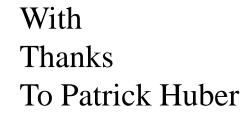
## Long-Baseline Neutrino Oscillation Phenomenology



Deborah Harris Fermilab NuFact09 IIT July 22, 2009



According to Wikipedia:

Particle physics phenomenology is the part of theoretical particle physics that deals with the application of theory to high energy particle physics experiments. Within the Standard Model, phenomenology is the calculating of detailed predictions for experiments, usually at high precision (e.g., including radiative corrections).

## Outline

- How to measure oscillation probabilities
- Experimental Regimes
  - Conventional Neutrino Beams
    - v<sub>e</sub> appearance
  - Beta Beams
    - $v_{\mu}$  appearance
  - Neutrino Factories
    - $v_{\tau}$  and  $v_{\mu}$  appearance
- Example from MINOS:  $v_e$  appearance
- "Phenomenology": making predictions for future
- What else will be needed in the future

Probabilities



$$N_{far} = \phi_{\nu_{\mu}} \sigma_{\nu_{x}} P(\nu_{\mu} \to \nu_{x}) \mathcal{E}_{x} M_{far} + B_{far}$$

 $\phi$ =flux,  $\sigma$ = cross section  $\epsilon$ =efficiency M=mass

$$P(\nu_{\mu} \rightarrow \nu_{x}) = \frac{N_{far} - B_{far}}{\phi_{\nu_{\mu}} \sigma_{\nu_{x}} \varepsilon_{x} M_{far}}$$

**B**<sub>far</sub>= Backgrounds at far detector, from any flux

$$B_{far} = \sum_{i=\mu,e} \phi_{v_i} (P) \sigma_{v_i} \varepsilon_{ix} M_{far}$$

NuINT matters for Signal and Background Cross sections, and indirectly for efficiencies!

### Probabilities, continued



$$\left(\frac{\delta P}{P}\right)^{2} = \frac{\left(N_{far} + \left(\delta B_{far}\right)^{2}\right)}{\left(\phi_{\nu_{\mu}}\sigma_{\nu_{x}}\varepsilon_{x}M_{far}\right)^{2}} + \frac{N_{far} - B_{far}}{\left(\phi_{\nu_{\mu}}\sigma_{\nu_{x}}\varepsilon_{x}\right)^{2}} \left[\delta(\phi_{\nu_{\mu}}\sigma_{\nu_{x}}\varepsilon_{x})\right]^{2}$$

$$\left(\frac{\delta P}{P}\right)^{2} = \frac{\left(N_{far} + \left(\delta B_{far}\right)^{2}\right)}{\left(\phi_{\nu_{\mu}}\sigma_{\nu_{x}}\varepsilon_{x}M_{far}\right)^{2}} + \left(N_{far} - B_{far}\right)\left[\left[\frac{\delta \phi_{\nu_{\mu}}}{\phi_{\nu_{\mu}}}\right]^{2} + \left(\frac{\delta \sigma_{\nu_{x}}}{\sigma_{\nu_{x}}}\right)^{2} + \left(\frac{\delta \varepsilon_{\nu_{x}}}{\varepsilon_{\nu_{x}}}\right)^{2}\right]$$

2 Regimes:

$$N_{far} >> B_{far}$$

$$N_{far} \approx B_{far}$$

### **Problem:**

Don't always know *a priori* which regime you are in ---depends on  $\Delta m^2$ , ---depends on  $\sin^2 2\theta_{13}$ 

### Near Detector Strategy



$$B_{far} = \sum_{i=\mu,e} \phi_{v_i far}(P) \sigma_{v_i} \varepsilon_{ix} M_{far}$$

Backgrounds come from several sources

$$N_{near} = \sum_{i=\mu,e} \phi_{v_i near} \sigma_{v_i} \varepsilon_{ix} M_{near}$$

Build near detector with same  $\varepsilon$ 

$$B_{far} = N_{near} \frac{\sum_{i=\mu,e} \phi_{v_i far}(P) \sigma_{v_i} \varepsilon_{ix} M_{far}}{\sum_{i=\mu,e} \phi_{v_i near} \sigma_{v_i} \varepsilon_{ix} M_{near}}$$

