



#### Mark Thomson University of Cambridge

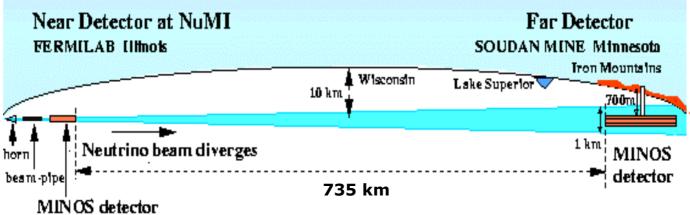




#### **This talk:**

- Overview
- NuMI Beam
- MINOS Far and Near Detectors
- Physics Capabilities
- First Data
  - cosmic muons
  - atmospheric Vs

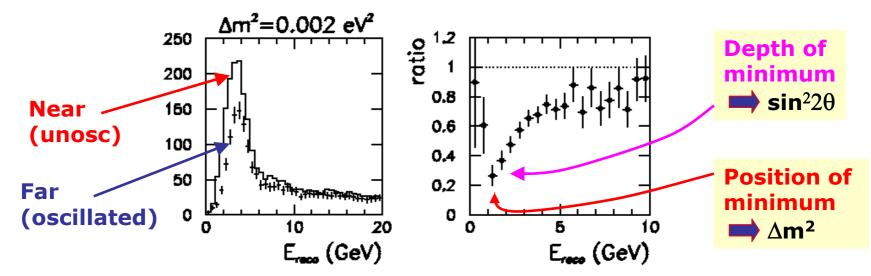
## **MINOS : Basic Idea**





Measure ratio of neutrino energy spectrum in far detector (oscillated) to that in the near detector (unoscillated)

**Partial cancellation of systematics** 



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## **MINOS Physics Goals**

#### **Demonstrate oscillation behaviour**

- confirm flavour oscillations describe data
- provide high statistics discrimination against alternative models:

decoherence, v decay, extra dimensions, etc.

- **\star** Precise Measurement of  $\Delta m_{23}^2$ 
  - ~10 %
- **Search for sub-dominant**  $v_{\mu} \rightarrow v_{e}$  oscillations
  - first measurements of  $\theta_{\textbf{13}}$  ?
- MINOS is the 1<sup>st</sup> large deep underground detector with a B-field
  - first direct measurements of  $\nu$  vs  $\overline{\nu}$  oscillations from atmospheric neutrino events



## The NuMI beam





- 120 GeV protons extracted from the MAIN INJECTOR in a single turn (8.7µs)
- ★ 1.9 s cycle time
- ★ i.e. ∨ beam `on' for 8.7µs every 1.9 s
- **\* 2.5x10<sup>13</sup> protons/pulse**
- ★ 0.3 MW on target !
- **\*** Initial intensity
  - 2.5x10<sup>20</sup> protons/year

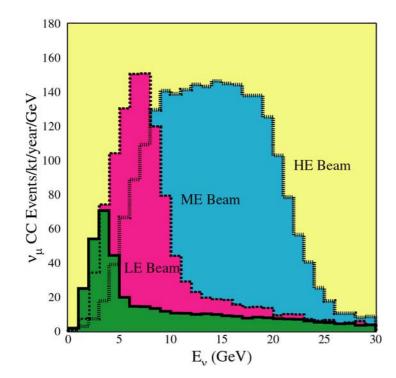
FERMILAB #98-765D

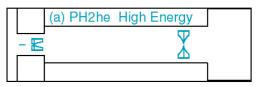


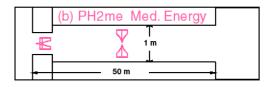
## **Tunable beam**

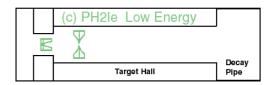


- Relative positions of the neutrino horns allow beam energy to be tuned. Act like a pair of (highly achromatic lenses)
- ★ Start with LE beam best for △m<sup>2</sup>~0.002 eV<sup>2</sup>

