


Open Neutrino Questions

A photograph of a historic European town square, likely Blois, France. The square is paved with cobblestones and surrounded by multi-story buildings with ornate facades, including a prominent tower with a conical roof. The sky is blue with scattered white clouds. The text "Open Neutrino Questions" is overlaid in a large, red, serif font with a white outline.

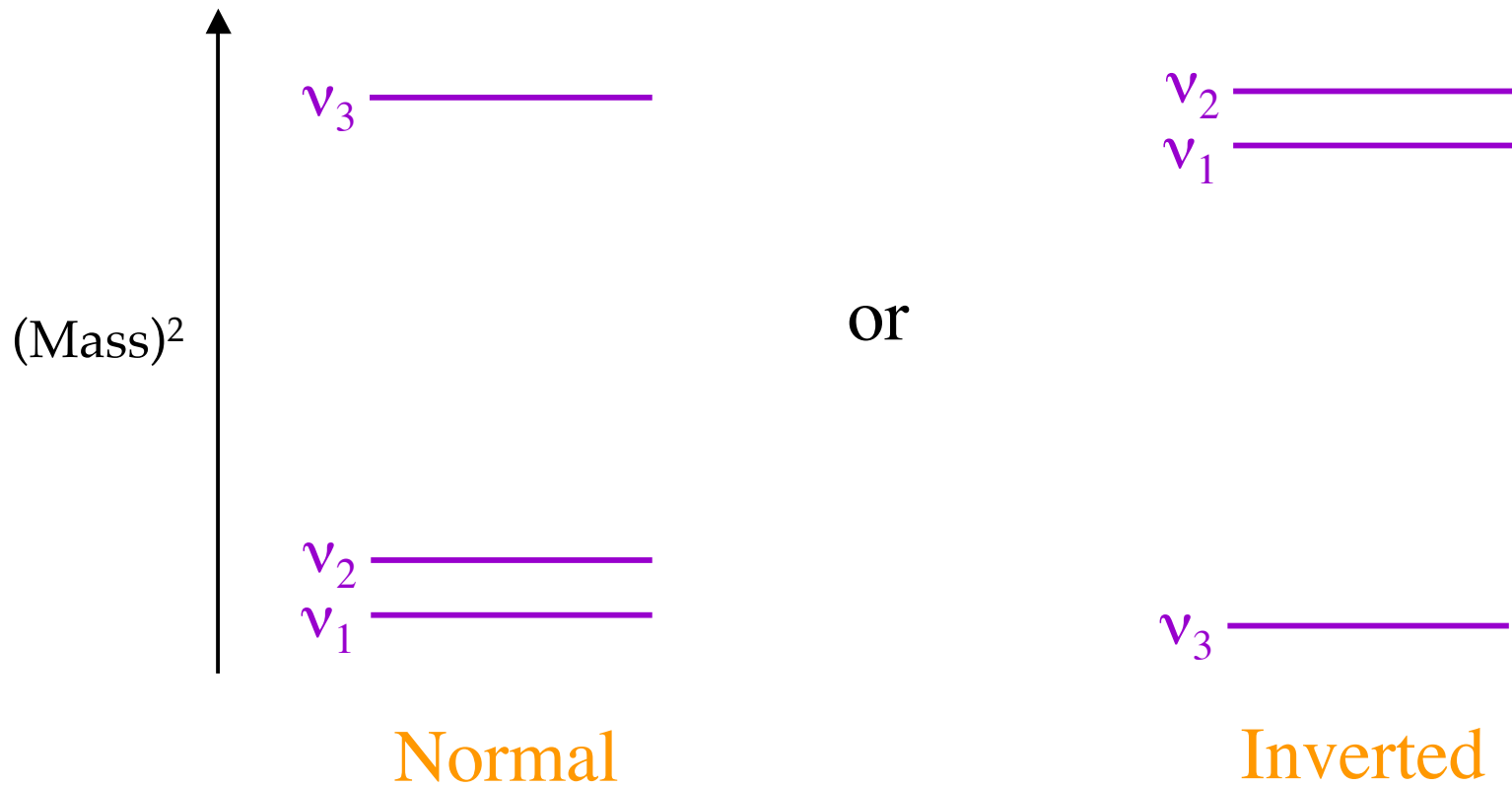
Boris Kayser, Blois, 19 July 2010



**What We
Have Learned**

Stan Wojcicki

The (Mass)² Spectrum



$$\Delta m_{21}^2 \cong 7.6 \times 10^{-5} \text{ eV}^2, \quad \Delta m_{32}^2 \cong 2.4 \times 10^{-3} \text{ eV}^2$$

Are there *more* mass eigenstates, as LSND suggests,
and MiniBooNE recently hints?

