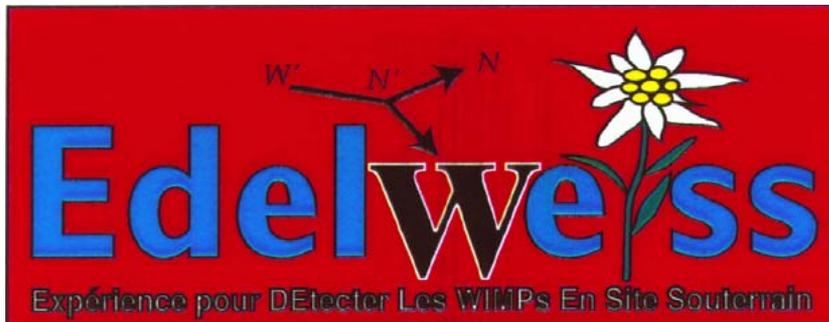


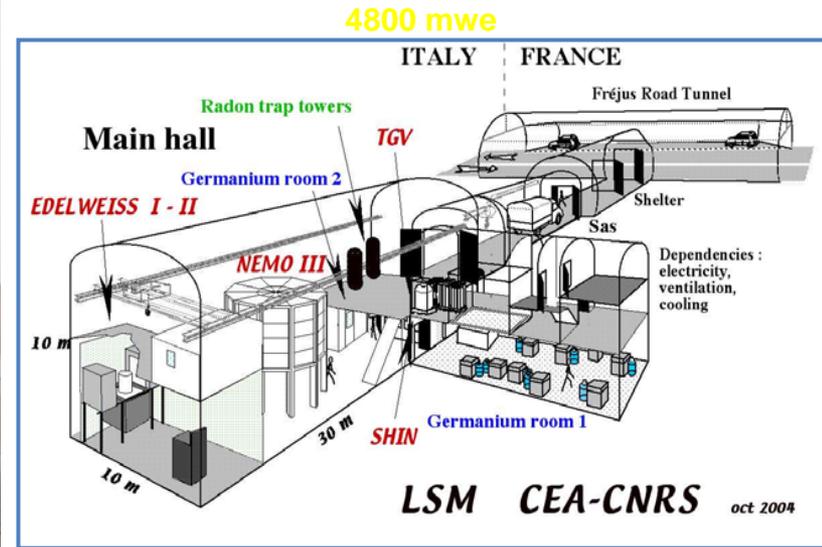
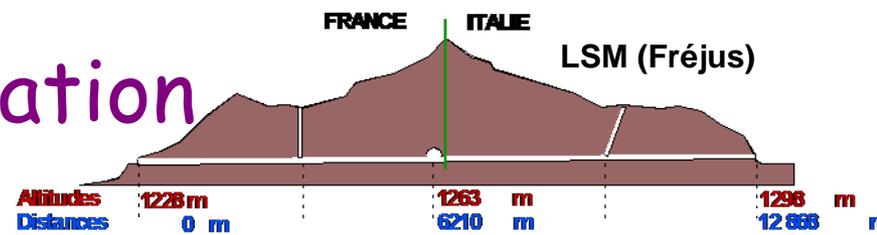
Results of the EDELWEISS-II WIMP dark matter direct search experiment



XXIInd RENCONTRES DE BLOIS
Château de Blois, July 15th-20th 2010
Corinne AUGIER, IPNL
UCBL Lyon 1 - IN2P3/CNRS



The EDELWEISS collaboration

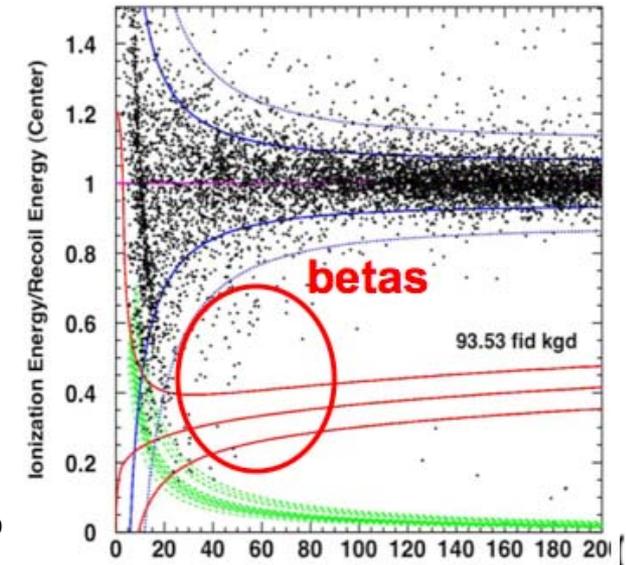
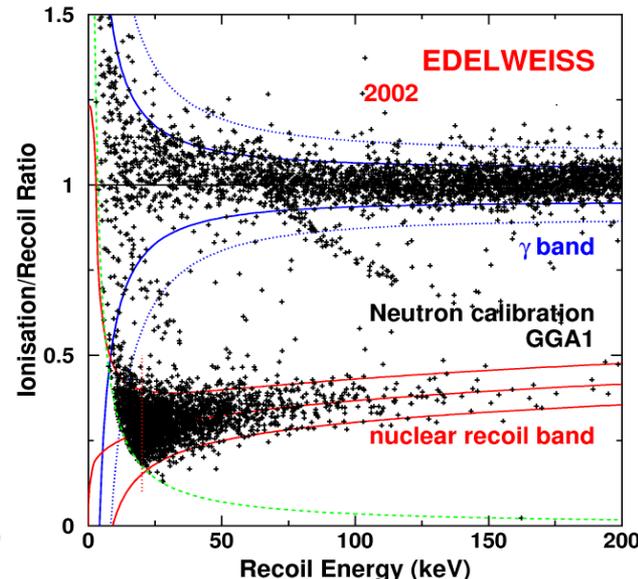
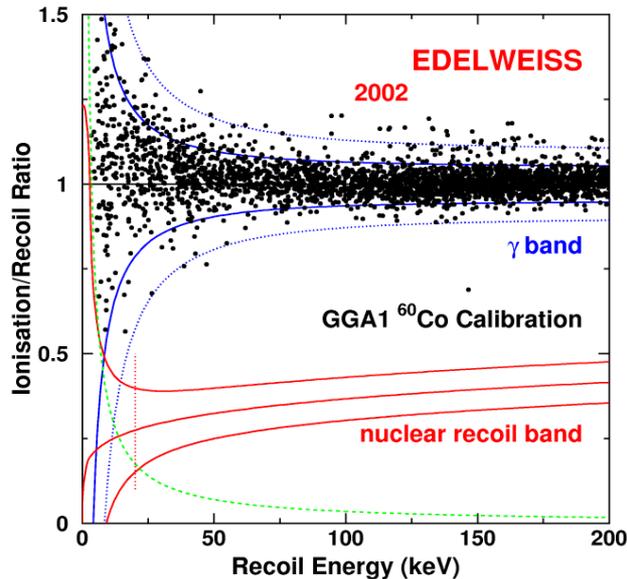
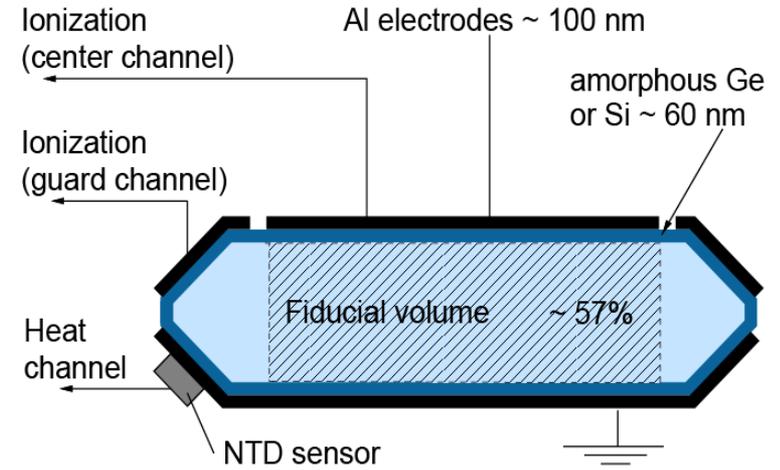


