# From the VLT to ALMA and to the E-ELT

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- Mission
  - Develop and operate world-class observing facilities for astronomical research
  - > Organize collaborations in astronomy
- Intergovernmental treaty-level organization
  - Founded in 1962 by 5 countries
  - Currently 14 member states
- Observatories in Chile
  - > Optical/infrared: La Silla and Paranal
  - Sub-mm: APEX and ALMA partnerships: Chajnantor
- HQ in Garching and Office in Santiago







### La Silla Paranal Observatory

Very Large Telescope (Paranal) 9 telescopes operational, one in commissioning 14 instruments in use, 7 in development Instrumentation covers the available optical infrared wavele gths from 300nm to 20µm Angular resolution from seeing limit (0.8 arcseconds) to 50 micro-arcseconds Imagers and wide-field imagers urveyitelescopes Spectrographs • resolution from  $-5 < \lambda/\Lambda\lambda < 150000$ Interferometric array (VLTI) aser Guide Star Facility

Blois, 20 July 2010



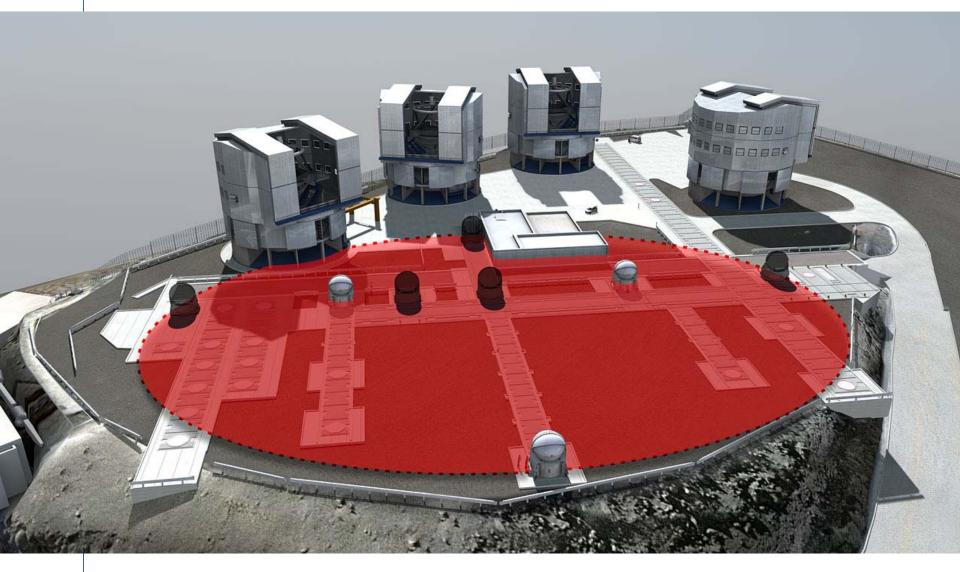


### **Laser Guide Star Facility**





### **The VLT Interferometer**





### **Top list of ESO science**

- Galactic Centre
  - Supermassive black hole
  - Measure gravity in its strong regime
- Extrasolar planets
  - Images and spectroscopy
  - Characterise other worlds
- Accelerating Universe
  - Spectroscopy of distant supernovae
- Gamma-Ray Bursts/Supernovae
  - Explosion physics
  - Tracers of the distant universe





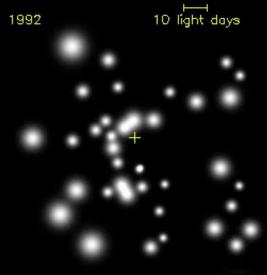
### Other top science from ESO

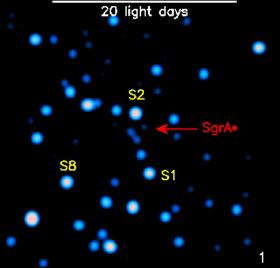
- Metal-poor stars
  - Finding the oldest known stars
  - Trace Big Bang nucleosynthesis
- Stellar populations in nearby galaxies
  - Stellar archeology
- Massive galaxies in the distant Universe
  - Formation and evolution of galaxies
- Varying physical constants?
  - Measure the fine-structure constant over time
- Testing the cosmological model
  - Cosmic background temperature