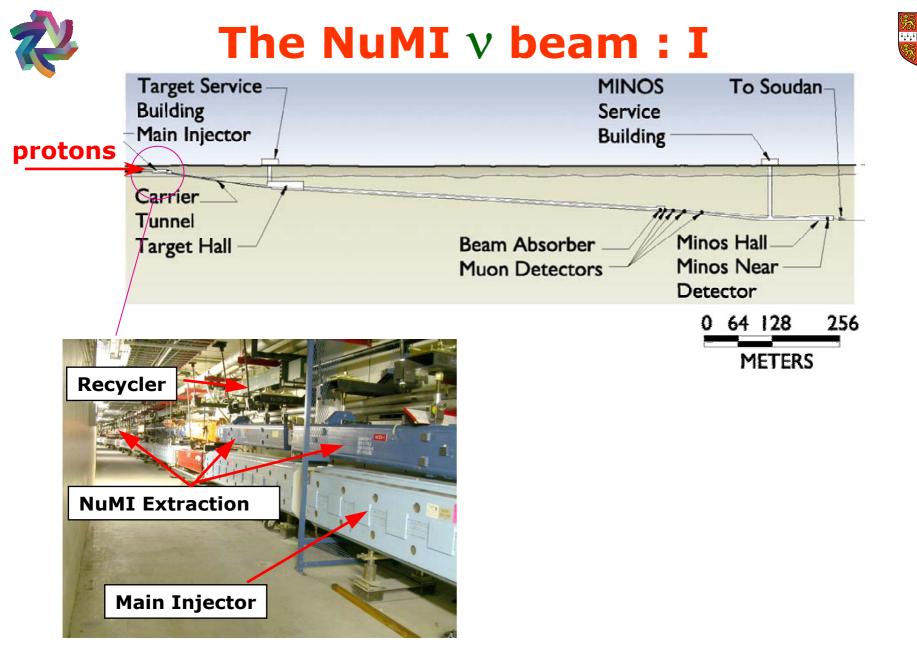


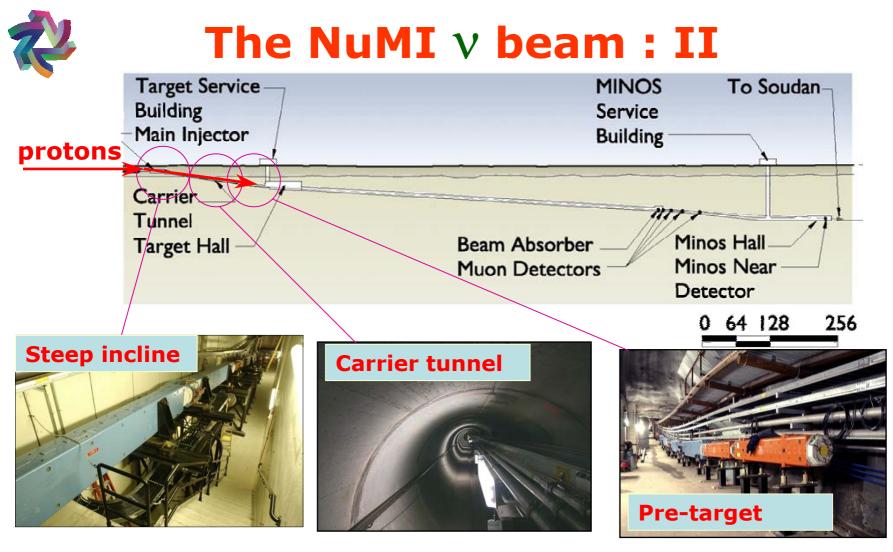




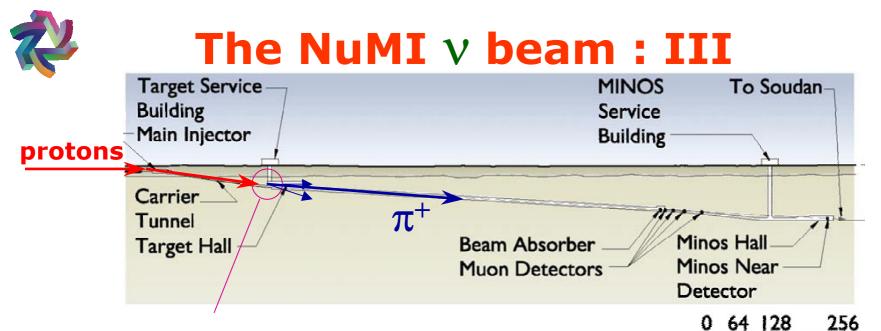








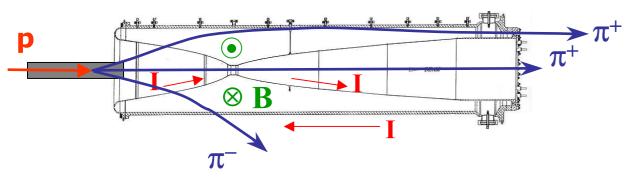
#### ★Beam points 3.3° downwards

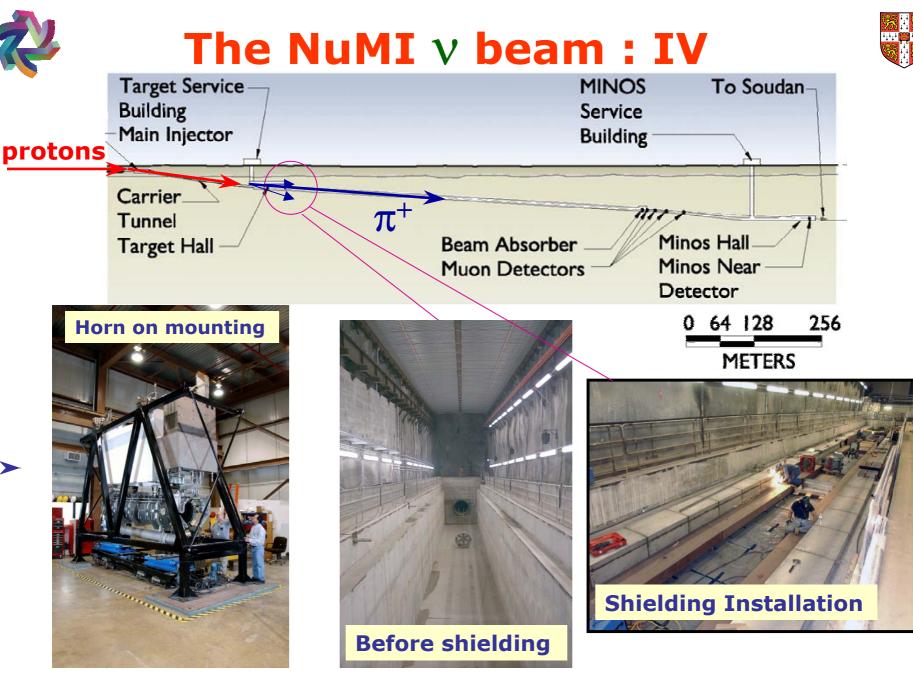


• Horn pulsed with 200 kA

METERS

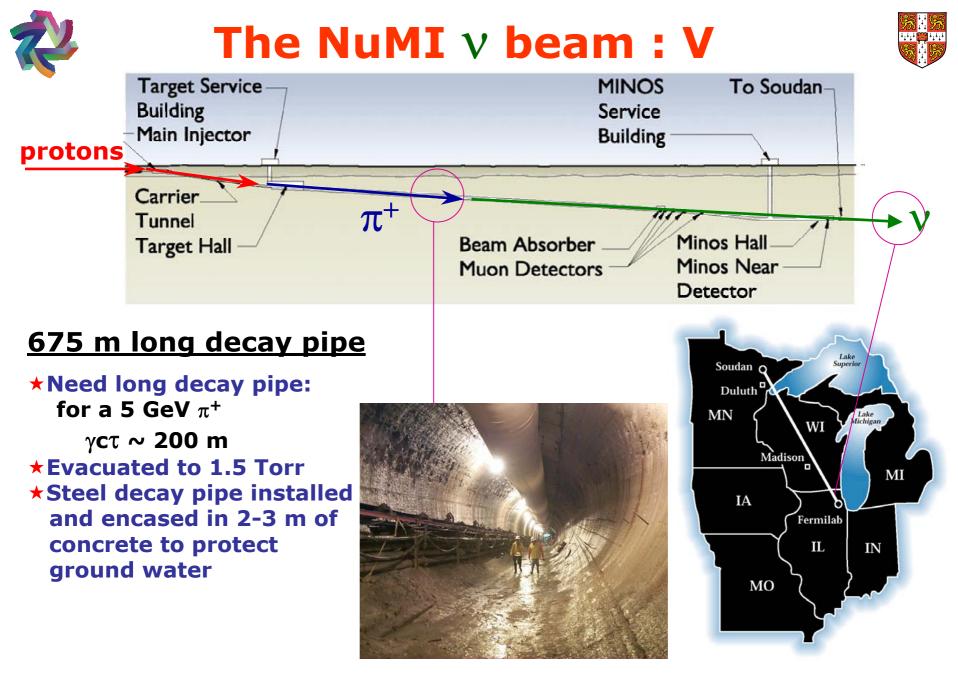
 Toroidal Magnetic field B ~ I/r between inner and outer conducters





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## **Going underground**