Simulations better at predicting ratios absolute levels

$$B_{far} = \sum_{i=\mu,e} N_{near,i} \frac{\phi_{v_i far}}{\phi_{v_i near}} \frac{\sigma_{v_i}}{\sigma_{v_i}} \frac{\varepsilon_{ix}}{\varepsilon_{ix}} \frac{M_{far}}{M_{near}}$$



$$B_{far} = \int dE_{v} \sum_{i=\mu,e} N_{near,i}(E_{v}) \left(\frac{\phi_{v_{i} far}}{\phi_{v_{i} near}}\right) (E_{v}) \left(\frac{\sigma_{v_{i}}}{\sigma_{v_{i}}}\right) (E_{v}) \left(\frac{\varepsilon_{ix}}{\varepsilon_{ix}}\right) (E_{v}) \frac{M_{far}}{M_{near}}$$

- But ratios don't cancel everything
- Underlying problem: fluxes are different
  - Near detector: line source, far detector: point source
  - But even if that is solved, still  $v_{\mu}CC$  oscillations for conventional beam case
- All of these terms are functions of energy
  - Uncertainties in energy dependence of cross sections translate into far detector uncertainties...

**D**. Harris Fermilab NuFact09 How low should Uncertainties be?

- Be careful what you wish for: at high  $\theta_{13}$ , you are looking for small differences in probabilities Neutrino-AntiNeutrino Asymmetry 1.0  $P(\nu_{\mu} \rightarrow \nu_{e}) \approx$  $\delta = \pi/2$  $|\sqrt{P_{atm}}e^{-i(\Delta_{32}+\delta)}+\sqrt{P_{sol}}|^2$ 0.8 Excluded 0.6  $\delta = \pi/6$  $\sqrt{P_{atm}} = \sqrt{P_{sol}}$ 0.4 But small if Choo  $\Delta_{32}=\pi/2$

 $\delta m_{21}^2 = 7.0 \times 10^{-5} \text{ eV}^2$ 0.2  $\delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2$ 10<sup>-3</sup> 10<sup>-2</sup> 0.0  $10^{-1}$  $\sin^2 2\theta_{13}$ 

 $|P-\overline{P}|/|P+\overline{P}|$ 

Graphics courtesy S. Parke

CP Asymmetry is largest when

 $\sqrt{P_{atm}} > \sqrt{P_{sol}}$ 

Asymmetry could be 20% or less with matter effects  $\rightarrow$ Systematics on difference < 7% !

### Systematic Uncertainties

- Neutrino Flux
  - Hadron Production
    - $\pi/K$  ratio
    - x and  $p_t$  spectrum of produced Pions/Kaons
  - Pion, Muon or Isotope Beam characteristics
    - Focusing uncertainties
    - Alignment Uncertainties
- Neutrino Interactions: Background and Signal!
  - Quasi-elastic Uncertainties
  - Resonance (low W) Uncertainties
  - DIS (high W)
  - Nuclear Effects
- Event Selection
- Event Energy Resolution
  - Important especially for measurements versus neutrino energy
  - Narrow Band beams: energy resolution is key to background rejection

Problem: uncertainties all affect the near and far detector both, you can't always separate one from the other

## Near Detector Design



- Far detector must be massive: the more instrumented it is, the more \$/kton...
- Tradeoff between segmentation and far detector mass
- Near Detector Design options:
  - "Identical" to far detector
    - Argue that detector efficiencies and cross sections are the "same", you just need independent flux measurements
    - Can't really be identical: different rates, different size detector
    - Also doesn't see the same flux near as far (line source, osc.)
  - Much more segmented and fine-grained
    - Try to measure fluxes and cross sections as best you can, make far detector prediction
- Ideally, you would do both...

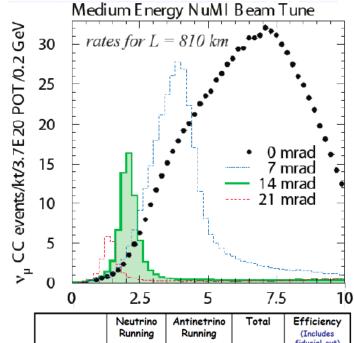


- Next Generation of accelerator searches for  $\theta_{13}$ 
  - T2K and NOvA
  - T2KK and LBNE
  - High Energy Beta Beam
  - Neutrino Factories low and high in energy
- First  $v_e$  Appearance in a Superbeam: MINOS
- Lessons learned
- What the future program will need

#### D. Harris Fermilab NuFact09 The Beamline Options Considered

- Conventional Neutrino Beams
  - Pion focusing and decay in long pipe
  - Mostly  $\nu_{\mu},$  few %  $\nu_{e}$
- Beta Beams
  - Collect radioactive isotope and focus and accelerate
  - Decay in long storage ring
  - All  $v_e$  or all anti-  $v_e$  (at given instant in time)
- Neutrino Factories
  - Collect muons, focus, accelerate
  - Decay in long storage ring
  - $\nu_{\mu}$  and anti-  $\nu_{e}~~or~~\nu_{e}$  and anti-  $\nu_{\mu}$