Absolute Mass Scale

Tritium β decay: $\langle m_i \rangle < 2\text{eV}$

Cosmology: $\Sigma m_i < (0.17 - 1.0)\text{eV}$

Mass(ν_i)

Leptonic Mixing

The neutrinos $\nu_{e,\mu,\tau}$ of definite flavor

($W \rightarrow e\nu_e$ or $\mu\nu_\mu$ or $\tau\nu_\tau$)

are **superpositions** of the neutrinos of definite mass:

$$|\nu_\alpha\rangle = \sum_i U^*_{\alpha i} |\nu_i\rangle .$$

Neutrino of flavor $\alpha = e, \mu, \text{ or } \tau$

Unitary (?) Leptonic Mixing Matrix

Neutrino of definite mass m_i

The Mixing Matrix

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \times \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \times \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Hints??

$$c_{ij} \equiv \cos \theta_{ij}$$

$$s_{ij} \equiv \sin \theta_{ij}$$

$$\times \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Doesn't affect oscillation

$$\theta_{12} \approx 34^\circ, \theta_{23} \approx 39-51^\circ, \theta_{13} \lesssim 11^\circ$$

δ would lead to $P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) \neq P(\nu_\alpha \rightarrow \nu_\beta)$. ~~CP~~

But note the crucial role of $s_{13} \equiv \sin \theta_{13}$.

Multiplied out, the mixing matrix is close to the simple **Tri-Bi-Maximal** mixing:

$$U_{TBM} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \\ \sqrt{\frac{1}{6}} & -\sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}$$

Is there a deep underlying reason, such as a symmetry, for this pattern?



The Open Questions

- What is the absolute scale of neutrino mass?
- Are neutrinos their own antiparticles?
- Are there *more* than 3 mass eigenstates?
 - Are there “sterile” neutrinos?
- What are the neutrino magnetic and electric dipole moments?

- What is the pattern of mixing among the different types of neutrinos?

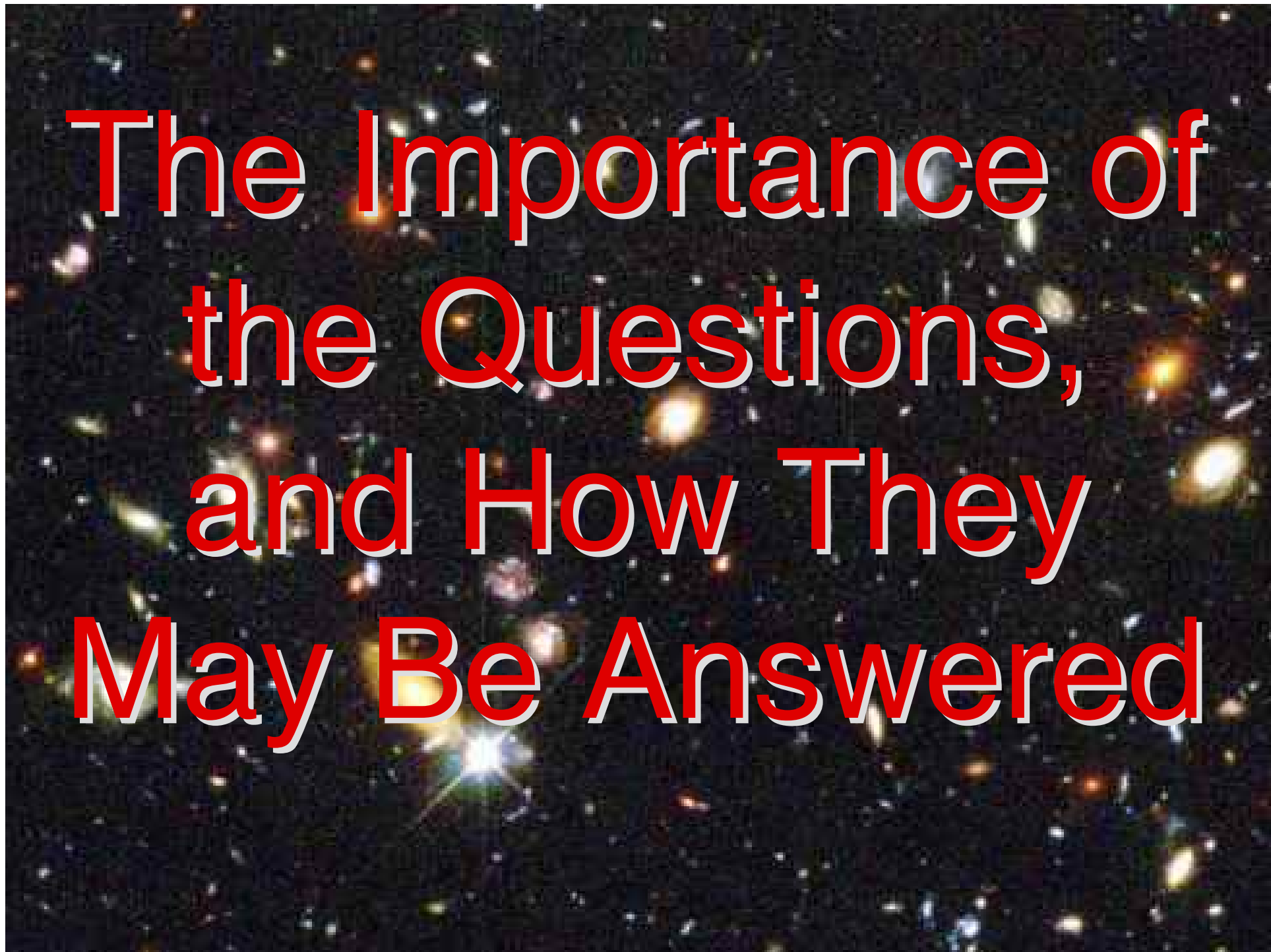
What is θ_{13} ?