- CEA Saclay (IRFU & IRAMIS) Detectors, electronics, acquisition, data handling, analysis
- CSNSM Orsay Detectors, cabling, cryogenics
- IPN Lyon Electronics, cabling, low radioactivity, analysis, detectors
- Institut Néel Grenoble Cryogenics, electronics
- Karlsruhe IK, IEKP (+ IPE 2011) Vetos, neutron detectors, background, electronics
- JINR Dubna Background, neutron and radon detectors
- Oxford Univ. New comer 2009: Detectors, cabling, cryogenics, analysis
- Sheffield Univ. New comer 2010: MC simulation

EDELWEISS-I detectors

Direct search of WIMP dark matter by energy measurement of detector nuclei recoils induced by the interaction of WIMPs from Milky Way galaxy

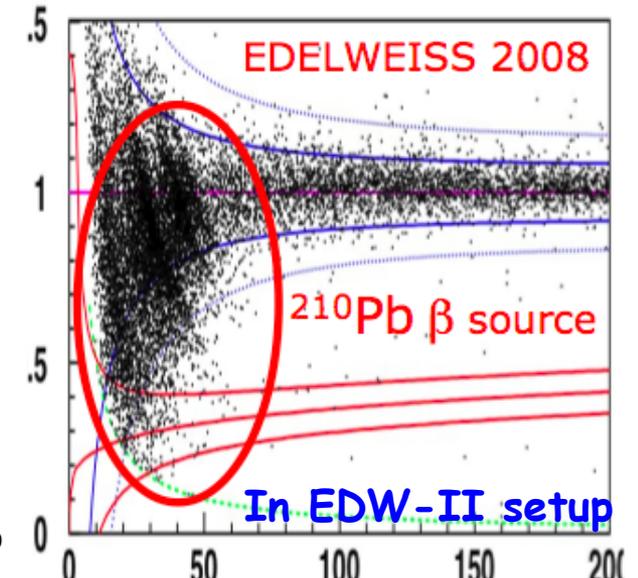
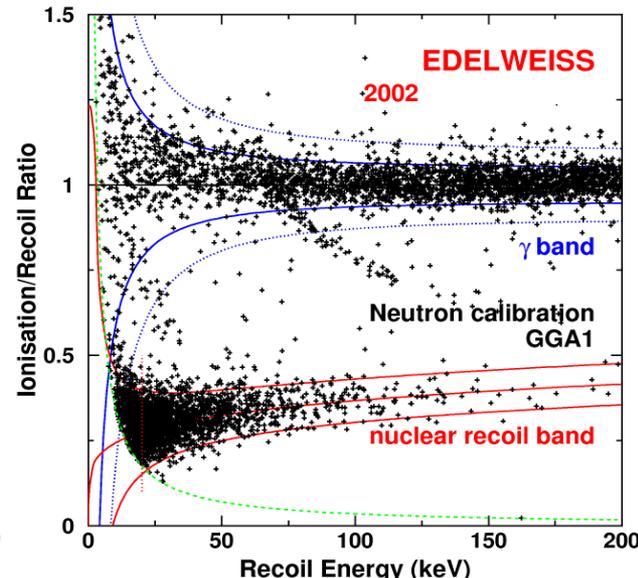
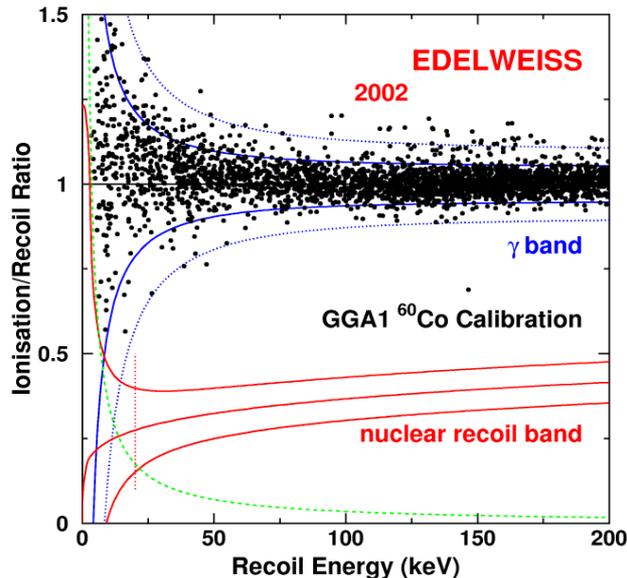
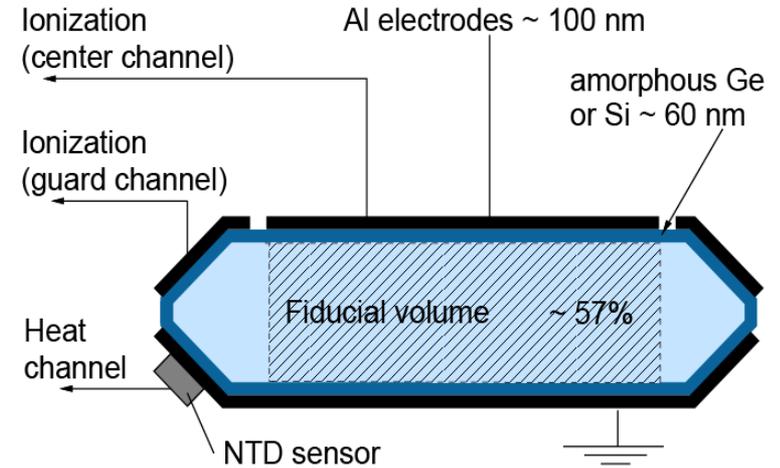
- Germanium crystals bolometers with both ionization and heat measurements → plain aluminum electrodes + NTD sensor at 20 mK
- WIMP and n bkg → nuclear recoils
- e-, γ, α bkg → electron recoils
- Discrimination using electron and nuclear recoils: $E_R = f(E_I, E_H)$ and $Q = E_I / E_R$
- Limitation: surface interactions