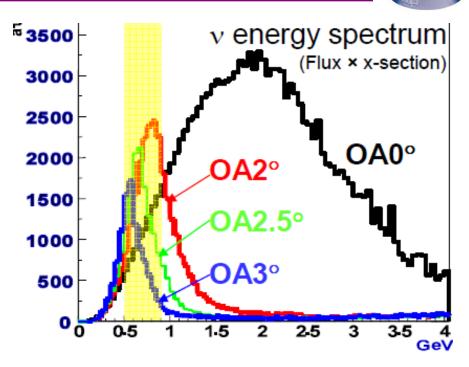


|                     | Running | Running | lotal | (Includes<br>fiducial cut) |
|---------------------|---------|---------|-------|----------------------------|
| $\nu_{e}$ signal    | 75.0    | 29.0    | 104   | 36%                        |
| Backgrounds:        | 14.4    | 7.6     | 22    |                            |
| $\nu_{\mu}$ NC      | 6.0     | 3.6     | 9.6   | 0.23%                      |
| $v_{\mu}$ CC        | 0.05    | 0.48    | 0.53  | 0.004%                     |
| Beam v <sub>e</sub> | 8.4     | 3.4     | 11.8  | 14%                        |
| FOM                 | 19.8    | 10.5    | 22.1  |                            |

Numbers generated assuming:

 $\sin^2(2\theta_{13}) = 0.10$ ,  $\sin^2(2\theta_{23}) = 1.0$ , and  $\Delta m_{32}^2 = 0.0024 \text{ eV}^2$ , no matter effects.

R. Ray, Neutrino 2008



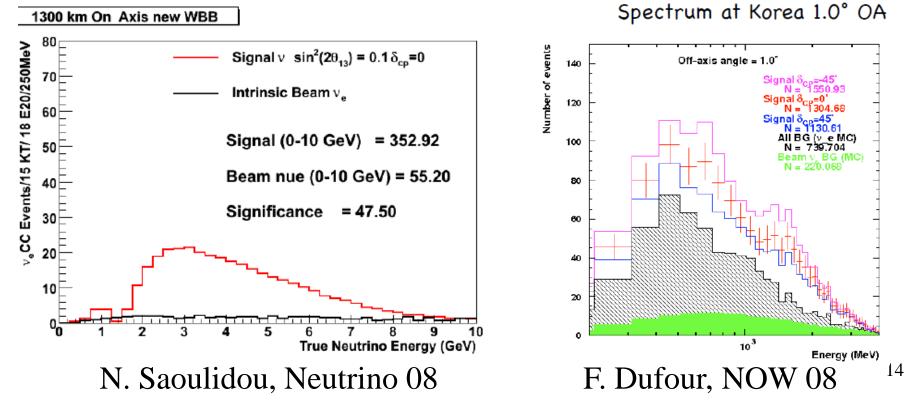
Expected number of events at SK (0.75kW beam x 5yr)

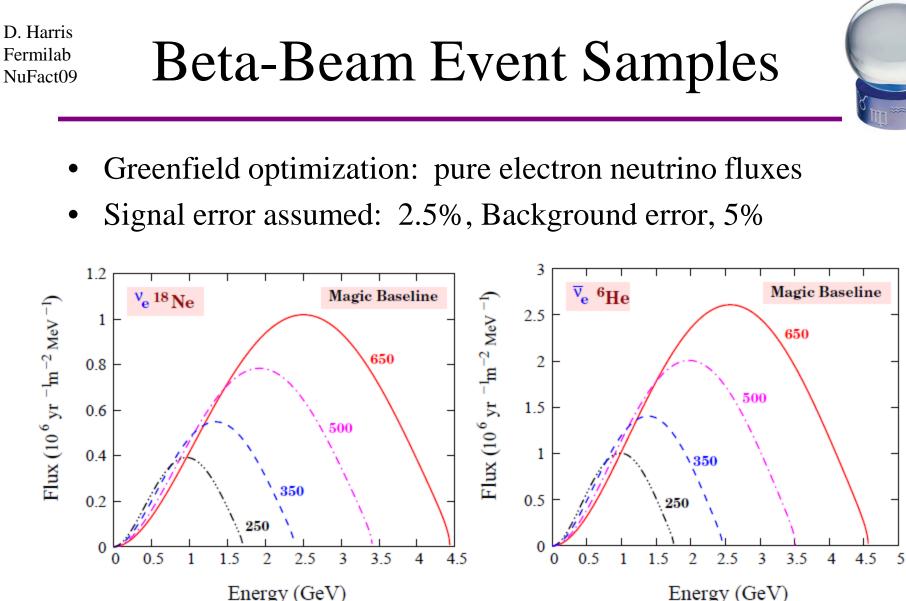
| $sin^22\theta_{13}$ | Backgrounds         |              |       | Signal |
|---------------------|---------------------|--------------|-------|--------|
|                     | $\nu_{\mu}$ induced | Beam $\nu_e$ | Total | Signal |
| 0.1                 | 10                  | 13           | 23    | 103    |
| 0.01                | 10                  |              |       | 10     |