IF



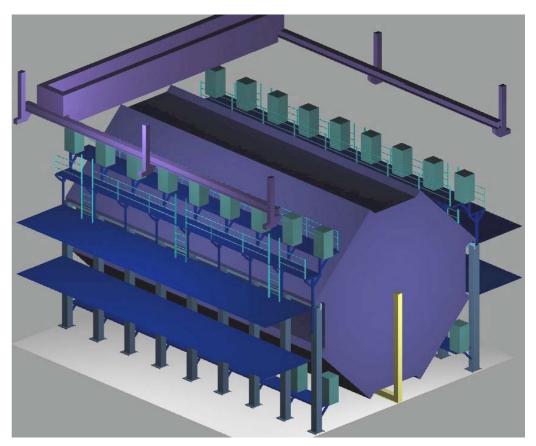




#### 8m octagonal steel & scintillator tracking calorimeter

- 2 sections, 15m each
- 5.4 kton total mass
- 55%/√E for hadrons
- 23%/ $\sqrt{E}$  for electrons

Magnetized Iron (B~1.5T) 484 planes of scintillator

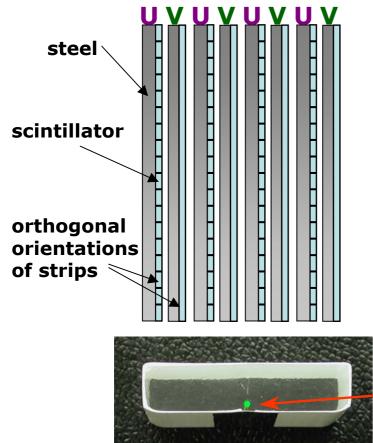


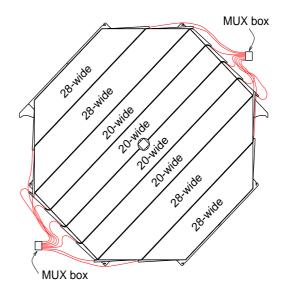
One Supermodule of the Far Detector... Two Supermodules total.