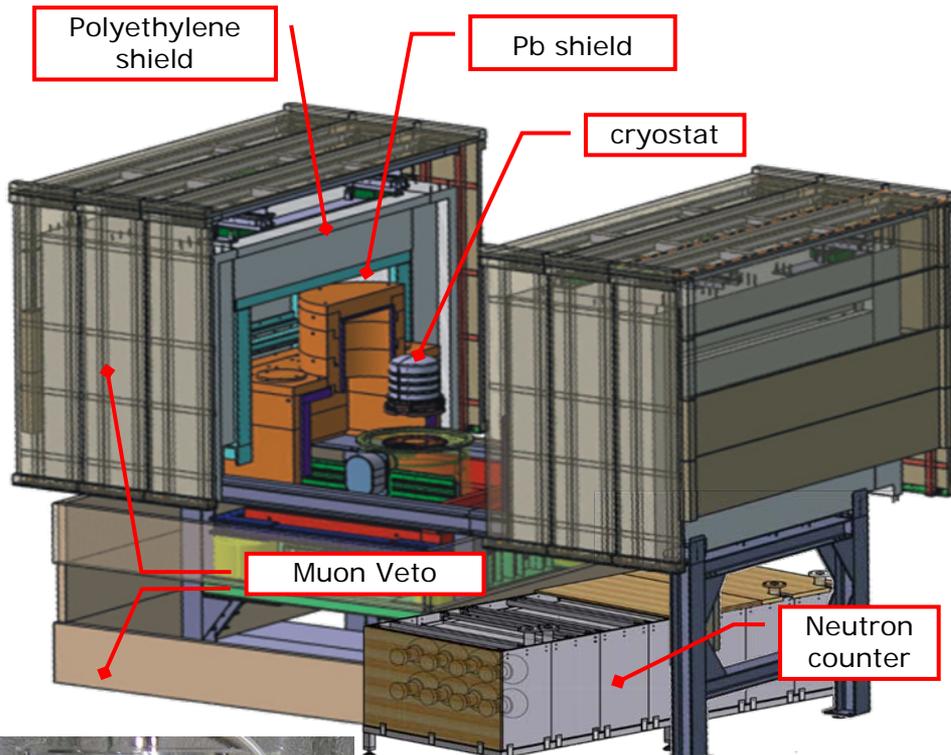
EDELWEISS-I detectors

Direct search of WIMP dark matter by energy measurement of detector nuclei recoils induced by the interaction of WIMPs from Milky Way galaxy

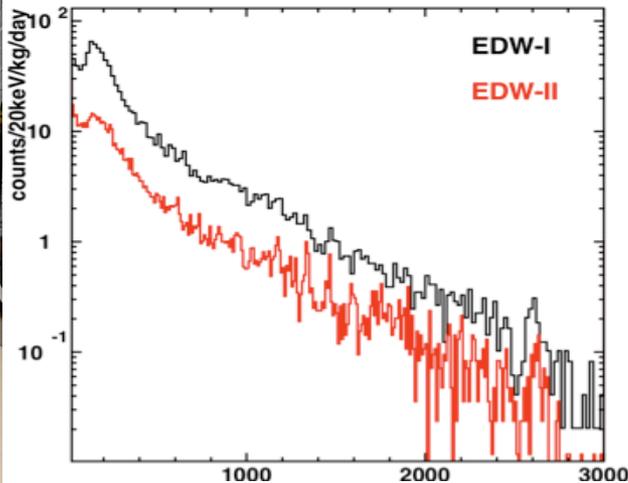
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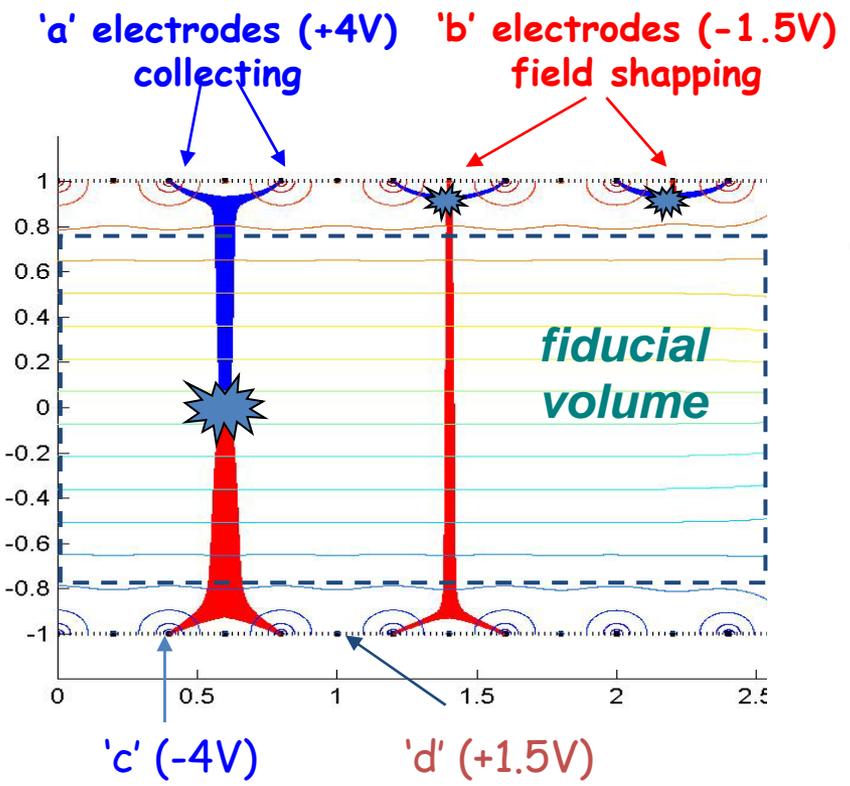
The Edelweiss-II setup