I. Kato, Neutrino 2008

#### D. Harris Fermilab NuFact09 T2KK and LBNE Event Samples

- Water Cerenkov in T2KK, assume Liquid Argon in LBNE (assumed in plot below)
- Few Hundred signal events, 5% systematic uncertainties assumed on signal and background predictions





Energy (GeV)

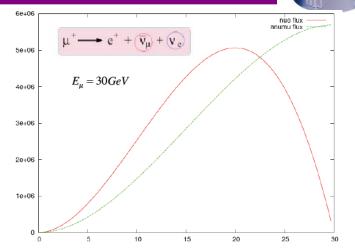
Agarwalla, Choubey, Raychaudhuri, Winter arxiv: 0802.3621 15 Fermilab NuFact09 Neutrino Factory Event Samples

- Gold Channel:  $v_e \rightarrow v_\mu$  in magnetized iron+readout (a la MINOS): MIND (50 kton)
- Silver Channels:  $v_e \rightarrow v_\tau$  in Possible magnetized OPERA detector (10 kton)
- Low energy neutrino factory option: TASD

D. Harris

1kton magnetized OPERA,  $\theta_{13}=5^{\circ}, \delta=90^{\circ}, 732$ km, 10<sup>21</sup> muon decays

Thousands of events, systematics assumed vary from 2 to 5%



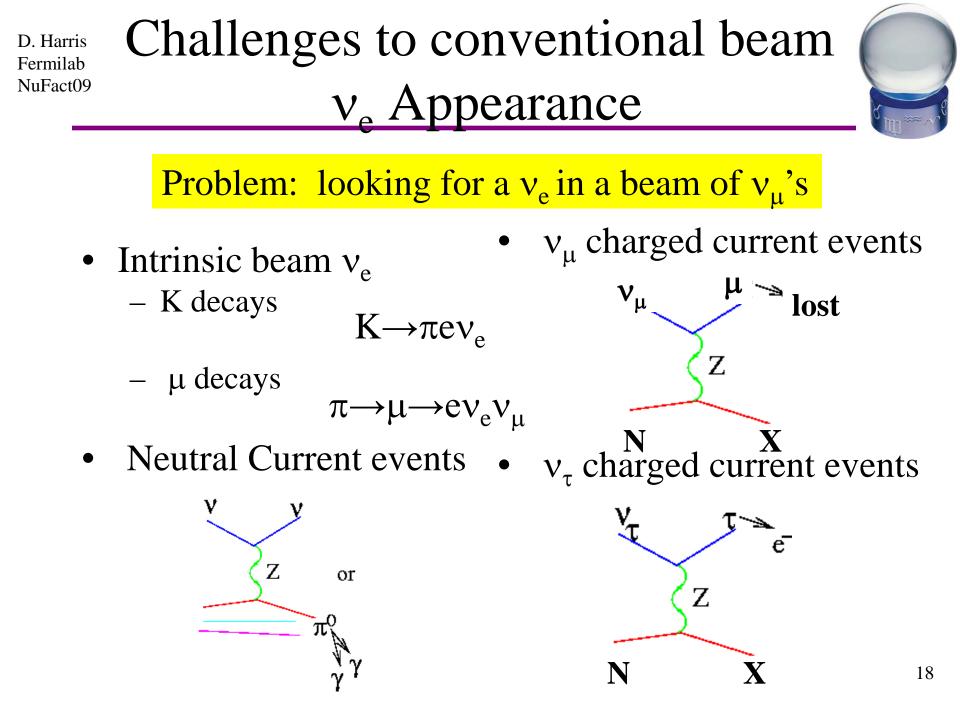
| $\mathcal{P}_{\mu^-}$ | $N_{\mu^{-}}/10^{4}$ | $N_{e^+}/10^4$ | $N_{\mu^+}$ | $N_{e^-}$ | $N_{\tau^+}$ | $N_{	au^-}/10^2$ |
|-----------------------|----------------------|----------------|-------------|-----------|--------------|------------------|
| 0                     | 172                  | 75             | 107         | 186       | 80.7         | 89.9             |
| 0.3                   | 150                  | 97.5           | 140         | 174       | 105          | 81.7             |
| 1                     | 97.5                 | 150            | 215         | 147       | 161          | 64.6             |
| $\mathcal{P}_{\mu^+}$ | $N_{\mu^+}/10^4$     | $N_{e^-}/10^4$ | $N_{\mu^-}$ | $N_{e^+}$ | $N_{	au}$ –  | $N_{	au^+}/10^2$ |
| 0                     | 87.4                 | 148            | 244         | 99        | 151          | 45.2             |
| -0.3                  | 76.1                 | 192            | 317         | 93.4      | 196          | 41.5             |
| -1                    | 49.5                 | 295            | 487         | 79.3      | 302          | 32.8             |