- **\* Steel-Scintillator sandwich : SAMPLING CALORIMETER**
- **★** Each plane consists of a 2.54 cm steel +1 cm scintillator
- **★** Each scintillator plane divided into 192 x 4cm wide strips
- **\*** Alternate planes have orthogonal strip orientations (U and V)

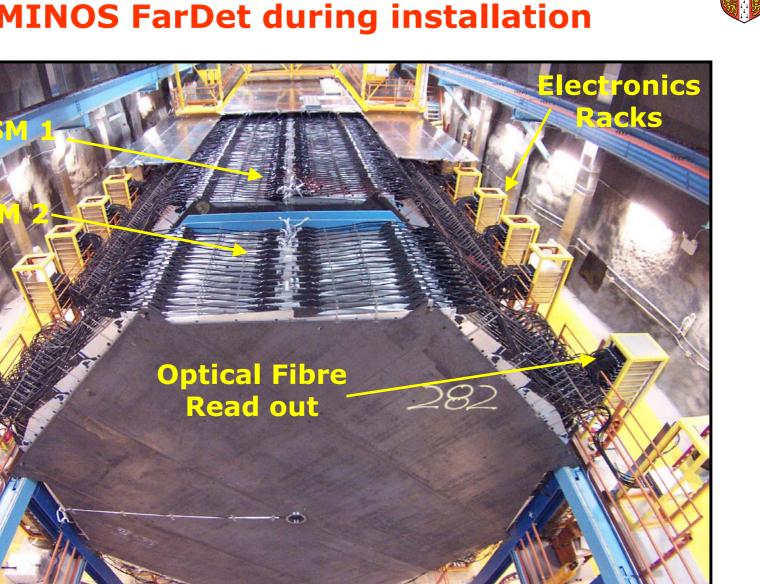




# Scintillation light collected by WLS fibre glued into groove Readout by multi-pixel PMTs



#### **MINOS FarDet during installation**

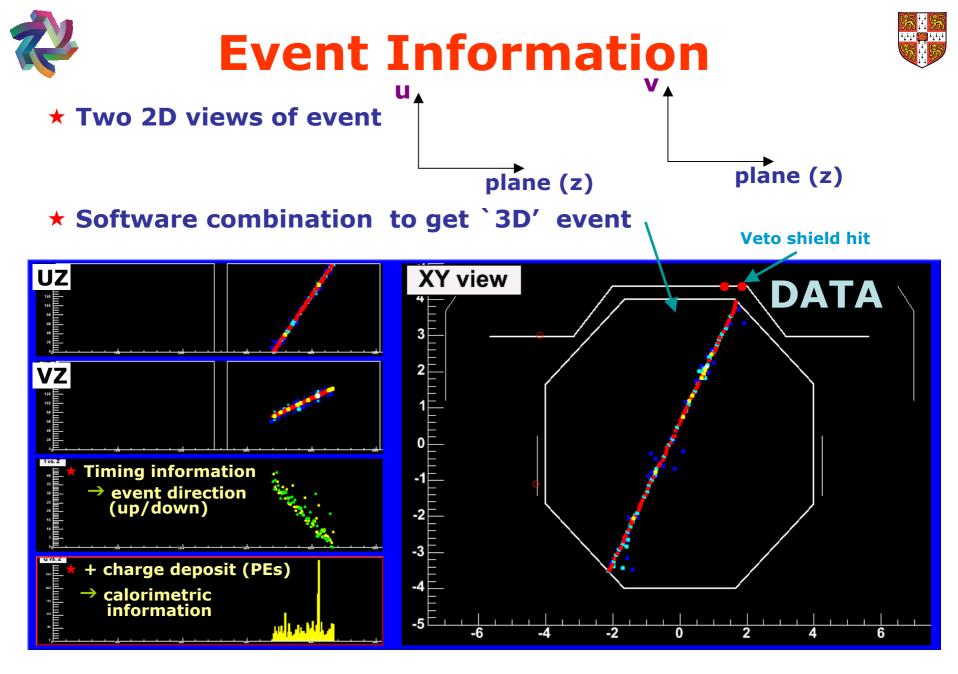




#### Far Detector fully operational since July 2003









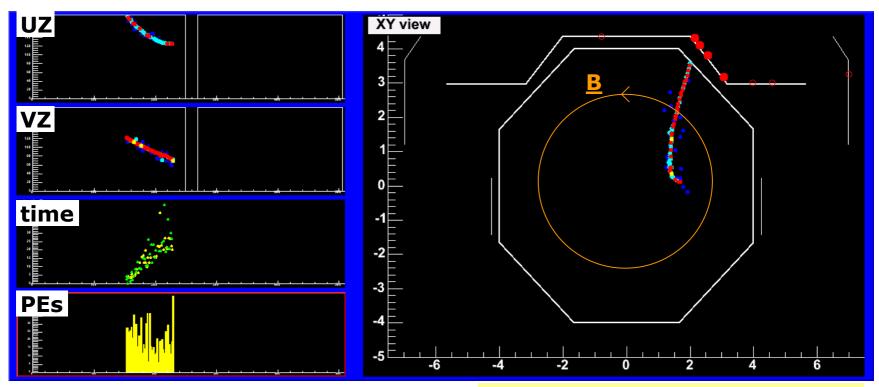




#### ~ 1.5 T Magnetic Field

**\*** Charge separation

#### Momentum measurement



#### Single Hit Resolution : 2.5 ns

<u>Stopping muon</u> P<sub>range</sub> = 3.86 GeV/c P<sub>curvature</sub> = 4.03 GeV/c