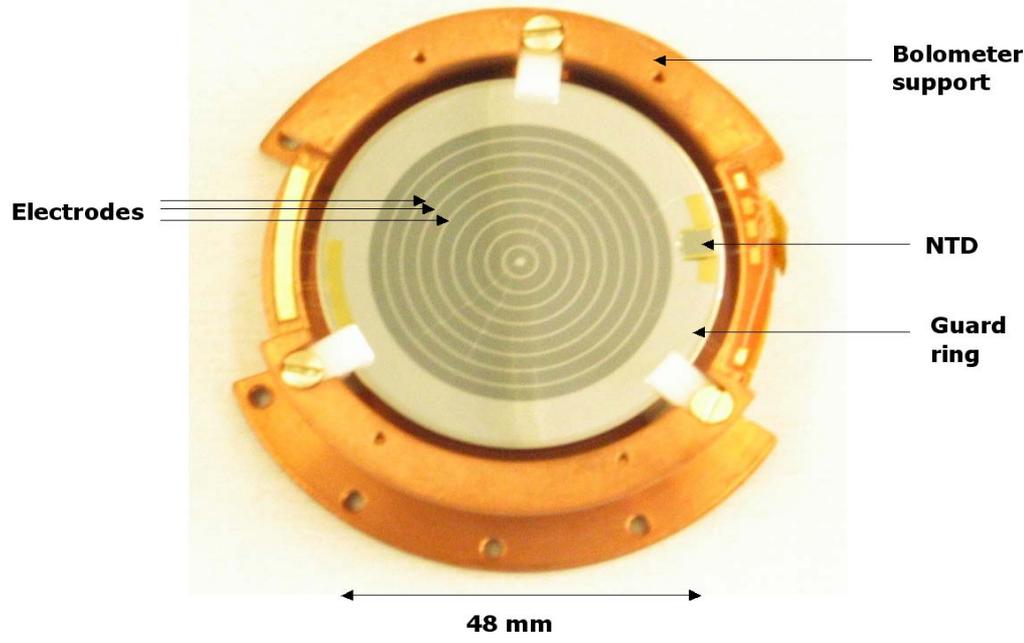
- Operated at the Fréjus Underground Laboratory in Modane (LSM, $4\mu/\text{day}/\text{m}^2$)
- **New goal 5×10^{-9} pb**
- Cryogenic installation (at 18 mK):
 - ◆ Reversed geometry cryostat, pulse tubes
 - ◆ Remotely controlled
 - ◆ **Can host up to 40 kg of detectors**
- Shieldings:
 - ◆ Clean room + deradonized air
 - ◆ Active muon veto (> 98% coverage)
 - ◆ PE shield 50 cm
 - ◆ Lead shield 20 cm
 - **γ background reduced by ~3 wrt EDW-I**
- Others:
 - ◆ Remotely controlled sources for calibrations + regenerations
 - ◆ Radiopurity selection of materials (HPGe)
 - ◆ Detector storage & repair within the clean room
 - ◆ **Radon detector down to few mBq/m³**
 - ◆ **He3 neutron detector (thermal neutron monitoring)**
 - ◆ **Liquid scintillator 1 T neutron counter (study of muon induced neutrons)**
- 12 cool-downs already operated since 2006



Rejecting surface events with interleaved electrodes



« ID » (InterDigit) detector



First prototype 200g built in 2007
 1x200g + 3x400g tested in 2008
 10x400g in run 12 since beginning 2009

- Modify the E field near the surfaces with interleaved electrodes
- Use 'b' and 'd' signals as vetos against surface events
- Rejection factor $> 10^4$ for low-energy surface events induced by e^- from ^{210}Pb
- Heat: keep the EDW-I NTD sensor

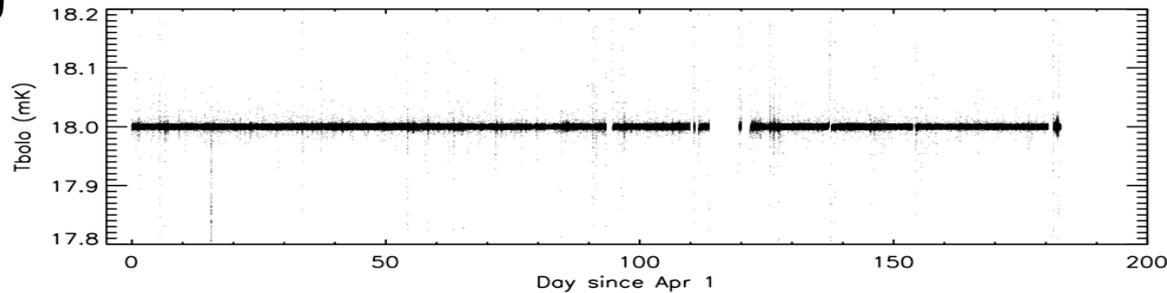
PLB 681 (2009) 305-309 [arXiv:0905.0753]

WIMP search with ID detectors: Run 12

Run 12 from April 1st 2009 to May 20th 2010

- 418 d total
- 322 d data (77% of 418)
- 305 d WIMP physics (73% of 418)
- All the 10 detectors working
- 90% electronics channels ok
- 9/10 bolos for Physics
- 8 d gamma calib
- 5 d neutron calib
- 4,5 d « other »

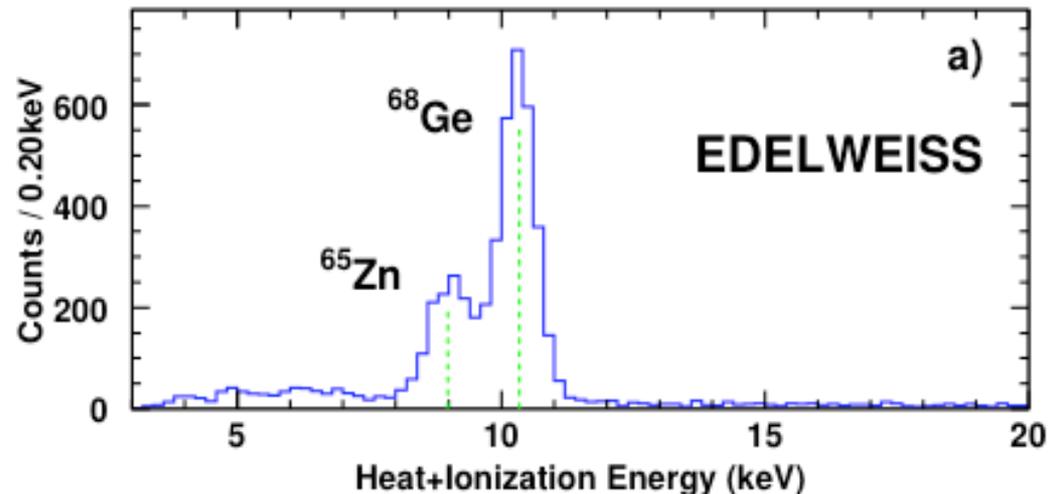
« One of the coldest place in the Universe » ...
Continuously at 18 mK during more
than 1 year !