Autiero et al, arXiv: 0305185

16

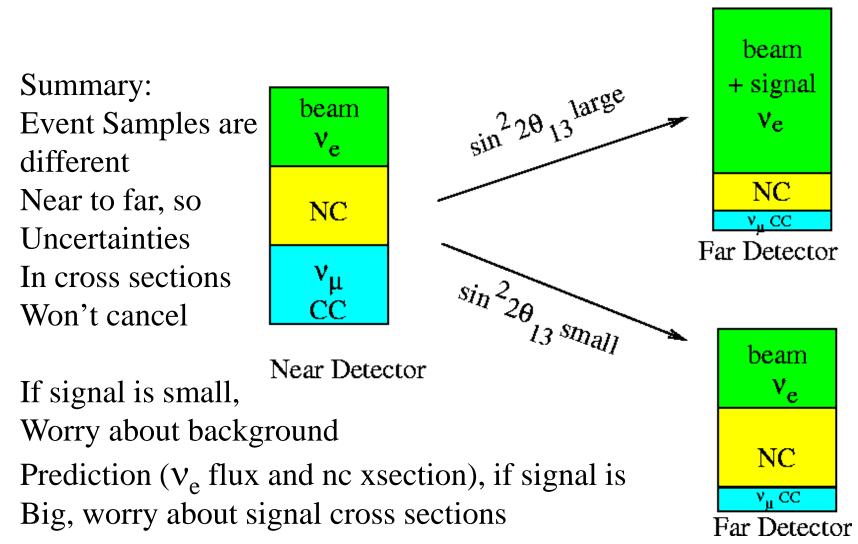
Phenomenology...

- These studies include allowances that there are uncertainties, often categorized as follows:
  - "signal" uncertainties
  - "background" uncertainties
- We need to start looking at how to achieve these 2-5% uncertainties, considering all the factors that go into them
  - Neutrino Flux
  - Detector Efficiencies for Signal and Backgrounds
  - Cross Sections
  - Detector Mass (i.e. Fiducial Volume)
- New Case Study since last NuFact: MINOS search for  $v_{\mu} \rightarrow v_{e}$



## $v_e$ Appearance Analysis



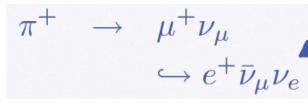


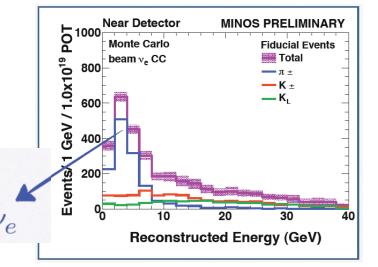
19

D. Harris Fermilab

<sup>NuFact09</sup> First "Superbeam"  $\nu_{\mu} \rightarrow \nu_{e}$ : MINOS

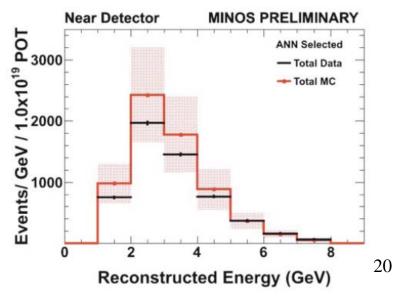
• Electron neutrinos dominated by muon decay

