nu_e$  appearance analysis, characterization of the  $\nu_\mu$  beam (which oscillates to signal) and the inherent  $\nu_e$  background component in the beam are important



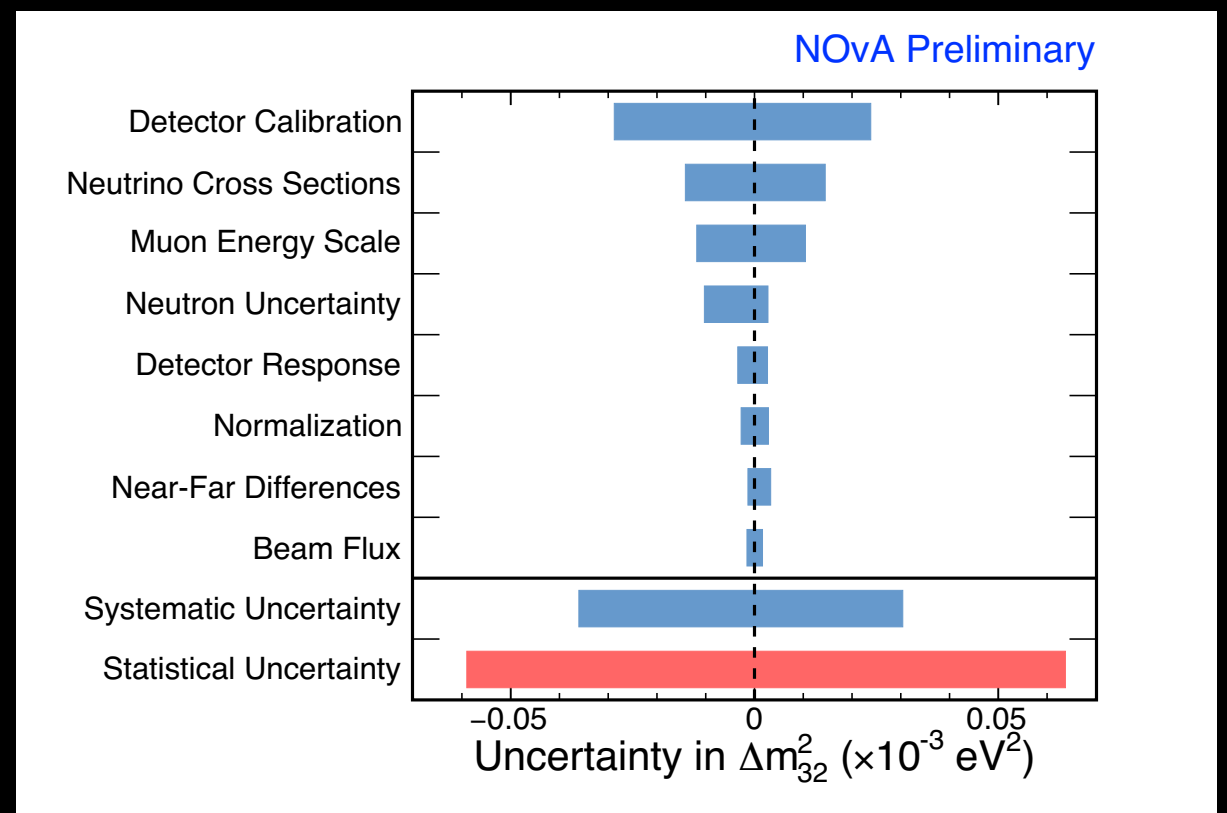
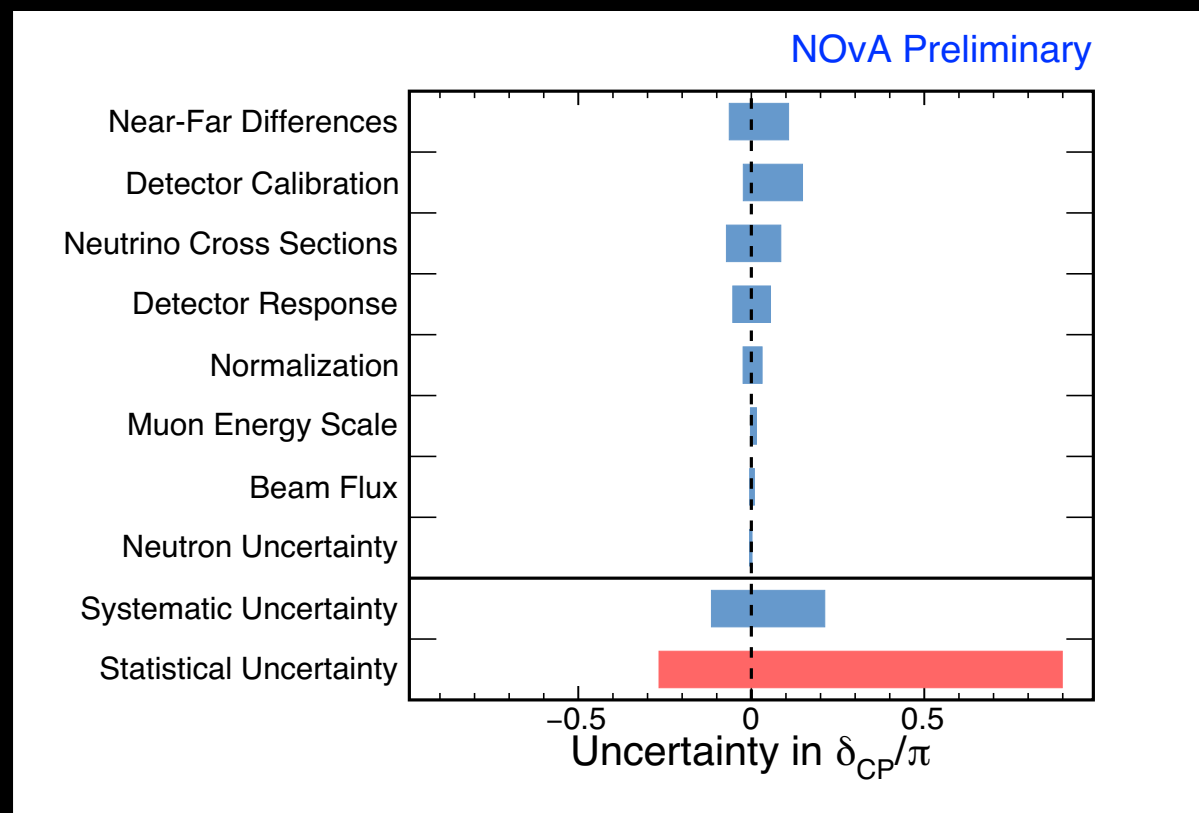
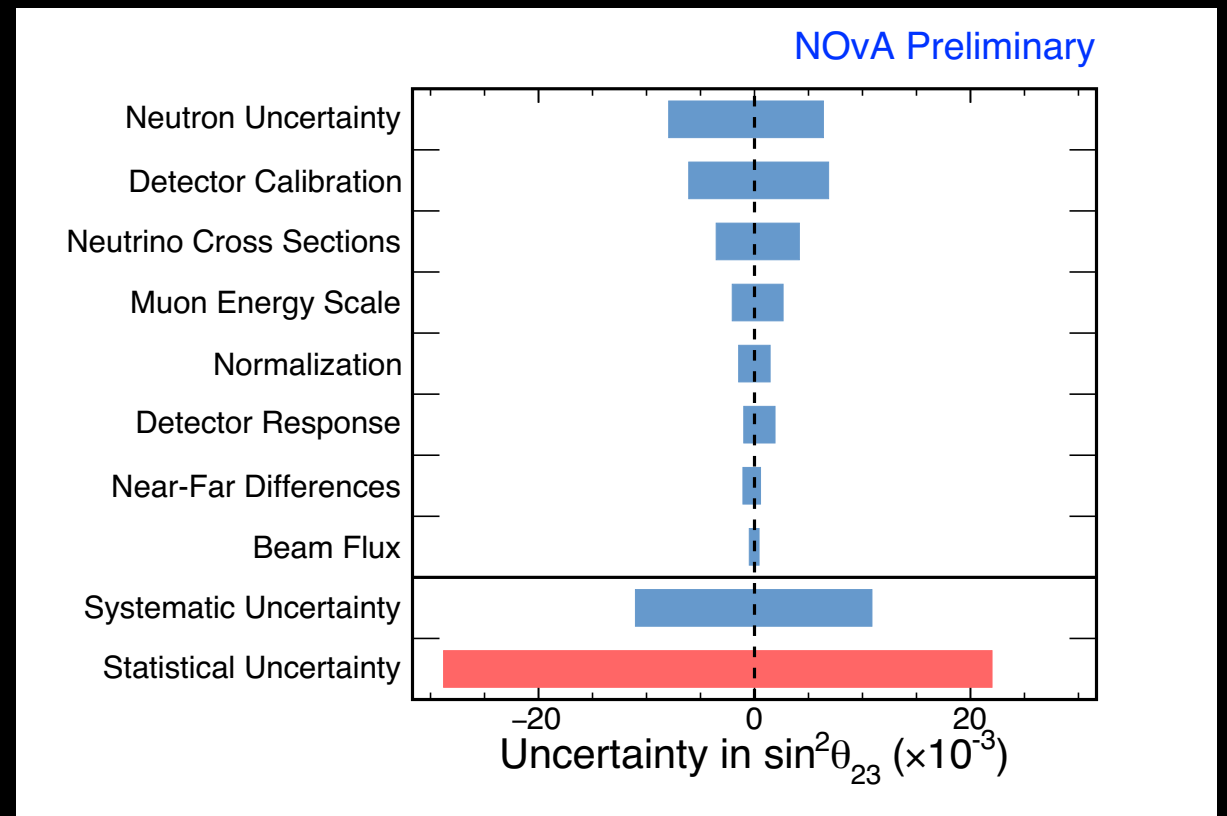
# Predicting far detector spectra

- Since NOvA has functionally similar near and far detector, extrapolation procedure reduces a number of systematic effects
  - Makes flux uncertainties quite small; greatly reduces cross-section uncertainties



# Joint $\bar{\nu}_\mu, \bar{\nu}_e$ Results

- Joint analysis is statistics limited in the major oscillation parameter results ( $\sin^2\theta_{23}$ ,  $\Delta m_{32}^2$ ,  $\delta_{CP}$ )





# Neutron systematic

- Scale amount of energy deposited by some neutrons to cover the discrepancy at low energy
- Shifts mean  $\nu_\mu$  energy by 1% in antineutrino beam and 0.5% in neutrino beam
  - Resolution changed by fractions of a percent
- Negligible impact found on selection efficiency