### Black hole at the Galactic Centre

Mass determination through stellar orbits Structure around the black hole revealed through flashes **Coordinated studies** with other wavelengths Multi-year study use of AO instruments (SHARP on NTT, ISAAC NACO, SINFONI on VLT)

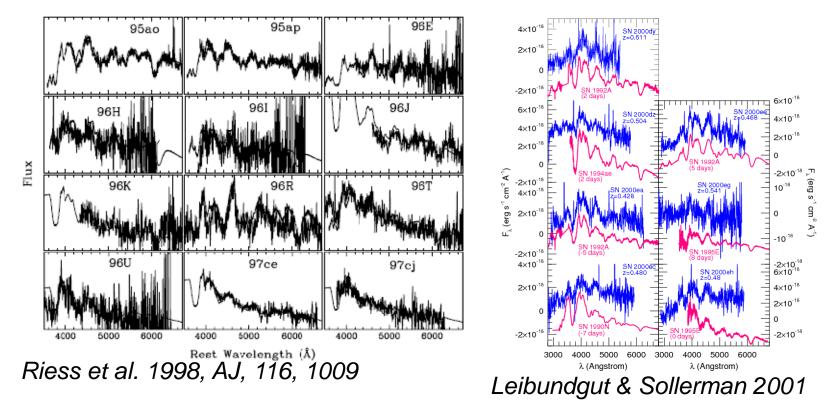






### **Accelerating Universe**

 Contribution of most of the early photometry and spectroscopy of High-z SN Search Team
 > difference between a 4m and a 8m telescope



European Southern Observatory



### Gamma-Ray Bursts

Identification relies on optical data

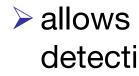
- redshifts, explosion energies, explosion physics
- Cosmological probes
  - the most distant observable stars
  - > light houses to measure the intergalactic medium
  - tracers of chemical enrichment?
- Very short duration
  - required special instrumentation and software to observe adequately

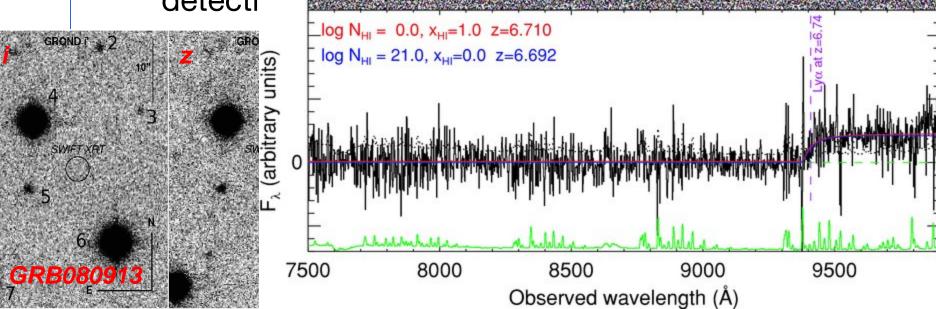


### Gamma-Ray Bursts

Most distant stellar objects ever observed

- redshifts 6.7 and 8.2 (tentative)
- Iookback time of nearly 12.5 billion years (or 95% of the age of the universe)
- VLT equipped with rapid response mode

