The Future of Particle Physics Chris Quigg Fermi National Accelerator Laboratory



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Three Cheers for Multimessenger Astronomy!

PRL 119, 161101 (2017)

Selected for a Viewpoint in *Physics* PHYSICAL REVIEW LETTERS

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GW170817: Observation of Gravitational Waves from a Binary Neutron Star Inspiral

B. P. Abbott et al.*

(LIGO Scientific Collaboration and Virgo Collaboration)

(Received 26 September 2017; revised manuscript received 2 October 2017; published 16 October 2017)

On August 17, 2017 at 12:41:04 UTC the Advanced LIGO and Advanced Virgo gravitational-wave detectors made their first observation of a binary neutron star inspiral. The signal, GW170817, was detected with a combined signal-to-noise ratio of 32.4 and a false-alarm-rate estimate of less than one per 8.0×10^4 years. We infer the component masses of the binary to be between 0.86 and 2.26 M_{\odot} , in agreement with masses of known neutron stars. Restricting the component spins to the range inferred in binary neutron stars, we find the component masses to be in the range 1.17–1.60 M_{\odot} , with the total mass of the system $2.74^{+0.04}_{-0.01}M_{\odot}$. The source was localized within a sky region of 28 deg² (90% probability) and had a luminosity distance of 40^{+8}_{-14} Mpc, the closest and most precisely localized gravitational-wave signal yet. The association with the γ -ray burst GRB 170817A, detected by Fermi-GBM 1.7 s after the coalescence, corroborates the hypothesis of a neutron star merger and provides the first direct evidence of a link between these mergers and short γ -ray bursts. Subsequent identification of transient counterparts across the electromagnetic spectrum in the same location further supports the interpretation of this event as a neutron star merger. This unprecedented joint gravitational and electromagnetic observation provides insight into astrophysics, dense matter, gravitation, and cosmology.

DOI: 10.1103/PhysRevLett.119.161101

GW + prompt short GRB, EM transients: test gravity theories, H₀ determination, heavy-element production (no UHE CRs, v)

week ending 20 OCTOBER 2017









Fermilab's Greatest Hits @DPF2017



50 years ago: How little we knew





Problems of High-Energy Physics (NAL Design Report, January 1968)

We would like to have answers to many questions. Among them are the following:

Which, if any, of the particles that have so far been discovered, is, in fact, elementary, and is there any validity in the concept of "elementary" particles?

What new particles can be made at energies that have not yet been reached? Is there some set of building blocks that is still more fundamental than the neutron and the proton?

Is there a law that correctly predicts the existence and nature of all the particles, and if so, what is that law?

Will the characteristics of some of the very short-lived particles appear to be different when they are produced at such higher velocities that they no longer spend their entire lives within the strong influence of the particle from which they are produced?

Do new symmetries appear or old ones disappear for high momentum-transfer events?

What is the connection, if any, of electromagnetism and strong interactions?

- Do the laws of electromagnetic radiation, which are now known to hold over an enormous range of lengths and frequencies, continue to hold in the wavelength domain characteristic of the subnuclear particles?
 - What is the connection between the weak interaction that is associated with the massless neutrino and the strong one that acts between neutron and proton?
- Is there some new particle underlying the action of the "weak" forces, just as, in the case of the nuclear force, there are mesons, and, in the case of the electromagnetic force, there are photons? If there is not, why not?
- In more technical terms: Is local field theory valid? A failure in locality may imply a failure in our concept of space. What are the fields relevant to a correct local field theory? What are the form factors of the particles? What exactly is the explanation of the electromagnetic mass difference? Do "weak" interactions become strong at sufficiently small distances? Is the Pomeranchuk theorem true? Do the total cross sections become constant at high energy? Will new symmetries appear, or old ones disappear, at higher energy?













Next for Fermilab: <u>CMS</u>, g–2, µ2e, DUNE, <u>astroparticle</u>



Sanford Underground **Research Facility**

800 miles (1300 kilometers)

UNDERGROUND PARTICLE DETECTOR

EXISTING LABS



DEEP UNDERGROUND NEUTRINO EXPERIMENT









Very-High-Rate Experiments



<u>The Allure of Ultrasensitive Experiments</u> Fermilab Academic Lectures

Standard Model Production Cross Section Measurements

Status: July 2017

Xe-Xe Day @LHC

Run:280235 Timestamp:2017-10-12 21:56:43(UTC) Colliding system:Xe-Xe Energy: 5.44 TeV

Run Number: 338037, Event Number: 657472305

Date: 2017-10-13 01:45:04 CEST

Event 23965322 Run 200429 Fri, 13 Oct 2017 03:26:47

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To-do / wish list for particle physics & friends, from 2005

In a decade or two, we can hope to ...