## **MINOS Near Detector**



- ★ 1 kton total mass
- Same basic design steel, scintillator, etc

#### ★ <u>Some differences, e.g.</u>

**Faster electronics** 

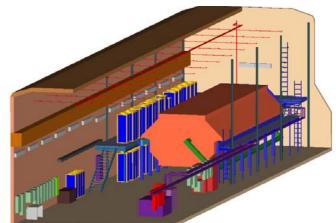
**Partially instrumented:** 

282 planes of steel

**153 planes of scintillator** 

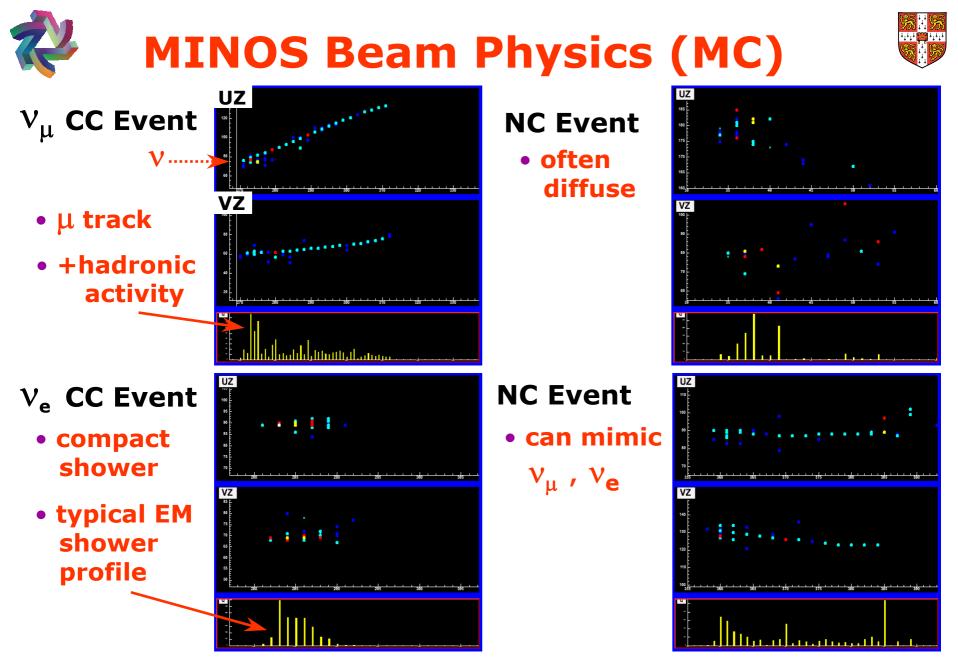
(Rear part of detector only used to track muons )

+.....



#### **Currently being installed at Fermilab**





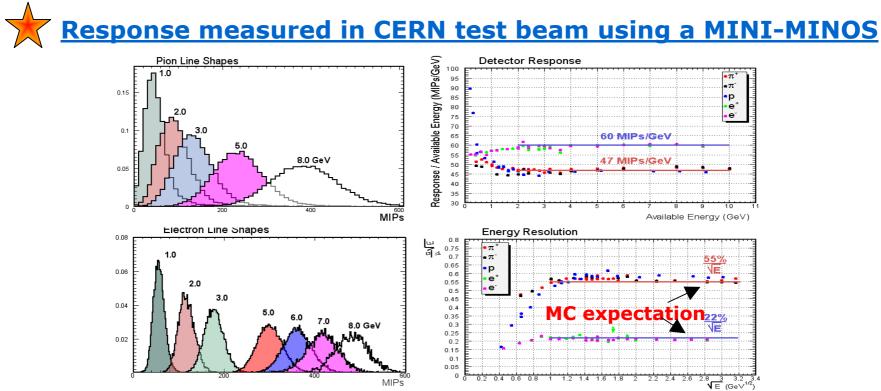


#### **Test Beam**



#### ★Energy response is important – know L, need Ev

- hadronic energy from pulse height (σ<sub>E</sub>/E ~ 55%/E<sup>1/2</sup>)
- $\mathbf{E}_{\mathbf{v}} = \mathbf{p}_{\mu} + \mathbf{E}_{had}$

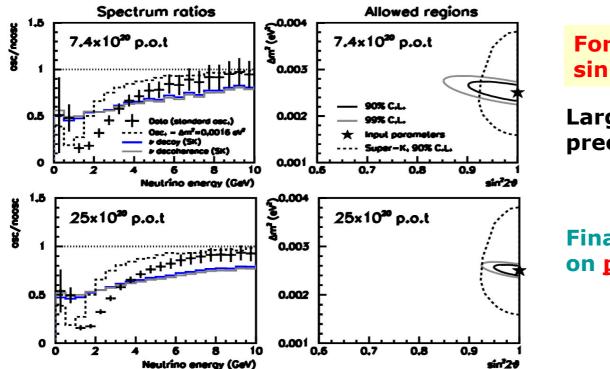


## Provides calibration information Test of MC simulation of low energy hadronic interactions

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# MINOS Physics Sensitivity ★ Measurement of Δm<sup>2</sup> and sin<sup>2</sup>2θ