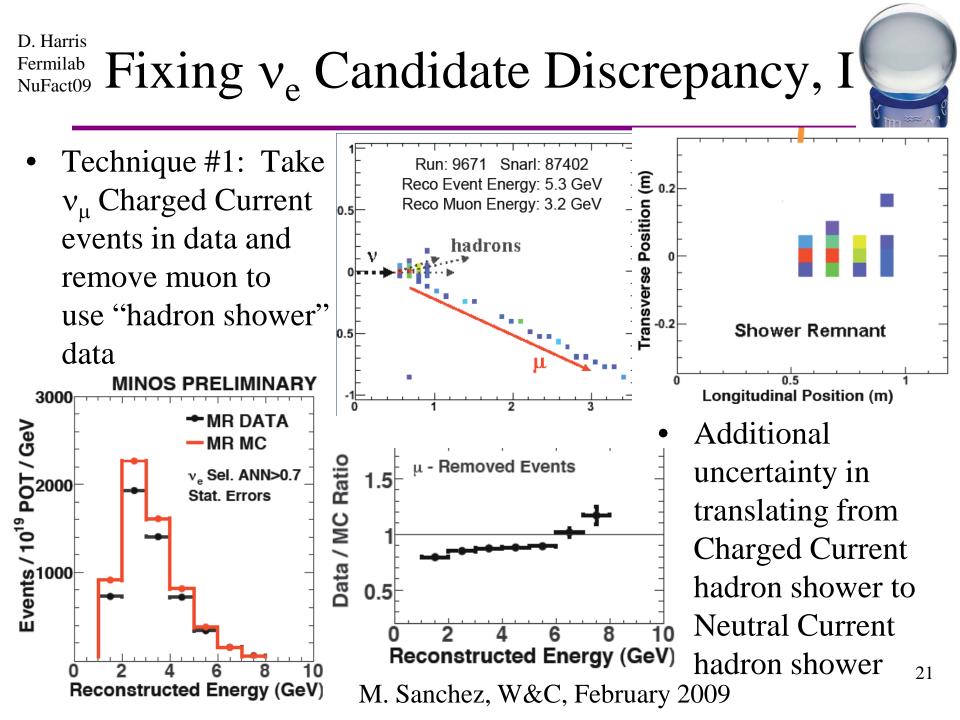




• Problem: when analysis was "done", the near detector data and prediction did not agree

M. Sanchez, W&C, February 2009





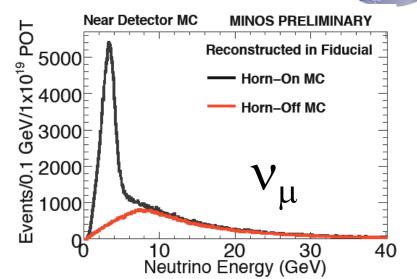
#### D. Harris Fermilab NuFact09 Beta-Beam and Neutrino Factory

- Question: how will we study hadron showers in beta beams and neutrino factory era?
- What if we just wait and fix it then?
  - Beta beam: will be much harder to "remove the electron" in the near detector event samples
  - Neutrino Factory: will be hard because if the signal will be from  $v_e$  to  $v_{\mu}$ , you only have anti- $v_{\mu}$  in the beam
- What can we do instead (and do now!)?
  - Measure hadron showers in fine grained well-understood detectors with good energy resolution and particle ID:
    - At the energies of interest
    - On the same nuclei as the future far detectors
  - Develop better models for effects of nucleus on outgoing hadrons in the interactions, then test them with data

### D. Harris Fermilab NuFact09 Fixing $v_{e}$ Candidate Discrepancy, II

Run beamline with very different incoming flux to change the near detector background fractions

PRELIMINARY Near Detector PRELIMINARY Near Detector Events / 1 GeV /1.0x10<sup>19</sup> POT 1000 GeV /1.0x10<sup>19</sup> POT Horn Off Monte Carlo Horn On Monte Carlo **ANN Selected** ANN Selected 800 -NC -NC - v., CC - v., CC 600 - beam v, CC - beam v, CC 400 v<sub>e</sub> Events / 1 e 200 **Reconstructed Energy (GeV)** Reconstructed Energy (GeV)