Baseline resolutions: Heat: 0.6 to 2 keV, average = 1.2 keV
Ionization: 0.8 to 1.2 keV, average = 0.9 keV

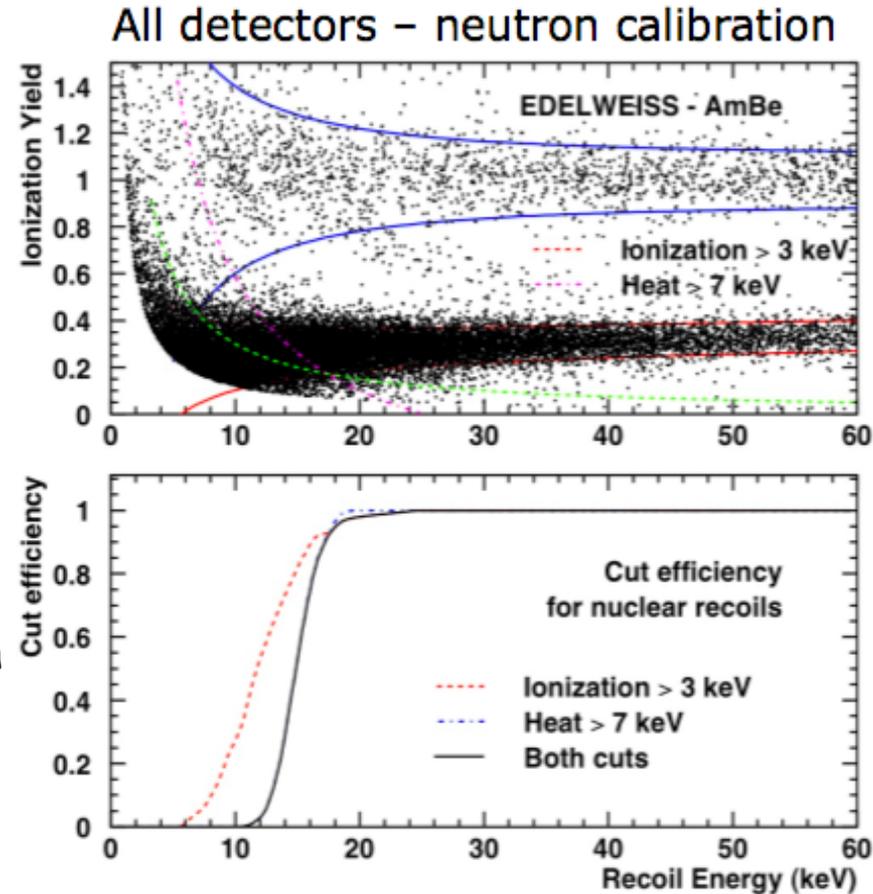
- γ -ray energy of fiducial events
- all detectors used in WIMP search
- online heat threshold 3.5 keV
- 8.98 and 10.34 keV peaks associated with e.c. decay of cosmogenically activated ⁶⁵Zn and ⁶⁸Ge.

→ 160 g fiducial mass for the 10 ID400 detectors.



Data analysis of first 6 months

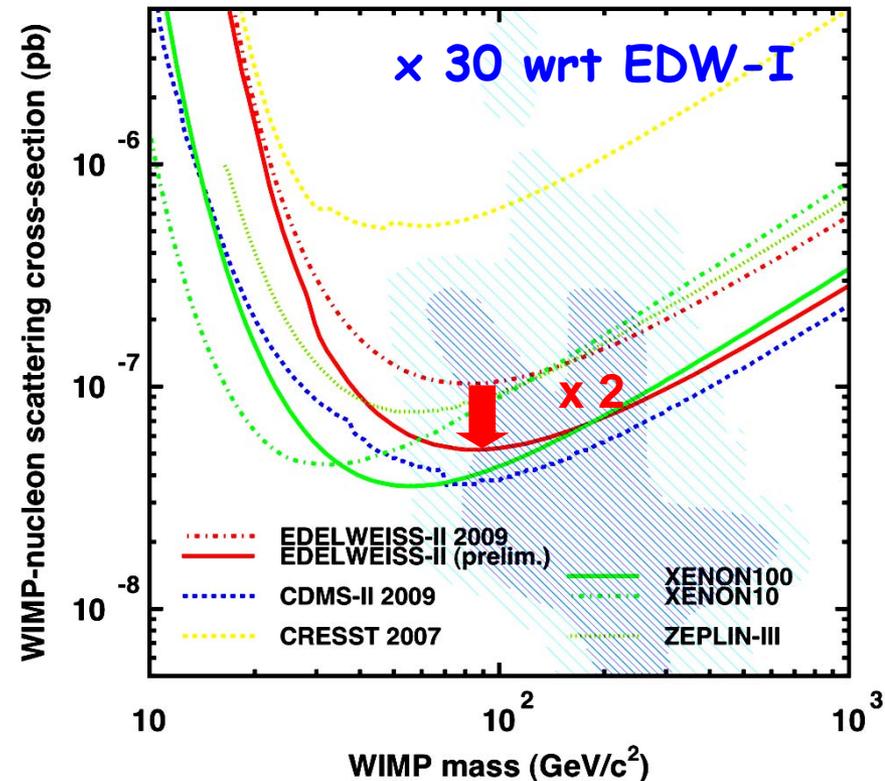
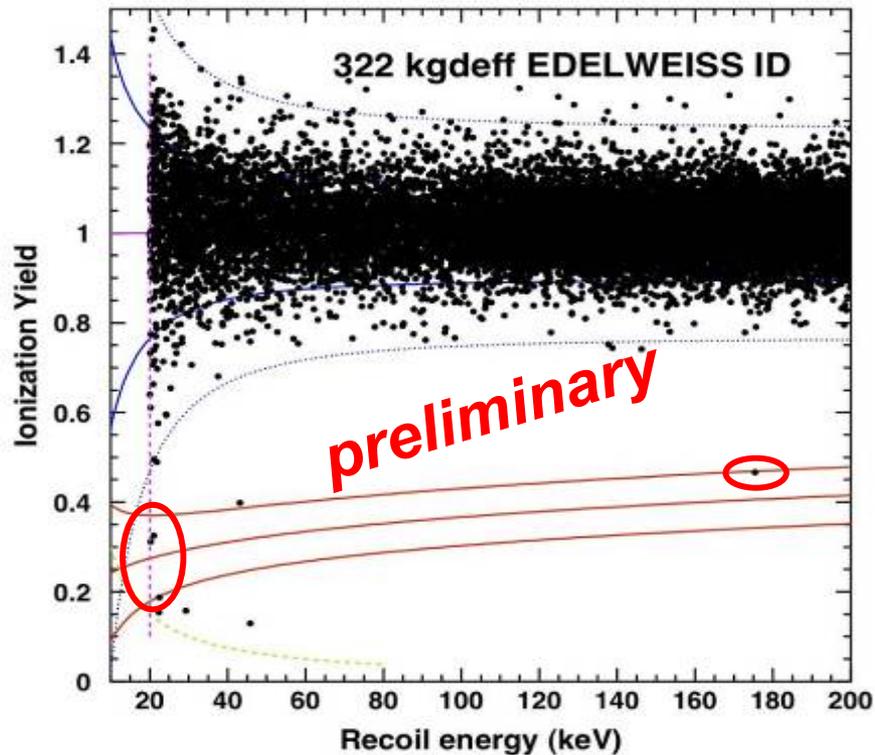
- 2 independent processing pipelines
- Pulse fits with optimal filtering using instantaneous noise spectra
- Period selection based on baseline noises
80% efficiency
- Pulse reconstruction quality (chi2)
97% efficiency
- **Fiducial cuts** based on ionization signals (160g \equiv 40%)
- **90% nuclear recoil acceptance**
- **99.99% gamma rejection**
- Bolo-bolo & bolo-veto coincidence rejection
- WIMP search threshold fixed a priori
Er > 20 keV (100 % acceptance)
- Agreement between the results of the two analyses