Understand electroweak symmetry breaking Observe the Higgs boson Measure neutrino masses and mixings Establish Majorana neutrinos ($\beta\beta_{0\nu}$) Thoroughly study CP violation in B decay Exploit rare decays (K, D, ...) Observe *n* EDM, pursue e^- EDM Use top as a tool Observe new phases of matter Understand hadron structure quantitatively Uncover QCD's full implications Observe proton decay Understand the baryon excess Catalogue matter & energy of universe Measure dark energy equation of state Search for new macroscopic forces Determine GUT symmetry

... learn the right questions to ask

Detect neutrinos from the universe Learn how to quantize gravity Learn why empty space is nearly weightless Test the inflation hypothesis Understand discrete symmetry violation Resolve the hierarchy problem Discover new gauge forces Directly detect dark-matter particles Explore extra spatial dimensions Understand origin of large-scale structure Observe gravitational radiation Solve the strong CP problem Learn whether supersymmetry is TeV-scale Seek TeV dynamical symmetry breaking Search for new strong dynamics Explain the highest-energy cosmic rays Formulate problem of identity

... and rewrite the textbooks!

. . .

Interactions: $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$ gauge symmetries

Antiscreening evolution of the strong coupling "constant"

The World's Most Powerful Microscopes nanonanophysics

sum of parts rest energy Nucleon mass (~940 MeV): exemplar of $m = E_0/c^2$ up and down quarks contribute few % $3 \frac{m_u + m_d}{2} = 10 \pm 2 \text{ MeV}$ χ PT: $M_N \rightarrow 870$ MeV for massless quarks

Lattice QCD: color-confinement origin of nucleon mass has explained nearly all visible mass in the Universe

(Quark masses ensure $M_p < M_n$)

NGC 1365 · DES

QCD could be complete*, up to MPlanck ... but that doesn't prove it must be Prepare for surprises!

(Breakdown of factorization) Free quarks / unconfined color New kinds of colored matter Quark compositeness Larger color symmetry containing QCD

How might QCD Crack?

New phenomena within QCD?

Look at events in informative coordinates. than from a few specimens!

- Multiple production beyond diffraction + short-range order?
 - High density of few-GeV partons ... thermalization?
 - Long-range correlations in y?
 - Unusual event structures ...
 - More is to be learned from the river of events

New spectroscopy of quarkonium-associated states

Stable doubly heavy tetraquark mesons

Eichten & CQ, PRL

Electroweak Symmetry Breaking

Interactions: $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$ gauge symmetries

<u>The Importance of the I-TeV Scale</u>

EW theory does not predict Higgs-boson mass Thought experiment: conditional upper bound

W+W-, ZZ, HH, HZ satisfy s-wave unitarity,

If not, weak interactions among W^{\pm} , Z, H become strong on I-TeV scale

New phenomena are to be found around I TeV

provided $M_H \leq (8\pi\sqrt{2}/3G_F)^{1/2} \approx 1 \text{ TeV}$

If bound is respected, perturbation theory is everywhere" reliable

Evolution of CMS 4-lepton Signal

LHC can study Higgs boson in many channels

 $\gamma\gamma$, WW^* , ZZ^* , $\tau^+\tau^-$, b pairs, ...

Evolution of ATLAS YY Signal

What the LHC has told us about H so far

Evidence is developing as it would for a "standard-model" Higgs boson

Unstable neutral particle near 125 GeV

Motivates HL-LHC, Motivates HL-LHC, Higgs factory electron-positron Higgs

dominantly spin-parity 0⁺ Hff couplings not universal

evidence for T^+T^- , *bb*, *tt*; $\mu^+\mu^-$ limited Only third-generation fermions tested

 $M_H = 125.09 \pm 0.24 \text{ GeV}$

decays to YY, W+W-, ZZ

Quantum corrections test electroweak theory

Why does discovering the agent matter?

Imagine a world without a symmetry-breaking (Higgs) mechanism at the electroweak scale

Electron and quarks would have no mass via Higgs QCD would confine quarks into protons, etc. Nucleon mass little changed Surprise: QCD would hide EW symmetry, give tiny masses to W, Z Massless electron: atoms lose integrity No atoms means no chemistry, no stable composite structures like liquids, solids, no template for life.