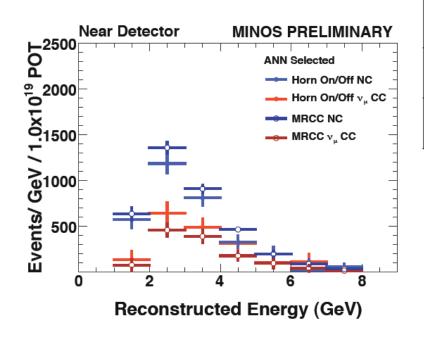


#### D. Harris Fermilab NuFact09 Betabeam and Neutrino Factory

- What changes can be made to the muons or radioactive isotopes to make changes in the flux?
  - Neutrino Factory: have investigated muon polarization to change relative amounts of  $v_e$  and  $v_{\mu}$  in the beam
    - Does this change the background or signal fractions enough?
- Beta Beams and neutrino factories: can energy of particles in storage ring be changed enough to do a systematic check in finite time?
- Aside on Superbeams:
  - NOvA will not have the capability to move target around, will changing horn current be enough?
  - T2K: beamline can take several off-axis angles, is this still a possibility for systematic studies?

### D. Harris Fermil Predicting Far Detector $v_e$ Candidates

Two strategies for "predicting" near detector v<sub>e</sub> candidates agree within errors, total background error quoted (after millions of ND events taken): 7.3%, statistical error is 23%

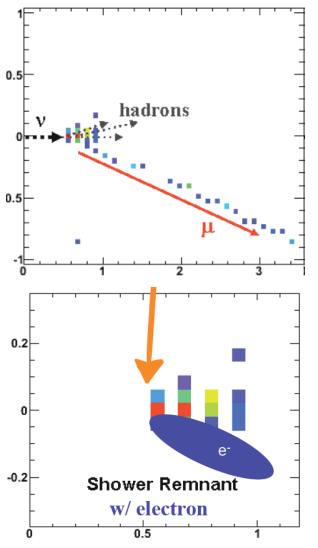


|                 | Total | NC   | ν <sub>µ</sub> CC | ν <sub>τ</sub> CC | $\nu_e$ beam |
|-----------------|-------|------|-------------------|-------------------|--------------|
| Horn on/<br>off | 27    | 18.2 | 5.1               | 1.1               | 2.2          |
| MRCC            | 28    | 21.1 | 3.6               | 1.1               |              |

How can MINOS predict Signal efficiency when it can't See individual  $v_e$  CC events?

# $v_e$ Signal efficiency at MINOS

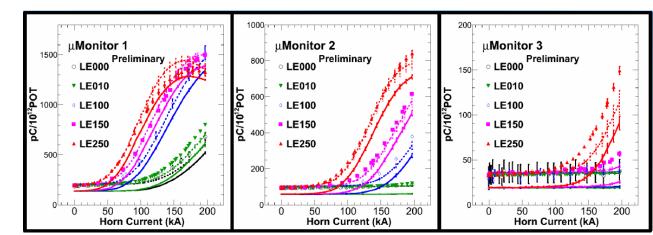
- Combine test beam information on electrons with neutrino beam information on hadron showers
- Note: this procedure assumes you know ratio of signal cross sections to normalization cross sections perfectly
- Note: this would be much harder in a near detector in a β beam: can't simply "remove" an electron shower and look at the remaining hadron shower
- Also note: in neutrino factory, the " $v_{\mu}$ " is an antineutrino...

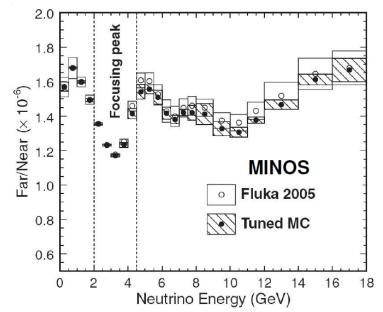




#### D. Harris Fermilab NuFact09 Need Absolute Flux Measurements

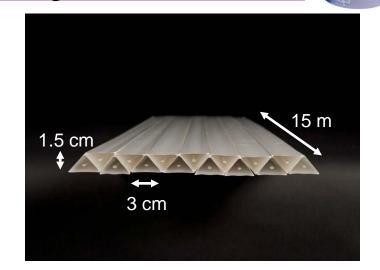
- MINOS ran with several different horn currents and target positions to best understand the far/near ratio (previous NuFact presentations)
- Currently evaluating muon monitor response versus focusing as well (Loiacono, Tuesday session)
- NuMI has flux measurements for 3 different populations of muons
- Plan for variable beam energy running resulted in much lower flux systematics