For  $\Delta m^2 = 0.0025 \text{ eV}^2$ , sin<sup>2</sup> 2 $\theta = 1.0$ 

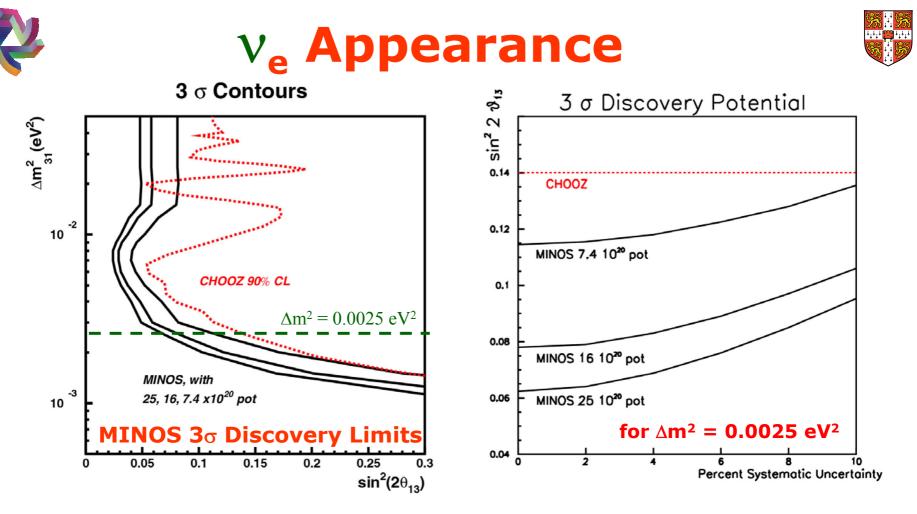
Large improvement in precision !

Final sensitivity depends on <u>protons</u> on target

#### \* Direct measurement of L/E dependence of $v_{\mu}$ flux \* Powerful test of flavour oscillations vs. alternative models

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 \* 3 σ discovery potential may significantly eat into current allowed region – exact reach depends on protons on target

\* reasonable chance of making the first measurement of  $\theta_{13}$  !



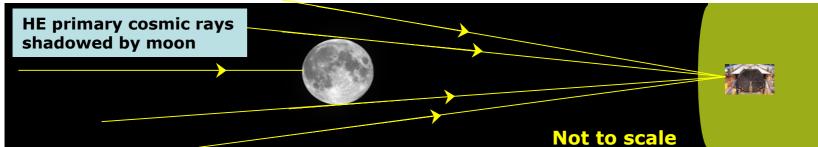


## First beam in December 2004 <u>BUT</u> Already Have Data....

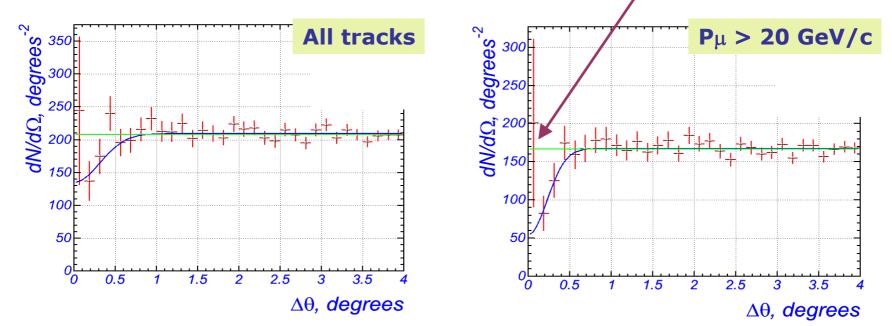


## **Moon Shadow**





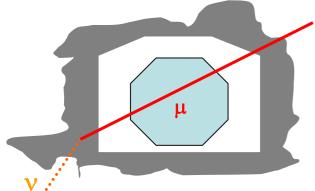
# \*Have recorded 10 M cosmic muons observed shadow of moon \*Angular res. improved by selecting high momenta muons (less multiple scattering) /



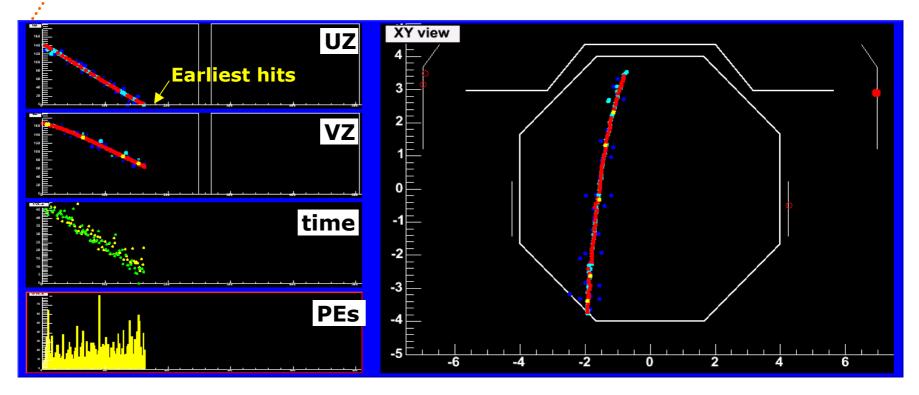
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## $\nu$ induced upward $\mu$