- Is the spectrum like \equiv or \equiv ?

- Do neutrino – matter interactions violate CP?

Is $P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) \neq P(\nu_\alpha \rightarrow \nu_\beta)$?

- What can neutrinos and the universe tell us about one another?
- Is CP violation involving neutrinos the key to understanding the matter – antimatter asymmetry of the universe?
- What physics is behind neutrino mass?
- What *surprises* are in store?



The Importance of
the Questions,
and How They
May Be Answered

Are Neutrinos Their Own Antiparticles?

For each *mass eigenstate* ν_i , and given helicity h , does —

- $\bar{\nu}_i(h) = \nu_i(h)$ (Majorana neutrinos)

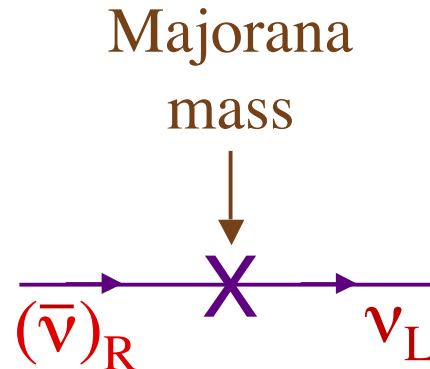
or

- $\bar{\nu}_i(h) \neq \nu_i(h)$ (Dirac neutrinos) ?

Equivalently, do neutrinos have *Majorana masses*? If they do, then the mass eigenstates are *Majorana neutrinos*.

Majorana Masses

An example of
their effect:



Majorana masses mix ν and $\bar{\nu}$, so they do not conserve the **Lepton Number L** that distinguishes leptons from antileptons:

$$L(\nu) = L(\ell^-) = -L(\bar{\nu}) = -L(\ell^+) = 1$$

A Majorana mass for any fermion f causes $f \leftrightarrow \bar{f}$.

Quark and *charged-lepton* Majorana masses are forbidden by electric charge conservation.

Neutrino Majorana masses would make the neutrinos *very* distinctive.

Majorana neutrino masses have a different origin than the quark and charged-lepton masses.

Why Majorana Masses \longrightarrow Majorana Neutrinos

As a result of $K^0 \longleftrightarrow \bar{K}^0$ mixing, the neutral K mass eigenstates are —

$$K_{S,L} \cong (K^0 \pm \bar{K}^0)/\sqrt{2} . \quad \overline{K_{S,L}} = K_{S,L} .$$

Majorana masses induce $\nu \longleftrightarrow \bar{\nu}$ mixing.