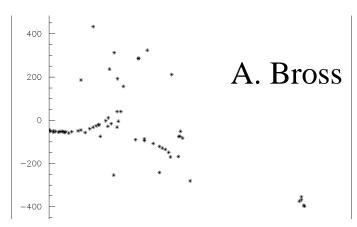




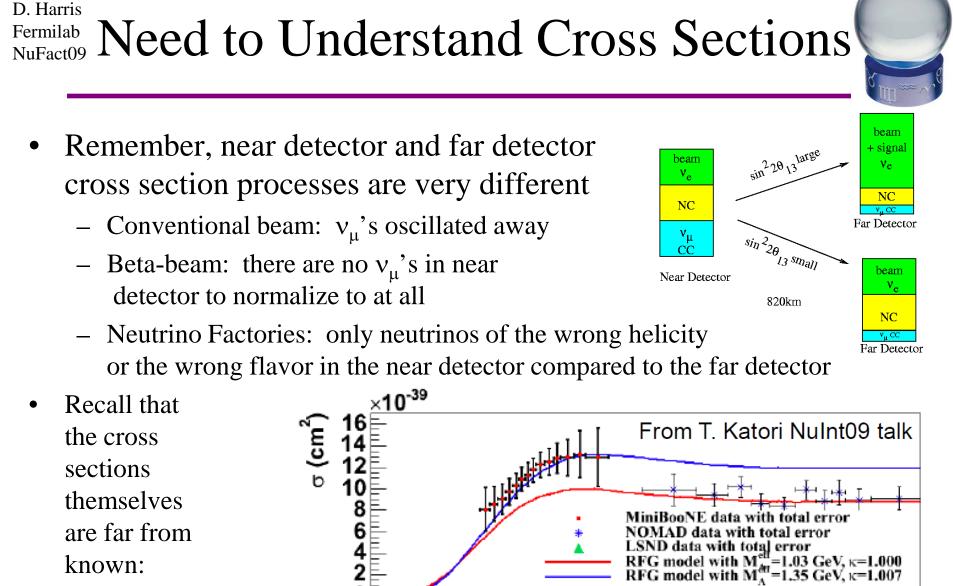
D. Harris Fermilab NuFact09 Understanding detector acceptance and background rejection

- Totally Active Scintillator Detector prototyping: trying to figure out how to put "MINERvA bars" in a large magnetic field
  - Can understand wrong sign backgrounds in low energy neutrino factory experiments
  - Need to understand ideal density for electron charge sign measurements  $(v_{\mu} \rightarrow v_{e})$
  - Hadron Shower studies in B field
- Liquid Argon Prototyping: MicroBooNE





- Can study hadron showers and MiniBooNE low energy excess



### Quasielastic news:

• See Dave Schmitz's Talk, Friday 9:45AM

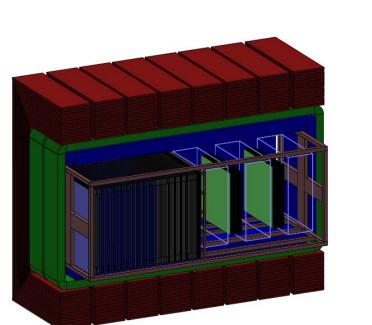
29

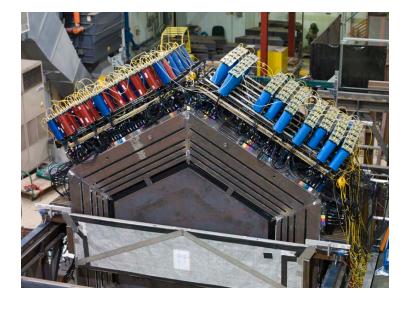
10

## Help is on the way...soon!

• MINERvA: exclusive final state measurements, 4 nuclear targets, to run in NuMI beamline in time for MINOS and NOvA (and T2K's) data

> T2K 280m Off axis detector to run in T2K beamline: inclusive π<sup>0</sup> measurements and some exclusive states, water target







#### D. Harris Fermilab NuFact09 Conclusions and Homework Assignments



- Many ideas for getting thousands of oscillated events in search of CP Violation & Mass Hierarchy
- To the Accelerator Physicists
  - Design in ability to cross-check the neutrino flux!
    - Muon monitoring stations for conventional beams
    - Varying polarization for muon storage rings
    - Measurements of ion and muon divergence in storage ring
- To the Theoretical Physicists
  - Develop/Improve description of neutrino interactions
    - Not just Quasi-elastic cross sections
    - Secondary particle production spectra:  $\pi^{\pm}$ ,  $\pi^{0}$
    - Signal and Background processes both important
- To the Experimental Particle Physicists
  - Measure your cross sections before you start, figure out hadron showers
  - Measure your detector response before you decide to build 100 ktons
  - Measure it again when design is final, in a test beam