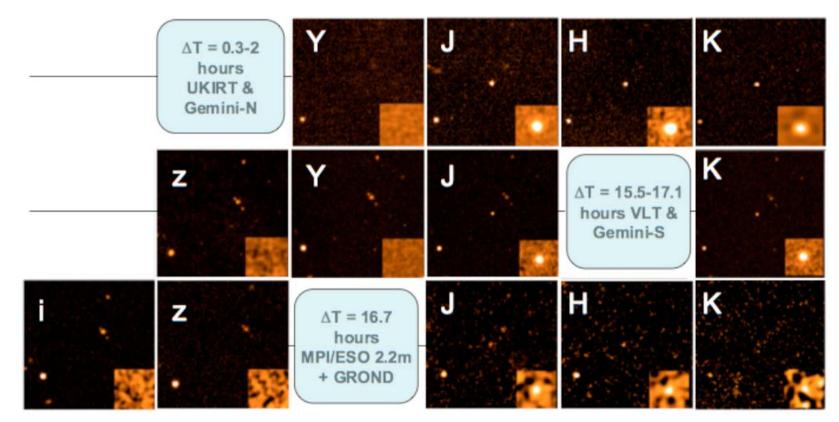






### Most distant stellar object yet observed – GRB 090423

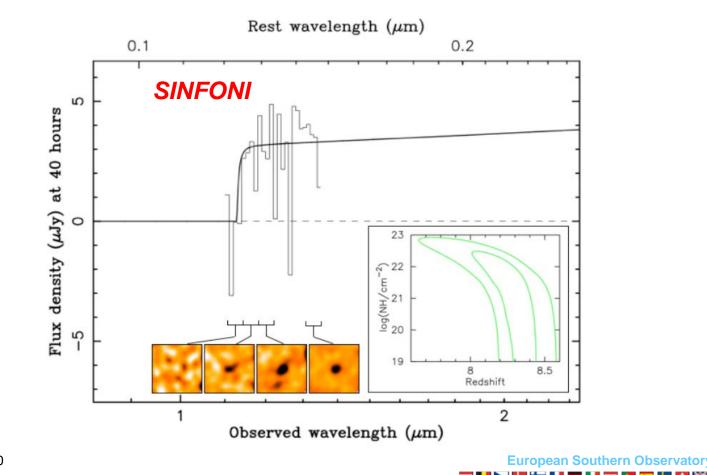
## Optical drop-out, bright in the near-infrared Rapid decline



Tanvir et al., Nature submitted

### GRB 090423

## Spectroscopy 17 hours after outburst Lyman break indicates a redshift of z≈8.2



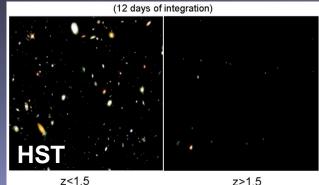
### **Atacama Large Millimeter Array**

Global project with Europe, North America and East Asia as partners

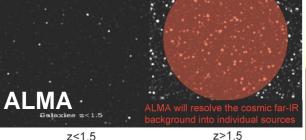
Science requirements
Detect CO and [CII] in Milky Way galaxy at z=3 in < 24 hr</li>
Dust emission, gas kinematics in proto-planetary disks
Resolution to match Hubble, JWST and 8-10m with AO
Complement to Herschel

#### Specifications

66 antennas (54x12m, 12x7m)
 14 km max baseline (< 10mas)</li>
 30-1000 GHz (10–0.3mm), up to 10 receiver bands