Effective exposure 144 kg.d - One nuclear recoil candidate above 20 keV

Best limit: $\sigma_{SI}(W-N) = 1.0 \times 10^{-7}$ pb at $M_W = 80$ GeV (90%CL) **x15 wrt EDW-I**

WIMP search: last result (end of run May 20th 2010)



Preliminary result : 1st analysis with same cuts as first 6 months,
2nd analysis ongoing \Rightarrow Increase in the sensitivity of factor 2 (scale with stat)

3 evts near threshold + 1 evt at 175 keV in the NRB x 30 wrt EDW-I

Best limit: $\sigma_{SI}(W-N) = 5.0 \times 10^{-8}$ pb at $M_W = 80$ GeV (90%CL)

\rightarrow background starts to appear?

Background studies

Gammas:

^{133}Ba calib rejection \times observed bulk γ < 1.0

Betas:

β source rejection \times observed surface evts < 0.2

Neutrons from μ 's:

μ veto efficiency \times observed muons < 0.25

Neutrons from Pb:

measured U limits \times Monte Carlo simu < 0.1

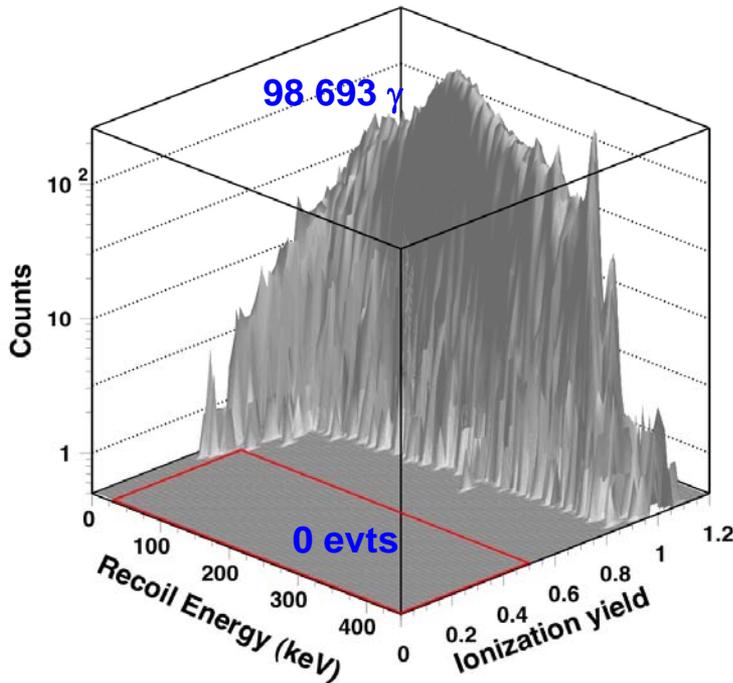
Neutrons from rock:

measured neutron flux \times Monte Carlo simu < 0.1
MC tuned with outside strong AmBe source

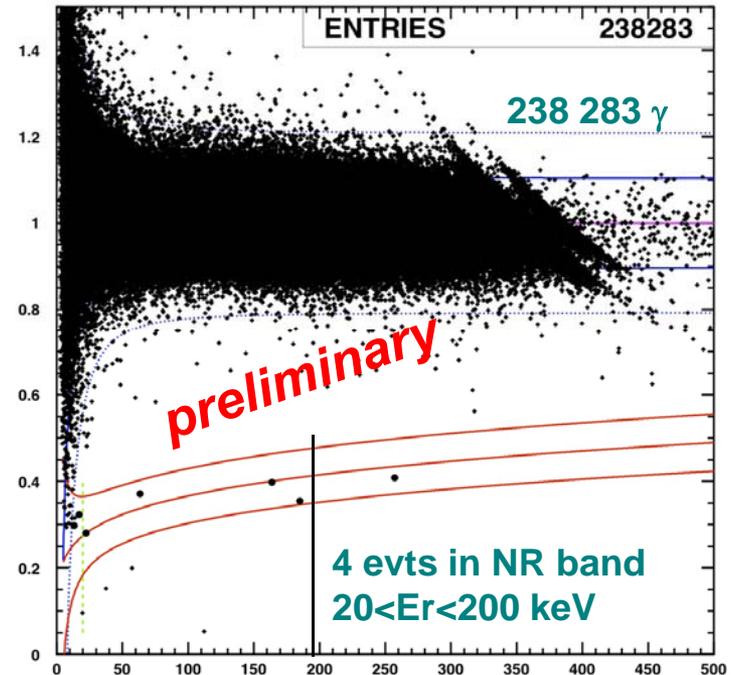
SUM < 1.6 for the whole WIMP run (90% CL upper limit)

Status of gamma ^{133}Ba calibrations

EDELWEISS - ^{133}Ba calibration (98693 γ)