As a result of $\nu \longleftrightarrow \bar{\nu}$ mixing, the neutrino mass eigenstate is —

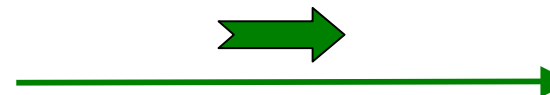
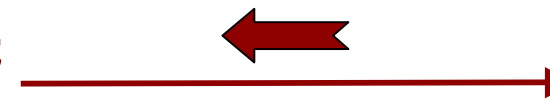
$$\nu_i = \nu + \bar{\nu} . \quad \bar{\nu}_i = \nu_i .$$

A Dirac Neutrino

We have 4 mass-degenerate states:

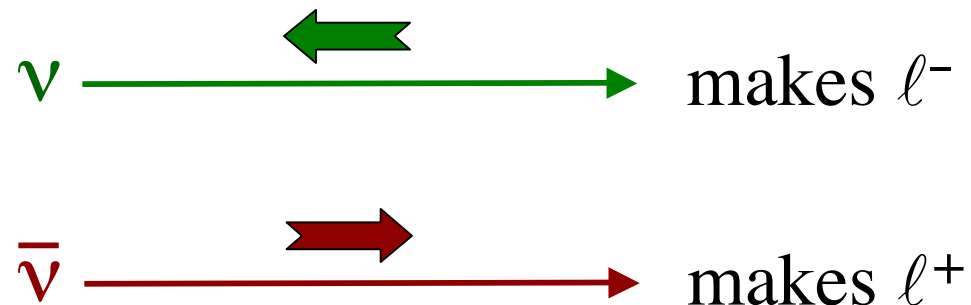
$\bar{\nu}$  makes ℓ^-

ν  makes ℓ^+

ν 
 $\bar{\nu}$ 
 These states do not interact.
 (The weak interaction is Left Handed.)

A Majorana Neutrino

We have only 2 mass-degenerate states:



The weak interactions violate *parity*.

An incoming left-handed neutral lepton makes ℓ^- .

An incoming right-handed neutral lepton makes ℓ^+ .

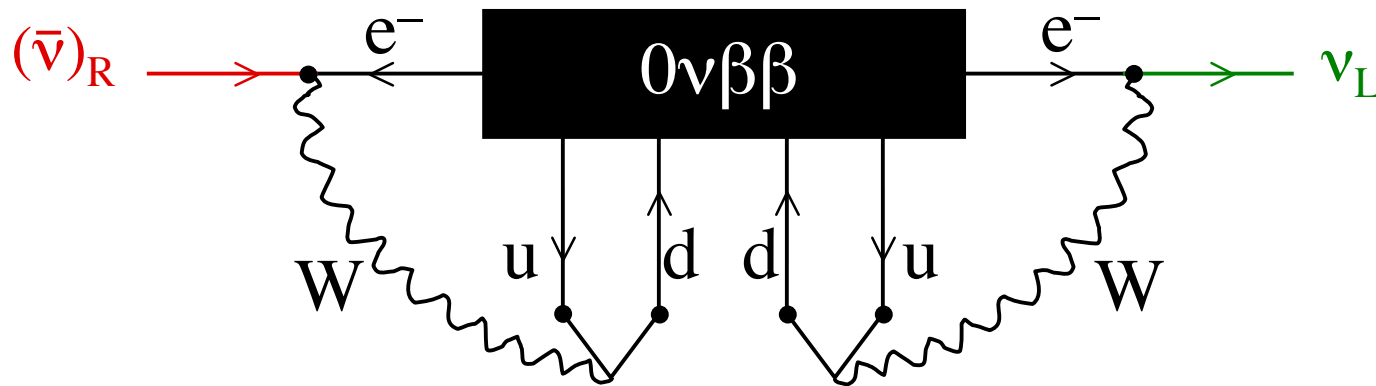
To Determine
Whether
Majorana Masses
Occur in Nature

The Promising Approach — Seek Neutrinoless Double Beta Decay [$0\nu\beta\beta$]



Whatever diagrams cause $0\nu\beta\beta$, its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)




$(\bar{\nu})_R \rightarrow \nu_L$: A (tiny) Majorana mass term

$\therefore 0\nu\beta\beta \rightarrow \bar{\nu}_i = \nu_i$

Mixing,
Mass Ordering,
and ~~CP~~

The Questions For
Oscillation Experiments

A desk lamp with a glowing shade, casting light on a surface. The lamp is positioned on the left side of the frame, and its light illuminates the text in the center. The background is a dark, gradient surface.

The theory of neutrino oscillations remains solid.