<u>arXiv:0901.3958</u>

What we expect of the standard-model Higgs sector

- Hide electroweak symmetry
 - Give masses to W, Z, H
- Regulate Higgs-Goldstone scattering
- Account for quark masses, mixings
- Account for charged-lepton masses
 - A role in neutrino masses?

Motivates VLHC

ΦBSM

Fully accounts for EWSB (W, Z couplings)? Couples to fermions? t from production, Htt need direct observation for b, T Accounts for fermion masses? Fermion couplings \propto masses? Are there others? Quantum numbers? $(I^{P} = 0^{+})$ SM branching fractions to gauge bosons? Decays to new particles? All production modes as expected? Implications of $M_H \approx 125$ GeV? Any sign of new strong dynamics?

More new physics on the TeV scale? WIMP dark matter "Naturalness" Hierarchy problem: EW scale « Planck scale Vacuum energy problem Clues to origin of EWSB

Direct searches for WIMP dark matter

Supersymmetry could respond to many SM problems, but (as we currently understand it) it is largely unprincipled!

> R-parity (overkill for proton stability) gives dark-matter candidate
> µ problem (getting TeV scale right)
> Taming flavor-changing neutral currents
> All these are added by hand!

Very promising: search in EW production modes reexamine squark + EWino, too.

How have we misunderstood the hierarchy problem?

If other physical scales are present, there is something to understand

We originally sought once-and-done remedies, such as supersymmetry or technicolor

Go in steps, or reframe the problem?

The unreasonable effectiveness of the standard model

arXiv:09053187 arXiv:1503.01756 arXiv:1507.02977

A Unified Theory?

Why are atoms so remarkably neutral?

Coupling constant unification?

Extended quark–lepton families: proton decay! $n-\overline{n}$ oscillations

Unification of Forces?

1/lpha

Might (HE-)LHC (or 100-TeV) see change in evolution?

log(Q [GeV])

Parameters of the Standard Model

coupling parameters $\alpha_s, \alpha_{em}, \sin^2 \theta_W$ parameters of the Higgs potential Flavor physics may be where we see, or diagnose, vacuum phase (QCD) quark masses the break in the SM. quark mixing angles CP-violating phase charged-lepton masses neutrino masses leptonic mixing angles leptonic CP-violating phase (+ Majorana ...) arbitrary parameters

- $\zeta_{e} \left[(\bar{e}_{L} \Phi) e_{R} + \bar{e}_{R} (\Phi^{\dagger} e_{L}) \right] \rightsquigarrow m_{e} = \zeta_{e} v / \sqrt{2}$
 - after spontaneous symmetry breaking
 - What does the muon weigh?
- ς_e : picked to give right mass, not predicted
- fermion mass implies physics beyond the standard model

Why does the muon weigh?

gauge symmetry allows

Charged Fermion Masses

Running mass $m(m) \dots m(U)$

Quark family patterns: generations

Rare Processes: Flavor-changing neutral currents

SM: BR(B_s $\rightarrow \mu^{+}\mu^{-}) = (3.65 \pm 0.30) \times 10^{-9}$

MSSM: BR(B_s $\rightarrow \mu^+ \mu^-) \propto \frac{m_b^2 m_t^2}{M_A^4} \tan^6 \beta$

LHCb: $BR(B_s \to \mu^+ \mu^-) = (3.0^{+0.7}_{-0.6}) \times 10^{-9}$

 $(\mathsf{B}^0,\mathsf{B}_s) \to \mu^+\mu^-$

Flavor anomalies LHCb sees several hints of flavor nonuniversality

$B^0 \rightarrow D^* T^+ V_T / B^0 \rightarrow D^* \mu^+ V_{\mu}$

BaBar hadronic tag PRD 88 (2013) 072012 $0.332 \pm 0.024 \pm 0.018$ Belle hadronic tag PRD 92 (2015) 072014 $0.293 \pm 0.038 \pm 0.015$ Belle SL tag PRD 94 (2016) 072007 $0.302 \pm 0.030 \pm 0.011$ Belle 1-prong PRL 118 (2017) 211801 $0.270 \pm 0.035 \pm 0.027$ LHCb muonic PRL 115 (2015) 111803 $0.336 \pm 0.027 \pm 0.030$ LHCb 3-prong LHCb-PAPER-2017-017 $0.285 \pm 0.019 \pm 0.028$ LHCb average $0.306 \pm 0.016 \pm 0.022$ Fajfer et al. (SM) PRD 85 (2012) 094025 0.252 ± 0.003 0.2 0.3 0.1 0.4

 $R(D^*)$

$B_c^+ \rightarrow J/\psi \tau^+ v_\tau / B_c^+ \rightarrow J/\psi \mu^+ v_\mu$

Too many T; other evidence for excess $\mu^+\mu^-/e^+e^-$

Some outstanding questions in V physics What is the composition of V_3 ?