simulation 3 days of integration 4'x4' arcmin



.21.5

850 GHz

5AU



### ALMA

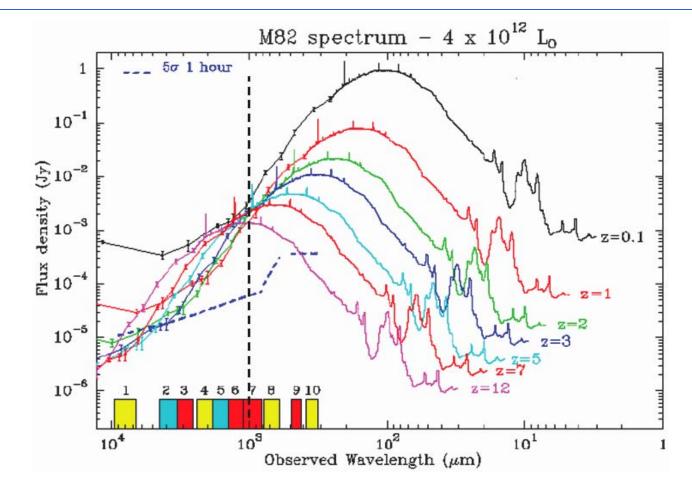
ALMA will explore the cold universe

- > typical gas temperatures` 10 K < T<sub>gas</sub> < 500 K</p>
- > molecular astrophysics
- ALMA has great capabilities to observe the distant universe
  - 'inverse' K-correction
  - detection of the cold gas in distant galaxies as tracers of star formation
    - [C II] 157µm is the strongest coolant in most galaxies
    - CO amongst the most abundant molecules in the universe

Sunyaev-Zeldovich effect directly measurable

### +ES+ 0

### K-corrections due to redshift



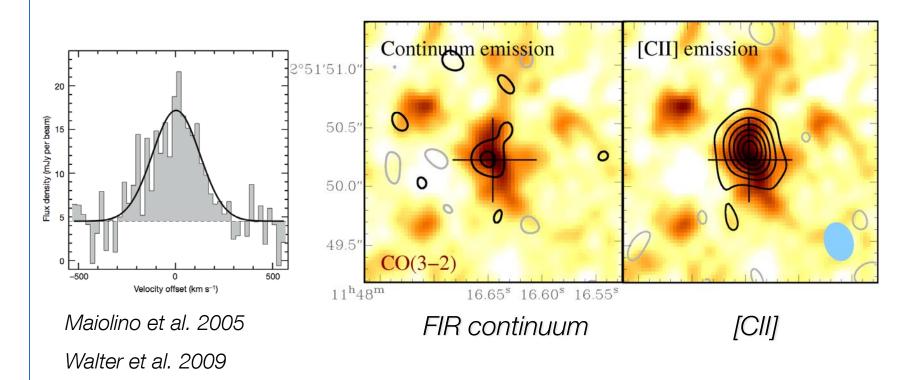
In the (sub-)millimeter the inverse K-correction compensates for the distance as redshift increases





### Starburst at z=6.4

#### Detection of [C II] in the distant universe

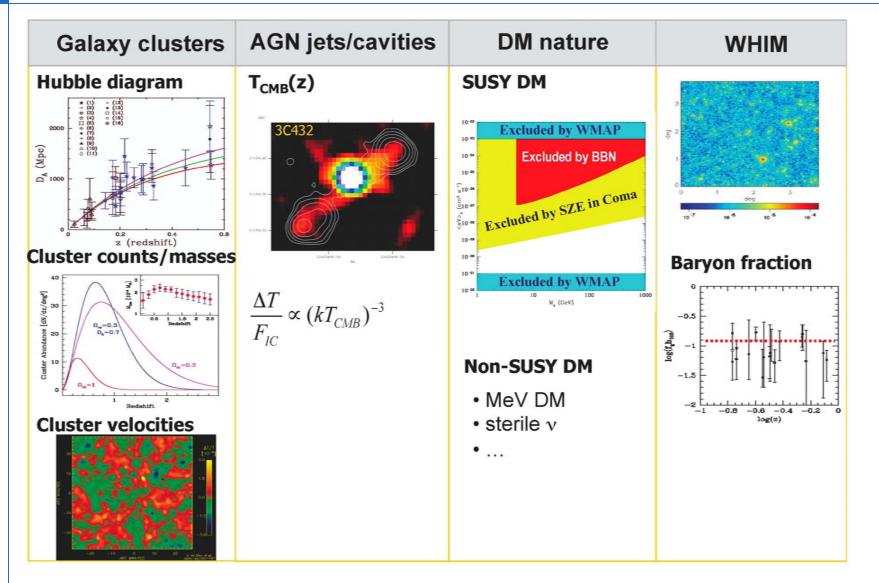


 $\geq$  [CII] size ~ 1.5 kpc => SFR/area ~ 1000 M<sub>o</sub> yr<sup>-1</sup> kpc<sup>-2</sup>