For a neutrino beam of energy E , the usual expression for the wavelength λ of the oscillation caused by a splitting Δm^2 is —

$$\lambda \equiv \text{Source to Detector distance for one oscillation} = \frac{4\pi E}{\Delta m^2} .$$

Recently, it was argued that this expression can be wrong by 27% to 57%!

The Δm^2 extracted from data using this expression would then be similarly wrong.

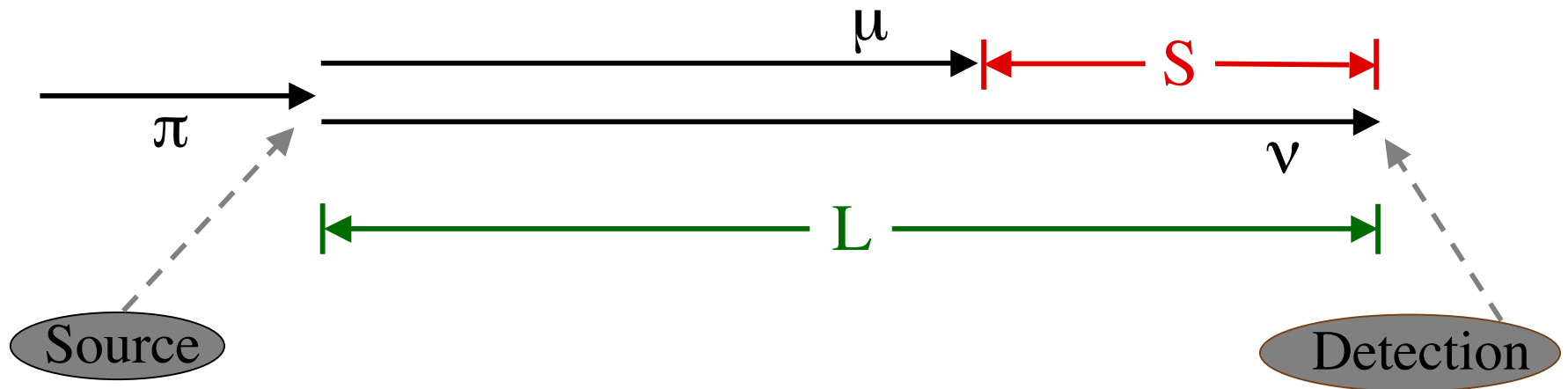
R. G. H. Robertson

Should one worry?

No!

The recent approach involves *entanglement* between the neutrino and its associated recoil — the muon in $\pi \rightarrow \mu\nu$.

The wavelength λ_S derived by this approach is in reality for oscillation as a function of the lab-frame $\nu - \mu$ separation S at the time in the π rest frame when the neutrino is detected:



One finds that —

$$\frac{\lambda_S}{\lambda_{Usual}} = \frac{S}{L}$$

That is, the “new” wavelength λ_S is physically equivalent to the usual one.

It just refers to a different distance variable.

B. K., Kopp, Robertson, and Vogel

No need to worry!

The Mass Spectrum: $\underline{\underline{=}}$ or $\underline{=}$?

Generically, grand unified models (GUTS) favor —

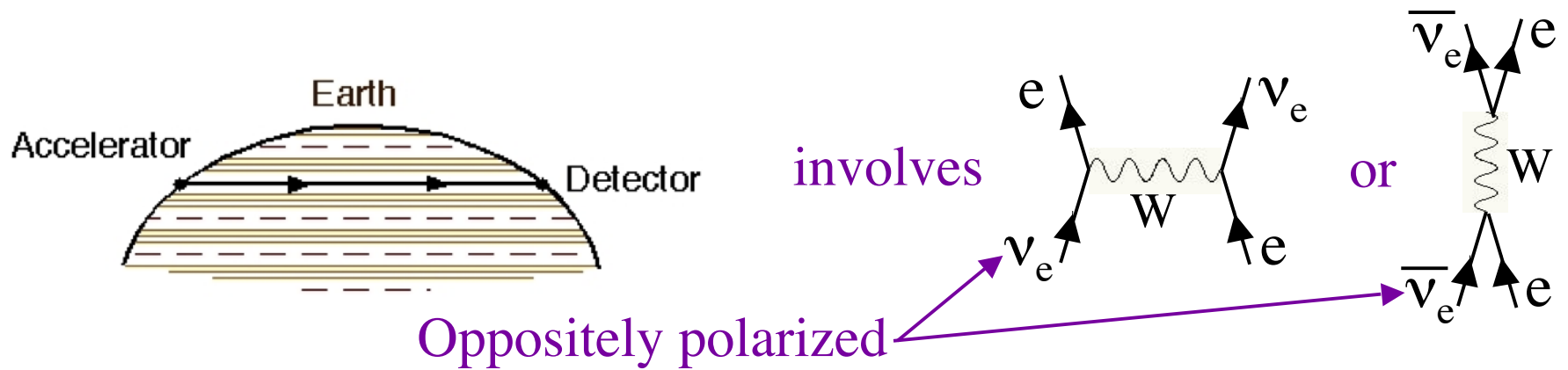
$\underline{\underline{=}}$

GUTS relate the **Leptons** to the **Quarks**.