Before most-recent experiments

Some outstanding questions in V physics What is the composition of V₃?

T2K favors maximal mixing, NOvA nonmaximal

Some outstanding questions in V physics

NOVA, T2K v_e appearance begin to hint normal hierarchy

Some outstanding questions in V physics CP Violation? T2K disfavors $0 < \delta < \pi$ at 90% CL NOVA shows some sensitivity

> Are neutrinos Majorana particles? Search for $(Z,A) \rightarrow (Z+2,A) + ee: \beta \beta_{0v}$

Do 3 light neutrinos suffice? Are there light sterile v? Short baseline v experiments test for light steriles

Might neutrinos decay? Can we detect the cosmic v background?

Tabletop precision experiments

 $|d_e| < 8.7 \times 10^{-29} e \cdot cm$ ACME Collaboration, ThO $|d_e| < 1.3 \times 10^{-28} e \cdot cm$ NIST, trapped 180Hf19F+

(<u>SM phases</u>: $d_e < 10^{-38} e \cdot cm$)

Electric dipole moment d_e : CP/T violation

Tabletop precision experiments

- (Anti)proton magnetic moments: CPT test
 - $\mu_{\overline{b}} = -2.792 847 344 (42) \mu_N$
 - VS. $\mu_{b} = + 2.792 847 350 (9) \mu_{N}$

BASE Collaboration @CERN Antiproton Decelerator

Accelerator and magnet R&D HE-LHC (x2 in energy) requires ~15 T magnets: NbTi \rightarrow Nb₃Sn ... Nuclear & particle physics consider e(p,A) electron positron, circular or linear Higgs factory high-energy lepton collider More attention to neutrino factory

Issues for the Future (Starting now!)

1. There is a Higgs boson! Might there be several? to the weak bosons? What sets the masses and mass related to the electroweak scale? into fermion masses and mixings?

2. Does the Higgs boson regulate WW scattering? 3. Is the Higgs boson elementary or composite? How does it interact with itself? What triggers EWSB? 4. Does the Higgs boson give mass to fermions, or only mixings of the quarks and leptons? (How) is fermion 5. Are there new flavor symmetries that give insights 6. What stabilizes the Higgs-boson mass below I TeV?

Issues for the Future (Now!)

7. Do the different CC behaviors of LH, RH fermions reflect a fundamental asymmetry in nature's laws? 8. What will be the next symmetry we recognize? Are there additional heavy gauge bosons? Is nature supersymmetric? Is EW theory contained in a GUT? 9. Are all flavor-changing interactions governed by the standard-model Yukawa couplings? Does "minimal flavor violation" hold? If so, why? At what scale? 10. Are there additional sequential quark & lepton generations? Or new exotic (vector-like) fermions? II. What resolves the strong CP problem?

Issues for the Future (Now!)

and electromagnetic interactions? dimensions? 15. What resolves the vacuum energy problem? inflation? ... for dark energy?

- 12. What are the dark matters? Any flavor structure? 13. Is EWSB an emergent phenomenon connected with strong dynamics? How would that alter our conception of unified theories of the strong, weak,
- 14. Is EWSB related to gravity through extra spacetime
- 16. (When we understand the origin of EWSB), what lessons does EWSB hold for unified theories? ... for

Issues for the Future (Now!)

17.What explains the baryon asymmetry of the universe? Are there new (CC) CP-violating phases? 18. Are there new flavor-preserving phases? What would observation, or more stringent limits, on electric-dipole moments imply for BSM theories? 19. (How) are quark-flavor dynamics and lepton-flavor dynamics related (beyond the gauge interactions)? 20. At what scale are the neutrino masses set? Do they speak to the TeV, unification, Planck scale, ...? 21. Could our laws of nature be environmental?

22. How are we prisoners of conventional thinking?