# ★Expect : 1 Event/6 Days ★Identified on basis of timing

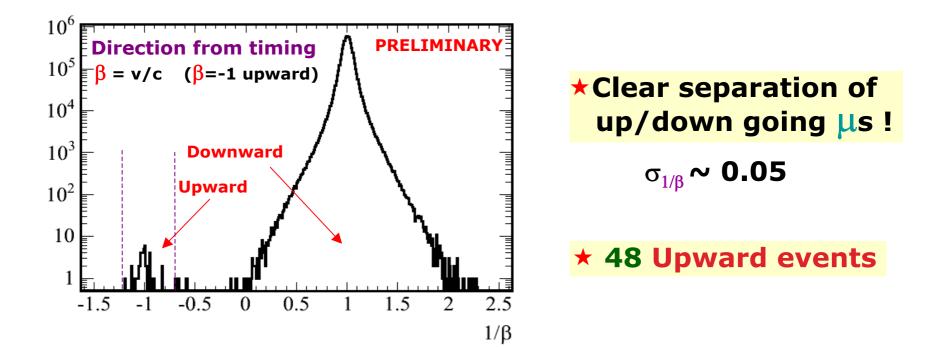




### $\boldsymbol{\nu}$ induced upward-going muons

- **★** Look for events coming from below horizon
- Require clear up/down resolution from timing
  - `Good track' > 2.0 m
  - >20 planes crossed

**\***Calculate muon velocity from hit times:  $\beta = v/c$ 



# Upward $\mu$ Analysis: Data vs. MC

#### **NUANCE generator:**

- Bartol '96 flux
- MC normalised to data (assuming no oscillations)

#### Charge-tagging:

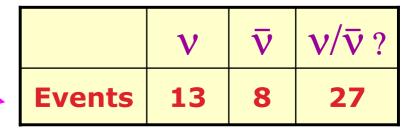
- Tag  $v/\overline{v}$  using muon charge
- Efficiency depends on:
  - muon momentum
  - track length
  - orientation wrt B-field
- Clean charge ID for approx. 50 % of events —

Understanding	systematics :	Work in progress
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10	No oscillations	
16	$\sin^2 2\theta = 1$ , $\Delta m^2 = 2.5 \times 10^{-3} (eV^2)$	
14		
12	1 .	
10		
8		
4-		
2	PRELIMINA	ARY
0 <u></u>	-0.9-0.8-0.7-0.6-0.5-0.4-0	).3-0.2-0.1 -0
		$\cos\theta$

Data



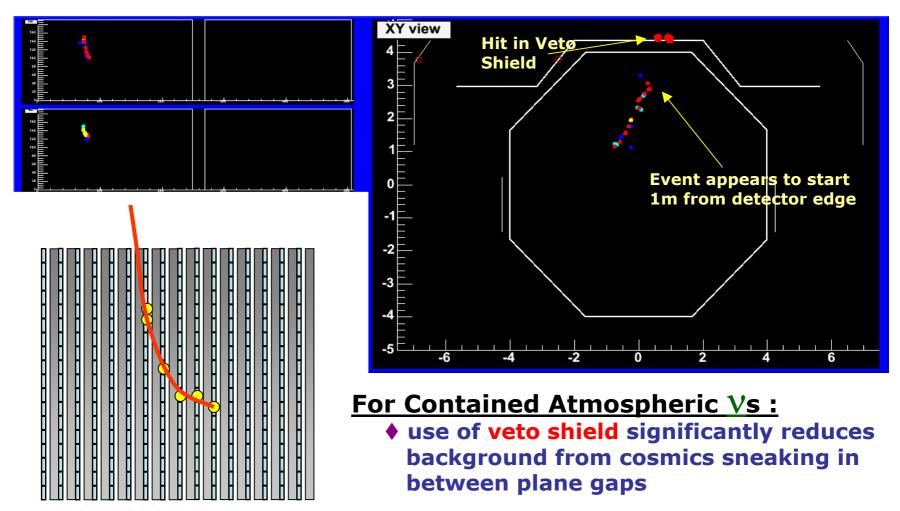




## **Contained Events**



# ★ MINOS Designed for Vs from FNAL – not atmospherics ★ Gaps between planes - potentially problematic





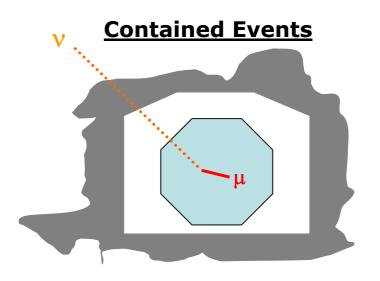




Have achieved rejection factor of ~ 1:10,000,000 !
 Efficiency ~ 75 % with 98 % purity