With 2 detectors: gaussian behaviour
No candidate event for $Q < 0.5$

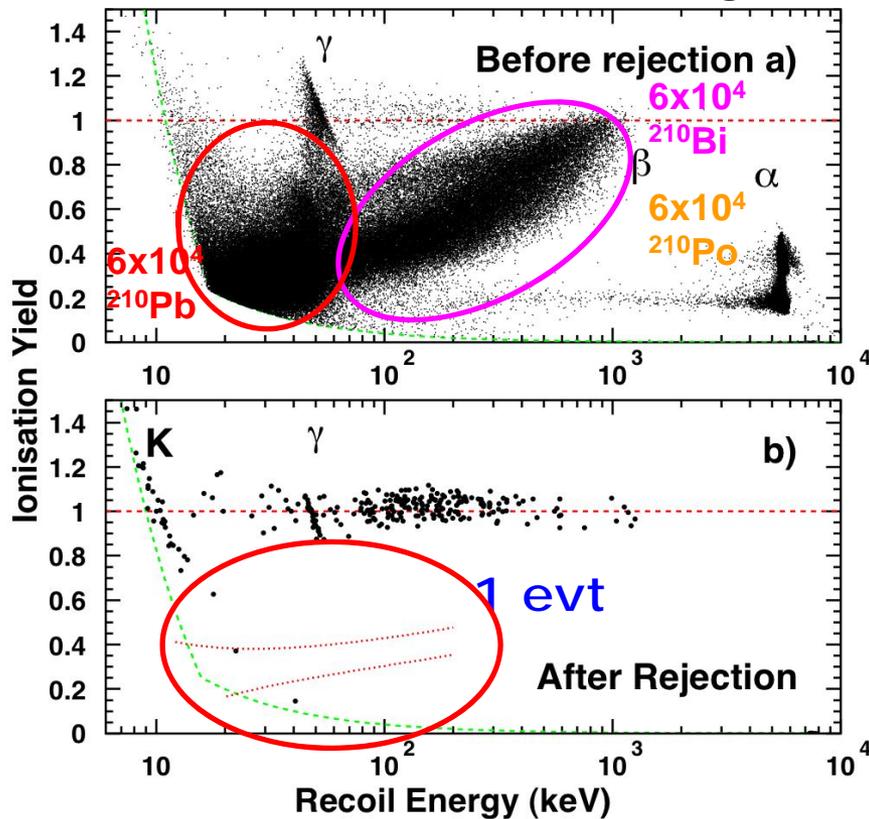


Statistics $\times 2.5$, all 10 detectors
 \Rightarrow 4 evts in the NRB

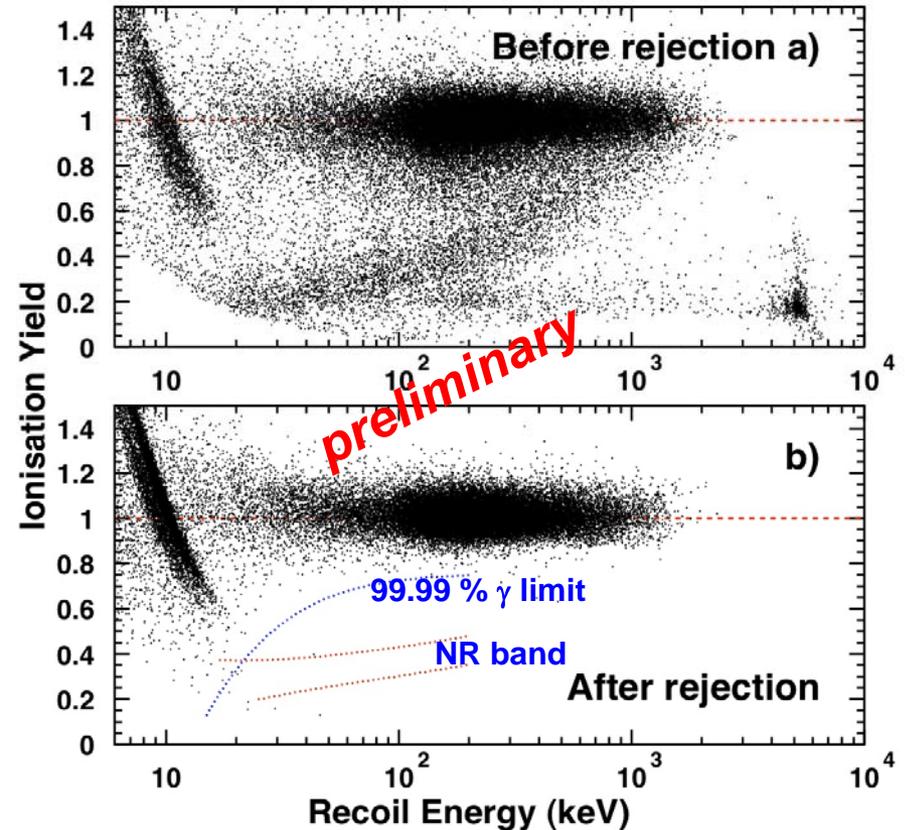
- ^{133}Ba calib: 134 000 evts in 20-200 keV \Rightarrow < 1 evt (90%CL) expected in 16 600 evts WIMP run
- Knobs to understand/improve:
 - Recombination e-h : optimize operation of polarization voltages, regeneration procedures
 - Pile up, multisite events : new fast readouts on heat and ionization
 - 2 NTD heat measurements for information redundancy

Beta calibrations & backgrounds

^{210}Pb calibration - 1st 200g ID



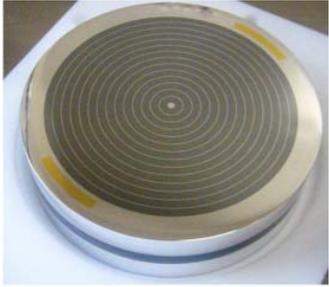
Data for WIMP search - Run 12



PLB 681 (2009) 305-309 [arXiv:0905.0753]

- Rejection factor + identified surface events in data
 - < 0.2 evt (90%CL) expected after rejection
- Knobs to improve:
 - change surface treatment
 - better E resolutions

New run: from IDs to FIDs

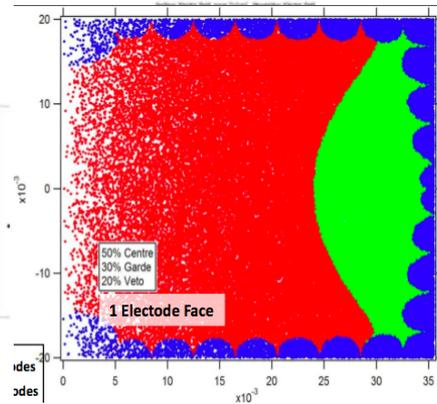
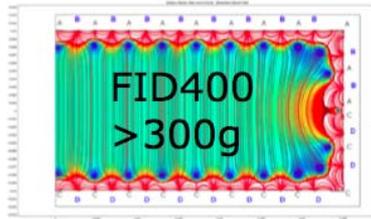
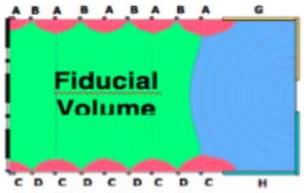


ID401 to 405:
Φ 70mm, H 20mm, 410g

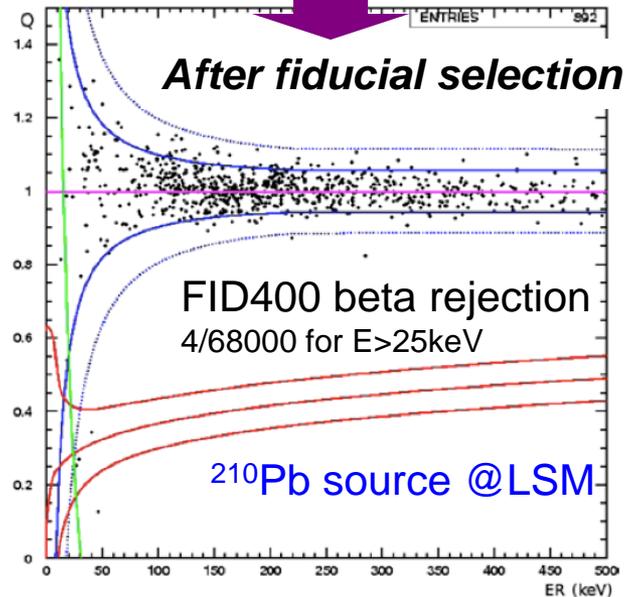
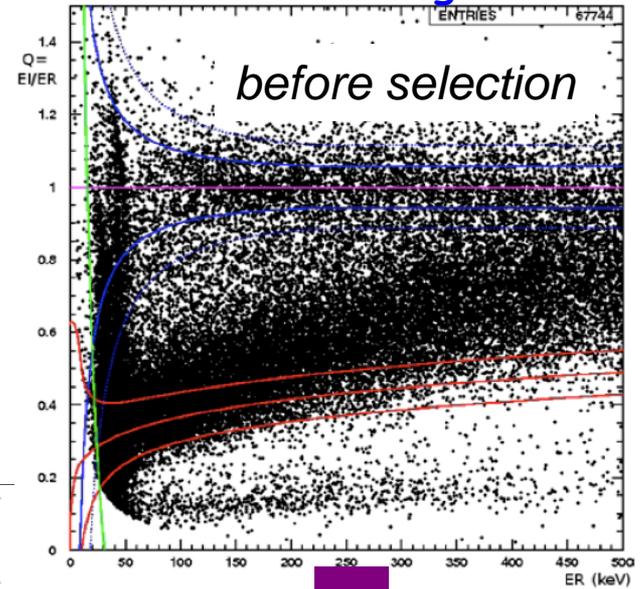
ID2 to ID5:
Φ 70mm, H 20mm, 410g