However, *Majorana masses*, with no quark analogues, could turn $\underline{\underline{=}}$ into $\underline{=}$.

How To Determine If The Spectrum Is Normal Or Inverted

Exploit the *matter effect*.



This raises the effective mass of ν_e , and lowers that of $\bar{\nu}_e$.

This leads to —

$$\frac{P(\nu_{\mu} \rightarrow \nu_e)}{P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)} \begin{cases} > 1 ; \equiv \\ < 1 ; \equiv \end{cases}$$




Do Neutrino Interactions
Violate CP?


*Are we descended
from heavy neutrinos?*

The Challenge – A Cosmic Broken Symmetry

The universe contains baryons,
but essentially no antibaryons.

Standard cosmology: Any initial
baryon – antibaryon asymmetry
would have been erased.

How did $n_B = n_{\bar{B}}$  $n_B \gg n_{\bar{B}}$?

Sakharov: $n_B = n_{\bar{B}}$  $n_B \gg n_{\bar{B}}$ requires \mathcal{CP} .

The \mathcal{CP} in the quark mixing matrix,
seen in B and K decays, leads to
much too small a $B-\bar{B}$ asymmetry.

If *quark* \mathcal{CP} cannot generate
the observed $B-\bar{B}$ asymmetry,
can some scenario involving *leptons* do it?

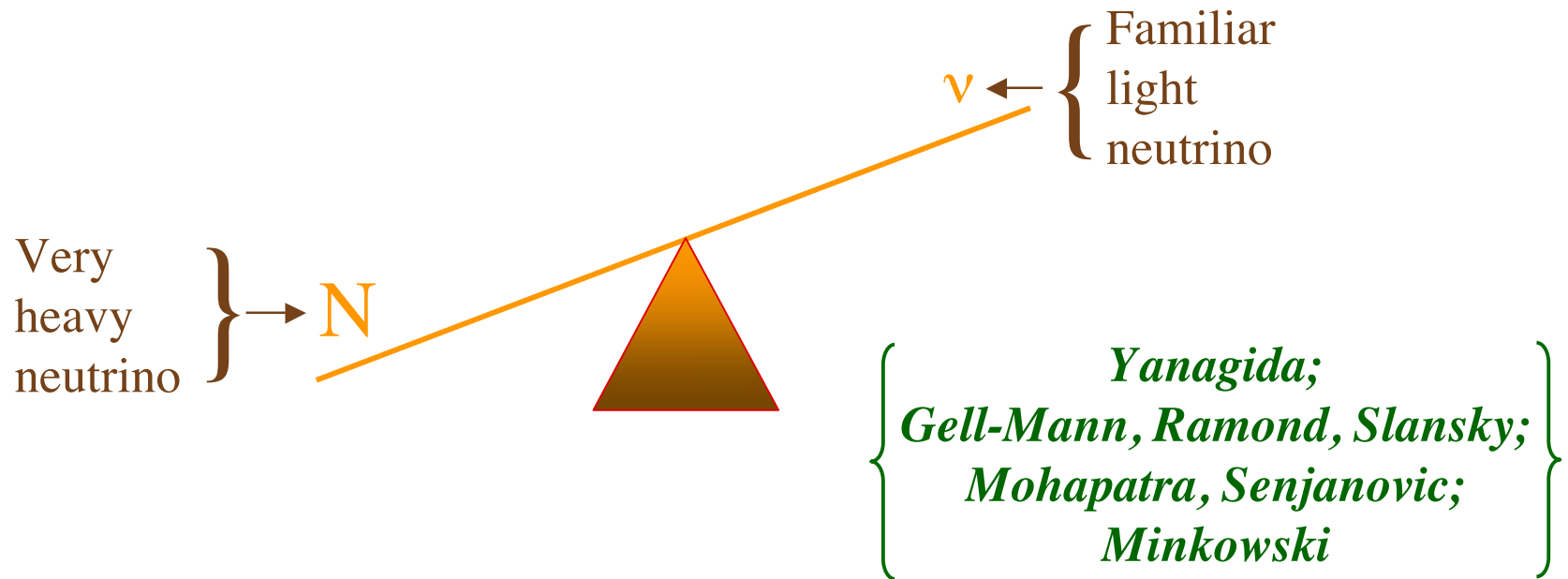
The candidate scenario: *Leptogenesis*.