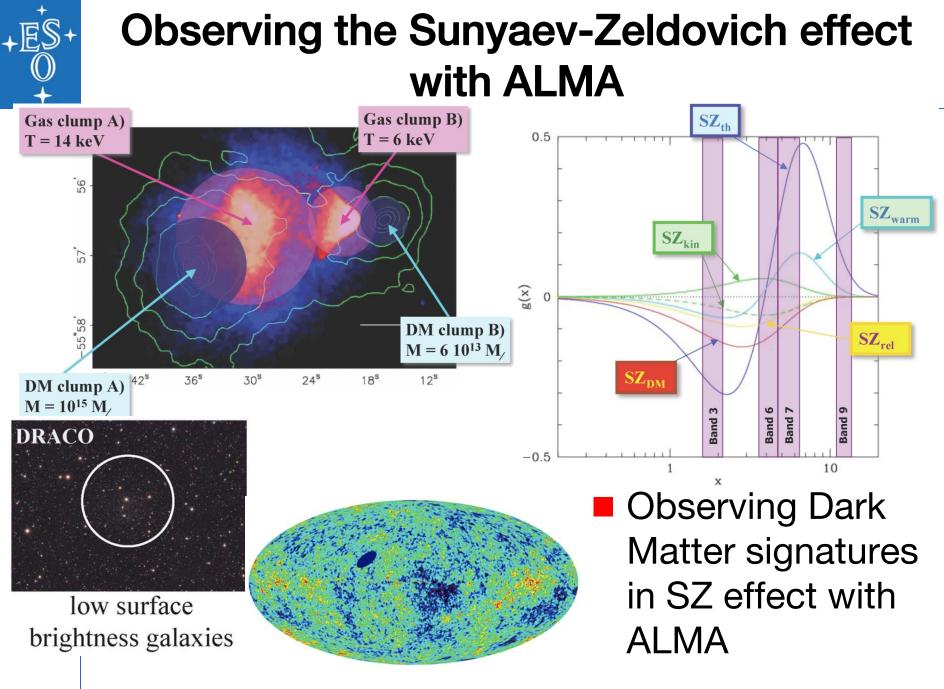
Blois, 20 July 2010



### **Astrophysical relevance**



Blois, 20 July 2010



### ALMA 2013





### E-ELT

Detailed design study
 Baseline 42m primary mirror
 Adaptive optics built-in
 8 instrument studies and 2 adaptive optics modules -> selection of the first generation of instruments (~ 6 instruments)
 Community strongly engaged
 Study complete in 2010

Project
Builds on entire expertise at ESO and in the member states
Construction 2011-2018
Synergy: JWST/ALMA/SKA



### Science cases for the E-ELT

Observe Earth-like planets in the habitable zone

- > characterisation, e.g. atmosphere
- > observe planetary systems
- Measure the dynamics of the cosmic expansion
  - > see Joe Liske's presentation
- Measure any evolution in the value of the finestructure constant α
  - ➤ also Joe's talk
- Observe the stellar population (individual stars) out to the Virgo cluster
- Explore the high-redshift universe
  - 'first objects'





### ESO telescopes have plenty to contribute to cosmology

- The La Silla Paranal Observatory continues to be upgraded
  - Second generation instruments (VLT/VLTI)
  - Key surveys with VST and VISTA
- Key science themes:
  - $\succ$  strong gravity  $\rightarrow$  black hole in the Galactic Centre
  - $\succ$  distant universe  $\rightarrow$  galaxy formation, first stars
- ALMA provides access to the cold stages of matter
- E-ELT will open the possibility to directly measure
  - $\succ$  dynamics of cosmic expansion
  - address the constancy of the fine-structure constant