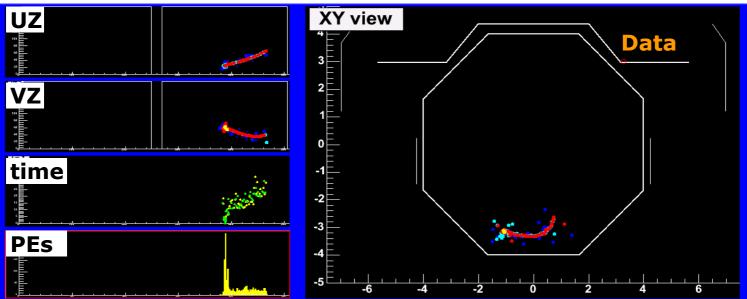
#### <u>CC $v_{\mu}$ EVENT SELECTION:</u>

- •Fiducial Volume: little activity within 50cm of detector edge
- •Reconstructed muon track track which crosses 8 planes
- •Cosmic muon rejection remove steep events
- Veto Shield no`in-time' Veto shield hit





# **Contained Event Selection**



#### **MINOS Preliminary**

	DATA	MC V no osc.*	MC Cosmic backgnd.
Before VETO	88	39	63±6
VETOED	51	1	61±6
v selection	37	38±8	2

Measure cosmic  $\mu$  bgd. from data using events solely rejected on basis of veto hit

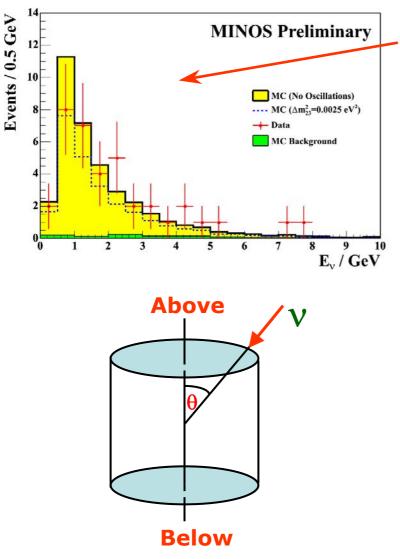
Vetoed background agrees with MC expectation !

#### $\nu$ MC : Battistoni et al

\* Does not include acceptance systematic uncertainties – work in progress



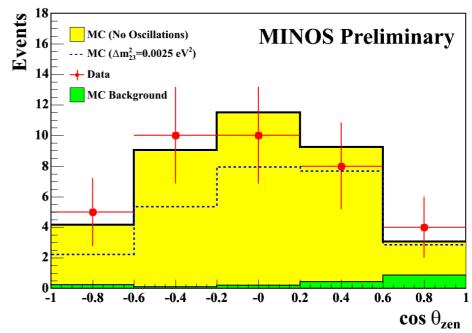
### **Event Distributions**





**MC normalised to data**(no oscillations)

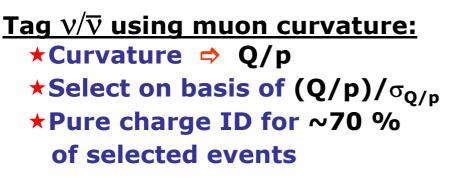
# Cosmic background from data from no. of vetoed events

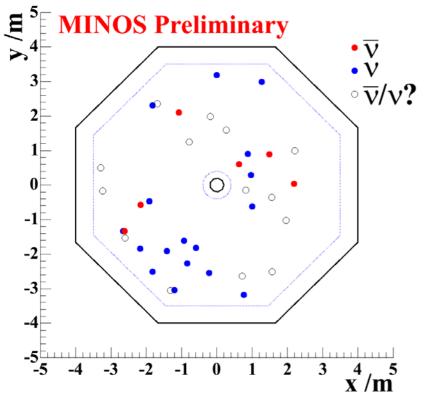


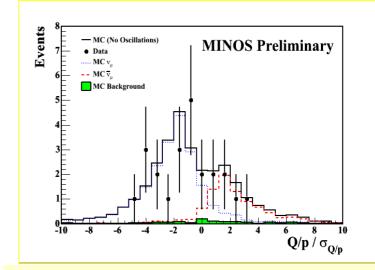
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## **Charge Reconstruction**







- **\*** 6  $\overline{v}$  events
- **★ 17** ∨ events
- **\* 14 too short to ID**  $\bar{v}$ /v

 $\Rightarrow N\overline{v}/Nv = 0.35 \pm 0.17$ 

(expect 0.51±? if  $\overline{\nu}/\nu$  oscillate with same parameters)

MINOS atmos v analysis underway ! just need more data.....

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## Conclusions





NuMI beam installation progressing well ! expect first protons on target December 2004 !



MINOS Near Detector currently being installed/ commisioned at FermiLab



MINOS Far Detector taking physics quality data since mid-2003



Atmospheric Vs already being seen in the MINOS Far Detector



First direct observation of  $v/\overline{v}$  separated atmospheric neutrinos



Eagerly awaiting first beam physics data, expected early 2005 ! Exciting times for MINOS.



## **MINOS en France**