FID401 and FID402:
Φ 70mm, H 20mm, 410g

ID400
160g



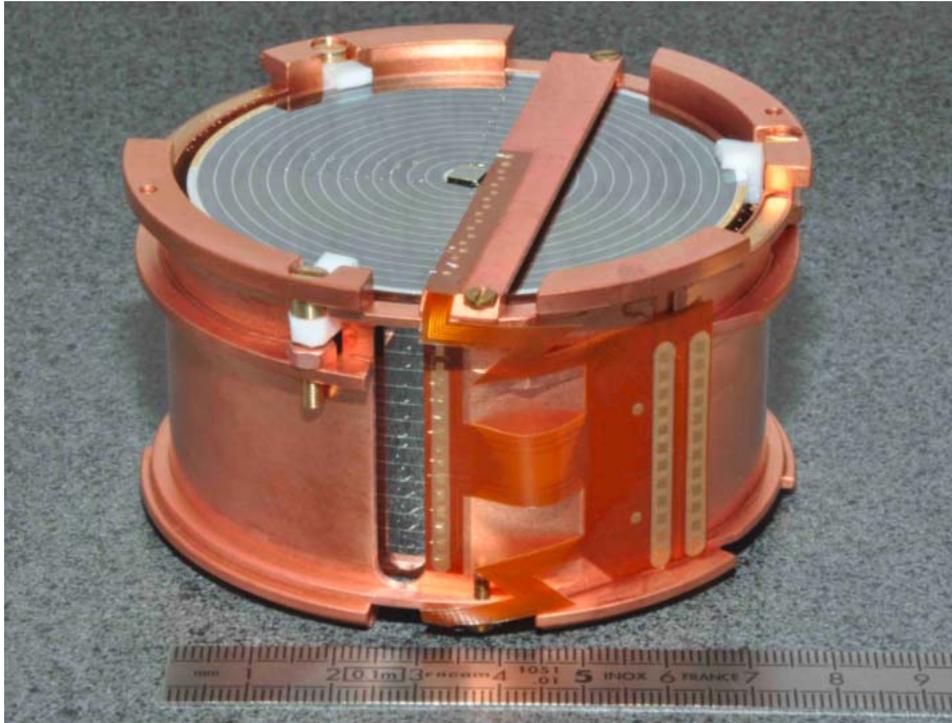
Run 12 - 400 g FID



Fiducial mass **x2** **x4**
ID200-ID400 ⇒ **FID400** ⇒ **FID800**

- Production of FID detectors performed at CSNSM Orsay using dedicated evaporation system

July 3rd: 4 FID800 installed in LSM cryostat Cool down of run 13 before end of July



FID800: 800 g « ID » detector, 2 NTD
6 ionization electrodes \Rightarrow detector segmentation
218 ultrasonics bondings / detector

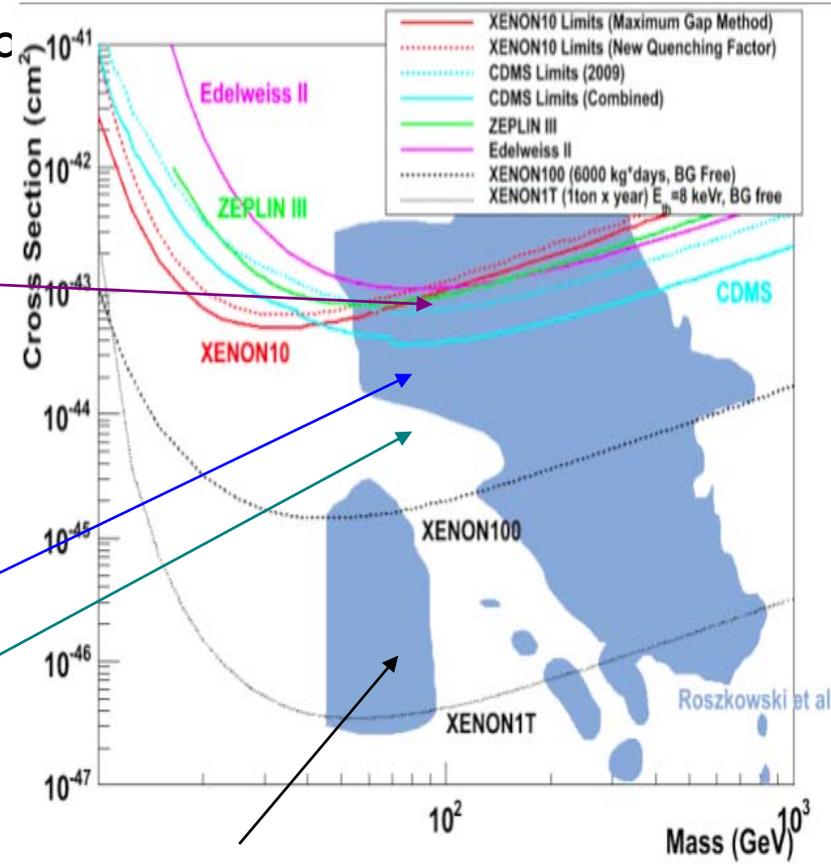
EDELWEISS: summary/prospects

- EDW new-generation ID detectors:

- Robust detectors with redundancy and very high beta rejection
- Preliminary analysis of 1 year data \Rightarrow
 - No hint of WIMPs
 - 5×10^{-8} pb sensitivity achieved

- New Goal: 5×10^{-9} pb

- Improvements wrt future bkg
 - Increased redundancy for ionization and heat measurements
 - Fast readout (multi-site evts, pile up)
 - Lower μ phonics, internal PE shield
- New prototypes: FID800
 - 2011: 1000 kg.d
- Build 40 detectors, upgrade setup
 - 2012: 3000 kg.d



EURECA