(*Fukugita, Yanagida*)

Leptogenesis – The General Idea

Leptogenesis is an outgrowth of the most popular theory of why neutrinos are so light —

The See-Saw Mechanism



The *very* heavy neutrinos **N** would have been made in the hot Big Bang.

Leptogenesis — Step 1

The heavy neutrinos N , like the light ones ν , are Majorana particles. Thus, an N can decay into e^- or e^+ .

If ν oscillation violates CP, then quite likely so does N decay. In the See-Saw, these two CP violations have a common origin: One Yukawa coupling matrix, y .

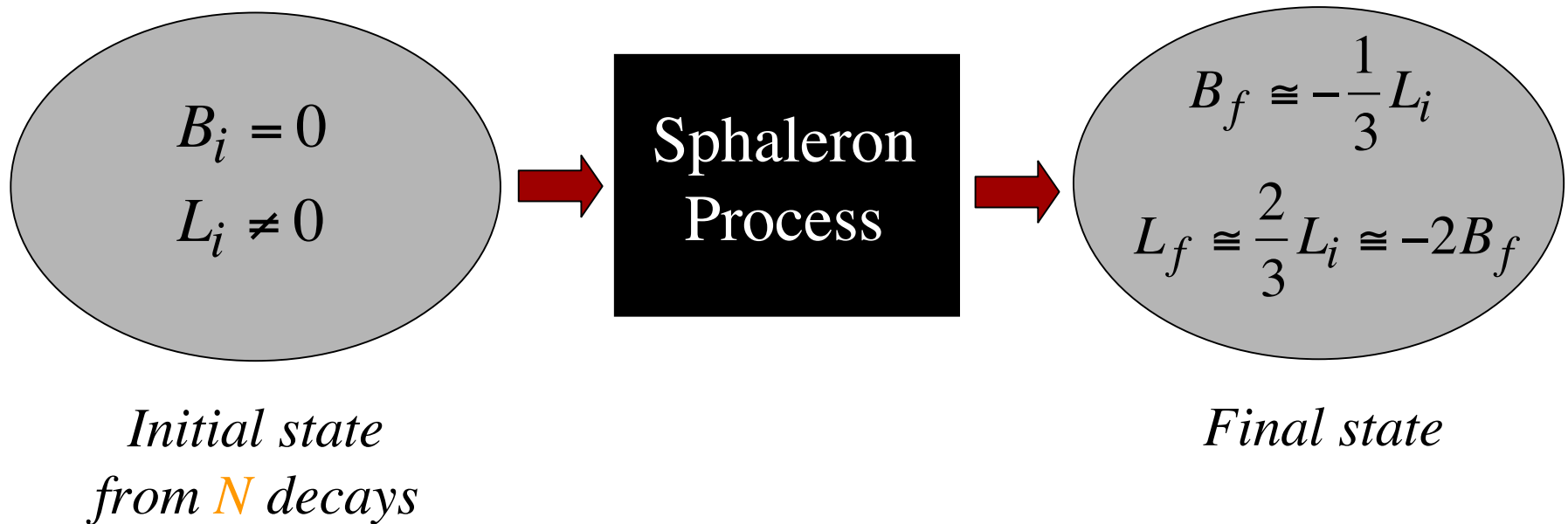
Then, in the early universe, we would have had different rates for the CP-mirror-image decays –



This produces a universe with unequal numbers of leptons and antileptons.

Leptogenesis — Step 2

The Standard-Model *Sphaleron* process, which does not conserve Baryon Number B , or Lepton Number L , but does conserve $B - L$, acts.



There is now a Baryon Asymmetry.

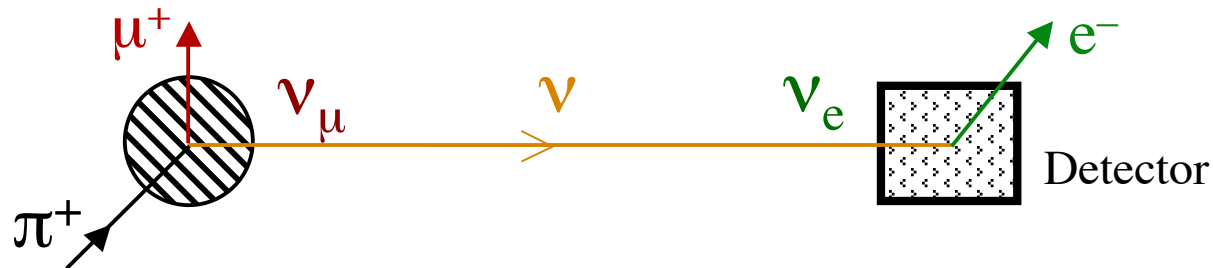
Evidence for the See-Saw and for Leptogenesis

The observation of $0\nu\beta\beta$ would be evidence in favor of the *See-Saw*.

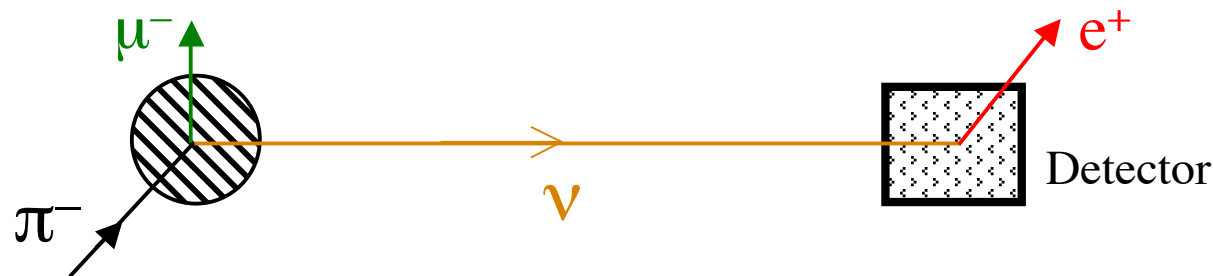
The observation of ~~CP~~ in neutrino oscillation would be evidence in favor of *Leptogenesis*.

Does Neutrino Oscillation Violate CP?

Compare

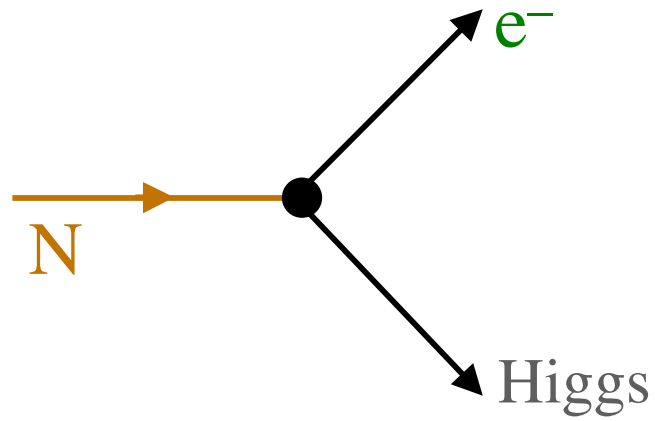


with

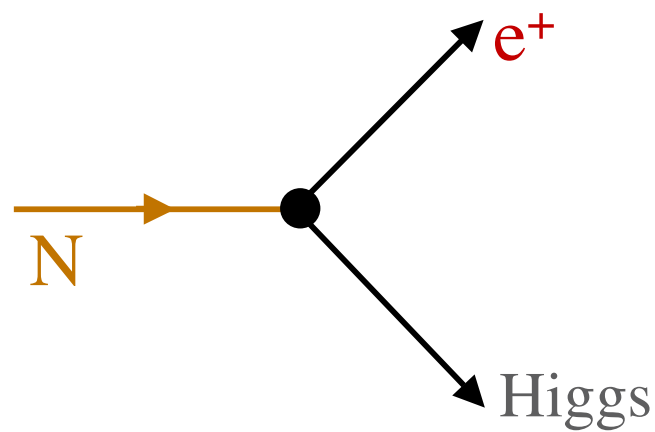


Do these two processes have different rates? (CP)

This is today's version of comparing —



with —



Physicists in Europe, Japan,
and the U.S. are working hard
to design an experimental facility
that can make this comparison,
or a related one.

Young-Kee Kim

Summary

*There are some very
interesting questions
to answer.*

Exciting times lie ahead.
